

Title: Finite States in Four Dimensional Quantized Gravity

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Abstract: The semiclassical-quantum correspondence (SQC) is a new principle which has enabled the explicit solution of the quantum constraints of GR in the full theory in the Ashtekar variables for gravity coupled to matter. The solutions, which constitute the physical space of states implementing the quantum dynamics of GR in the Dirac procedure, include a special class of states known as the generalized Kodama states (GKod). The GKodS can be seen as an analogue of the pure Kodama state (Kod) when quantum gravity (QGRA) is coupled to matter fields quantized on the same footing. The criterion for finiteness stems from a precise cancellation of the ultraviolet singularities stemming from the quantum Hamiltonian constraint, allowing for an exact solution. This signifies the following developments for 4D QGRA: (i) Equivalence among the Dirac, reduced phase, geometric and path integration approaches to quantization for GKods; (ii) A generalization of topological field theory to include matter fields via the instanton representation of GKod; (iii) A possible mechanism to establish 4D QGRA, via tree networks, as a renormalizable theory (iv) A direct link from QGRA to Minkowski spacetime physics, which would enable tests of 4D QGRA without the necessity to access the Planck scale (v) A third-quantized analogy to second quantized spin network states implementing the quantum dynamics of GR. The aforementioned algorithm is designed to construct explicit solutions to the constraints of the full theory by inspection, while implementing any desired  $\hat{\epsilon}^{\sim}$ boundary $\hat{\epsilon}^{\text{TM}}$  conditions on the states necessary to reduce to the appropriate semiclassical limit. Conversely, the finite states of 4D QGRA can place severe restrictions on phenomena occurring in the weak gravitational limit below the Planck scale. While we demonstrate this for the GKodS in this talk, the procedure can be applied to obtain a family of states labeled by two arbitrary functions of position, which possess the requisite Hilbert space structure in the limit where the matter fields are turned off. Remaining areas of research in progress include the illumination of the Hilbert space structure of the GKodS, analysis of various models for which the SQC can produce tractable solutions, in the full theory and in minisuperspace, and the addressal of any issues of interest regarding the mathematical rigor of the states.

A photograph of a university campus. In the background, there is a large, modern building with a green roof and many windows. The foreground shows a green lawn and a path where several people are walking. Some are wearing dark uniforms and hats, while others are in civilian clothes. The sky is blue with some clouds. The text is overlaid on the image in a large, white, serif font.

# Finite States in Four-dimensional Quantized Gravity. Eyo Eyo Ita III

# Implications of a finite state of QGRA

- Formal equivalence of Dirac, RPS, path integration and geometric quantization procedures
  - Would imply convergence of the path integral for GR
  - Hence the effective action for gravity would be finite
- No need to access the Planck scale to test for or measure quantum gravitational effects
  - QGRA would place constraints on phenomena below the Planck scale consistent with a finite state, which cannot be deduced based on classical GR or QFT alone
  - This would imply a verifiable semiclassical limit
- No more cosmological constant problem
  - $\Lambda$  is a fundamental expansion parameter of the theory
  - Nonperturbative QGRA would explain why  $\Lambda$  is so small

# The pure Kodama state

$$\Psi_{Kod}[A] = \exp[-6(G\Lambda\hbar)^{-1}I_{CS}[A]]$$

- Chern-Simons functional of a self-dual SU(2) connection in the Ashtekar variables  $A_i^a(\mathbf{x}, t), \tilde{\sigma}_a^i(\mathbf{x}, t)$
- Exact solution to the quantum constraints of vacuum GR with cosmological term, discovered by Hideo Kodama: (Hideo Kodama: Phys. Rev. D42 (1990) 2548)
- Has a well-defined semiclassical limit below the Planck scale of quantum DeSitter spacetime (Lee Smolin: hep-th/0207079)
- Candidate for the ground state wavefunction of the universe
- Open issues with the Kodama state
  - Normalizability and Unitarity (Ed Witten objections by analogy to Yang-Mills theory: gr-qc/0306083)

# Why is $\Psi_{Kod}[A]$ special?

- Only known nontrivial exact solution to the constraints of quantum GR in the **full** theory
- Both a semiclassical and **exact** quantum state
- Resembles TQFT, which is renormalizable/finite

$$\begin{aligned}\Psi_{Kod}[A] &= \exp\left[-6(\hbar G\Lambda)^{-1} \int_{\partial M} \text{tr}\left(A \wedge dA + \frac{2}{3}A \wedge A \wedge A\right)\right] \\ &= \exp\left[-6(\hbar G\Lambda)^{-1} \int_M \text{tr}F \wedge F\right] \sim \exp\left[\int_M \text{tr}\left(\Sigma \wedge F + \frac{\hbar G\Lambda}{24}\Sigma \wedge \Sigma\right)\right]\end{aligned}$$

- But if our universe also includes quantized matter fields coupled to quantized gravity:
- How can  $\Psi_{Kod}[A]$  be generalized accordingly?

