

Title: General Relativity 1 - Introduction of a New Idea

Date: Aug 01, 2006 09:00 AM

URL: <http://pirsa.org/08080071>

Abstract: We shift our ideas from Newton's law of gravity to a new set of equations that describe how gravity is a consequence of the curvature of spacetime.
Learning Outcomes:
• John Michell and his hypothetical object called a dark star.
• How to determine the mass of a planet required for the escape velocity of an object to be the speed of light.
• Einstein's Equivalence Principle.





Lecture 10

Black Holes





Lecture 10

Black Holes



Lecture 10

Black Holes



Lecture 10

Black Holes





Centaurus A: Feeding a Black Hole

Ground-based image of Centaurus A, ZOOM into the Hubble Space Telescope WFPC2 Camera image of Centaurus A and its Nucleus, DISSOLVE in the HST NICMOS instrument infrared image. The animation shows what is seen from the NICMOS image: inside the galaxy are a hot-gas disk and a twisted jet and disk around a feeding black hole.

*Ground-based image: NOAO/CTIO
WFPC2 and NICMOS images: Ethan Schreier,
The Space Telescope Science Institute and NASA
Animation: Thomas Goertel*

Lecture 10

Black Holes

Dark stars

- **Rev. John Michell (1783)**

A British born "natural philosopher" dared to combine the corpuscular description of light with Newton's gravitation laws to predict what large compact stars should look like.

- He showed that a star, that has the same density of the sun, but 500 time as big, would have such a gravity, that "All light emitted from such a body would be made to return towards it". He said we wouldn't be able to see such a body, but we sure will feel it's gravitational pull.
- We could fly close to this "Dark star" and look around and describe the features of the object.
- A novelty, world lost interest when light was shown to be waves in 1803 by Thomas Young.

