

Title: The Volume of the Universe after Inflation and de Sitter Entropy

Date: Dec 05, 2008 11:00 AM

URL: <http://pirsa.org/08110034>

Abstract: I will show the calculation of the probability distribution for the volume of the Universe after slow-roll inflation both in the eternal and the non-eternal regime. Far from the eternal regime the probability distribution for the number of e-foldings, defined as one third of the logarithm of the volume, is sharply peaked around the number of e-foldings of the classical inflaton trajectory. At the transition to the eternal regime this probability is still peaked (with the width of order one e-foldings) around the average, which however gets twice larger at the transition point. As one enters the eternal regime the probability for the volume to be finite rapidly becomes exponentially small. In addition to developing techniques to study eternal inflation, these results allow us to establish the quantum generalization of the recently proposed bound on the number of e-foldings in non-eternal regime: the probability for slow-roll inflation to produce a finite volume larger than $\text{Exp}[S_{dS}/2]$, where S_{dS} is the de Sitter entropy at the end of the inflationary stage, is smaller than the uncertainty due to non-perturbative quantum gravity effects. The existence of such a bound provides a consistency check for the idea of de Sitter complementarity.

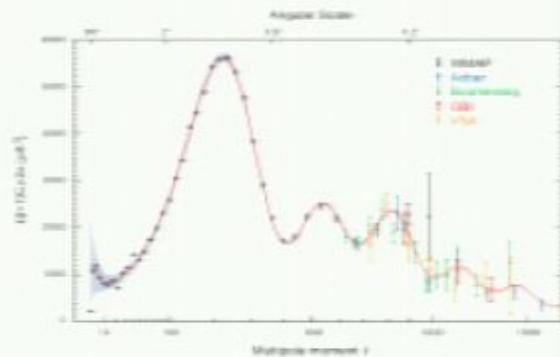
de Sitter and us

- Quasi dS phase in our present and in our past
- Possibly both are eternal

- Landscape in String Theory



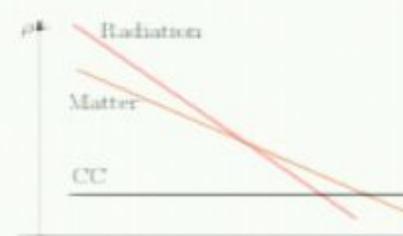
- Weinberg's solution to CC (and to other parameters)



- Eternal Inflation has become extremely important

- Difficult task

- Puzzles with de Sitter Entropy

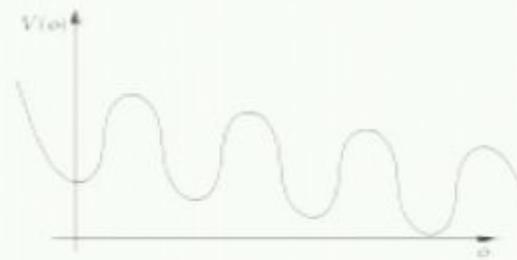


- This motivates detailed studies

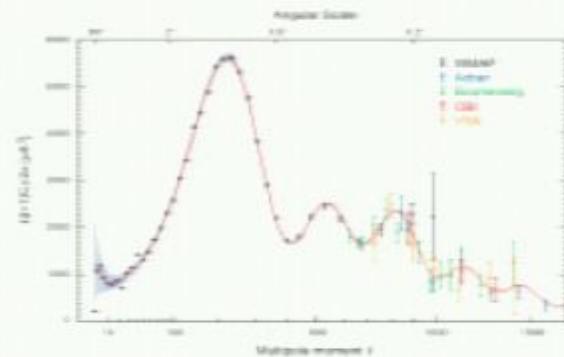
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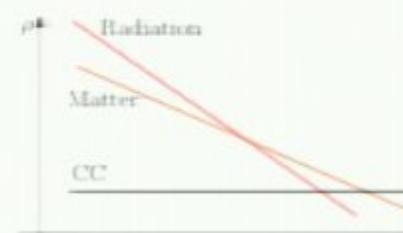
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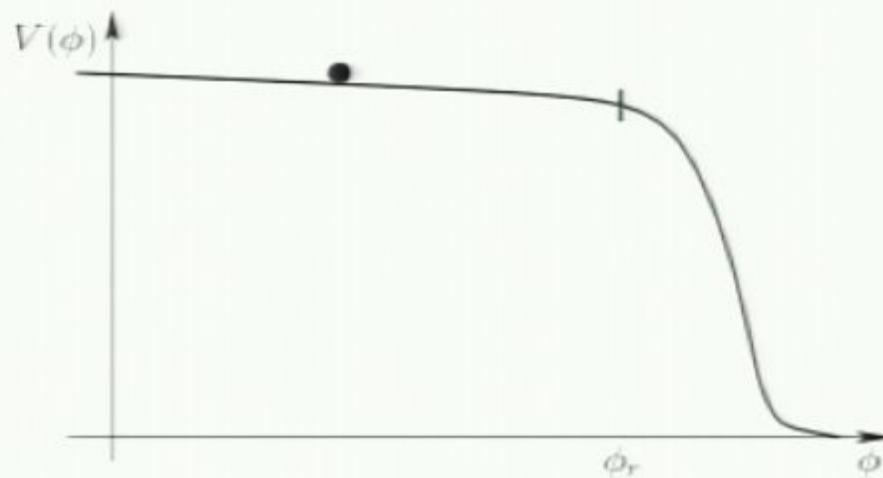


- This motivates detailed studies

What is Eternal Inflation?

What is Inflation?

$$a \sim e^{Ht}$$



What is Eternal Inflation?

Classical Motion Vs Quantum Motion

$$\Delta\phi_{\text{Cl}} \sim \dot{\phi} H^{-1} \quad \text{vs} \quad \Delta\phi_{\text{Q}} \sim H$$

Quantum dominates for $\frac{\dot{\phi}}{H^2} \lesssim 1 \Rightarrow$ Eternal Inflation

Make Eternal Inflation sharp?

- No Semiclassical

- No FRW $\delta a/a \sim H^2/\dot{\phi} \sim 1$

Perturbativity of the system

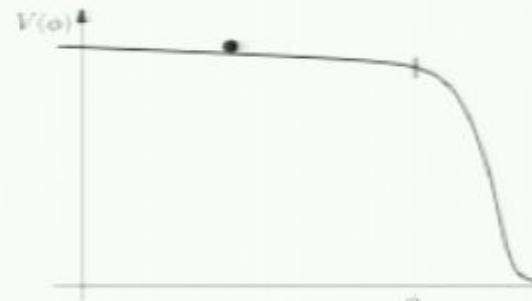
Close to de-Sitter

$$\epsilon \simeq \frac{\dot{\phi}^2}{V(\phi)} \ll 1$$

- Still unperturbed before reheating: $\delta g \sim \sqrt{\epsilon} \frac{H}{M_{\text{Pl}}}$

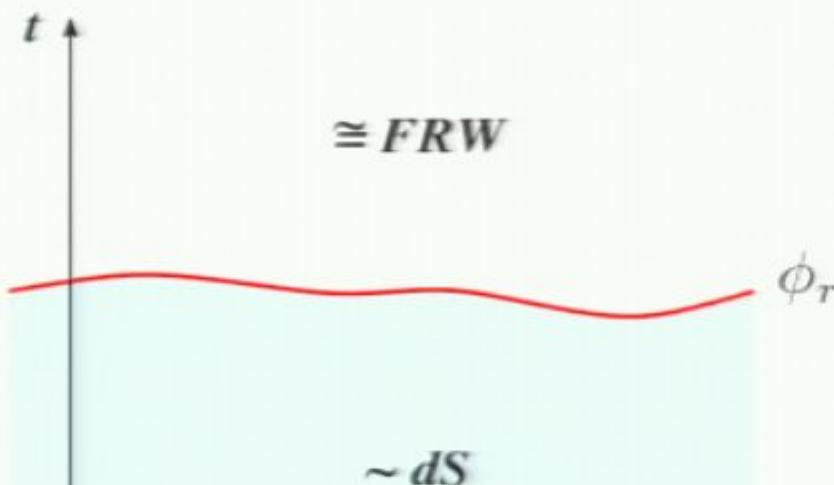
No big interactions:

$$\frac{S_3}{S_2} \sim \sqrt{\epsilon} \frac{H}{M_{\text{Pl}}}$$

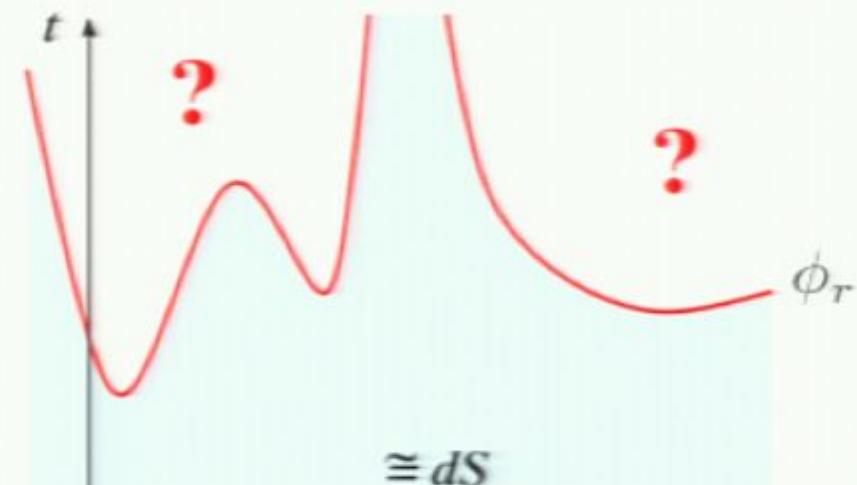


→ Study the volume of the Reheating surface $\phi = \phi_r$

Standard Infl.



Eternal Infl.



A random walk

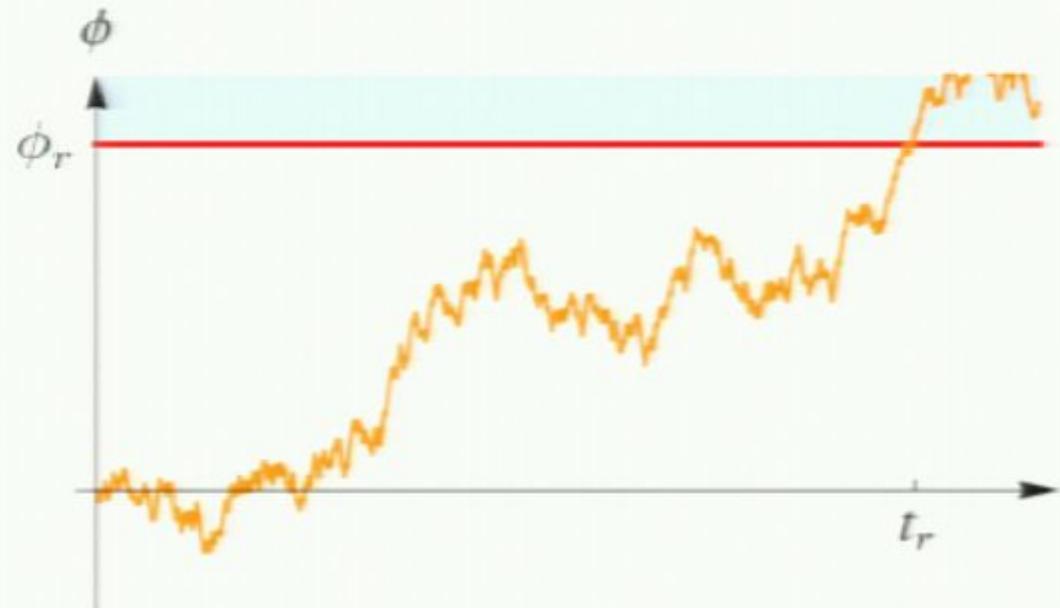
smoothing the field: $\Lambda \ll H$:

- » $\Delta t \gtrsim H^{-1}$ for reheating
- » $[\delta\phi_k, \dot{\delta\phi}_{-k}] \rightarrow 0$

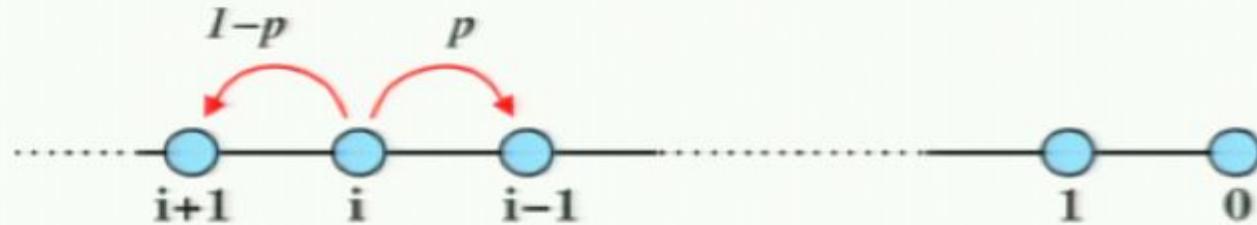
Inflaton \sim Classical stochastic system with Gaussian statistics

Probability distribution follows a diffusion equation

$$P(\bar{\phi}, \phi, t) \sim e^{-\frac{(\bar{\phi} - \phi - \dot{\phi}t)^2}{H^3 t}}$$



Bacteria Model: a discretization



Reproduction $N \Rightarrow \sim e^3$ New Hubble volumes

Number of dead bacteria \Rightarrow Reheated Volume

Classical motion $\dot{\phi} \Rightarrow p - \frac{1}{2}$,

$$j = -\frac{\phi}{\Delta\phi}, \quad n = \frac{t}{\Delta t}$$

Continuum Limit: $\bar{P}(j, n+1) = (1-p)\bar{P}(j-1, n) + p\bar{P}(j+1, n)$

$$\partial_{\sigma^2} P(\psi, \sigma^2) = \frac{4\pi^2}{H^3} \left((1-2p)\frac{\Delta\phi}{\Delta t} + \dot{\phi} \right) \partial_\psi P(\psi, \sigma^2) + \frac{1}{2} \frac{4\pi^2}{H^3} \frac{\Delta\phi^2}{\Delta t} \partial_\psi^2 P(\psi, \sigma^2)$$

Identifications:

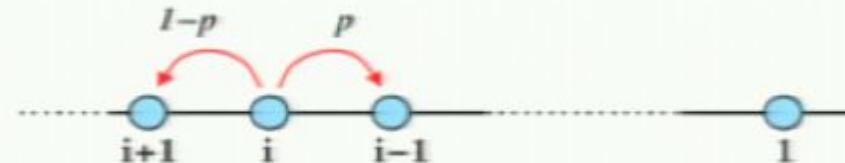
$$-(1-2p)\frac{\Delta\phi}{\Delta t} = \dot{\phi} \quad \frac{4\pi^2}{H^3} \frac{\Delta\phi^2}{\Delta t} = 1 \quad N = 1 + 3H\Delta t$$

The Extinction Probability

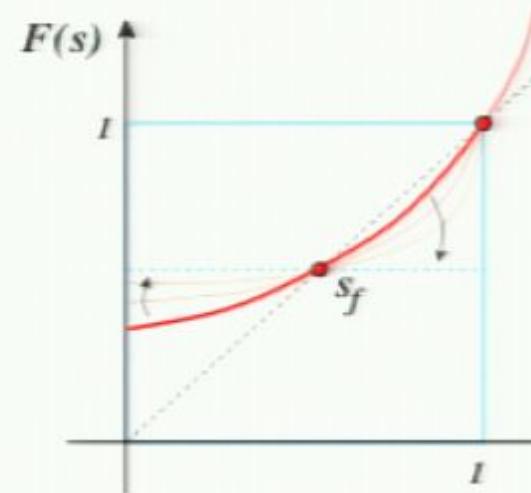
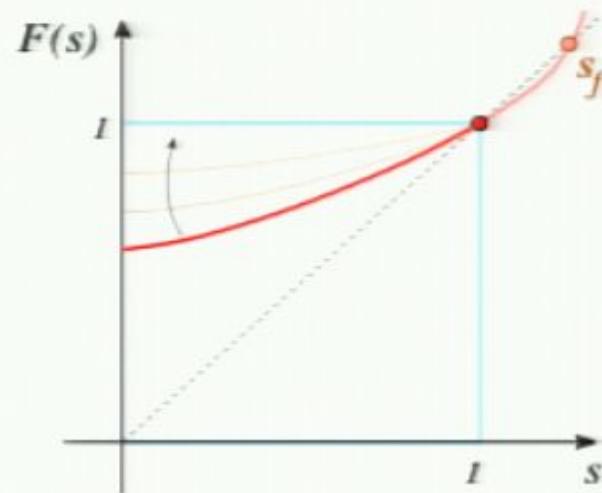
generating function:

$$f_i^{(n)}(s_j) = \sum_{k_1 \dots k_L} p_{i;k_0 \dots k_L}^{(n)} s_0^{k_0} \dots s_L^{k_L} \quad 0 \leq s_i \leq 1$$

Recursion: $F_{n+1} = F_1(F_n)$, $F_1(F_\infty) = F_\infty$



is only a function of s_0

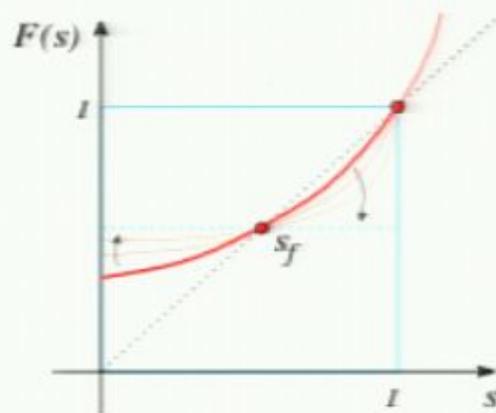
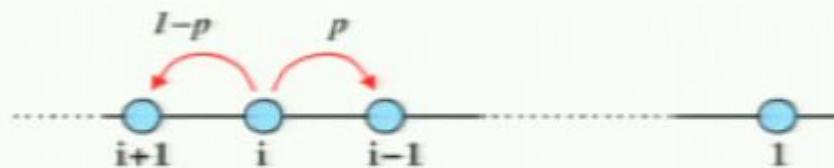


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$$\Rightarrow f_j^{(\infty)}(s_0) = \sum_{k=0}^{\infty} p_{j,k} s_0^k$$

concentrating on F_∞ :

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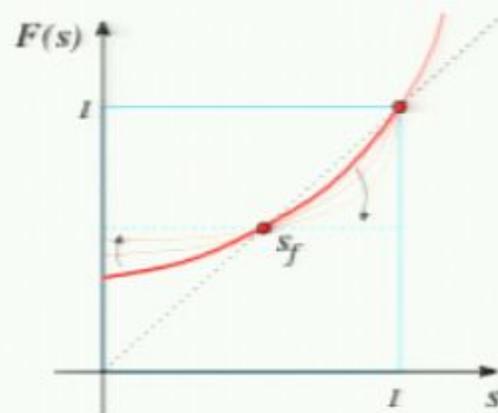
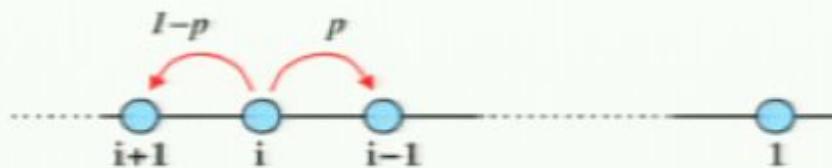