# Generalized Kodama states $\Psi_{GKod}[A_i^a(x), \psi^A(x)]$

- Is there an analogous wavefunction of the universe that
  - Exactly solves the quantum constraints of 4D general relativity, for the full theory, coupled to matter fields quantized on the same footing?
  - Bears an analogous topological representation in the presence of matter fields as the pure Kodama state does to 4D TQFT devoid of matter?
  - Is both a semiclassical and an exact quantum state to all orders?
  - Has a well-defined semiclassical limit below the Planck scale (e.g. QFT of matter fields on Minkowski spacetime, or whatever could be consistent with accelerator experiments or observation)?
- This leads to the principle of the Semiclassical/Quantum correspondence (SQC), which requires a consistent and exhaustive application of the CCR in the functional Schrodinger representation as a necessary condition for a finite state of full quantum gravity (Eyo Ita: *Class. Quantum Grav.* (2008) 125001 and 125002, and articles on gr-qc archive)

# General procedure

- CDJ Ansatz (Capovilla/Dell/Jacobson):
  - Solves classical constraints of GR with matter fields  
(Thiemann: Class. Quantum Grav. 10 (1993)1907-1921)

$$\tilde{\sigma}_a^i(x) = \Psi_{ae} [A_i^a(x), \psi^A(x)] B_e^i(x)$$

- But we want the quantum constraints to be solved
  - Quantize the CDJ matrix subject to the SQC
  - Expand quantum statistical fluctuations about  $\Psi_{Kod}$

$$\Psi_{ae} = -\frac{6}{\Lambda} \left( \delta_{ae} + \frac{\Lambda}{6} \epsilon_{ae} \right)$$

- Use the D.O.F. of the CDJ matrix to eliminate all matter-induced UV singularities, not as a Lagrange multiplier

Dit t

$$\Psi_{ae} \equiv \Psi_{(ae)}$$

$$\epsilon_{ijk} B_a^i B_e^k \Psi_{ae} = 0 \implies$$

$$\epsilon_{ijk} \tilde{\sigma}_a^i B_a^k =$$



Dit

$$\Psi_{ac} \equiv \Psi_{(ac)}$$

$$\epsilon_{ijk} B_a^i B_c^k \Psi_{ac} = 0 \Rightarrow$$

Hamiltonian

$$\epsilon_{ijk} \tilde{\sigma}_a^i B_a^k = 0$$

$$\frac{1}{6} \epsilon_{ijk} \epsilon^{abc} \tilde{\sigma}_a^i$$

$$\epsilon_{ijk} \tilde{\sigma}_a^i B_a^k = 0$$

$$\frac{\Lambda}{6} \epsilon_{ijk} \epsilon^{abc} \tilde{\sigma}_a^i \tilde{\sigma}_b^j \tilde{\sigma}_c^k$$

$$\tilde{\sigma}_a^i = \Psi_{ae} B_e^i$$

$$+ \epsilon_{ijk} \epsilon^{abc} \tilde{\sigma}_a^i \tilde{\sigma}_b^j B_c^k = 0$$

$$\Rightarrow \det B \left( \Lambda \det \Psi + \text{Var} \Psi \right)$$
$$\text{Var} \Psi = (\text{tr} \Psi)^2 - \text{tr} \Psi^2$$

$$\Psi_{ae} = \begin{pmatrix} 0 & & \\ & \Lambda & \\ & & \Lambda \end{pmatrix} \begin{pmatrix} \lambda_1 & & \\ & \lambda_2 & \\ & & \lambda_3 \end{pmatrix} \begin{pmatrix} 0 & & \\ & 1 & \\ & & 1 \end{pmatrix}$$

$$\underline{\Psi}_{ae} = \left[ \mathbf{O} \mathbf{O} \mathbf{diag}(\lambda_1, \lambda_2, \lambda_3) \mathbf{O}^T \right]_{ae}$$

$$\mathbf{O}_{ae} = \left( e^{\mathbf{A} \cdot T} \right)_{ae}$$

$$D_i \psi = \dots = 0$$

$$(D_i \psi_{ae}) B_e^i + \psi_{ae} D_i B_e^i$$

Branch  $\Rightarrow 0$

$$\Rightarrow B_e^i D_i \psi_{ae}$$

$$B_e^i \frac{\partial}{\partial x^i}$$

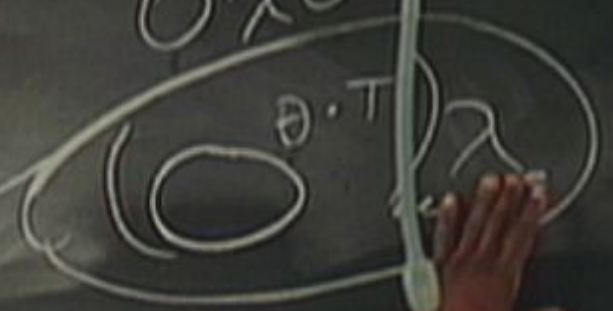
$$\frac{\partial}{\partial x^i} +$$

$\Rightarrow B' v_i \Psi_{ac}$

$B' \frac{\partial}{\partial x^i}$

$$\left[ \frac{\partial}{\partial t} + (AB) \right] \Psi_{ac} = 0$$

$$0 \vec{\lambda} 0^T$$



$\Rightarrow B e^{\lambda t} \Psi_{ac}$

$B e^{\frac{\partial}{\partial x^i}}$   $\left[ \frac{\partial}{\partial t} + (AB) \right] \Psi_{ac} = 0$