Calculation of Escape Velocity

$$\frac{1}{2}mv^2 = \frac{GMm}{r}$$



Calculation of Escape Velocity

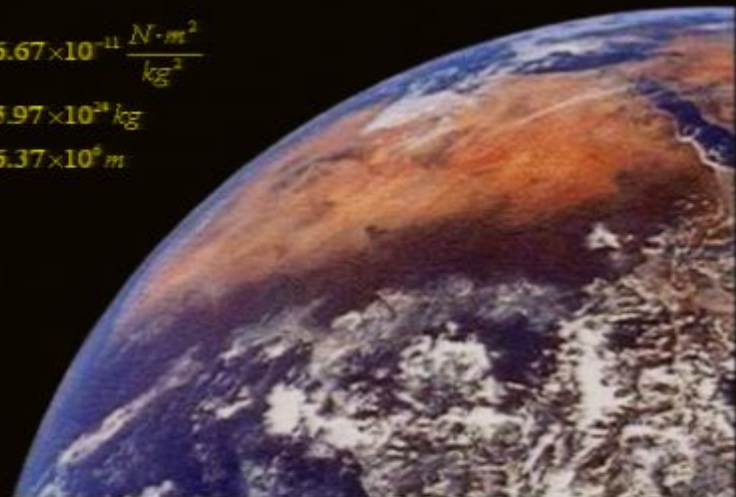
$$\frac{1}{2}mv^2 = \frac{GMm}{r}$$

$$G = 6.67 \times 10^{-11} \frac{N \cdot m^2}{kg^2}$$

$$M = 5.97 \times 10^{24} kg$$

$$r = 6.37 \times 10^6 m$$

Calculate Escape Velocity



Calculation of Escape Velocity

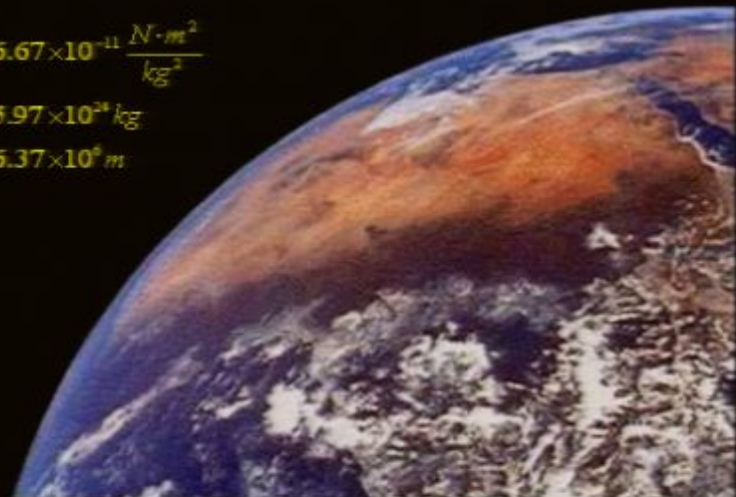
$$\frac{1}{2}mv^2 = \frac{GMm}{r}$$

$$G = 6.67 \times 10^{-11} \frac{N \cdot m^2}{kg^2}$$

$$M = 5.97 \times 10^{24} kg$$

$$r = 6.37 \times 10^6 m$$

Calculate Escape Velocity



Calculation of Escape Velocity

$$\frac{1}{2}mv^2 = \frac{GMm}{r}$$

$$G = 6.67 \times 10^{-11} \frac{N \cdot m^2}{kg^2}$$

$$M = 5.97 \times 10^{24} kg$$

$$r = 6.37 \times 10^6 m$$

Calculate Escape Velocity



Calculation of Escape Velocity

$$\frac{1}{2}mv^2 = \frac{GMm}{r}$$

$$G = 6.67 \times 10^{-11} \frac{N \cdot m^2}{kg^2}$$

$$M = 5.97 \times 10^{24} kg$$

$$r = 6.37 \times 10^6 m$$

Calculate Escape Velocity



Calculation of Escape Velocity

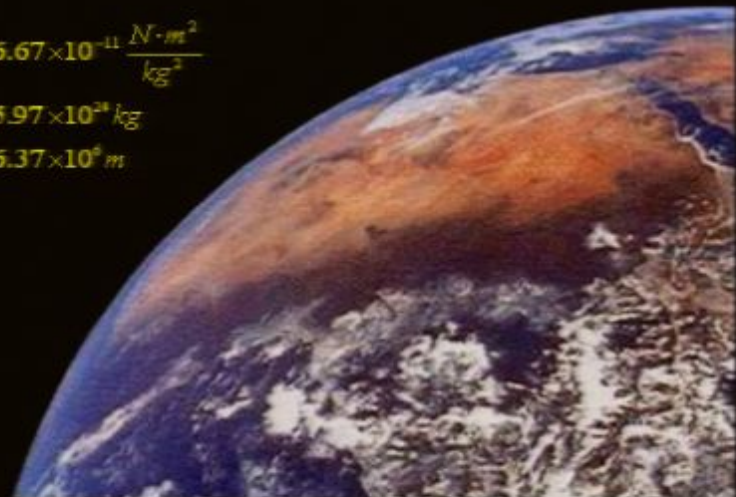