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$$\begin{aligned}
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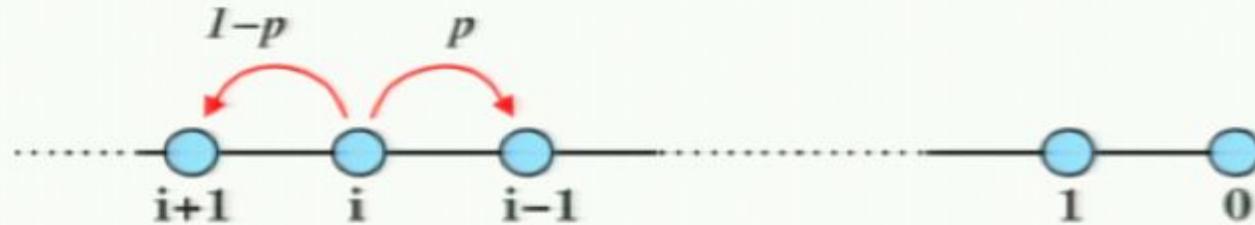
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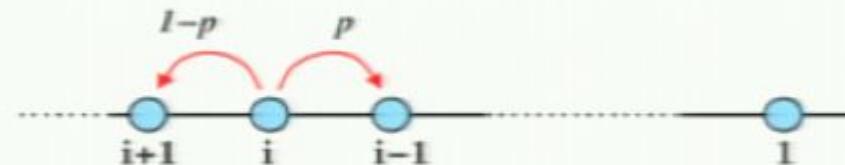
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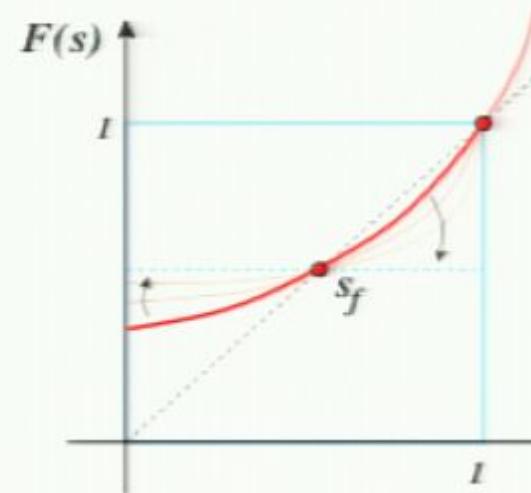
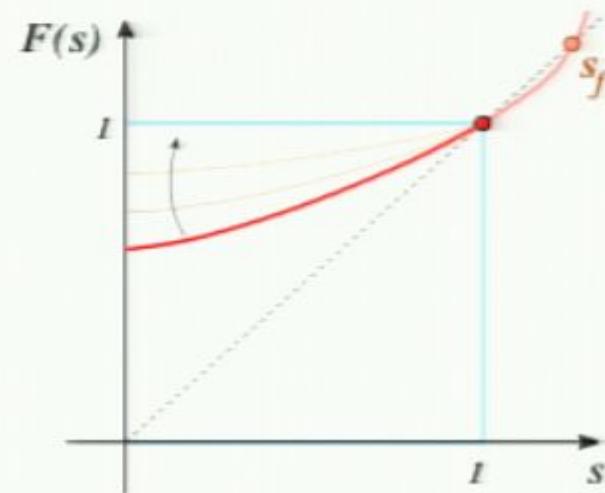
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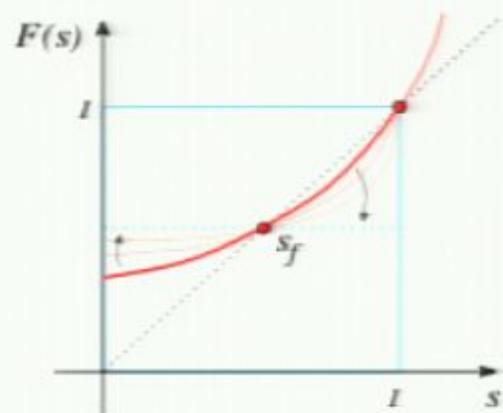
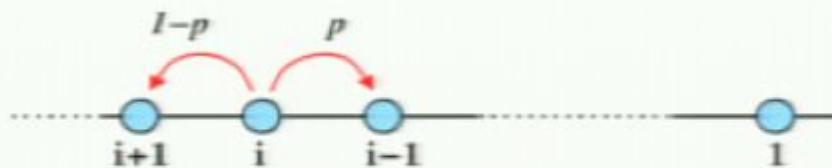


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king the continuum limit:

on the sites: $i \Rightarrow \phi$

A differential equation:

$$f^{(\infty)}(\phi; s_0) = ((1-p)f^{(\infty)}(\phi + \Delta\phi; s_0) + p f^{(\infty)}(\phi - \Delta\phi; s_0))^{N_r} \Rightarrow$$

$$\frac{\partial^2}{\partial \phi^2} f^{(\infty)}(\phi; s_0) - \frac{2\pi\sqrt{6\Omega}}{H} \frac{\partial}{\partial \phi} f^{(\infty)}(\phi; s_0) + \frac{12\pi^2}{H^2} f^{(\infty)}(\phi; s_0) \log [f^{(\infty)}(\phi; s_0)] = 0$$

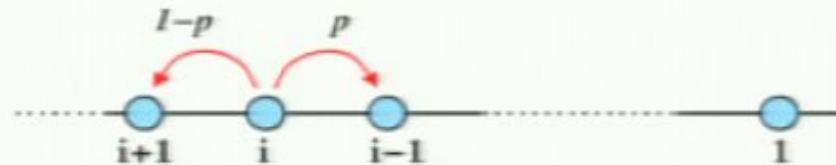
$$\Omega \equiv \frac{2\pi^2}{3}$$

with boundary conditions: $f^{(\infty)}(\phi = 0, s_0) = s_0$, $\left. \frac{\partial}{\partial \phi} f^{(\infty)}(\phi; s_0) \right|_{\phi_b} = 0$

or the quantity: $f_j^{(\infty)}(s_0) = \sum_{k=0}^{\infty} p_{j,k} s_0^k \Rightarrow f^{(\infty)}(\phi; s_0) = \int_0^{\infty} dV \rho(\phi, V) s_0^k$

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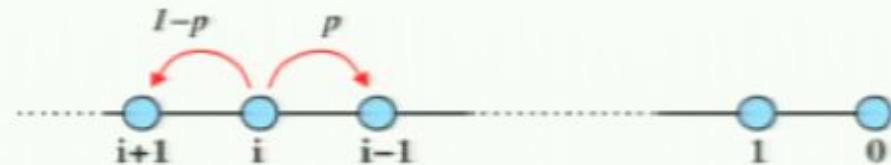
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Study the differential eq.

iff. Eq. :

$$\ddot{f}(\tau; z) - 2\sqrt{\Omega}\dot{f}(\tau; z) + f(\tau; z) \log [f(\tau; z)] = 0$$

here $\tau \propto \phi \propto N_c$, $z = -\log(s_0)$, $\Omega \equiv \frac{2\pi^2}{3} \frac{\dot{\phi}^2}{H^4}$

Boundary conditions:

$$f(0; z) = s_0 = e^{-z},$$

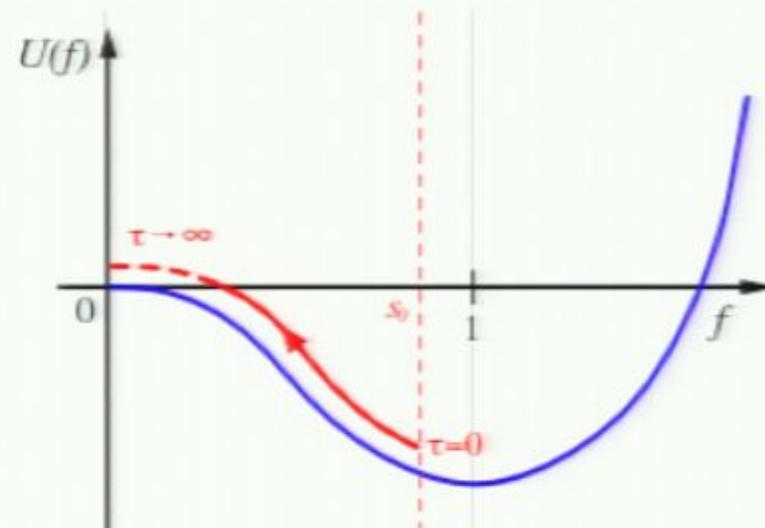
$$\dot{f}(\tau_b; z) = 0,$$

$$f(\tau; z) \in [0, 1]$$

Mechanical Problem: Anti-friction and potential

$$U(f) = \frac{f^2}{4} (\log f^2 - 1)$$

Barrier at infinity:



e Phase Transition

$$\ddot{f}(\tau; z) - 2\sqrt{\Omega}\dot{f}(\tau; z) + f(\tau; z) \log [f(\tau; z)] = 0$$

$$S_{\text{ext}} \equiv \int_0^\infty dV \rho(V, \tau) = f(\tau; 0)$$

$$s_0 \rightarrow 1 \quad \Rightarrow \quad z \rightarrow 0$$

⇒ Linearized solution:

$$\ddot{f} - 2\sqrt{\Omega}\dot{f} + f - 1 = 0,$$

$$\Rightarrow$$

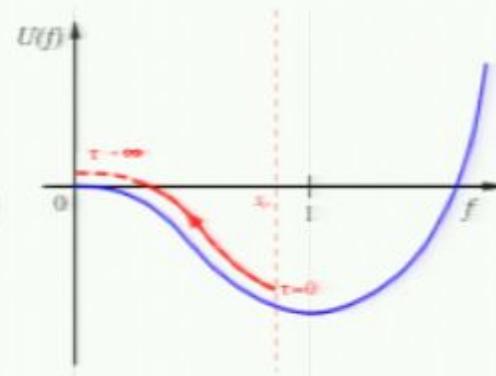
$$f = 1 - e^{\sqrt{\Omega}\tau} \left(A e^{\sqrt{\Omega-1}\tau} + B e^{-\sqrt{\Omega-1}\tau} \right)$$

Different behaviors for

$$\Omega \gtrless 1$$

$$\Omega \equiv \frac{2\pi^2}{3} \frac{\dot{\phi}^2}{H^4}$$

$$z = -\log(s_0)$$



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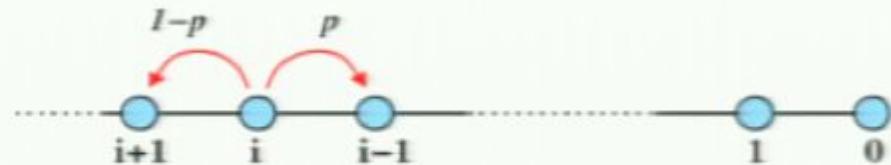
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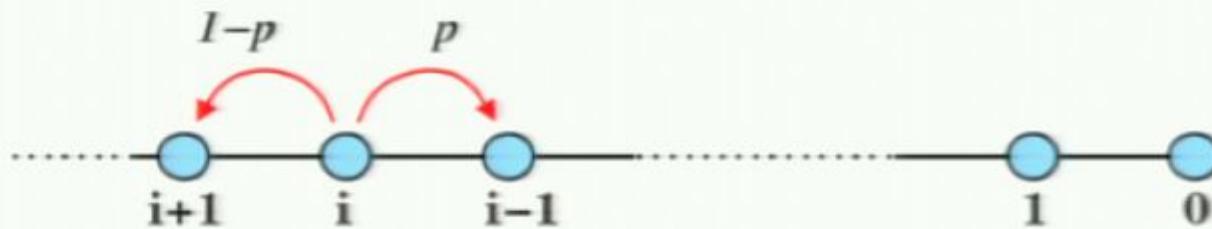
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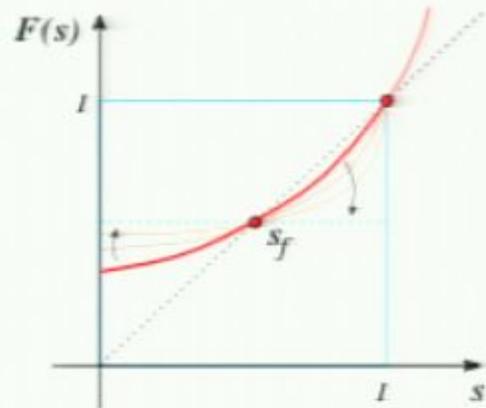
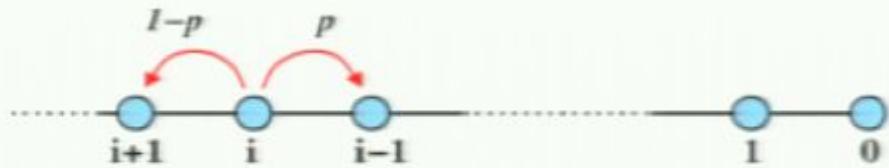
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$$\frac{\partial^2}{\phi^2} f^{(\infty)}(\phi; s_0) - \frac{2\pi\sqrt{6\Omega}}{H} \frac{\partial}{\partial\phi} f^{(\infty)}(\phi; s_0) + \frac{12\pi^2}{H^2} f^{(\infty)}(\phi; s_0) \log [f^{(\infty)}(\phi; s_0)] = 0$$

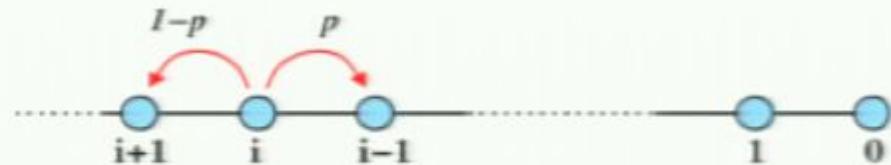
$$\Omega \equiv \frac{2\pi^2}{3} \frac{\dot{\phi}^2}{H^4}$$

with boundary conditions: $f^{(\infty)}(\phi = 0, s_0) = s_0$, $\left. \frac{\partial}{\partial\phi} f^{(\infty)}(\phi; s_0) \right|_{\phi_b} = 0$

or the quantity: $f_j^{(\infty)}(s_0) = \sum_{k=0}^{\infty} p_{j,k} s_0^k \Rightarrow f^{(\infty)}(\phi; s_0) = \int_0^{\infty} dV \rho(\phi, V) s_0^V$

e. for the Laplace transform of $\rho(\phi, V)$

$$\rho(\phi, V) = \frac{1}{2\pi i} \int_{\gamma-i\infty}^{\gamma+i\infty} d(-\log(s_0)) f^{(\infty)}(\phi; s_0) e^{-V \log(s_0)}$$



Identifying the differential eq.

iff. Eq. :

$$\ddot{f}(\tau; z) - 2\sqrt{\Omega}\dot{f}(\tau; z) + f(\tau; z) \log [f(\tau; z)] = 0$$

here $\tau \propto \phi \propto N_c$, $z = -\log(s_0)$, $\Omega \equiv \frac{2\pi^2}{3} \frac{\dot{\phi}^2}{H^4}$

boundary conditions:

$$f(0; z) = s_0 = e^{-z},$$

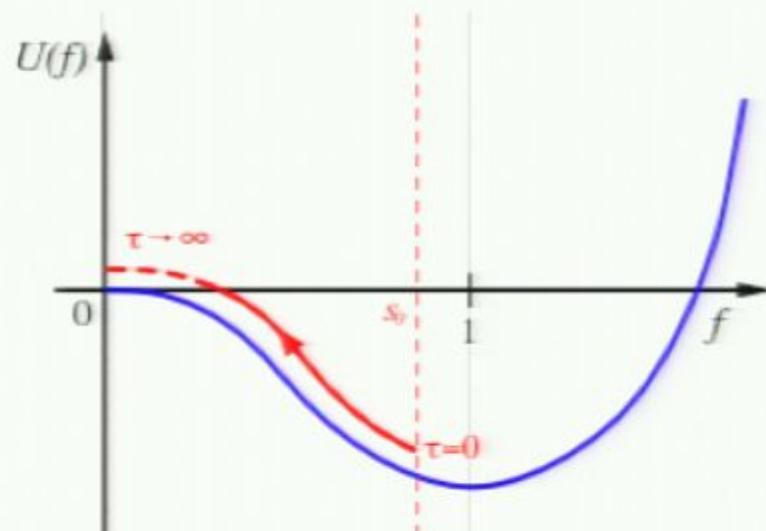
$$\dot{f}(\tau_b; z) = 0,$$

$$f(\tau; z) \in [0, 1]$$

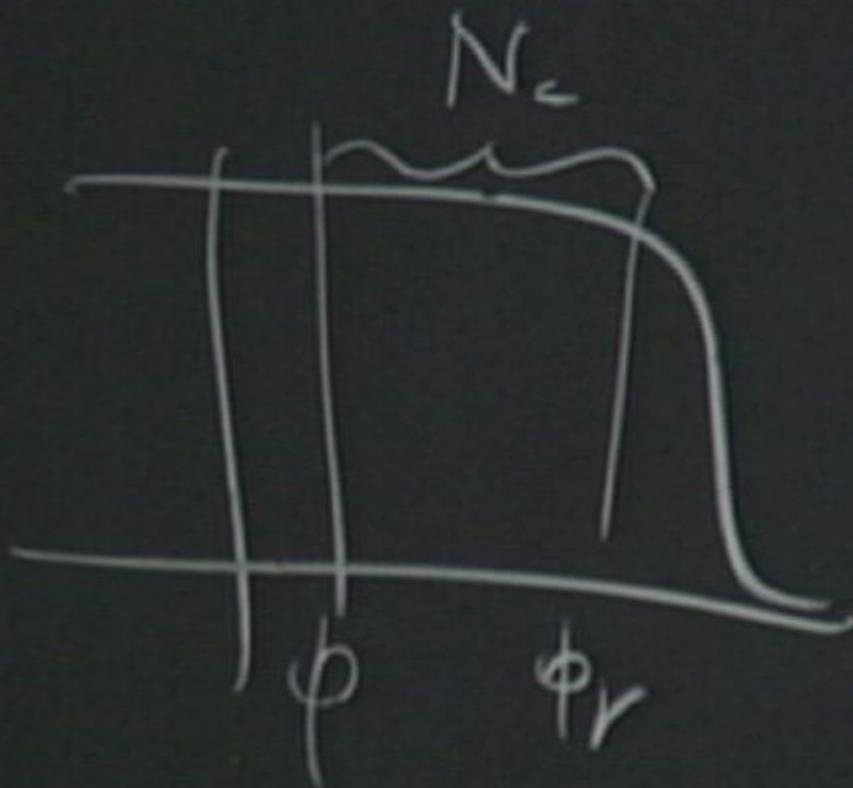
mechanical Problem: Anti-friction and potential

$$U(f) = \frac{f^2}{4} (\log f^2 - 1)$$

Barrier at infinity:



$$V = e^{3N}$$



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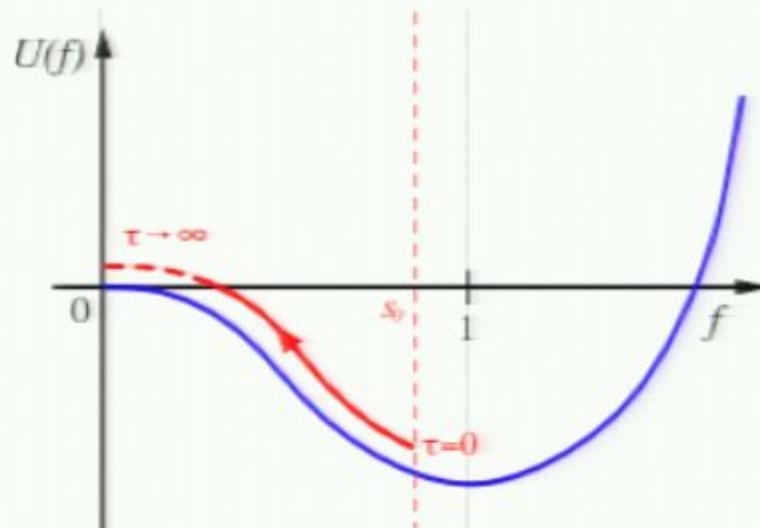
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$$s_0 \rightarrow 1 \quad \Rightarrow \quad z \rightarrow 0$$

⇒ Linearized solution:

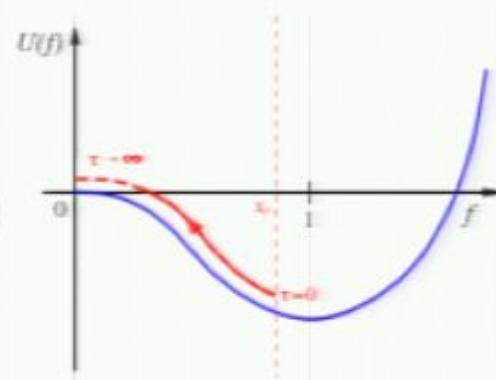
$$\ddot{f} - 2\sqrt{\Omega}\dot{f} + f - 1 = 0, \quad \Rightarrow$$

$$f = 1 - e^{\sqrt{\Omega}\tau} \left(A e^{\sqrt{\Omega-1}\tau} + B e^{-\sqrt{\Omega-1}\tau} \right)$$

Different behaviors for $\Omega \gtrless 1$

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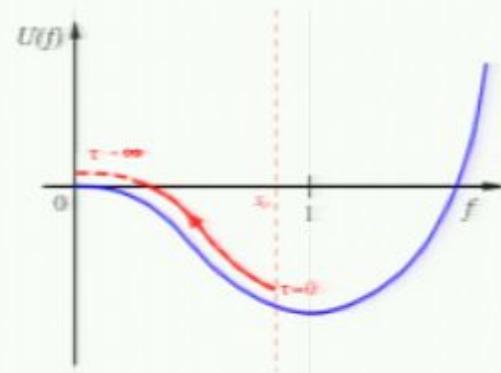
e Phase Transition