$0 \vec{\lambda} 0^T$

$(0 \quad \theta \cdot T)$

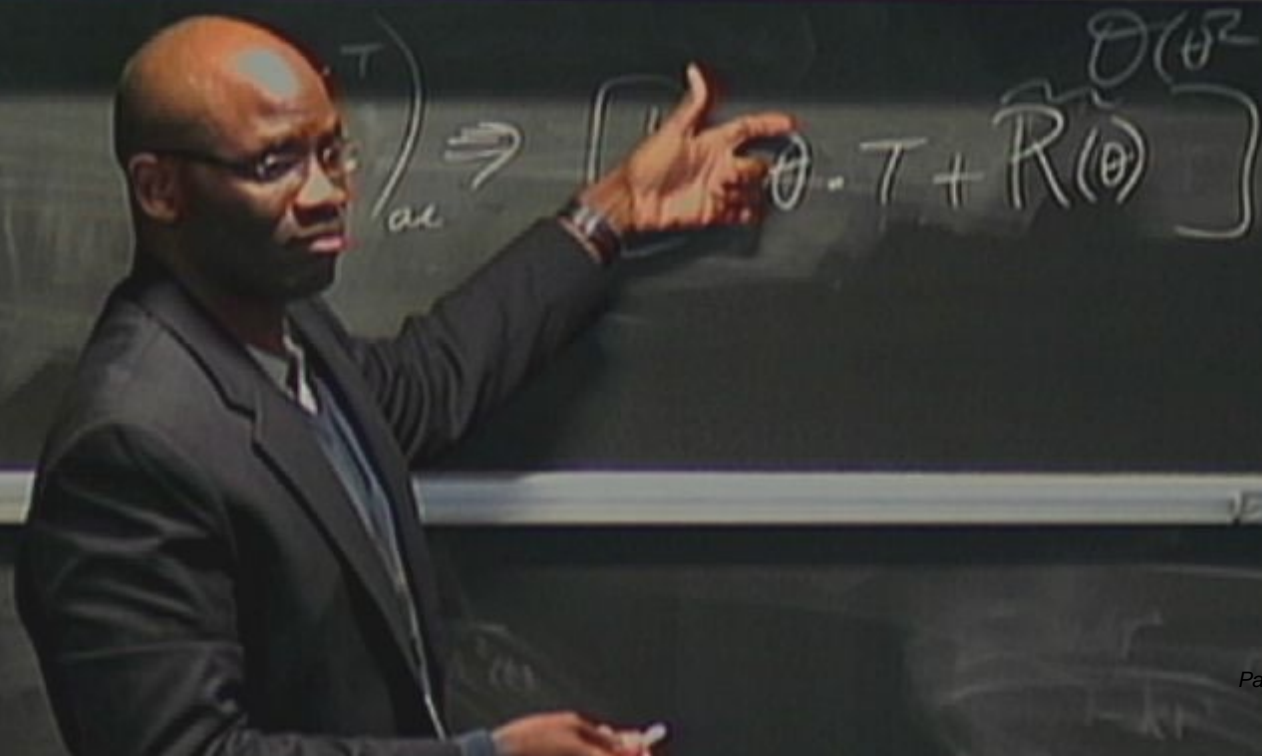
$(e^{\theta \cdot T})$

$$\Rightarrow B_{e_i} \Psi_{ae}$$

$$B_{e_i} \frac{\partial}{\partial x^i} \left[ \frac{\partial}{\partial t} + (A \cdot B) \right] \Psi_{ae} = 0$$

$$0 \vec{\lambda} 0^T$$

$$\left( \begin{array}{c} 0 \\ \theta \cdot T \\ \lambda_e \end{array} \right)$$



$$\Rightarrow B e^{A t} \Psi_{ac}$$

$$B e^{A t} \frac{\partial}{\partial x} \left[ \frac{\partial}{\partial t} + (A B) \right] \Psi_{ac} = 0$$

$$0 \vec{\lambda} 0^T$$

$$\begin{pmatrix} 0 & \theta \cdot T \\ \theta \cdot T & R(\theta) \end{pmatrix} \lambda$$

