

Title: Weclome and introduction

Date: Jul 14, 2008 10:00 AM

URL: <http://pirsa.org/08070005>

Abstract:

# Variation of Fundamental Constants

V.V. Flambaum

School of Physics, UNSW, Sydney, Australia

Co-authors:

Atomic calculations V.Dzuba, M.Kozlov,

E.Angstmann, J.Berengut, M.Marchenko, Cheng Chin, S.Karshenboim, A.Nevsky

Nuclear and QCD calculations E.Shuryak, V.Dmitriev, D.Leinweber, A.Thomas,  
R.Young, A.Hoell, P.Jaikumar, C.Roberts, S.Wright, A.Tedesco, W.Wiringa

Cosmology J.Barrow

Quasar data analysis

J.Webb, M.Murphy, M.Drinkwater, W.Walsh, P.Tsanavaris, S.Curran

Quasar observations C.Churchill, J.Prochazka, A.Wolfe, S.Muller, C.Henkel, F.Combes,  
T.Wiklind, thanks to W.Sargent, R.Simcoe

Laboratory measurements S.J. Ferrel, A.Cingoz, A.Lappiere, A.-T.Nguyen, N.Leefer,  
D.Budker, S.K.Lamoreaux, J.R.Torgerson, S.Blatt, A.D.Ludlow, G.K.Cambell,

J.W.Thomsen, T.Zelevinsky, M.M.Boyd, J.Ye, X.Baillard, M.Fouche, R.LeTargat, A.Brush,  
P.Lemonde, M.Takamoto, F.-L.Hong, H.Katori

# Motivation

- **Extra space dimensions** (Kaluza-Klein, Superstring and M-theories). Extra space dimensions is a common feature of theories unifying **gravity** with other interactions. Any change in size of these dimensions would manifest itself in the 3D world as variation of fundamental constants.
- **Scalar fields** . Fundamental constants depend on scalar fields which vary in space and time (variable vacuum dielectric constant  $\epsilon_0$  ). May be related to “dark energy” and accelerated expansion of the Universe..
- **“Fine tuning”** of fundamental constants is needed for humans to exist. Example: low-energy resonance in production of carbon from helium in stars ( $\text{He}+\text{He}+\text{He}=\text{C}$ ). Slightly different coupling constants — no resonance — no life.

Variation of coupling constants in space provide natural explanation of the “fine tuning”: we appeared in area of the Universe where values of fundamental constants are suitable for our existence.

# Search for variation of fundamental constants

- Big Bang Nucleosynthesis
- Quasar Absorption Spectra <sup>1</sup>
- Oklo natural nuclear reactor
- Atomic clocks <sup>1</sup>
- Enhanced effects in atoms <sup>1</sup>, molecules<sup>1</sup> and nuclei
- Dependence on gravity

# Search for variation of fundamental constants

- Big Bang Nucleosynthesis
- Quasar Absorption Spectra <sup>1</sup>
- Oklo natural nuclear reactor
- Atomic clocks <sup>1</sup>
- Enhanced effects in atoms <sup>1</sup>, molecules<sup>1</sup> and nuclei
- Dependence on gravity

evidence?

evidences?

# Dimensionless Constants

Since variation of dimensional constants cannot be distinguished from variation of units, it only makes sense to consider variation of dimensionless constants.

- **Fine structure constant**  $\alpha = e^2 / hc = 1/137.036$
- **Electron or quark mass/QCD strong interaction scale**,  $m_{e,q} / \Lambda_{\text{QCD}}$

$$\alpha_{\text{strong}}(r) = \text{const} / \ln(r \Lambda_{\text{QCD}} / ch)$$

$m_{e,q}$  are proportional to Higgs vacuum (weak scale)

---

Relation between variations of  
different coupling constants  
Grand unification models (Calmet, Fritzsch;  
Langecker, Segre, Strasser; Dent...)

$$\alpha_i^{-1}(v) = \alpha_{GUT}^{-1} + b_i \ln(v / v_0)$$

*Variation of GUT const  $\alpha_{GUT}$*

$$d\alpha_1^{-1} = d\alpha_2^{-1} = d\alpha_3^{-1} = d\alpha_{GUT}^{-1}$$

$$d\alpha_3 / \alpha_3 = (\alpha_3 / \alpha_1) d\alpha_1 / \alpha_1$$

$$\alpha_3^{-1}(m) = \alpha_{\text{strong}}^{-1}(m) = b_3 \ln(m / \Lambda_{\text{QCD}})$$

$$\alpha^{-1}(m) = 5/3 \alpha_1^{-1}(m) + \alpha_2^{-1}(m)$$

$$\frac{\Delta(m / \Lambda_{\text{QCD}})}{m / \Lambda_{\text{QCD}}} = \frac{1}{b_3 \alpha_3} \frac{\Delta \alpha_3}{\alpha_3} = \frac{\text{const}}{\alpha} \frac{\Delta \alpha}{\alpha} \sim 35 \frac{\Delta \alpha}{\alpha}$$

Proton mass  $M_p \sim 4 \Lambda_{\text{QCD}}$ , measure  $m_e / M_p$

. Nuclear magnetic moments

$$\mu = g e \hbar / 4 M_p c, \quad g = g(m_q / \Lambda_{\text{QCD}})$$

. Nuclear energy levels and resonances



# Dimensionless Constants

Since variation of dimensional constants cannot be distinguished from variation of units, it only makes sense to consider variation of dimensionless constants.