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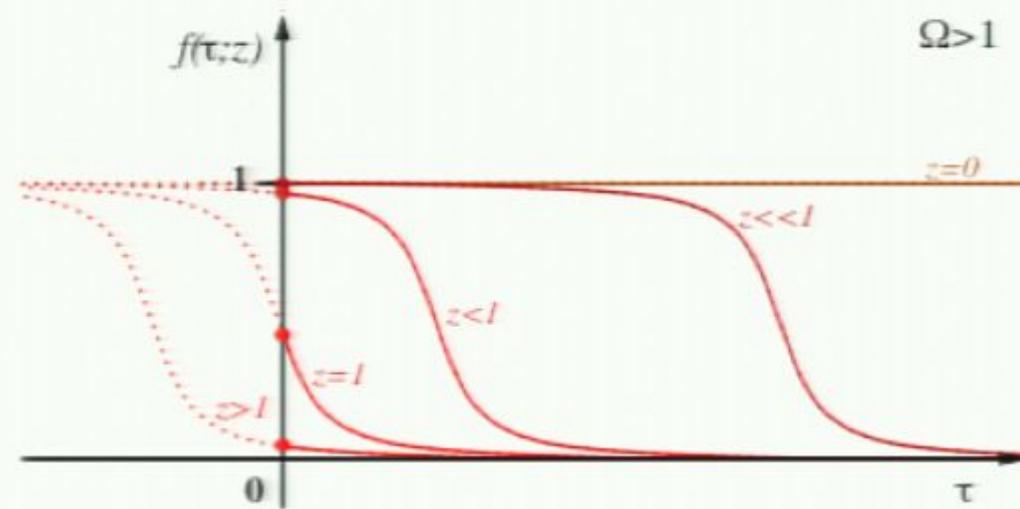
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total Extinction

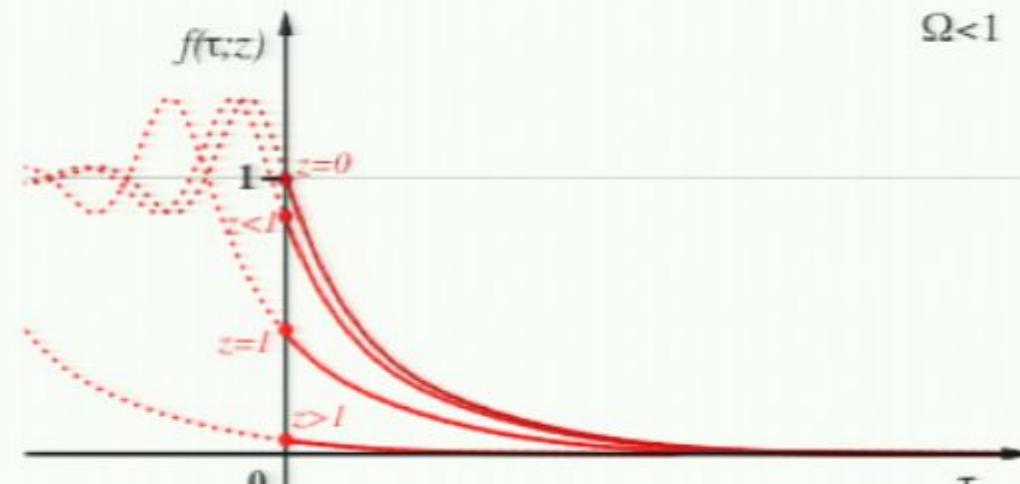
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Phase transition:

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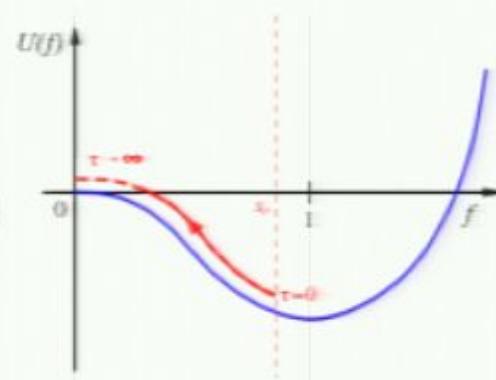
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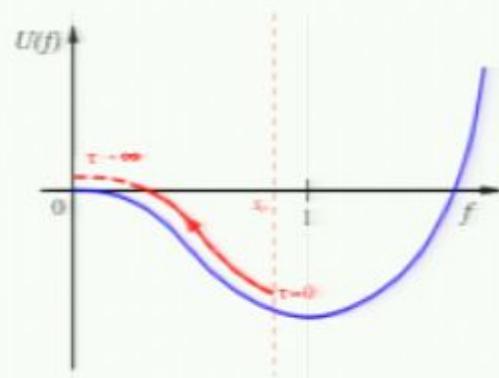
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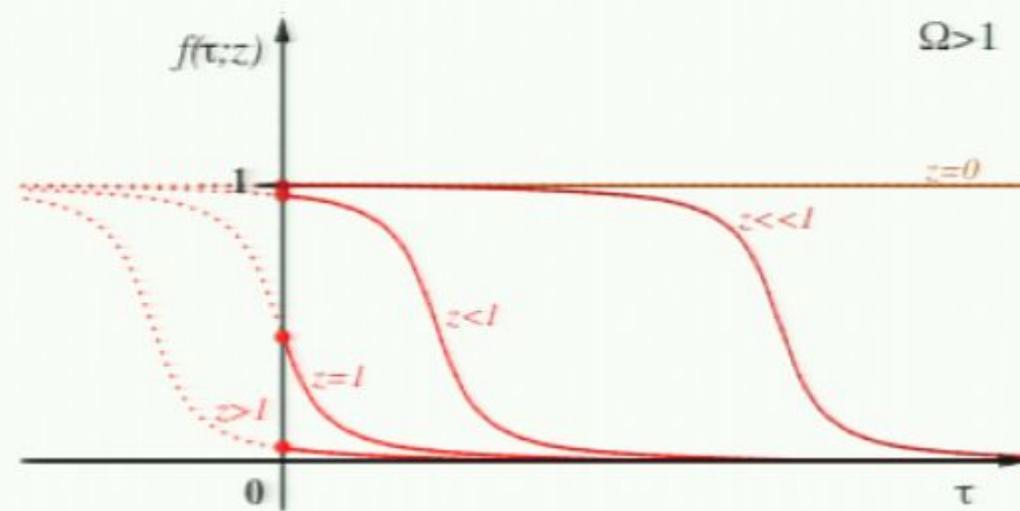
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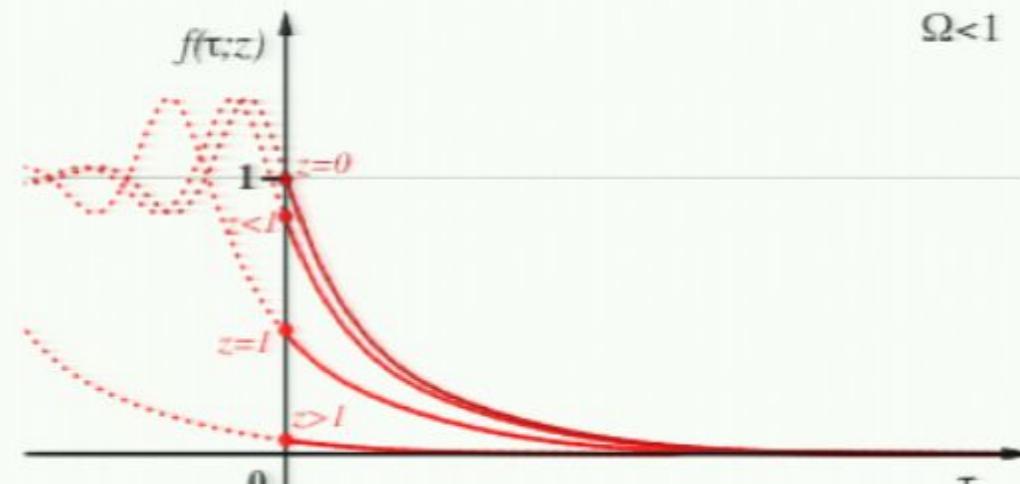
$$\Omega > 1$$



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partial Extinction

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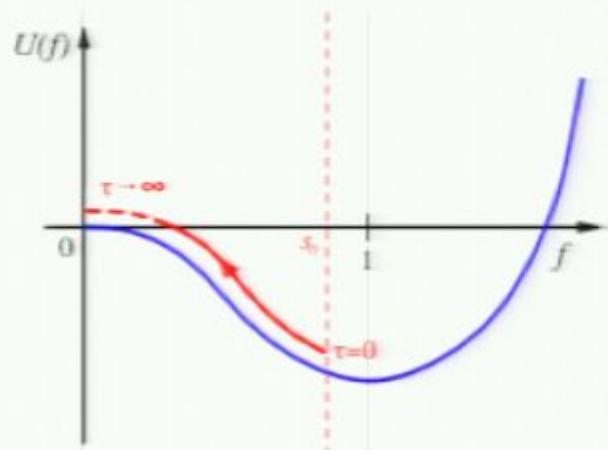
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saddle point

$$\rho(V, \tau) \approx \frac{1}{\sqrt{2\pi|S''(z_0)|}} e^{-S(z_0)} \equiv \mathcal{N} e^{-S(z_0)},$$

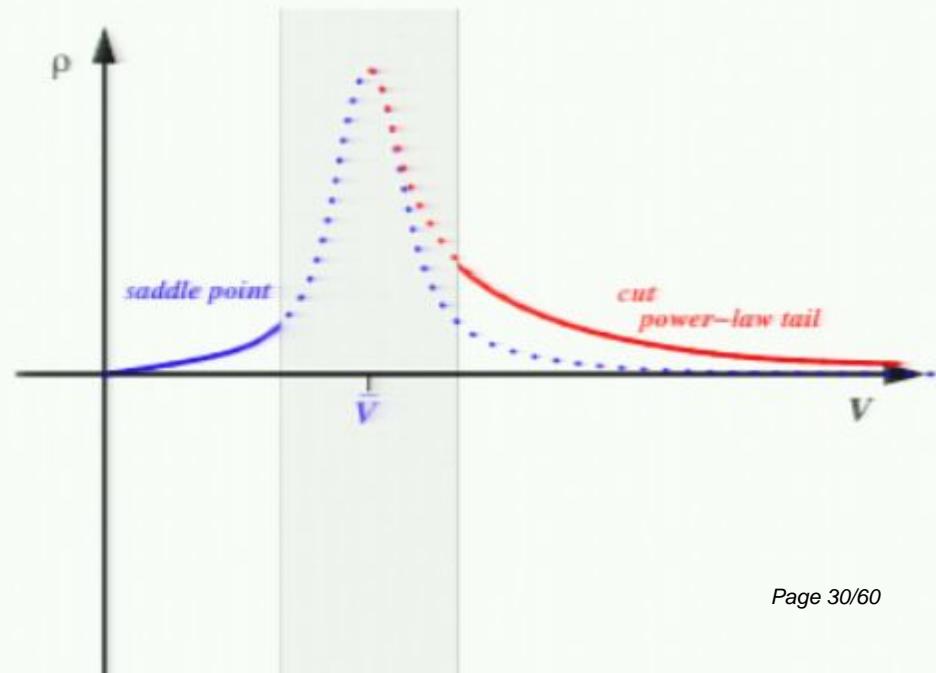
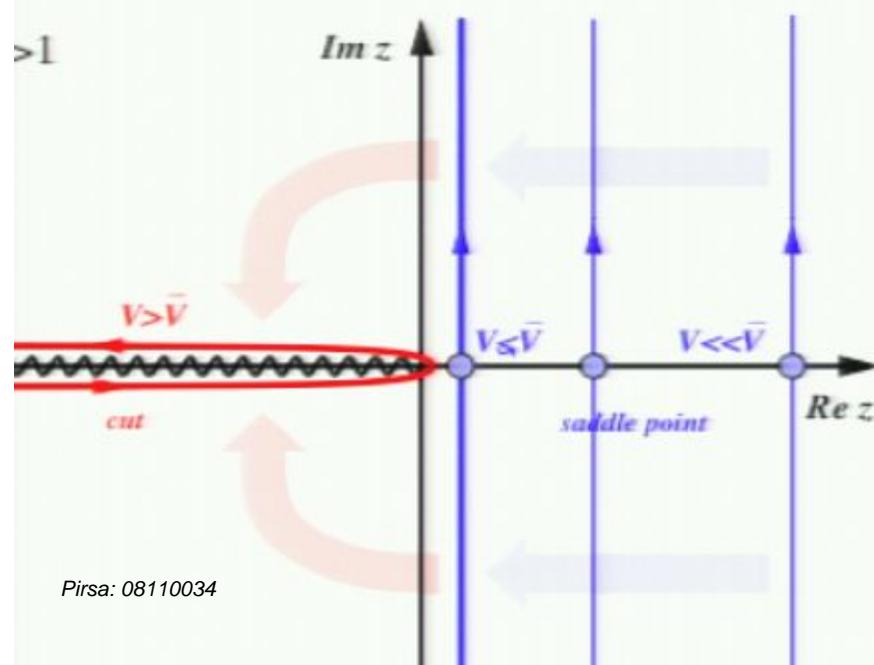
here $S(z) \simeq (\tau + \tau_0)^2 - zV \simeq \left(\tau - \frac{1}{\omega_-} \log z\right)^2 - Vz$

idle point $z_0 \approx \frac{1}{2\omega_- V} \left[\tau - \frac{1}{\omega_-} \log \left(\frac{2\omega_- V}{\tau} \right) \right]$

$$\rho(V, \tau) \approx \mathcal{N} e^{-\frac{1}{4}\Omega\left(1+\sqrt{1-\frac{1}{\Omega}}\right)^2 \left[\log\left(\frac{V}{\bar{V}}\right)\right]^2} = \mathcal{N} e^{-\Omega\left[\frac{3N}{2}\left(1+\sqrt{1-\frac{1}{\Omega}}\right) - 3N_c\right]^2}, \quad V \lesssim \bar{V},$$

where

$$\bar{V} \equiv e^{\omega_- \tau} = e^{3N_c \frac{2}{1+\sqrt{1-1/\Omega}}}.$$



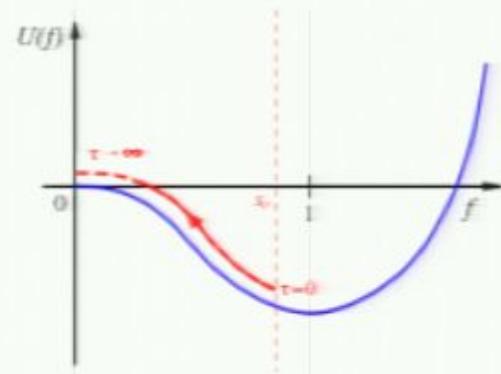
e Phase Transition

different behaviors for

$$\Omega \gtrless 1$$

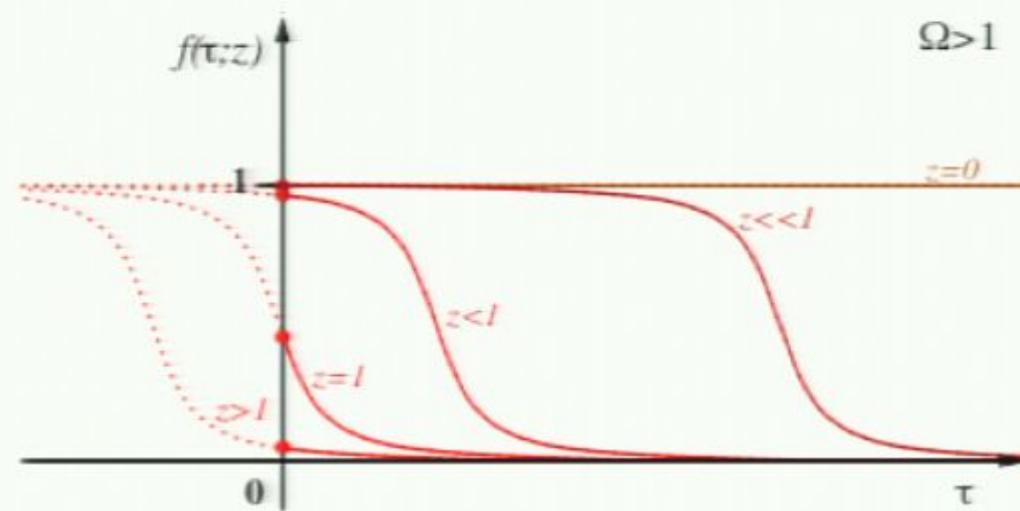
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total Extinction

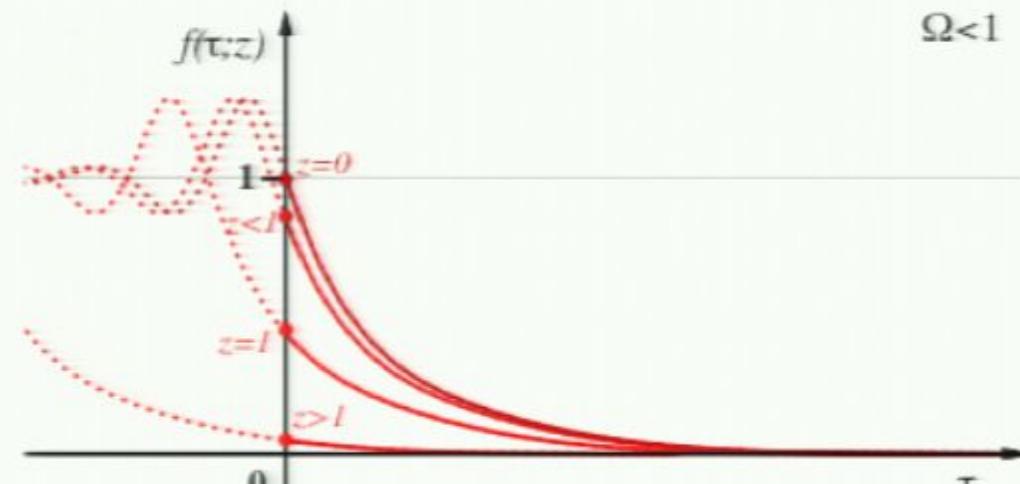
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$$\Omega > 1$$