$$\frac{1}{2}mv^2 = \frac{GMm}{r}$$

$$G = 6.67 \times 10^{-11} \frac{N \cdot m^2}{kg^2}$$

$$M = 5.97 \times 10^{24} kg$$

$$r = 6.37 \times 10^6 m$$

Calculate Escape Velocity



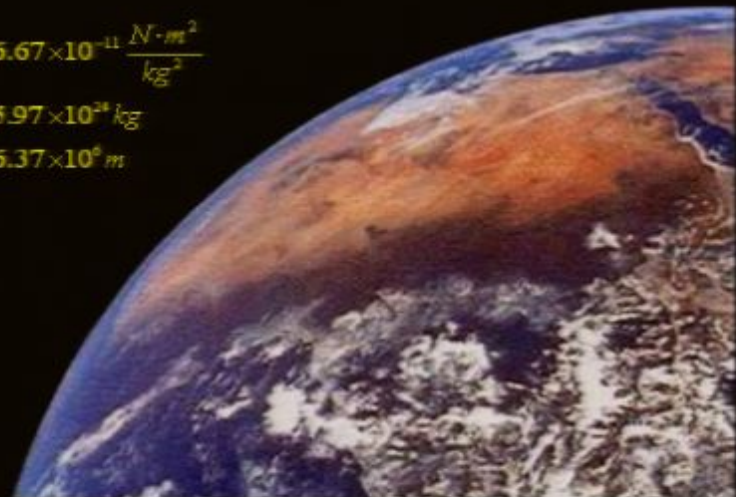
Calculation of Escape Velocity

$$\frac{1}{2}mv^2 = \frac{GMm}{r}$$

$$G = 6.67 \times 10^{-11} \frac{N \cdot m^2}{kg^2}$$

$$M = 5.97 \times 10^{24} kg$$

$$r = 6.37 \times 10^6 m$$



Calculation of Escape Velocity

$$\frac{1}{2}mv^2 = \frac{GMm}{r}$$

$$\frac{1}{2}v^2 = \frac{GM}{r}$$

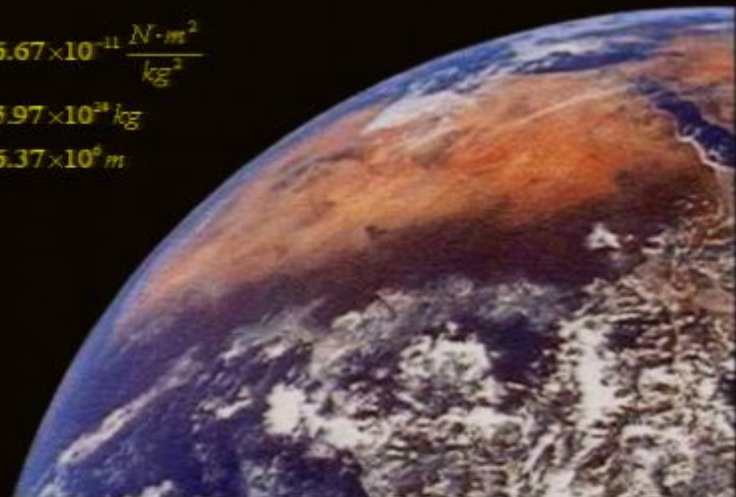
$$v = \sqrt{2 \frac{(6.67 \times 10^{-11})(5.97 \times 10^{24})}{6.37 \times 10^6}}$$

$$v \approx 11181 \text{ m/s}$$

$$G = 6.67 \times 10^{-11} \frac{\text{N} \cdot \text{m}^2}{\text{kg}^2}$$

$$M = 5.97 \times 10^{24} \text{ kg}$$

$$r = 6.37 \times 10^6 \text{ m}$$



Calculation

$$\frac{1}{2}mv^2 = \frac{GMm}{r}$$

$$\frac{1}{2}v^2 = \frac{GM}{r}$$

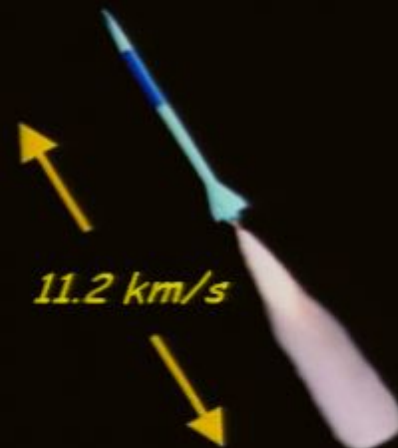
$$v = \sqrt{2 \frac{(6.67 \times 10^{-11})(5.97 \times 10^{24})}{6.37 \times 10^6}}$$