$(D\theta)$

$$e^{\theta \cdot T} \Rightarrow \left[ I + \theta \cdot T + R(\theta) \right]$$



$$\Rightarrow B e^{\lambda t} \Psi_{ac}$$

$$B e^{\lambda t} \frac{\partial}{\partial x^i} \left[ \frac{\partial}{\partial t} + (A B) \right] \Psi_{ac} = 0$$

$$(\mathbb{D}\theta)_e \Rightarrow A(\vec{\lambda}) + R(\theta)$$

$$\begin{pmatrix} 0 & \vec{\lambda} & 0^T \\ 0 & \theta \cdot \tau & \\ & & \lambda \end{pmatrix}$$

$$(e^{\theta \cdot \tau})_{ac} \Rightarrow \left[ I + \theta \cdot \tau + R(\theta) \right]$$

$$(\theta_e)_{(a)} = (\mathbb{D}^{-1} A(\vec{\lambda})) \Rightarrow R_{(a)}(\theta; \vec{\lambda})$$

$$A_{n+1} = D^{-1}A(\vec{x}) + R(\vec{x}, \vec{\theta}_n)$$

$$\theta_{n+1} = D^{-1} A(\vec{x}) + R(\vec{x}, \vec{\theta}_n)$$

$$\vec{\theta} = \lim_{n \rightarrow \infty} \vec{\theta}_n$$

$$\mathbf{A}_{n+1} = \mathbf{D}^{-1} \mathbf{A}(\vec{\lambda}) + \mathbf{R}(\vec{\lambda}, \vec{\theta}_n)$$

$$\vec{\theta} = \lim_{n \rightarrow \infty} \vec{\theta}_n$$

$$\vec{\sigma}_e^i = \frac{\Psi}{\lambda_e} \mathbf{B}_e^i$$

$(\lambda_1, \lambda_2)$

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# Kinematic constraints of 4D QGRA

- Diffeomorphism constraint:

$$\epsilon_{aef}\epsilon_{ef}(x) - G \frac{B_a^i(x)}{\det B(x)} \pi_A(x) (D_i \psi(x))^A = 0$$

- Gauss' law constraint:

$$\left( \delta^{fg} B_a^i(x) \frac{\partial}{\partial x^i} + C_a^{fg}(x) \right) \epsilon_{fg}(x) + G \pi_A(x) (T_a)^A_B \psi^B(x) = 0$$

- Mixed partials condition (new, due to SQC):

$$\pi_A(x) - \frac{i}{G} \int_{\Gamma} B_e^i(x) \delta A_i^a(x) \frac{\partial \epsilon_{ae}}{\partial \psi^A} = f_A(\psi^B(x))$$

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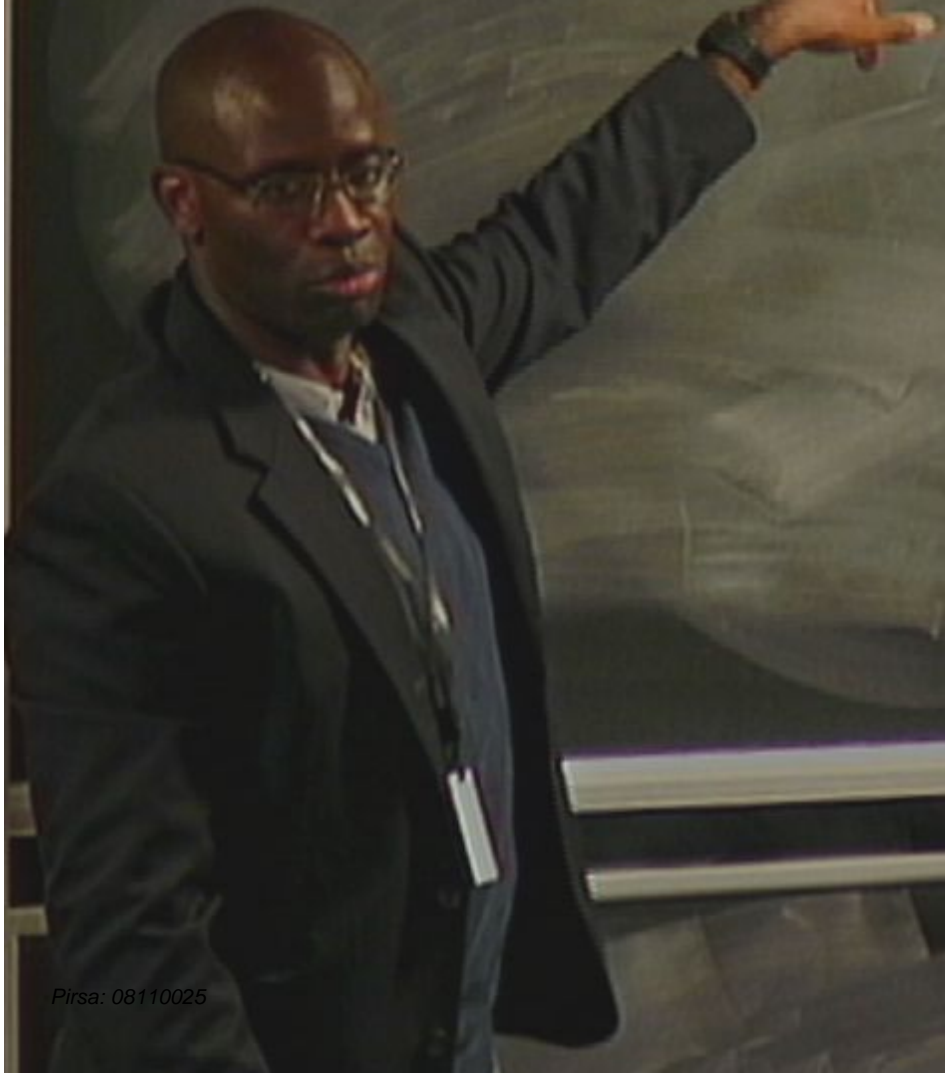
# Quantum Hamiltonian constraint