Extinction

$$\Omega < 1$$



$$\Omega < 1$$

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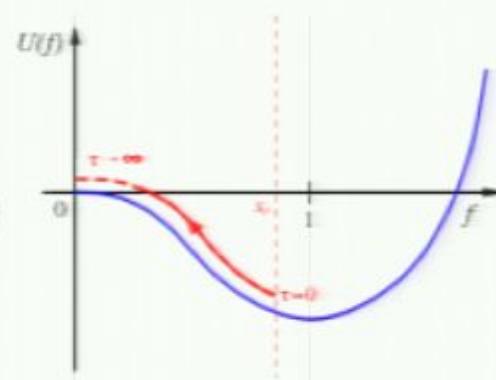
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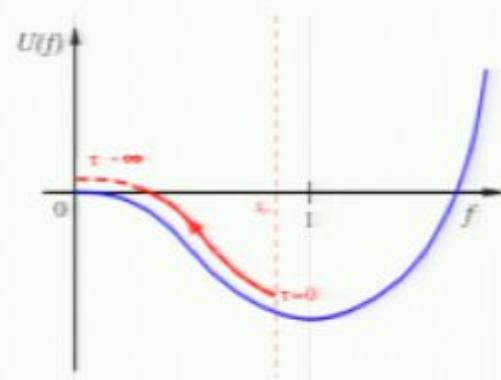
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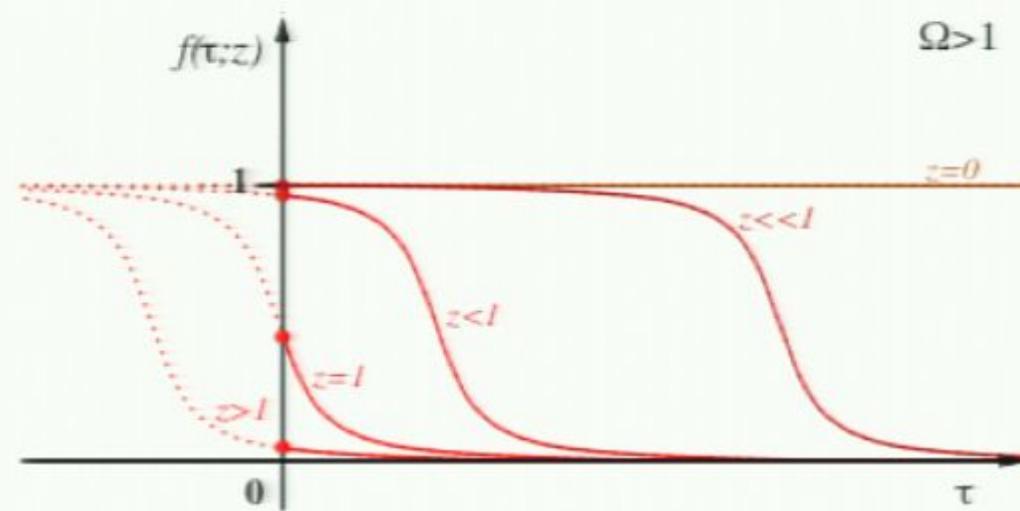
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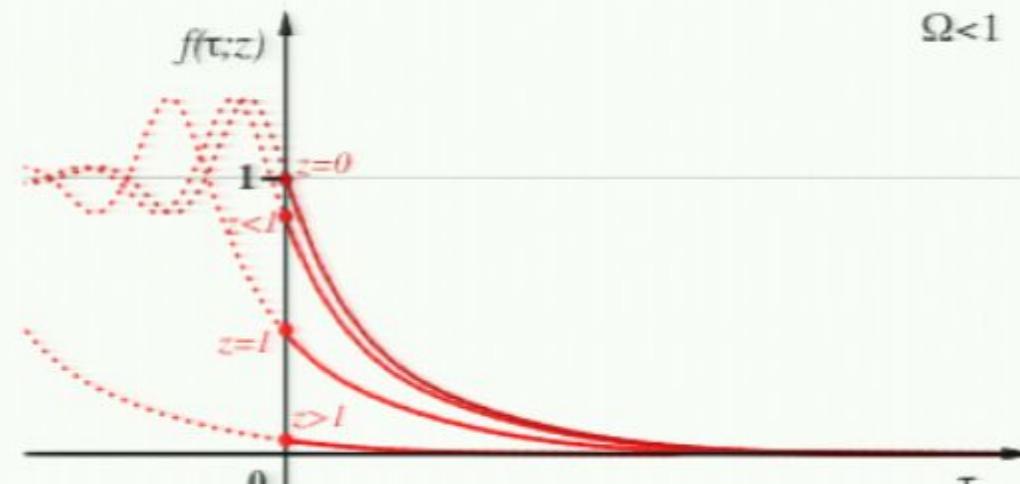
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$$\Omega > 1$$

Extinction

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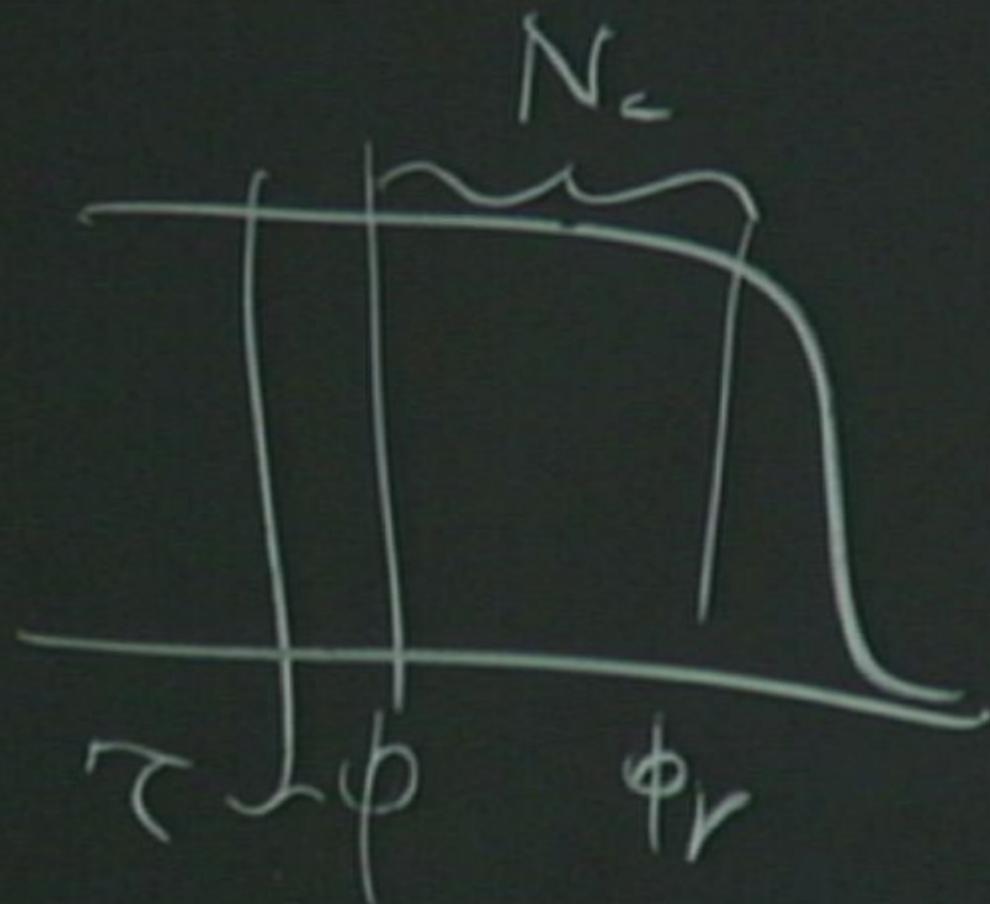
$$\Omega < 1$$

Phase transition:

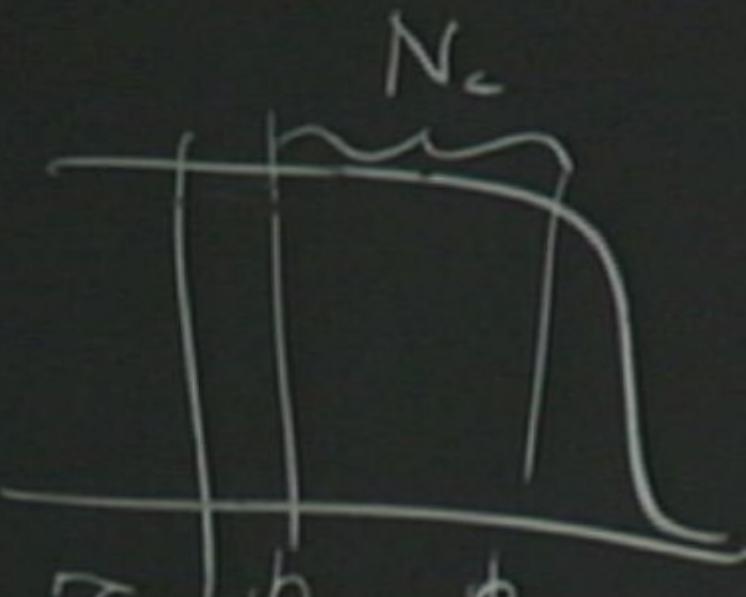
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$$V = e^{3N}$$



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$$\left(\frac{\Delta\phi}{H}\right)^2$$

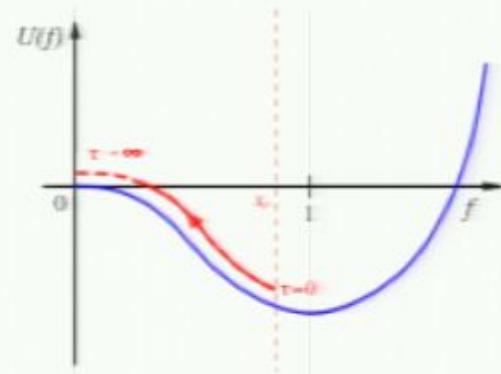
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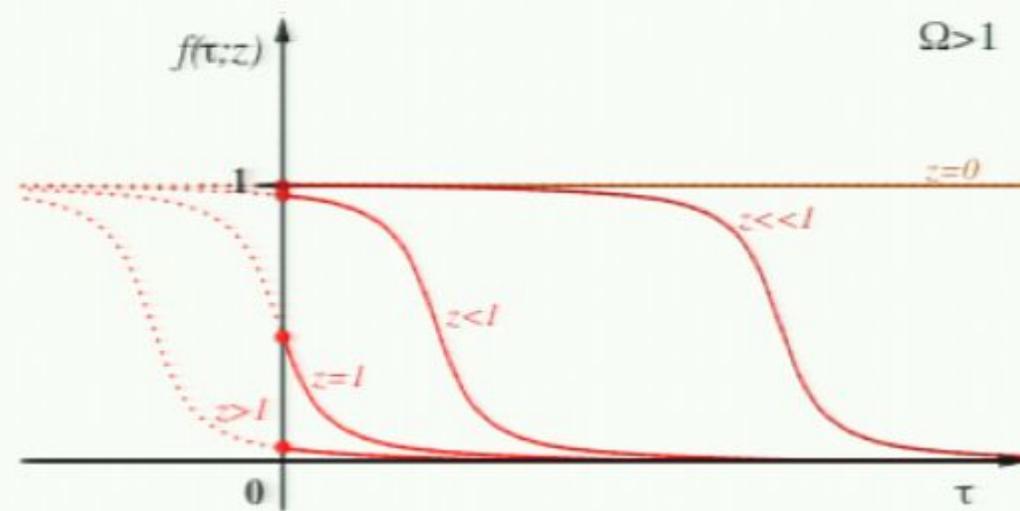
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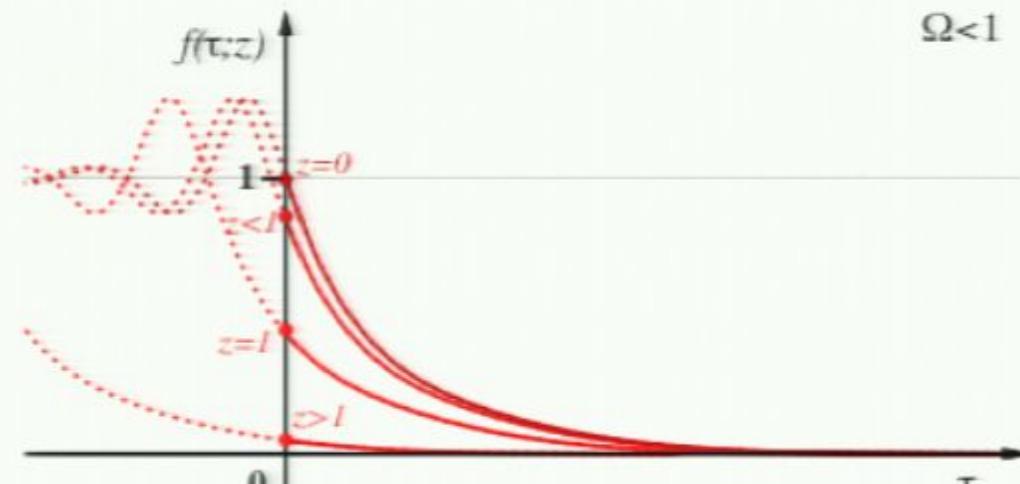
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> 1 : Approaching the Transition from above:

Solutions in two regimes:

$$\ddot{f}(\tau; z) - 2\sqrt{\Omega}\dot{f}(\tau; z) + f(\tau; z) \log[f(\tau; z)] = 0$$

Gaussian

$$f \simeq 0 \quad \Rightarrow \quad f \approx f_g = e^{-\frac{(\tau+\tau_1)^2}{4}} \quad |\tau + \tau_1| \gg 1$$

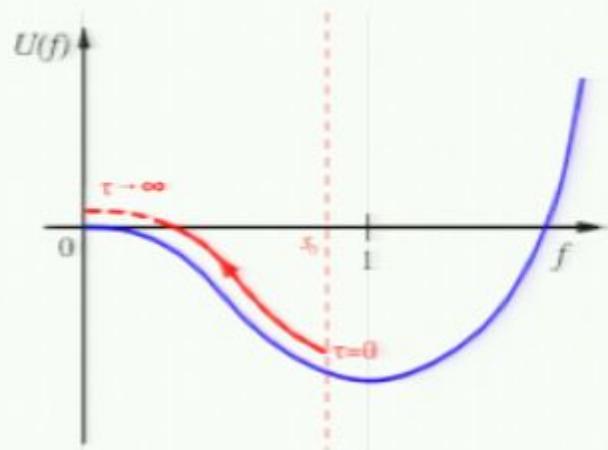
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Cut at

$$z = 0. \quad \Rightarrow \quad f_{\text{lin}}(\tau; z) \simeq 1 - z e^{\omega_- \tau} - \sigma z^{\omega_+^2} e^{\omega_+ \tau}$$



Saddle point or integral along the cut: $\rho(V, \tau) = \frac{1}{2\pi i} \int_{0^+ - i\infty}^{0^+ + i\infty} dz f(\tau; z) e^{zV}$

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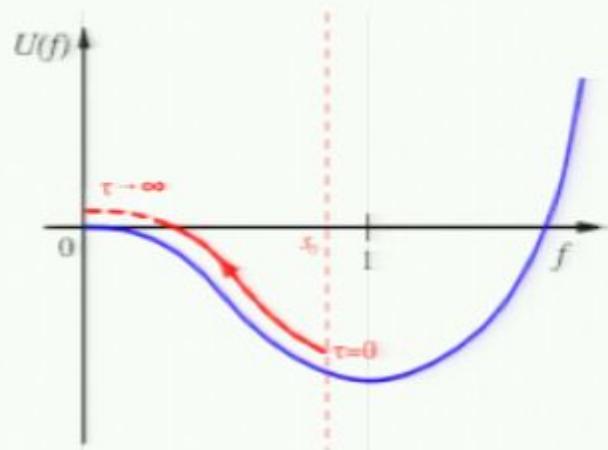
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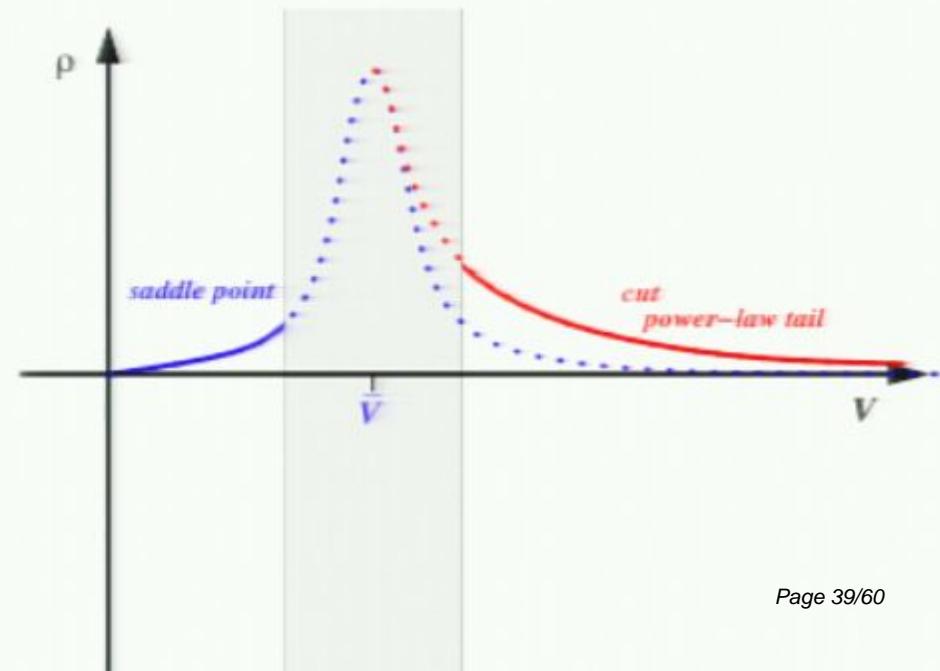
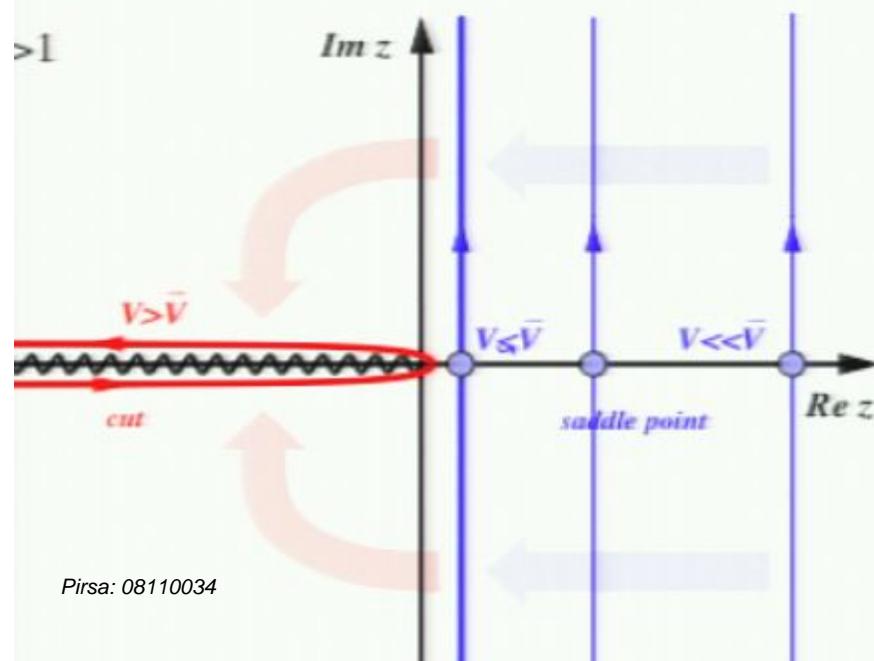
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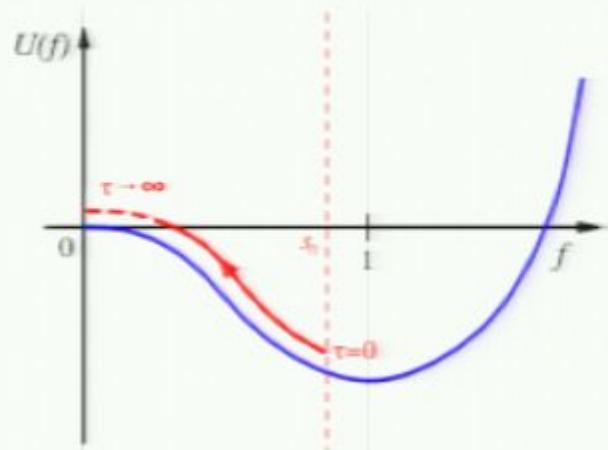
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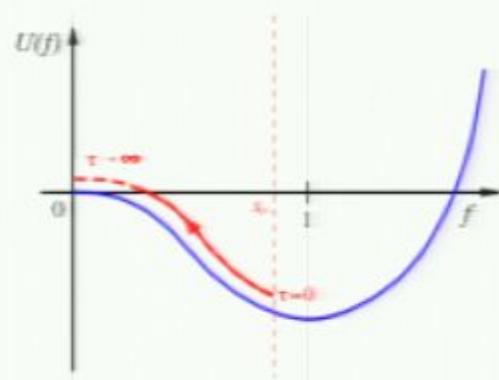
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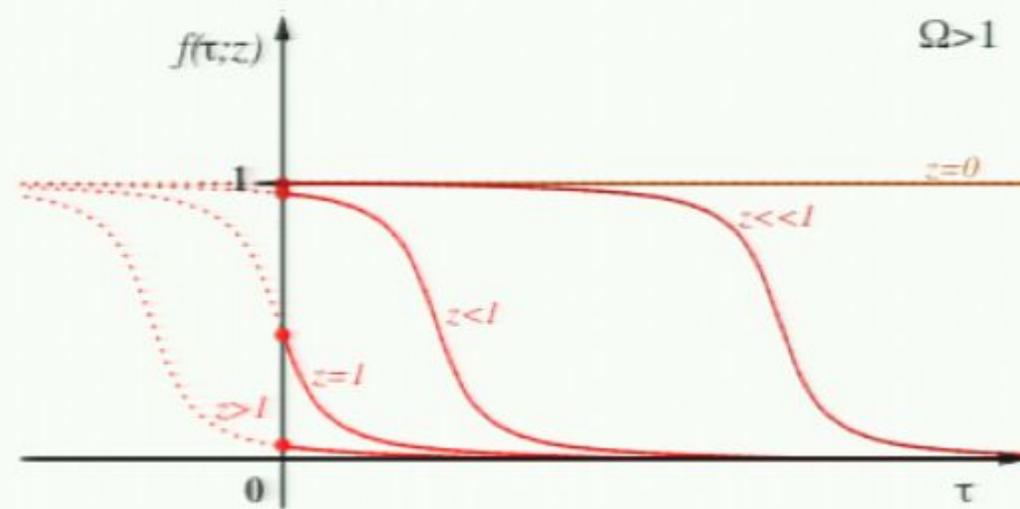
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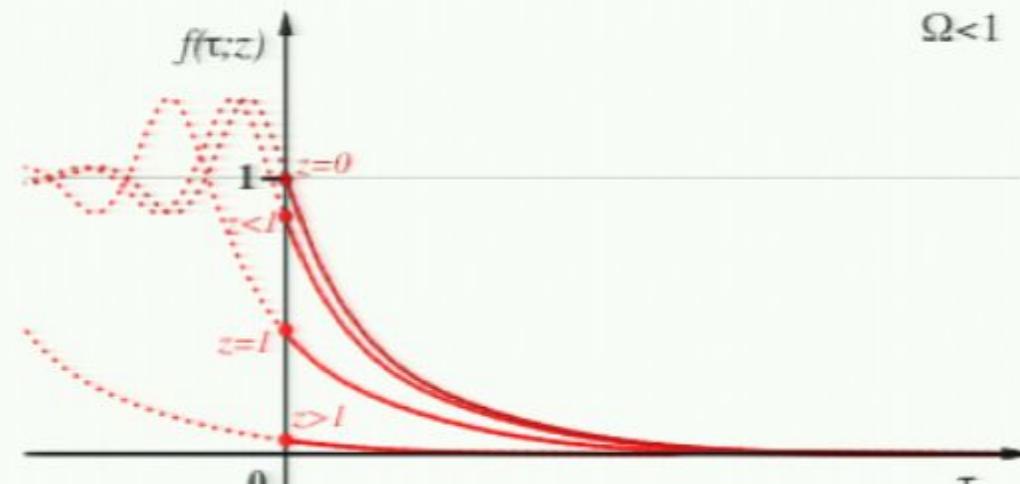
$$\Omega > 1$$



$$\Omega > 1$$

Extinction

$$\Omega < 1$$



$$\Omega < 1$$

Phase transition:

$$\Omega \equiv \frac{2\pi^2}{3} \frac{\dot{\phi}^2}{H^4} < 1$$

> 1 : Approaching the Transition from above:

Solutions in two regimes:

$$\ddot{f}(\tau; z) - 2\sqrt{\Omega}\dot{f}(\tau; z) + f(\tau; z) \log[f(\tau; z)] = 0$$

Gaussian

$$f \simeq 0 \quad \Rightarrow \quad f \approx f_g = e^{-\frac{(\tau+\tau_1)^2}{4}} \quad |\tau + \tau_1| \gg 1$$

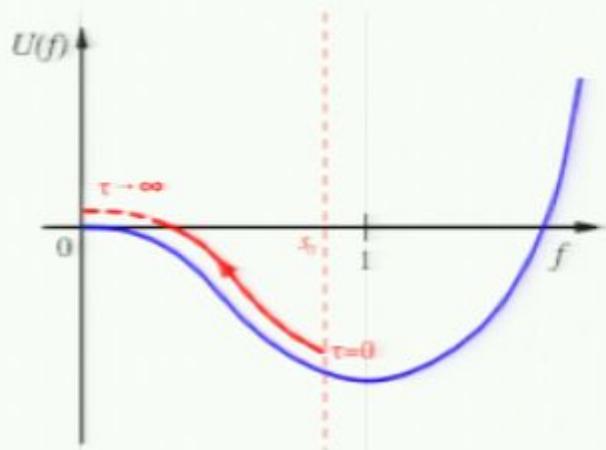
Linear

$$f \simeq 1 \quad \Rightarrow \quad f_{\text{lin}}(\tau; z) = 1 - e^{\omega_- (\tau + \tau_0)} - \sigma e^{\omega_+ (\tau + \tau_0)} \\ \omega_{\pm} \equiv \sqrt{\Omega} \pm \sqrt{\Omega - 1} \quad \tau + \tau_0 \ll -1$$

Boundary Condition: $f_{\text{lin}}(0; z) = 1 - e^{\omega_- \tau_0} - \sigma e^{\omega_+ \tau_0} = e^{-z} \Rightarrow \tau_0 \approx \log(z)/\omega_-$

at

$$z = 0. \quad \Rightarrow \quad f_{\text{lin}}(\tau; z) \simeq 1 - z e^{\omega_- \tau} - \sigma z^{\omega_+^2} e^{\omega_+ \tau}$$



Saddle point or integral along the cut: $\rho(V, \tau) = \frac{1}{2\pi i} \int_{0^+ - i\infty}^{0^+ + i\infty} dz f(\tau; z) e^{zV}$

> 1 : Approaching the Transition

$$\rho(V, \tau) = \frac{1}{2\pi i} \int_{0^+ - i\infty}^{0^+ + i\infty} dz f(\tau; z) e^{zV}$$

saddle point

$$\rho(V, \tau) \approx \frac{1}{\sqrt{2\pi|S''(z_0)|}} e^{-S(z_0)} \equiv \mathcal{N} e^{-S(z_0)},$$

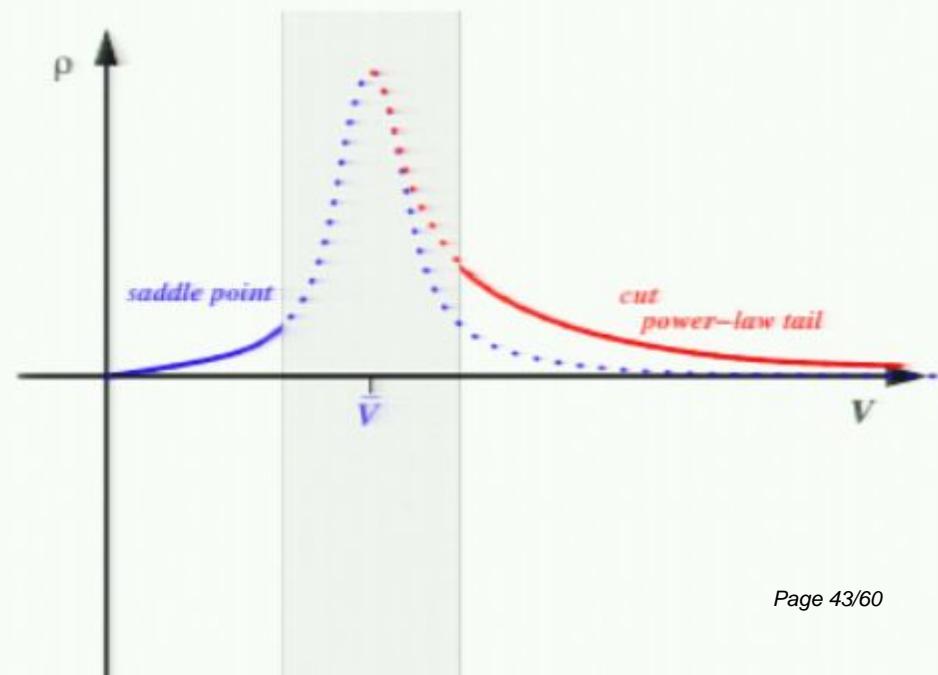
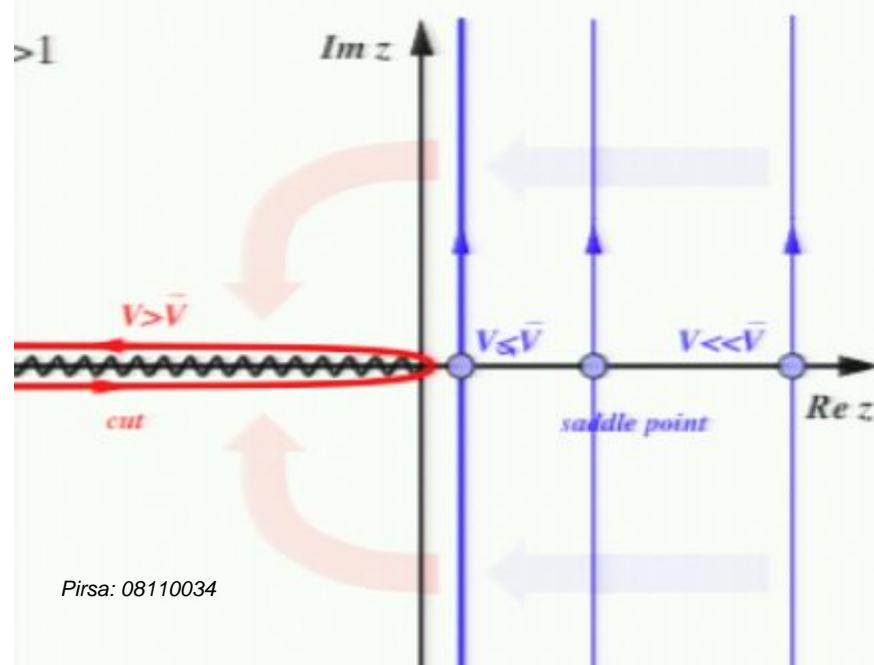
here $S(z) \simeq (\tau + \tau_0)^2 - zV \simeq \left(\tau - \frac{1}{\omega_-} \log z\right)^2 - Vz$

idle point $z_0 \approx \frac{1}{2\omega_- V} \left[\tau - \frac{1}{\omega_-} \log \left(\frac{2\omega_- V}{\tau} \right) \right]$

$$\rho(V, \tau) \approx \mathcal{N} e^{-\frac{1}{4}\Omega\left(1+\sqrt{1-\frac{1}{\Omega}}\right)^2 \left[\log\left(\frac{V}{V}\right)\right]^2} = \mathcal{N} e^{-\Omega\left[\frac{3N}{2}\left(1+\sqrt{1-\frac{1}{\Omega}}\right)-3N_c\right]^2}, \quad V \lesssim \bar{V},$$

where

$$\bar{V} \equiv e^{\omega_- \tau} = e^{3N_c \frac{2}{1+\sqrt{1-1/\Omega}}}.$$



> 1 : Approaching the Transition

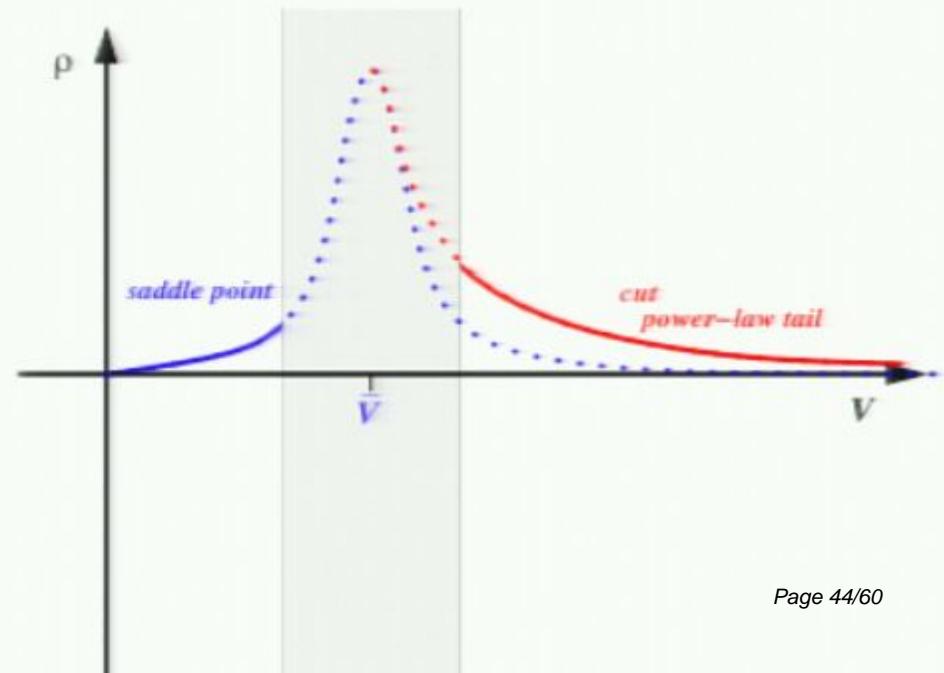
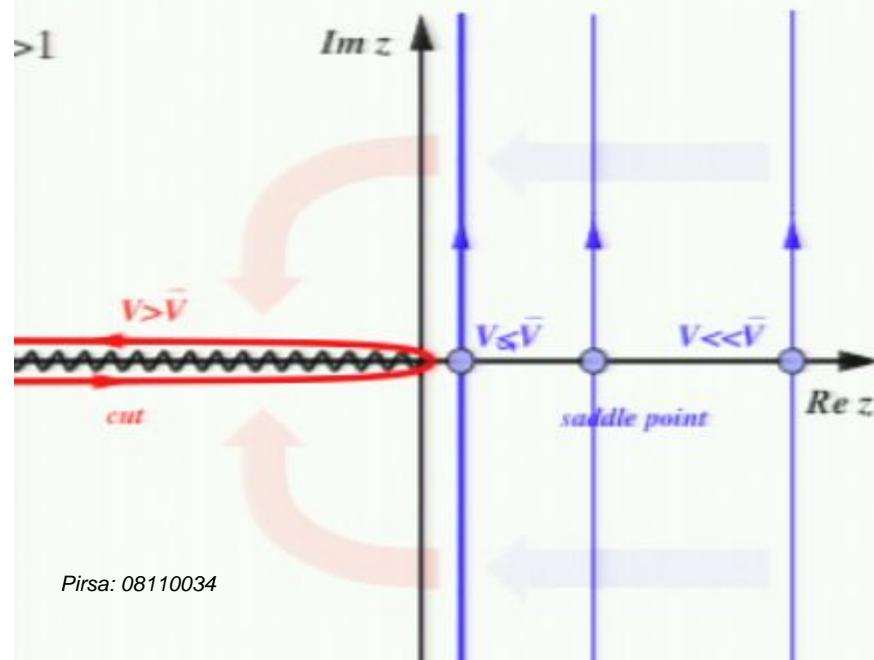
$$\rho(V, \tau) = \frac{1}{2\pi i} \int_{0^+ - i\infty}^{0^+ + i\infty} dz f(\tau; z) e^{zV}$$

long the cut $\rho(V, \tau) = \frac{1}{2\pi i} \int_0^{+\infty} d|z| 2i \operatorname{Im}[f(\tau; -|z|)] e^{-V|z|}$

where $\operatorname{Im}[f_{\text{lin}}(\tau; z)]_{\text{cut}} \sim e^{\omega_+ \tau} z^{\frac{\omega_+}{\omega_-}}$

integral in linear regime: $\rho(V, \tau) \sim \left(\frac{\bar{V}}{V}\right)^{\frac{\omega_+}{\omega_-} + 1}$ for $V \gtrsim \bar{V} \frac{\omega_+}{\omega_-}$
 $\rho(V, \tau) \sim e^{-\Omega\left[\frac{3N}{2}\left(1 + \sqrt{1 - \frac{1}{\Omega}}\right) - 3N_c\right]^2}$, for $V \lesssim \bar{V}$,

here $\bar{V} \equiv e^{\omega_+ \tau} = e^{\frac{3N_c}{1 + \sqrt{1 - 1/\Omega}}}$. Agrees with classical limit $\Omega \gg 1$, and momenta



$\lesssim 1$: Inside Eternal Inflation: $\Omega = 1 - \epsilon$

solution in two regimes:

Gaussian

$$f \simeq 0 \quad \Rightarrow \quad f_g(\tau; z) = e^{-\frac{(\tau + \tau_0)^2}{4}}, \quad \tau + \tau_0 \gg 1,$$

Linear

$$f \simeq 1 \quad \Rightarrow \quad$$

$$f(\tau; z) = 1 - \sigma e^{\sqrt{\Omega}(\tau + \tau_0)} \cos(\sqrt{\Omega - 1}(\tau + \tau_0)) \approx 1 - \sigma e^{\tau + \tau_0} \cos(\sqrt{\epsilon}(\tau + \tau_0))$$

Boundary Condition:

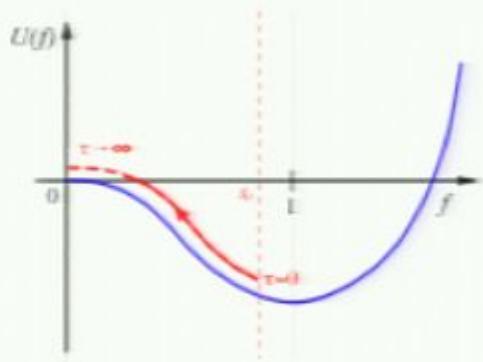
$$e^{-z} = f_{\text{lin}}(0; z) = 1 - \sigma e^{\tau_0} \cos(\sqrt{\epsilon}\tau_0) \quad \Rightarrow \quad z \approx \sigma e^{\tau_0} \cos(\sqrt{\epsilon}\tau_0)$$

at

$$0 = \frac{dz}{d\tau_0} \quad \Rightarrow \quad \tau_0 \simeq -\frac{\pi}{2\sqrt{\epsilon}} - 1 \quad \Rightarrow \quad z_{\text{cut}} \simeq -\frac{\sigma\sqrt{\epsilon}}{eV_\epsilon}, \quad V_\epsilon \equiv e^{\frac{\pi}{2\sqrt{\epsilon}}}$$

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$$\rho(V, \tau) = \frac{1}{2\pi i} \int_{0^+ - i\infty}^{0^+ + i\infty} dz f(\tau; z) e^{zV}$$



> 1 : Approaching the Transition

$$\rho(V, \tau) = \frac{1}{2\pi i} \int_{0+-i\infty}^{0++i\infty} dz f(\tau; z) e^{zV}$$

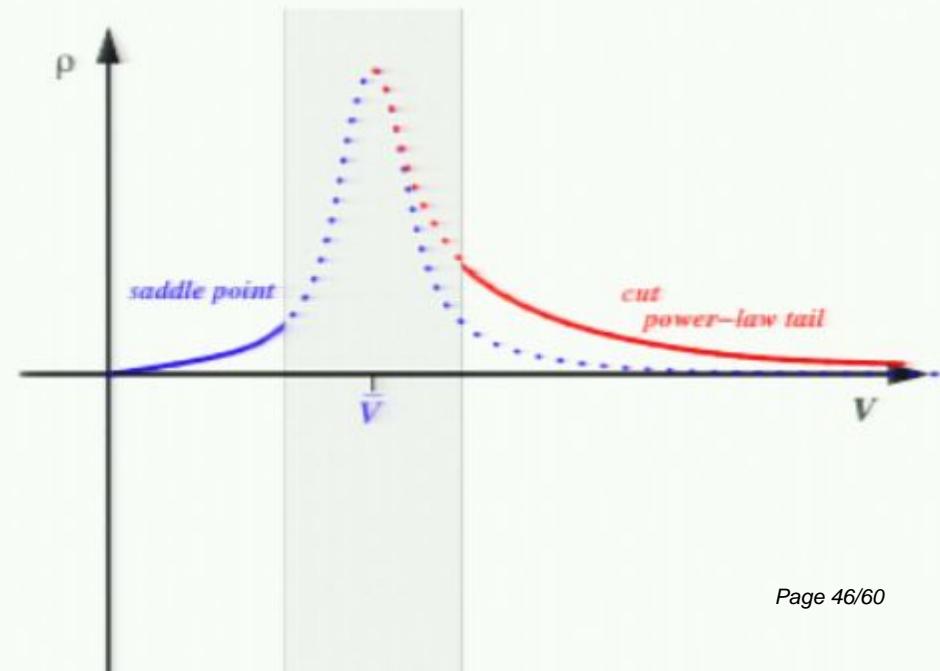
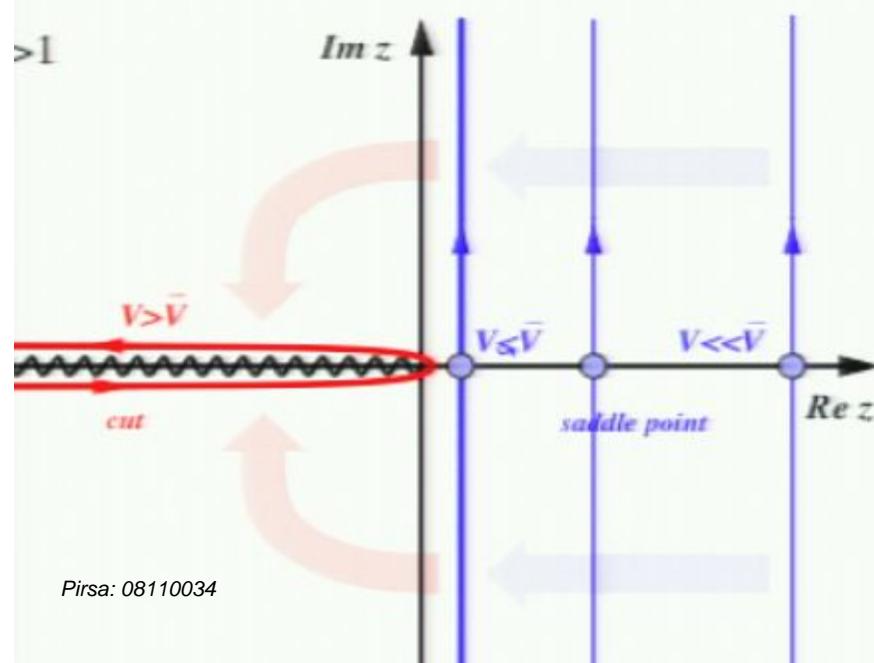
long the cut $\rho(V, \tau) = \frac{1}{2\pi i} \int_0^{+\infty} d|z| 2i \operatorname{Im}[f(\tau; -|z|)] e^{-V|z|}$