$$v \approx 11181 \text{ m/s}$$

$$G = 6.67 \times 10^{-11} \frac{\text{N} \cdot \text{m}^2}{\text{kg}^2}$$

$$M = 5.97 \times 10^{24} \text{ kg}$$

$$r = 6.37 \times 10^6 \text{ m}$$



What is r when $v=3 \times 10^8$ m/s

$$\frac{1}{2}mv^2 = \frac{GMm}{r}$$

$$\frac{1}{2}v^2 = \frac{GM}{r}$$

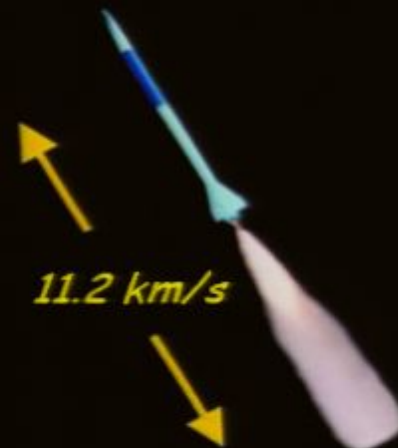
$$v = \sqrt{2 \frac{(6.67 \times 10^{-11})(5.97 \times 10^{24})}{6.37 \times 10^6}}$$

$$v \approx 11181 \text{ m/s}$$

$$G = 6.67 \times 10^{-11} \frac{\text{N} \cdot \text{m}^2}{\text{kg}^2}$$

$$M = 5.97 \times 10^{24} \text{ kg}$$

$$r = 6.37 \times 10^6 \text{ m}$$



What is r when $v=3 \times 10^8 \text{ m/s}$

8.8 mm

$$\frac{1}{2}mv^2 = \frac{GMm}{r}$$

$$\frac{1}{2}v^2 = \frac{GM}{r}$$

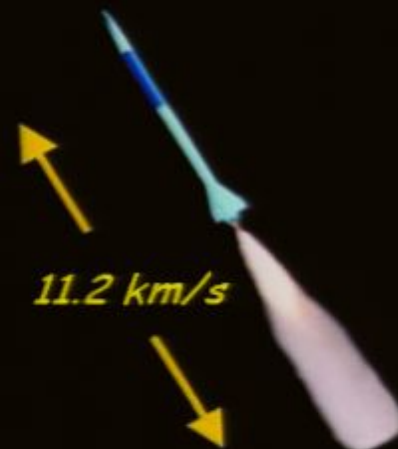
$$v = \sqrt{2 \frac{(6.67 \times 10^{-11})(5.97 \times 10^{24})}{6.37 \times 10^6}}$$

$$v \approx 11181 \text{ m/s}$$

$$G = 6.67 \times 10^{-11} \frac{\text{N} \cdot \text{m}^2}{\text{kg}^2}$$

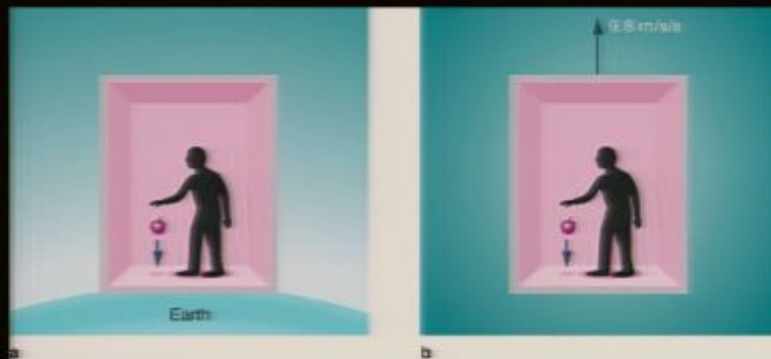
$$M = 5.97 \times 10^{24} \text{ kg}$$

$$r = 6.37 \times 10^6 \text{ m}$$



Einstein's Equivalence Principle

- There is no experiment that you can perform that will distinguish these two diagrams







Einstein's Equivalence Principle

- There is no experiment that you can perform that will distinguish these two diagrams