- SQC broken due to matter fields: Dilemma
  - (i) No physical states if constraints not satisfied
  - (ii) Problem of time in QGRA if constraints are satisfied

$$\hat{H}\Psi_{GKod}[A_i^a, \psi^A] = \left[ \frac{\Lambda}{6} \hbar^3 G^3 \epsilon^{abc} \epsilon_{ijk} \frac{\delta}{\delta A_i^a(x)} \frac{\delta}{\delta A_j^b(x)} \frac{\delta}{\delta A_k^c(x)} \right. \\ \left. \hbar^2 G^2 \epsilon^{abc} \epsilon_{ijk} \frac{\delta}{\delta A_j^b(x)} \frac{\delta}{\delta A_k^c(x)} B_a^i(x) + \hat{\Omega}[\psi^A, \delta/\delta\psi^A, \delta/\delta A_i^a] \right] \Psi_{GKod}[A_i^a, \psi^A] \\ = (q_0(x) + \hbar G \delta^{(3)}(0) q_1(x) + (\hbar G \delta^{(3)}(0))^2 q_2(x)) | \Psi_{GKod}[A_i^a, \psi^A] = 0$$



$$\mathcal{F}'_a = \Psi_{ae} B'_e$$



$$\vec{\sigma}_a^i = \Psi_{ac} B_c^i \quad \text{Quantize}$$

$$\hbar G \frac{\delta}{\delta A_i^a(x)} \Psi_{\text{class}}[A, \Psi] = \Psi$$

$$\vec{\sigma}_a^i = \Psi_{ae} B_e^i \quad \text{Quantize}$$

$$\hbar G \int_{\text{Grod}} \delta A_a^i(x) \Psi_{\text{Grod}}[A, \Psi] = \left( \Psi_{ae}[A(x), \Psi(x)] B_e^i(x) \right) \Psi_{\text{Grod}}[A, \Psi]$$

$$\tilde{\sigma}_a^i = \Psi_{ac} B_c^i \quad \text{Quantize}$$

$$\hbar G \int_{\text{Grod}} \frac{\delta \Psi[A, \Psi]}{\delta A_i^a(x)} = \left( \Psi_{ac}[A(x), \Psi(x)] B_c^i(x) \right) \Psi[A, \Psi] \Big|_{\text{Grod}}$$

$$\bar{\sigma}_a^i = \Psi_{ac} B_c^i \quad \text{Quantize}$$

$$\hbar G \int_{\text{Grod}} \delta \Psi_{\text{Grod}} [A, \Psi] = \left( \Psi_{ac} [A(x), \Psi(x)] B_c^i(x) \right) \Psi [A, \Psi]$$

$$\delta A_i^a(x)$$



$$\Psi = e^{\int d^3x C[A(x), \Psi(x)]}$$

$$\vec{\sigma}_a^i = \Psi_{ac} B_c^i \quad \text{Quantize}$$

$$\hbar G \frac{\delta \Psi_{\text{Grav}}[A, \Psi]}{\delta A_i^a(x)} = \left( \Psi_{ac}[A(x), \Psi(x)] B_c^i(x) \right) \Psi_{\text{Grav}}[A, \Psi]$$



$$\Psi = e^{\int d^3x \mathcal{L}(A(x), \Psi(x))}$$

$$\frac{\delta \Psi}{\delta \Psi(x)} = \mathcal{H}(x)$$

$$\bar{\sigma}_a^i = \Psi_{ac} B_c^i \quad \text{Quantize}$$

$$\hbar G \frac{\delta}{\delta A_i^a(x)} \Psi_{\text{Grod}}[A, \Psi] = \left( \Psi_{ac}[A(x), \Psi(x)] B_c^i(x) \right) \Psi_{\text{Grod}}[A, \Psi]$$



$$\Psi = e^{\int d^3x C[A(x), \Psi(x)]}$$

$$\frac{\delta \Psi}{\delta \Psi(x)} = \Pi(x) \Psi$$



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$$\frac{\delta \Psi}{\delta \Psi(x)} = \Pi(x) \Psi \quad \Psi @ e^{\int}$$