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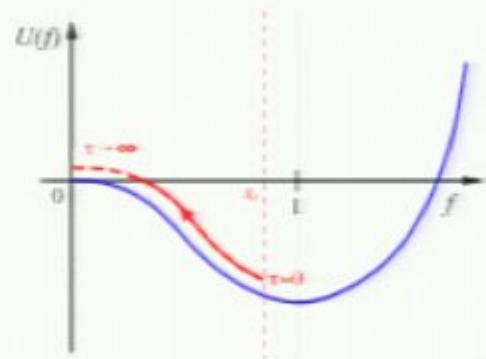
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Cut at

$$0 = \frac{dz}{d\tau_0} \quad \Rightarrow \quad \tau_0 \simeq -\frac{\pi}{2\sqrt{\epsilon}} - 1 \quad \Rightarrow \quad z_{\text{cut}} \simeq -\frac{\sigma\sqrt{\epsilon}}{eV_\epsilon}, \quad V_\epsilon \equiv e^{\frac{\pi}{2\sqrt{\epsilon}}}$$

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$\lesssim 1$: Inside Eternal Inflation

saddle point

$$\rho(V, \tau) \approx \frac{1}{\sqrt{2\pi|S''(z_0)|}} e^{-S(z_0)} \equiv \mathcal{N} e^{-S(z_0)},$$

here $S(z) = \frac{1}{4}(\tau + \tau_0)^2 - zV$

idle point $V = \frac{(\tau + \tau_0)e^{-\tau_0}}{2\sigma [\cos(\sqrt{\epsilon}\tau_0) - \sqrt{\epsilon}\sin(\sqrt{\epsilon}\tau_0)]}$

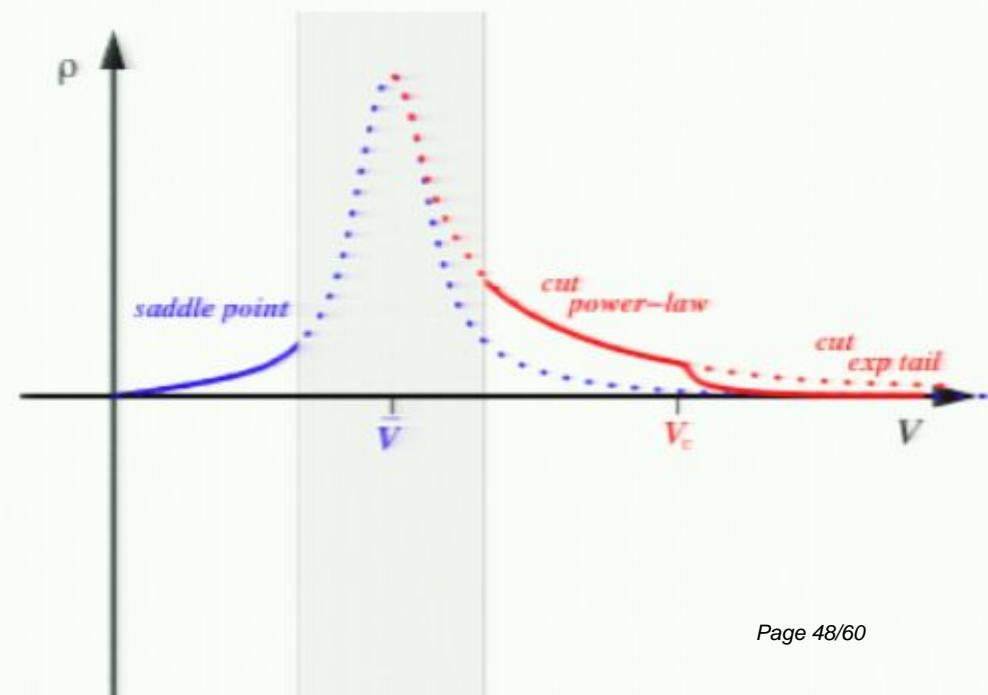
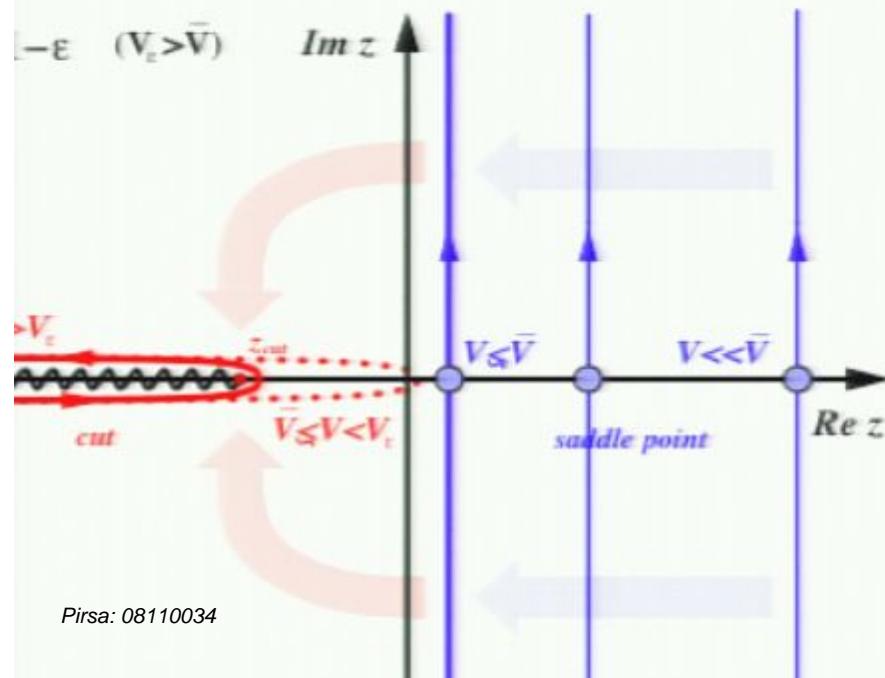
$$\rho(V, \tau) = \frac{1}{2\pi i} \int_{0^+ - i\infty}^{0^+ + i\infty} dz f(\tau; z) e^{zV}$$

where $z \approx \sigma e^{\tau_0} \cos(\sqrt{\epsilon}\tau_0)$,

as long as $\tau_0 \gtrsim -\log V_\epsilon$, we have $\tau_0 \approx -\log V$

$\rho \approx \mathcal{N} e^{-\frac{1}{4}(\tau - \log V)^2}$ for $1 \ll V \lesssim \text{Min}\{\bar{V}, V_\epsilon\}$

where $\bar{V} = e^{6N_c}$



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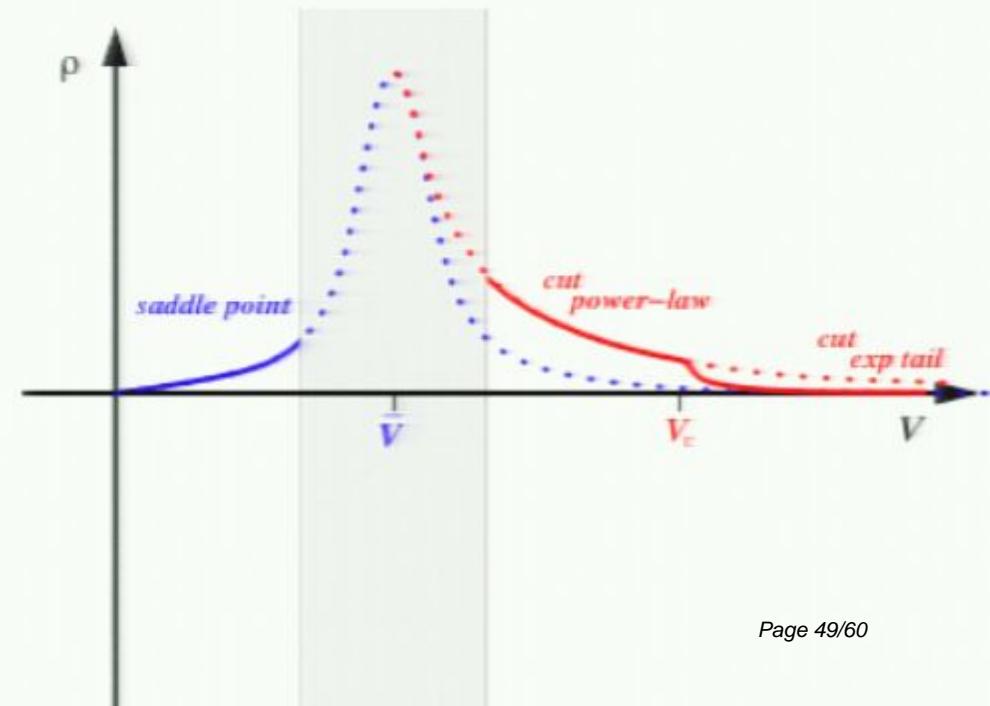
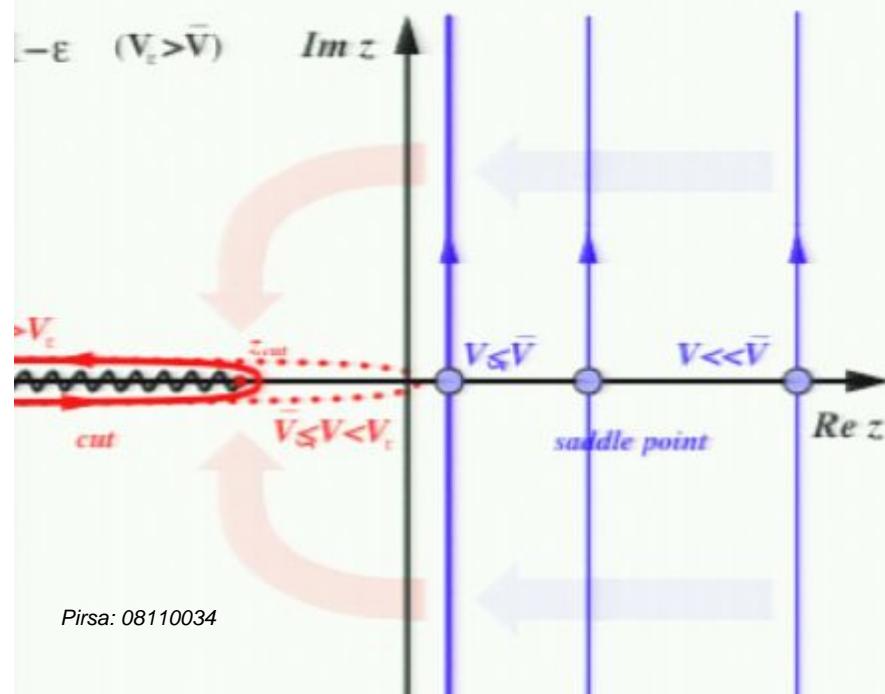
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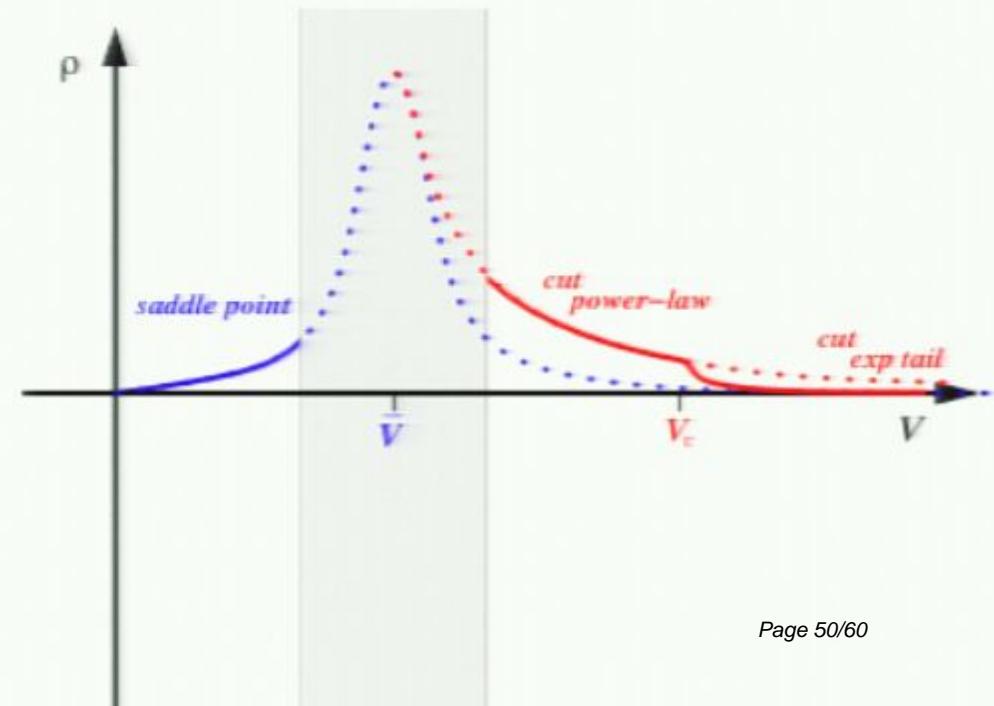
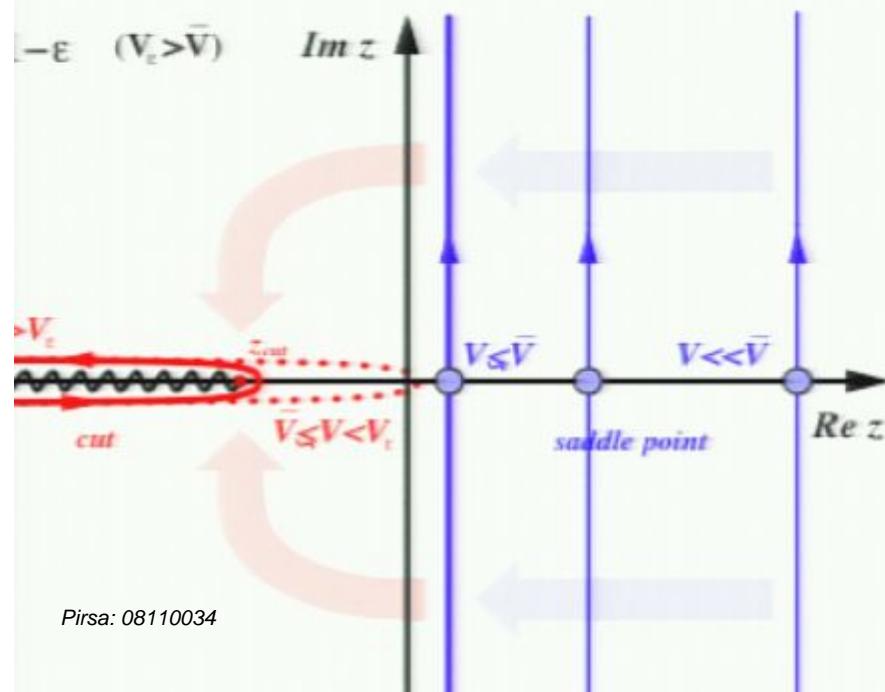
$\lesssim 1$: Inside Eternal Inflation $\bar{V} \lesssim V_\epsilon$

long the cut $\rho(V, \tau) = \frac{1}{2\pi i} \int_{z_{\text{cut}}}^{+\infty} d|z| 2i \text{Im}[f(\tau; -|z|)] e^{-V|z|}$ $z_{\text{cut}} \simeq -\frac{\sigma\sqrt{\epsilon}}{eV_\epsilon}$.
 or $V < V_\epsilon$, we have $|z| \sim 1/V$, and $f_{\text{lin}}(\tau; z) \approx 1 - \sigma e^{\tau + \tau_0} \cos(\sqrt{\epsilon}(\tau + \tau_0))$

Power law: $\rho(V, \tau) \sim \int_{z_{\text{cut}}}^{+\infty} d|z| e^\tau z e^{-V|z|} \sim \frac{\bar{V}}{V^2}, \quad \bar{V} \lesssim V \lesssim V_\epsilon$

or $V > V_\epsilon$, exponential

$$\rho(V, \tau) \sim \int_{z_{\text{cut}}}^{+\infty} d|z| e^\tau \frac{1}{V_\epsilon} e^{-V|z|} \sim \frac{\bar{V}}{V_\epsilon V} e^{-\frac{\sigma}{\epsilon}\sqrt{\epsilon}V/V_\epsilon}, \quad \bar{V} \lesssim V_\epsilon \lesssim V,$$



$\lesssim 1$: Inside Eternal Inflation $\bar{V} \gtrsim V_\epsilon$

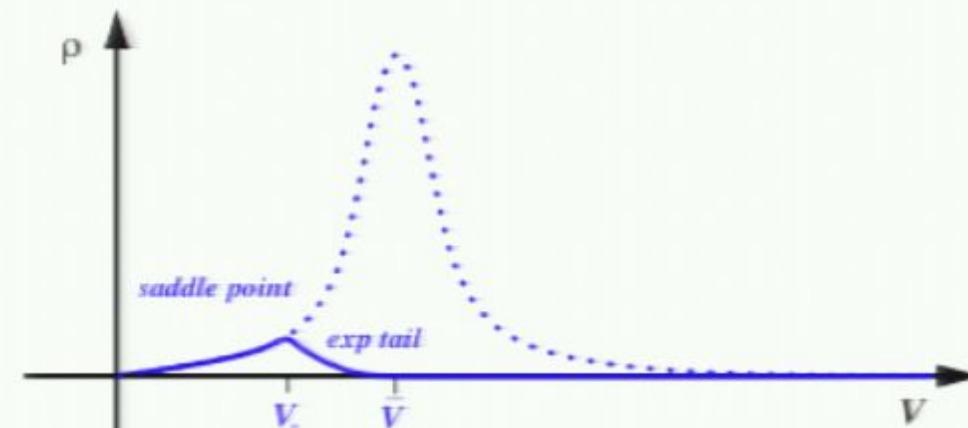
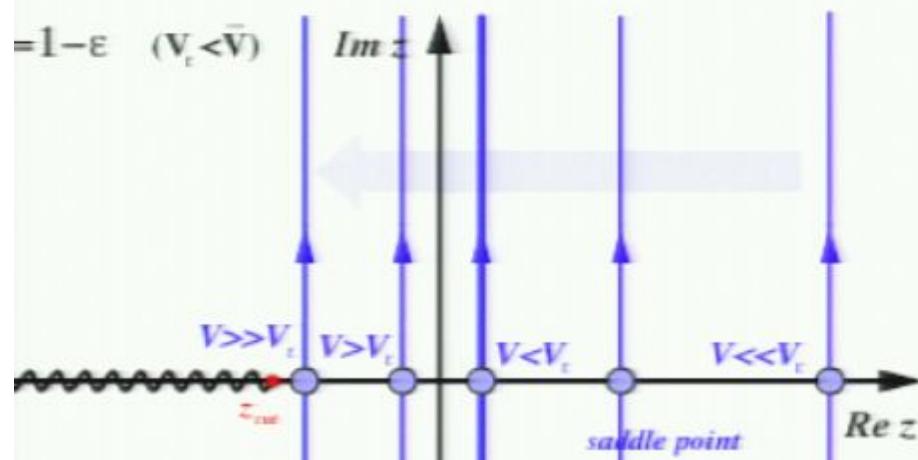
saddle point $\rho(V, \tau) \approx \frac{1}{\sqrt{2\pi|S''(z_0)|}} e^{-S(z_0)} \equiv \mathcal{N} e^{-S(z_0)}$,