$$\bar{\sigma}_a^i = \Psi_{ac} B_c^i \quad \text{Quantize}$$

$$\hbar G \frac{\delta}{\delta A_a^i(x)} \Psi_{\text{Grod}}[A, \Psi] = \left( \Psi_{ac} [A(x), \Psi(x)] B_c^i(x) \right) \Psi_{\text{Grod}}[A, \Psi]$$



$$\Psi = e^{\int d^3x \mathcal{L}(A(x), \Psi(x))}$$

$$\frac{\delta \Psi}{\delta \Psi(x)} = \mathcal{H}(x) \Psi \quad \Psi \in \mathcal{E}^T$$

$$\left( \frac{SI}{\int A(x)} \right) e^I$$

# Quantum Hamiltonian constraint

- SQC broken due to matter fields: Dilemma
  - (i) No physical states if constraints not satisfied
  - (ii) Problem of time in QGRA if constraints are satisfied

$$\hat{H}\Psi_{GKod}[A_i^a, \psi^A] = \left[ \frac{\Lambda}{6} \hbar^3 G^3 \epsilon^{abc} \epsilon_{ijk} \frac{\delta}{\delta A_i^a(x)} \frac{\delta}{\delta A_j^b(x)} \frac{\delta}{\delta A_k^c(x)} \right. \\ \left. \hbar^2 G^2 \epsilon^{abc} \epsilon_{ijk} \frac{\delta}{\delta A_j^b(x)} \frac{\delta}{\delta A_k^c(x)} B_a^i(x) + \hat{\Omega}[\psi^A, \delta/\delta\psi^A, \delta/\delta A_i^a] \right] \Psi_{GKod}[A_i^a, \psi^A] \\ = (q_0(x) + \hbar G \delta^{(3)}(0) q_1(x) + (\hbar G \delta^{(3)}(0))^2 q_2(x)) | \Psi_{GKod}[A_i^a, \psi^A] = 0$$

Standard form for large class of models coupled to gravity: Regularization-independent at canonical level

- 9+N system for CDJ deviation matrix

$$O_{gh}^{ae} \epsilon_{ae} + \Lambda \Sigma_{gh}^{aebf} \epsilon_{ae} \epsilon_{bf} + \Lambda^2 E_{gh}^{abcdef} \epsilon_{ad} \epsilon_{be} \epsilon_{cf} = GQ_{gh}$$

- Identical structure to EOM of Yang Mills theory, which is renormalizable and finite

$$(\delta_\nu^\gamma \partial^2 - \partial^\gamma \partial_\nu) A_\nu^a + g (f^{ade} (\partial^\mu A_\mu^d) A_\nu^e - f^{abc} A_\mu^c \partial_\nu A^{b\mu}) + g^2 f^{abc} f^{bde} A_\mu^c A^{d\mu} A_\nu^e - J_\nu^a = 0$$

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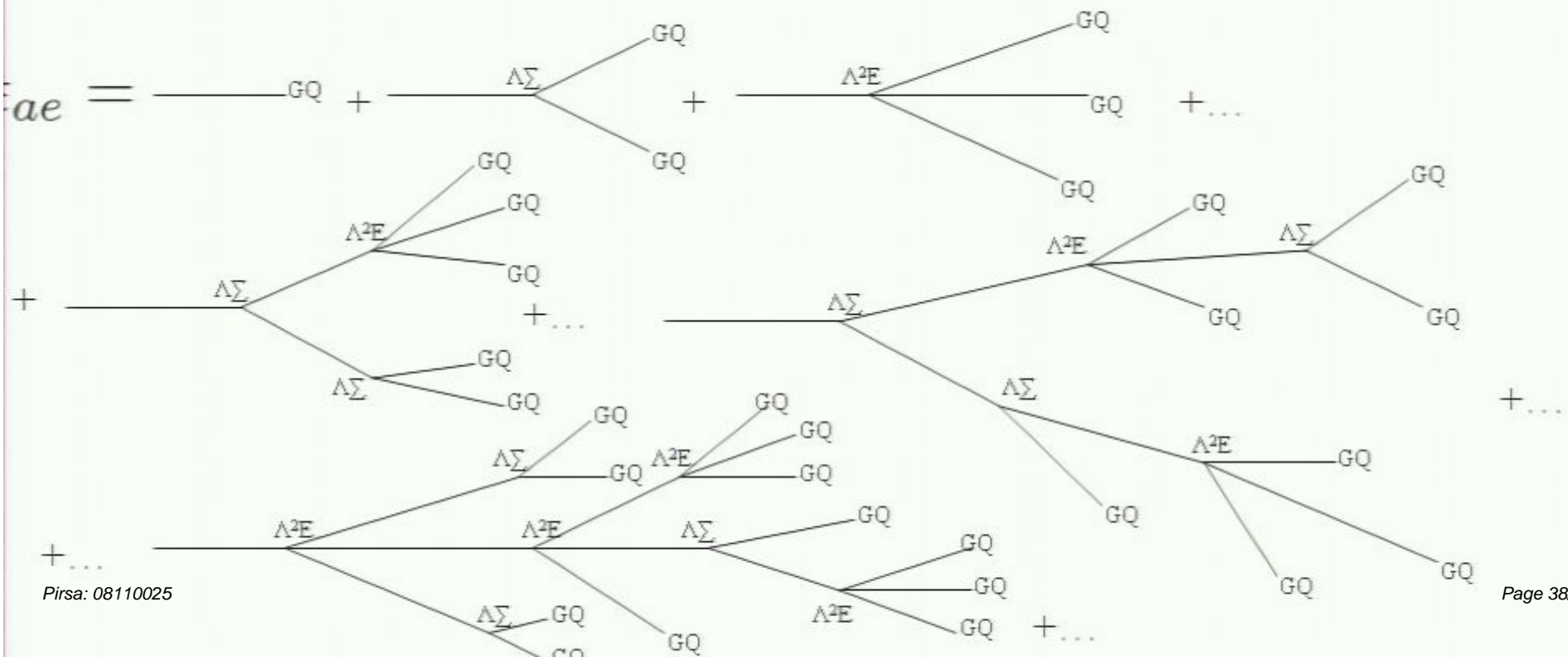
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# General solution by Feynman-like diagrammatic tree networks

- Third-quantized generalization of spin network state solving the quantum constraints of GR



# Wavefunction of the Universe

- Perturbative expansion in dimensionless coupling constant  $G\Lambda$  or in  $G\Lambda' = G\Lambda + G^2V(\psi)$

$$\Psi_{ae} = - \left[ \frac{6}{\Lambda} \delta_{ae} + \sum_{n=0}^{\infty} (G\Lambda')^n \left( \prod_i \Sigma_{g_i h_i}^{a_i e_i b_i f_i} \prod_j E_{a_j e_j b_j f_j c_j d_j}^{g_j h_j} \prod_k O_{a_k b_k}^{c_k d_k} Q_{c_k d_k} [f] \right) \right]_{ae}$$

$$\mathcal{U}_{GKod} [A_i^a, \psi^A] = \exp \left[ \int_M d^4x \left( (\hbar G)^{-1} \Psi_{ae} B_e^i \dot{A}_i^a + \frac{i}{\hbar} f_A \dot{\psi}^A \right) \right]$$

- Choose  $f$  to reproduce SR for matter below the Planck scale: sets boundary condition on QGRA

# Representations of $\Psi_{GKod}$

- Boundary term on final spatial hypersurface:  
Problem of time has disappeared

$$\exp \left[ (\hbar G)^{-1} \int_{\partial M} d^3x \int_{\Gamma_A} \Psi_{ae} B_e^i \delta A_i^a \right] \exp \left[ \frac{i}{\hbar} \int_{\partial M} d^3x \int_{\Gamma_\psi} f_A \delta \psi^A \right]$$

- Instanton representation: TQFT plus matter

$$\exp \left[ (\hbar G)^{-1} \int_M \Psi_{ae} F^a \wedge F^e \right] \exp \left[ \frac{i}{\hbar} \int_M d^4x f_A (D_0 \psi)^A \right]$$



# Wavefunction of the Universe

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# In progress research

- Formalizing Hilbert space structure of  $\Psi_{GKod}$
- Generalization to include gravitational DOF (done)
- Develop interpretation as a third-quantized theory
- Semigroup interpretation of the Hamiltonian constraint functional Green's functions
- Yang-Mills theory requires special treatment since non-polynomial in the basic variables
- Creation of a library of states for different models in the full theory: Ultimate goal is the Standard Model