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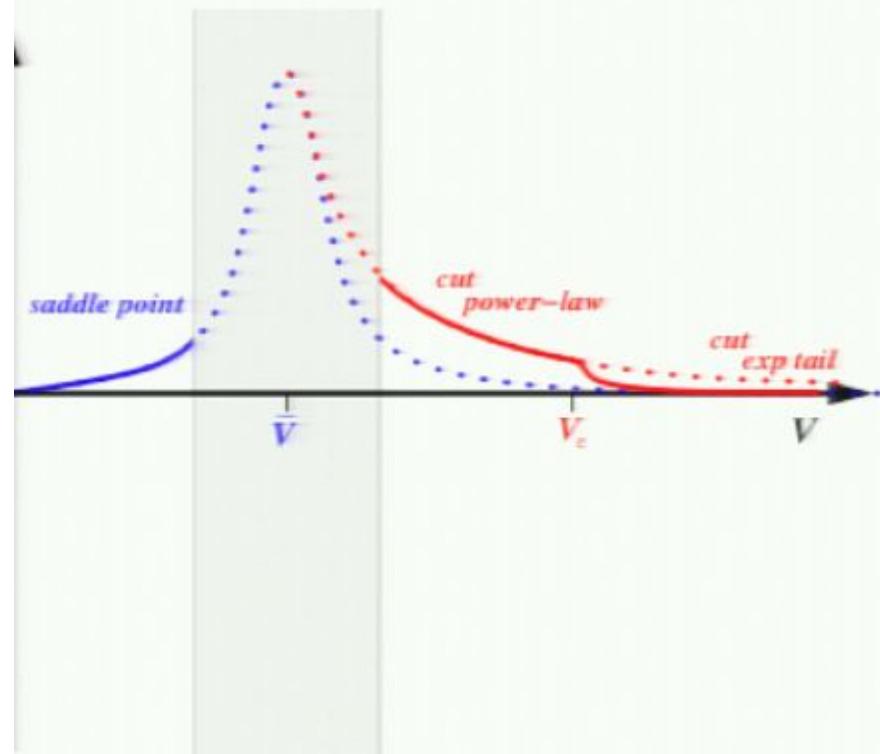
as long as $\tau_0 \gtrsim -\log V_\epsilon$, we have $\tau_0 \approx -\log V$. When $V \gtrsim V_\epsilon$, $\tau_0 \simeq -\frac{\pi}{2\sqrt{\epsilon}} - 1$

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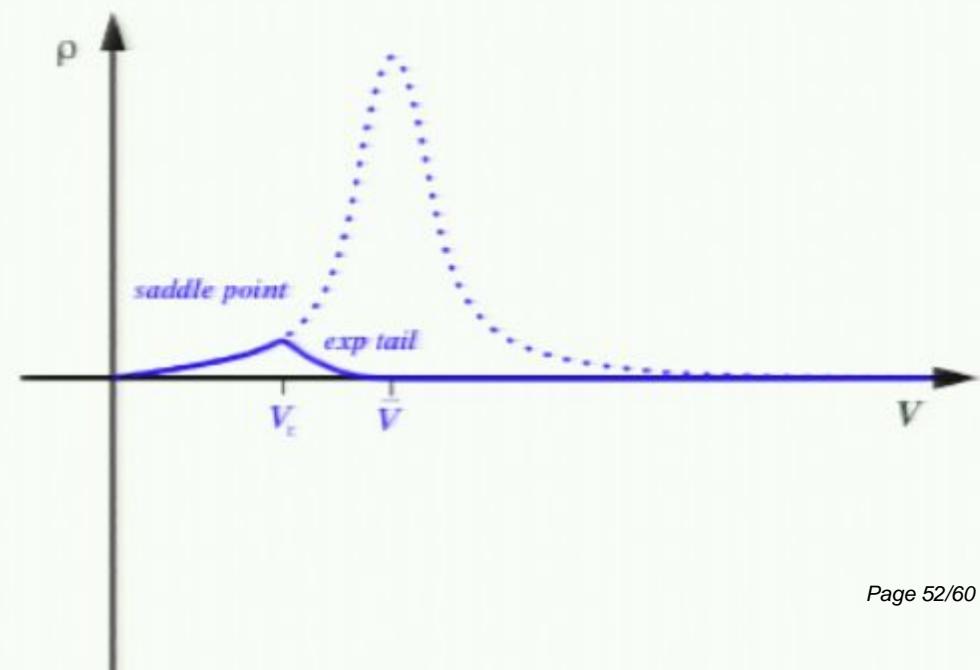
$\lesssim 1$: Inside Eternal Inflation: Summary

Gaussian, power law, and exponential



$$\bar{V} \lesssim V_\epsilon$$

$$\bar{V} \gtrsim V_\epsilon$$



$\simeq 0$: Deeply inside eternal inflation

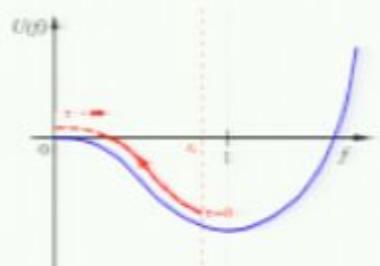
gain exact solution

$$\ddot{f}(\tau; z) - 2\cancel{\sqrt{\Omega}} \dot{f}(\tau; z) + f(\tau; z) \log [f(\tau; z)] = 0$$

$$z) = e^{\frac{1}{2} - \frac{1}{4}(\tau + \sqrt{2+4z})^2} \quad \Rightarrow$$

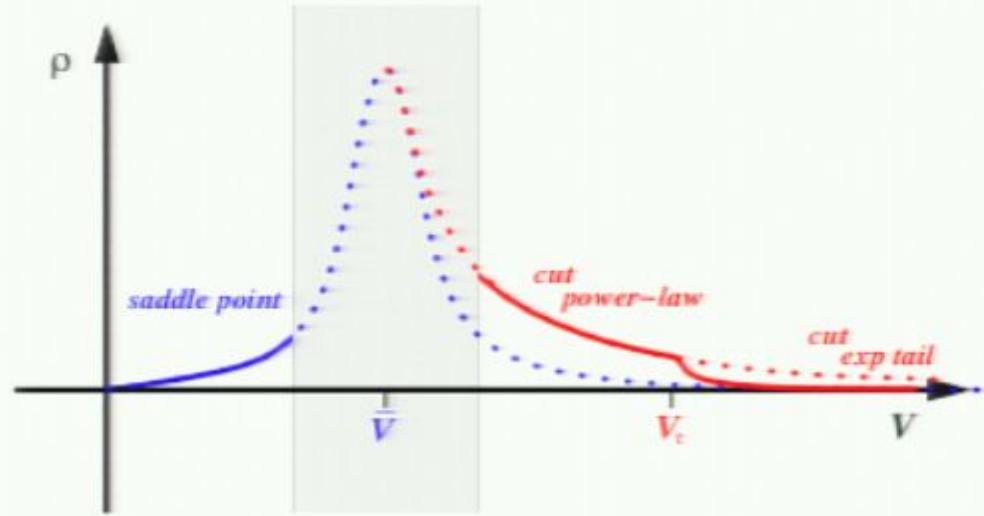
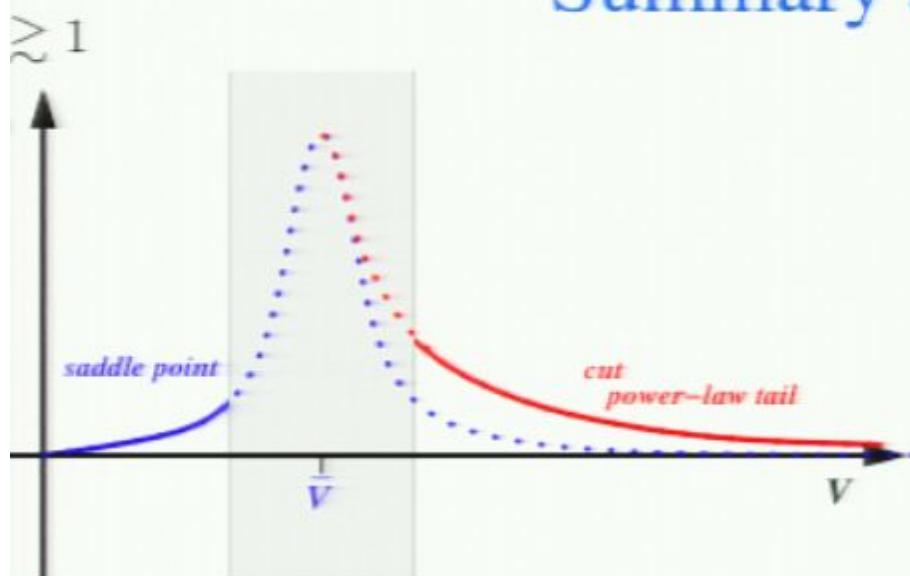
middle point

$$\rho(V, \tau) \sim e^{-\frac{V-1}{2} - \frac{V}{(V-1)} \frac{\tau^2}{4}}$$



Summary and intuition

$$\Omega \lesssim 1$$



Gaussian behavior $N \ll N_c$

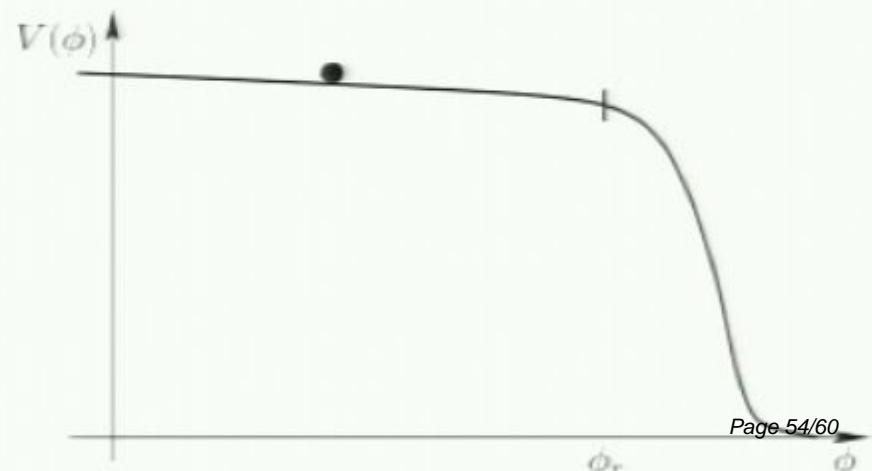
$$p(\phi, t) \sim e^{-\frac{(\bar{\phi} - \phi - \dot{\phi}t)^2}{H^3 t}} \sim e^{-\Omega \frac{(N - N_c - Ht)^2}{Ht}} \Rightarrow p \propto e^{-c(N - N_c)^2}, \quad \text{for } N \ll N_c.$$

Exponential $N \gg N_c$

$$p \propto e^{-c_1 \Delta \phi^2 / t} \propto e^{-c_2 N}$$

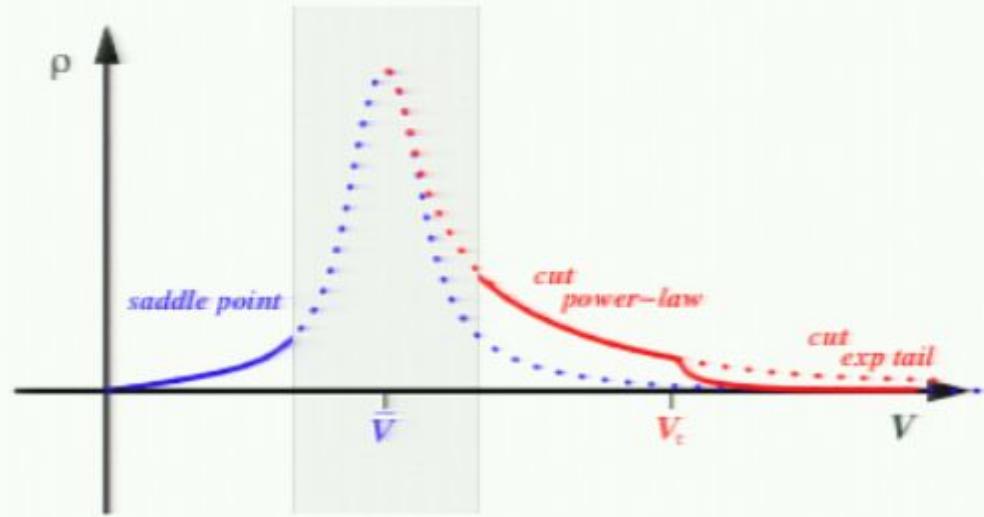
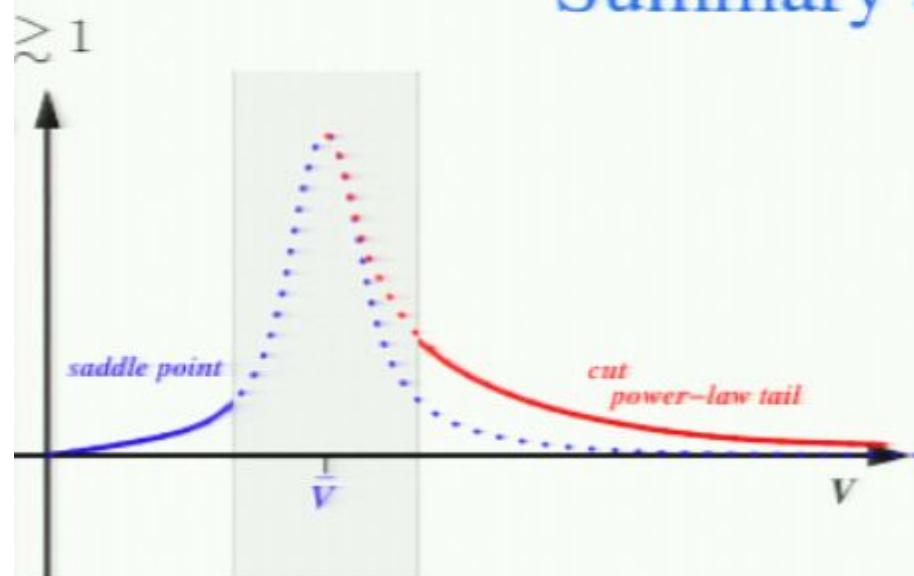
Super-Exponential: Eternal Inflation

$$p \propto \tilde{p} e^{3N}$$



Summary and intuition

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Gaussian behavior $N \ll N_c$

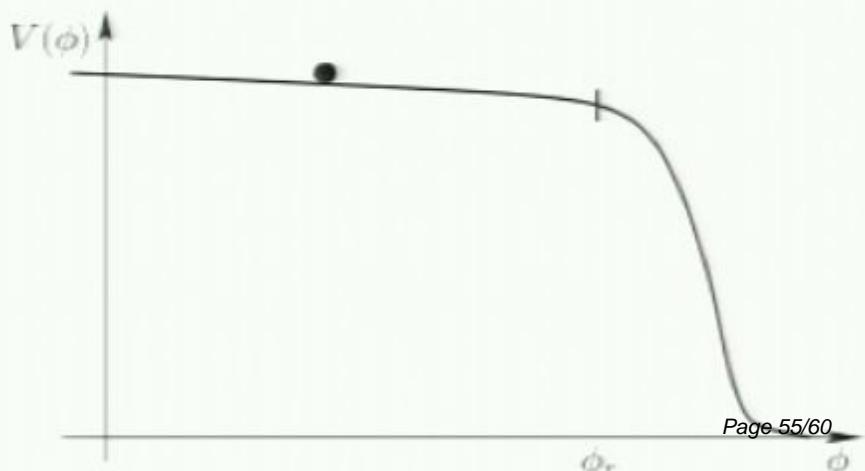
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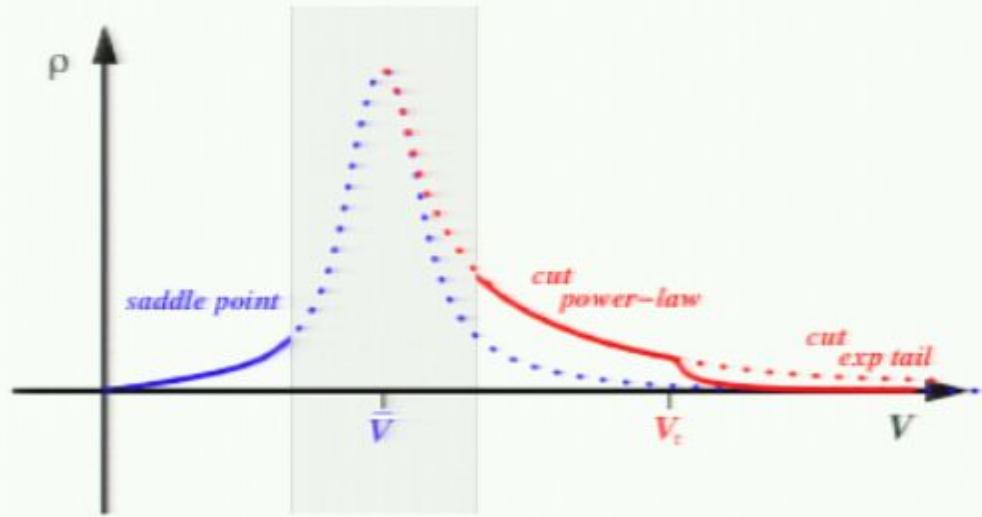
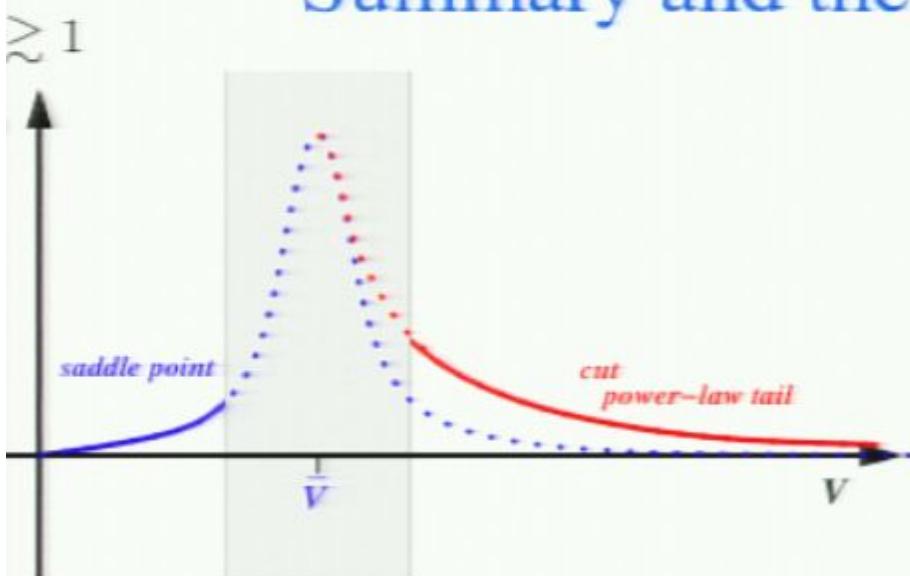
Hyper-Exponential: Eternal Inflation

$$p \propto \tilde{p} e^{3N}$$



Summary and the de Sitter Entropy

$$\Omega \lesssim 1$$



entropy and classical number of e-foldings

$$\frac{S_{ds}}{V_c} = \pi M_{Pl}^2 \frac{dH^{-2}}{Hdt} = 8\pi \frac{\dot{\phi}^2}{H^4} \quad \Rightarrow \quad N_e \lesssim \frac{S_{ds}}{12}$$