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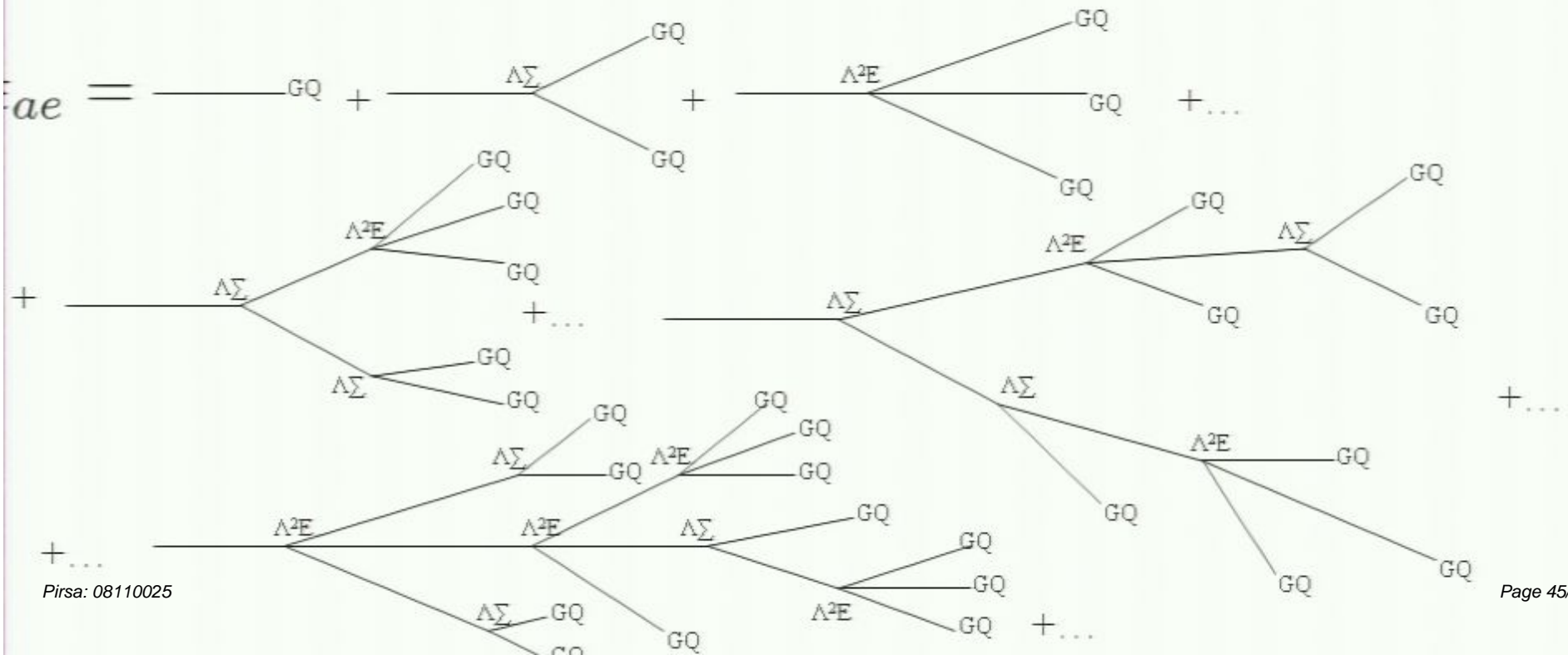
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- Third-quantized generalization of spin network state solving the quantum constraints of GR



Solution: Set  $q_0 = q_1 = q_2 = 0$

$$\left( \delta_{ae} - \frac{G}{2} \frac{B_a^i B_e^j}{\det B} \partial_i \phi \partial_j \phi \right) \epsilon_{ae} + \frac{\Lambda}{6} \left( \delta_{ae} \delta_{bf} - \delta_{af} \delta_{be} - \frac{G}{4} \frac{B_e^i B_f^j}{\det B} \partial_i \phi \partial_j \phi \right) \epsilon_{ae} \epsilon_{bf} + \frac{\Lambda^2}{72} \epsilon_{abc} \epsilon_{def} \epsilon_{ad} \epsilon_{be} \epsilon_{cf} - \frac{G\Lambda}{12 \det B} \left( \frac{\pi^2}{2} + \frac{1}{2} B_a^i B_a^j \partial_i \phi \partial_j \phi \right) = 0;$$

$$\left[ \det B \left( \delta_{bf} (B^{-1})_i^a - (B^{-1})_i^b \delta_{af} + \frac{G}{4} \delta_{ab} \tau_{if} \right) \frac{\partial}{\partial A_i^a} + 2 \delta_{bf} \text{tr} C \right] \epsilon_{bf} - \frac{i}{4 \det B} \frac{\partial \pi}{\partial \phi}$$

$$+ \frac{\Lambda}{8} \left( \epsilon^{abd} \epsilon^{fec} (B^{-1})_i^c \frac{\partial}{\partial A_i^d} + 4 (\delta_e^a C_f^b - \delta_e^b C_f^a) \right) \epsilon_{ae} \epsilon_{bf} = 0;$$

$$\left[ \epsilon_{ijk} \epsilon^{abc} B_e^k \frac{\partial}{\partial A_i^a} \frac{\partial}{\partial A_j^b} + 4 (\delta^{ce} A_k^a - \delta^{ae} A_k^c) \frac{\partial}{\partial A_k^a} + 12 \delta^{ce} \right] \epsilon_{ce} = 0;$$

$$\Psi_{ae} = \underbrace{\lambda_{ae}} + \epsilon_{ae}$$

$$\int_{\mathcal{D}} d^3x \lambda_{ae}(x) X^{ae}(x)$$

$$\text{Diag}(\lambda_1, \lambda_2, \lambda_3, (\lambda, \lambda))$$

$$B_c^i \delta A_i^a = \delta X^{ae}$$

$$\Psi_{ae}(X) = e^{-iX^2} \int DX$$

$$\epsilon_r \Rightarrow B_c^i \delta A_i^a = \delta L_{CS}[A]$$