$$S_{ds} = \pi \frac{M_{Pl}^2}{H^2}$$

N. Arkani-Hamed et al.
JHEP 0705:055 2007

every finite volume realization

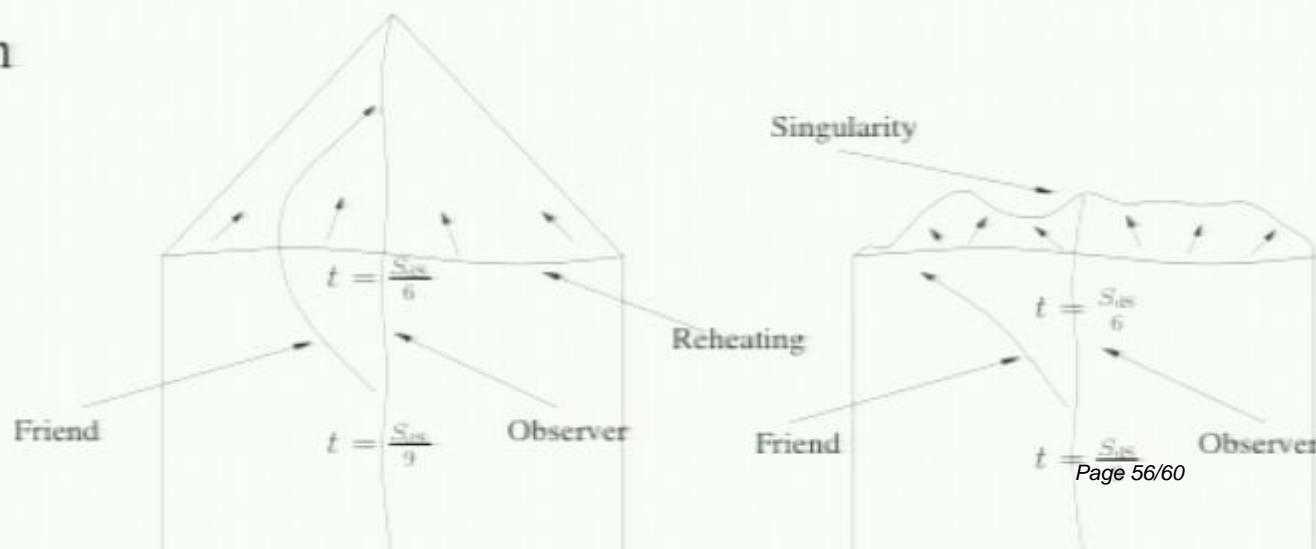
$$V > e^{\frac{S_{ds}}{2}} \quad \text{and} \quad V < e^{-\alpha S_{ds}}$$

$$\text{Finite Realization} < e^{\frac{S_{ds}}{2}}$$

Check for de Sitter entropy

Pirsa: 08110034

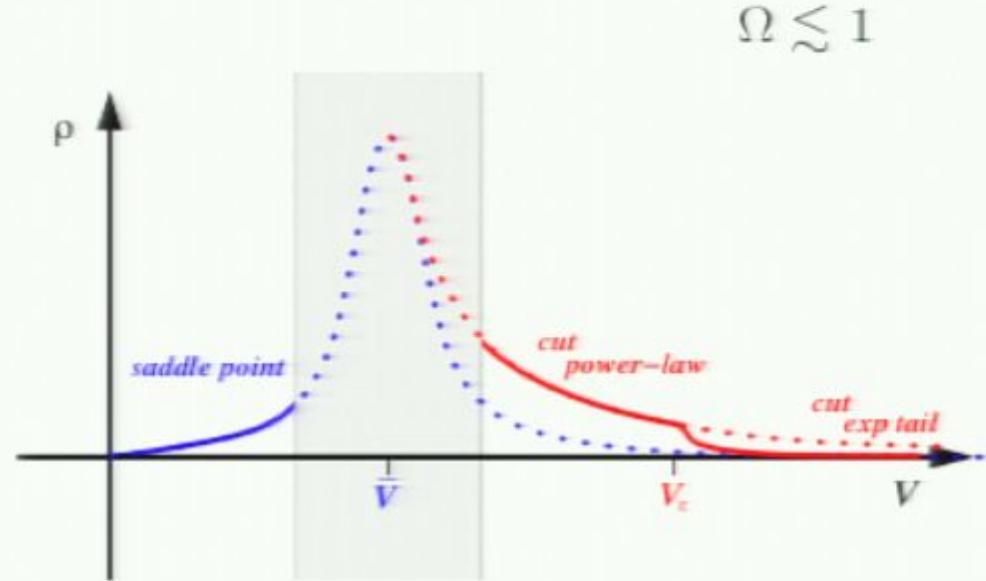
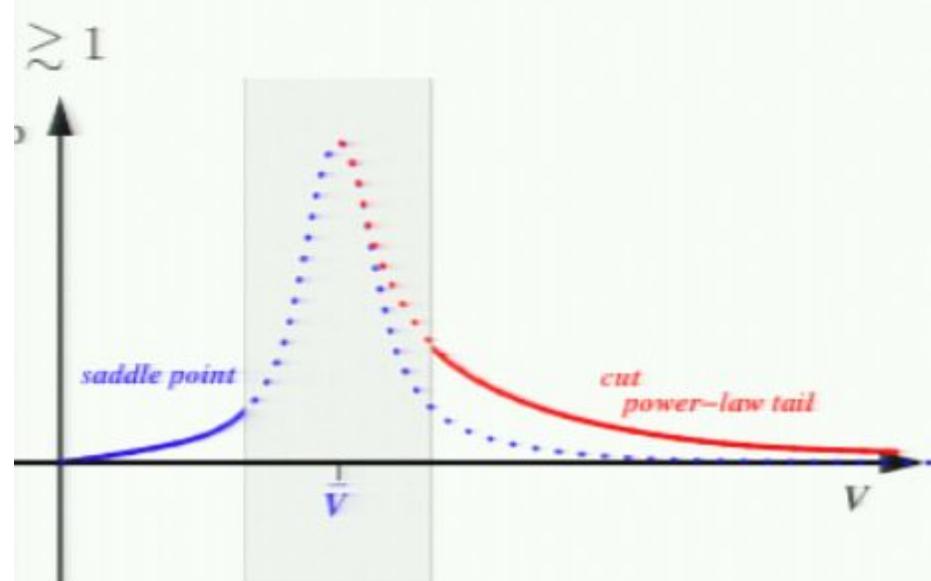
void Xerox Paradox)



Conclusions

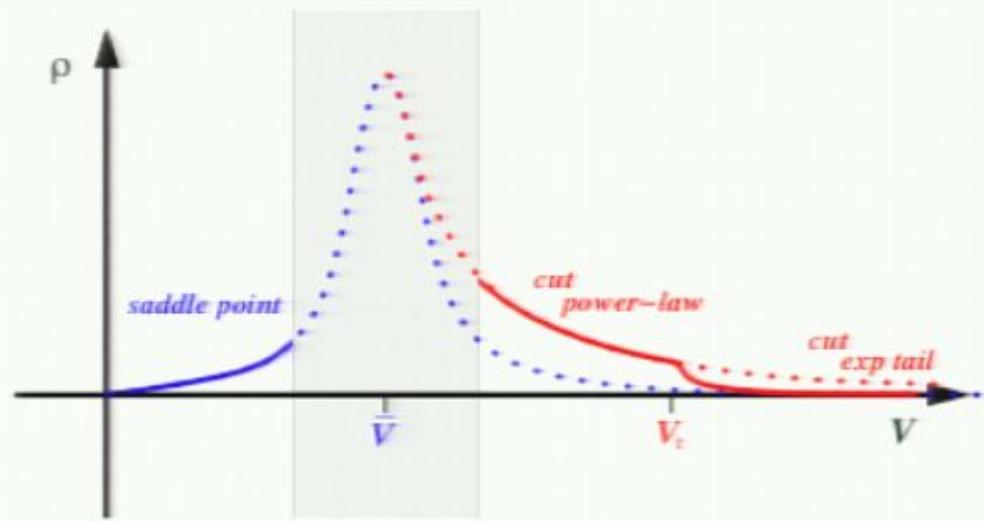
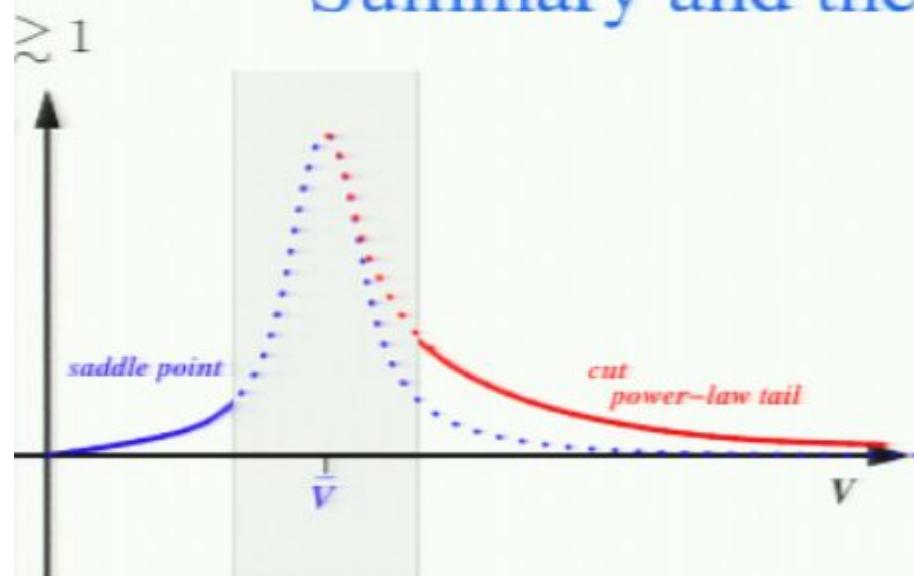
Inflation and the Landscape: detailed calculations
Volume of the Universe after Slow Roll Inflation

- Detailed study of the volume in Eternal and non-Eternal Inflation
- A bound of the finite volume produced by Inflation: $V_{\text{Finite Realization}} < e^{\frac{S_{\text{dS}}}{2}}$
- Educated speculations about de Sitter



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N. Arkani-Hamed et al.
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every finite volume realization

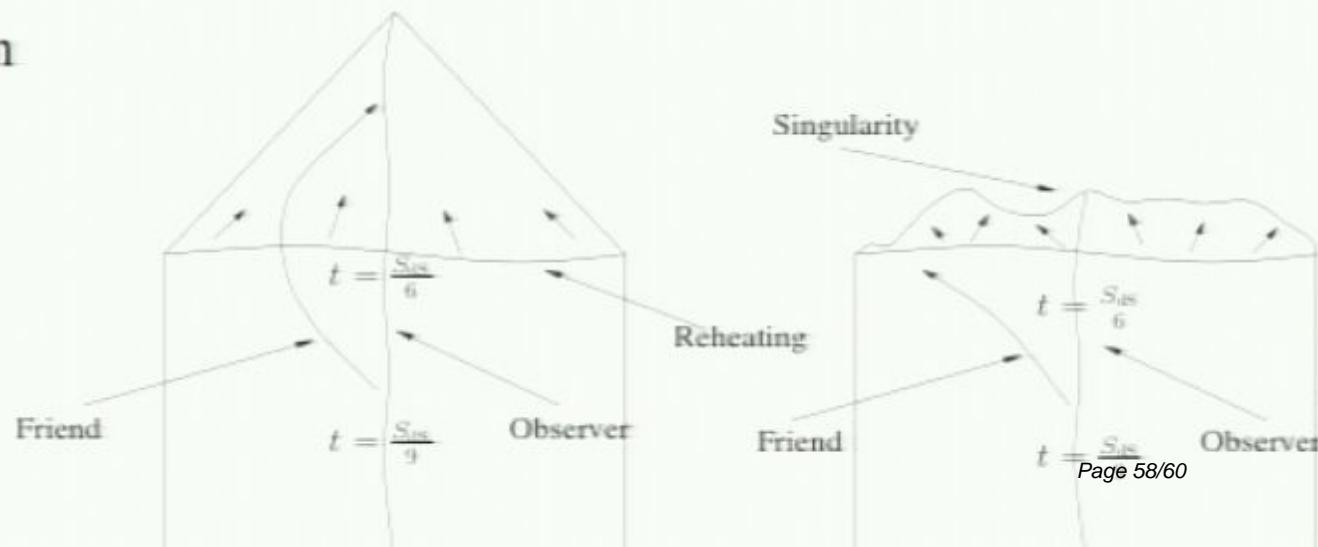
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Check for de Sitter entropy

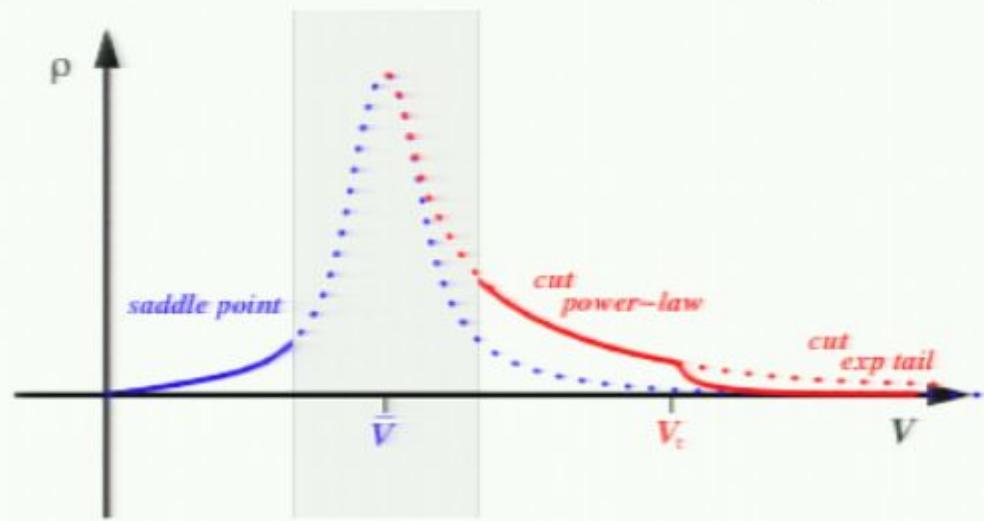
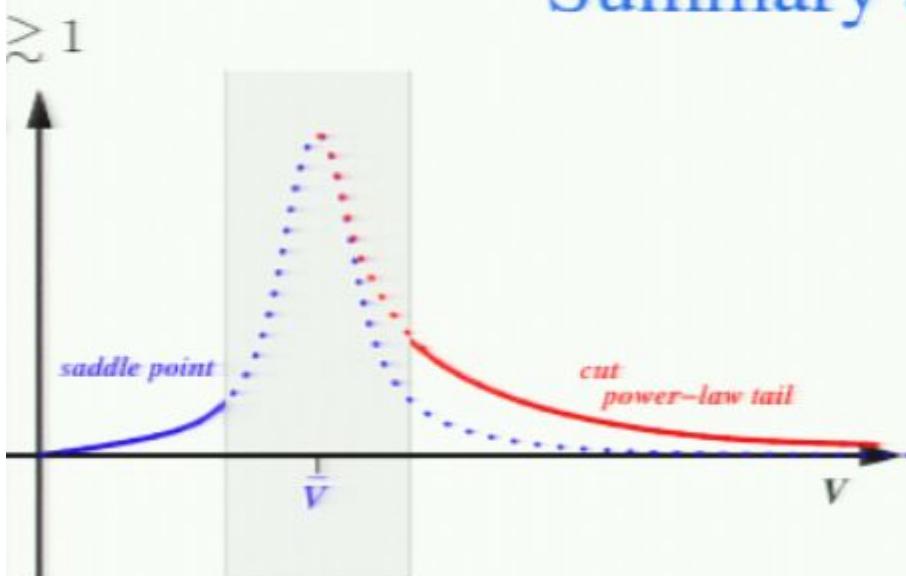
Pirsa: 08110034

void Xerox Paradox)



Summary and intuition

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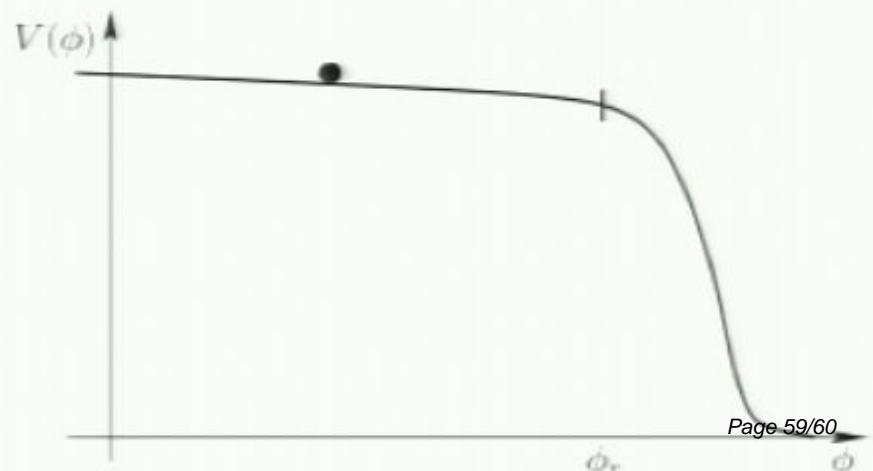
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Exponential $N \gg N_c$

$$p \propto e^{-c_1 \Delta \phi^2 / t} \propto e^{-c_2 N}$$

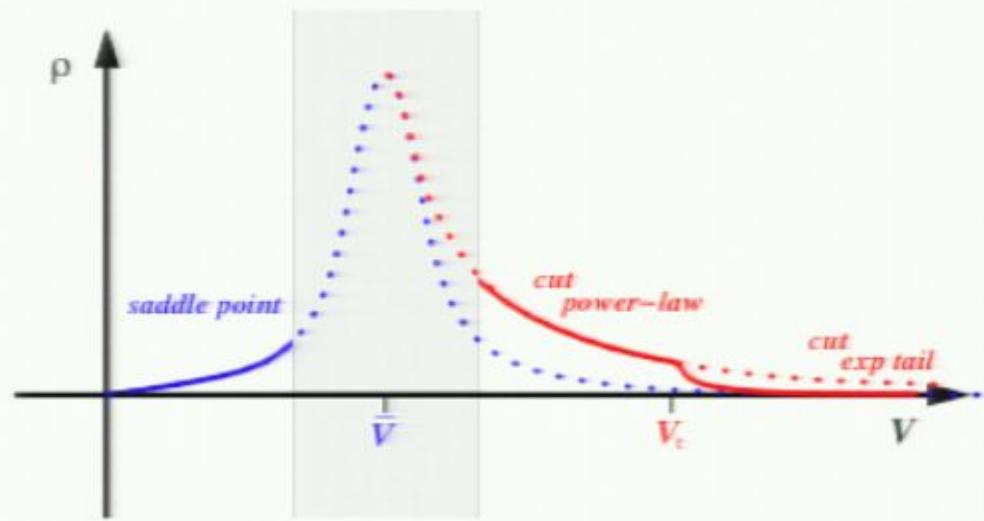
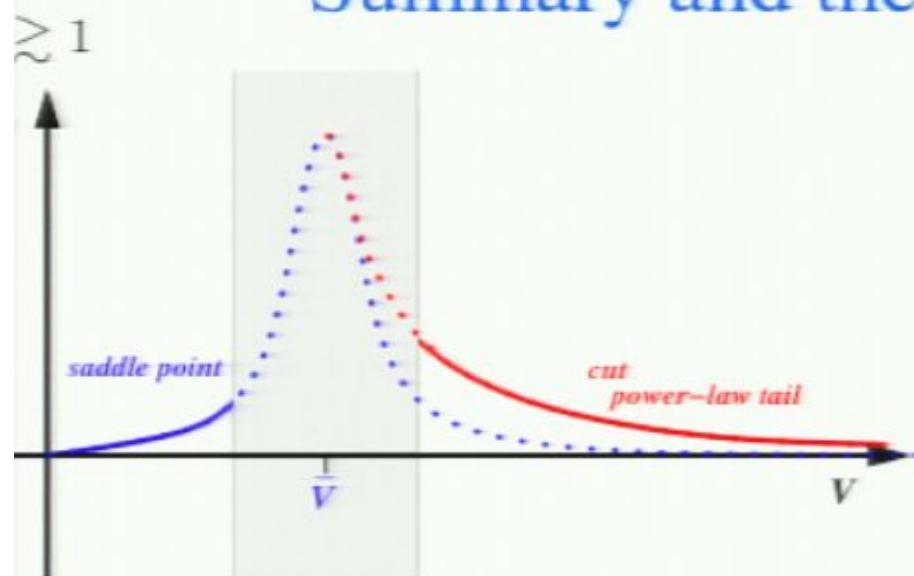
Hyper-Exponential: Eternal Inflation

$$p \propto \tilde{p} e^{3N}$$



Summary and the de Sitter Entropy

$$\Omega \lesssim 1$$



entropy and classical number of e-foldings

$$\frac{S_{ds}}{N_e} = \pi M_{Pl}^2 \frac{dH^{-2}}{Hdt} = 8\pi \frac{\dot{\phi}^2}{H^4} \quad \Rightarrow \quad N_e \lesssim \frac{S_{ds}}{12}$$

$$S_{ds} = \pi \frac{M_{Pl}^2}{H^2}$$

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every finite volume realization

$$V > e^{\frac{S_{ds}}{2}} \quad \text{and} \quad V < e^{-\alpha S_{ds}}$$

$$\text{Finite Realization} < e^{\frac{S_{ds}}{2}}$$

Check for de Sitter entropy