- **Fine structure constant**  $\alpha = e^2 / hc = 1/137.036$
- **Electron or quark mass/QCD strong interaction scale**,  $m_{e,q} / \Lambda_{\text{QCD}}$

$$\alpha_{\text{strong}}(r) = \text{const} / \ln(r \Lambda_{\text{QCD}} / ch)$$

$m_{e,q}$  are proportional to Higgs vacuum (weak scale)

$$\alpha_3^{-1}(m) = \alpha_{\text{strong}}^{-1}(m) = b_3 \ln(m / \Lambda_{\text{QCD}})$$

$$\alpha^{-1}(m) = 5/3 \alpha_1^{-1}(m) + \alpha_2^{-1}(m)$$

$$\frac{\Delta(m / \Lambda_{\text{QCD}})}{m / \Lambda_{\text{QCD}}} = \frac{1}{b_3 \alpha_3} \frac{\Delta \alpha_3}{\alpha_3} = \frac{\text{const}}{\alpha} \frac{\Delta \alpha}{\alpha} \sim 35 \frac{\Delta \alpha}{\alpha}$$

Proton mass  $M_p \sim 4 \Lambda_{\text{QCD}}$ , measure  $m_e / M_p$

. Nuclear magnetic moments

$$\mu = g e \hbar / 4 M_p c, \quad g = g(m_q / \Lambda_{\text{QCD}})$$

. Nuclear energy levels and resonances

# Dependence on quark mass

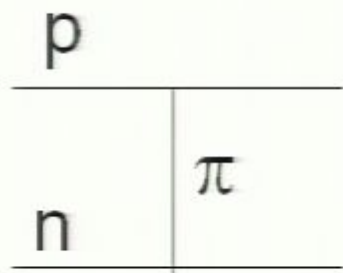
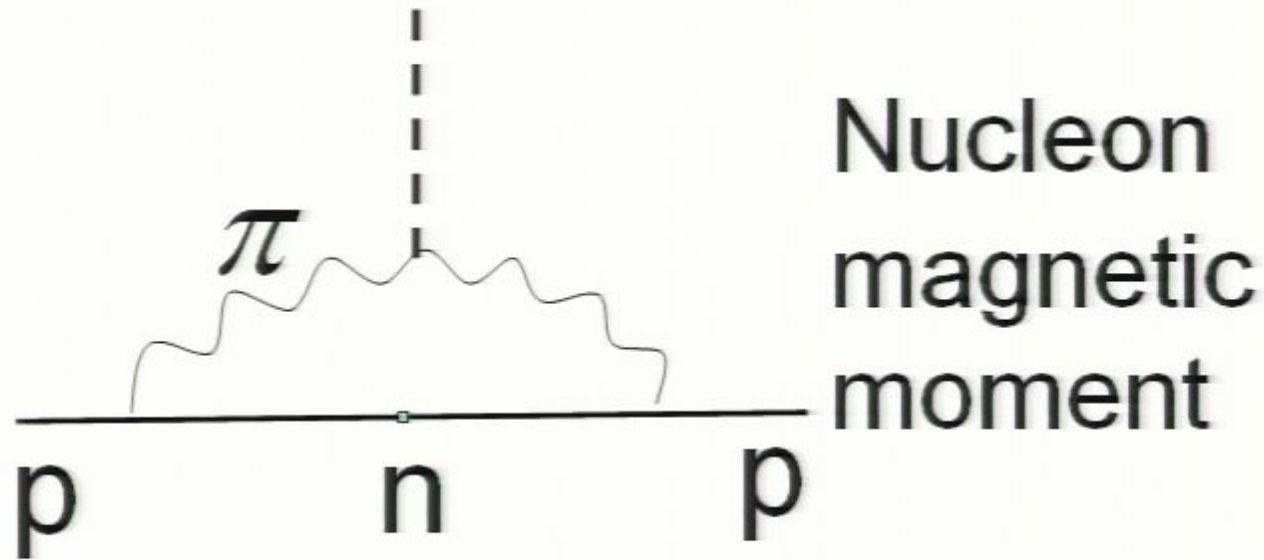
- Dimensionless parameter is  $m_q/\Lambda_{\text{QCD}}$ . It is convenient to assume  $\Lambda_{\text{QCD}} = \text{const}$ , i.e. measure  $m_q$  in units of  $\Lambda_{\text{QCD}}$
- $m_\pi$  is proportional to  $(m_q \Lambda_{\text{QCD}})^{1/2}$   
 $\Delta m_\pi / m_\pi = 0.5 \Delta m_q / m_q$
- Other meson and nucleon masses remains finite for  $m_q = 0$ .  $\Delta m / m = K \Delta m_q / m_q$

Argonne: K are calculated for p, n,  $\rho$ ,  $\omega$ ,  $\sigma$ .

$$m_q = \frac{m_u + m_d}{2} \approx 4 \text{MeV}, \quad \Lambda_{\text{QCD}} = 220 \text{MeV} \rightarrow K = 0.02 - 0.06$$

Strange quark mass  $m_s = 120 \text{MeV}$

Nuclear magnetic moments depends on  $\pi$ -meson mass  $m_\pi$



Spin-spin interaction between valence and core nucleons

# Nucleon magnetic moment

$$\mu = \mu_0 (1 + am_\pi + \dots) = \mu_0 (1 + b\sqrt{m_q} + \dots)$$

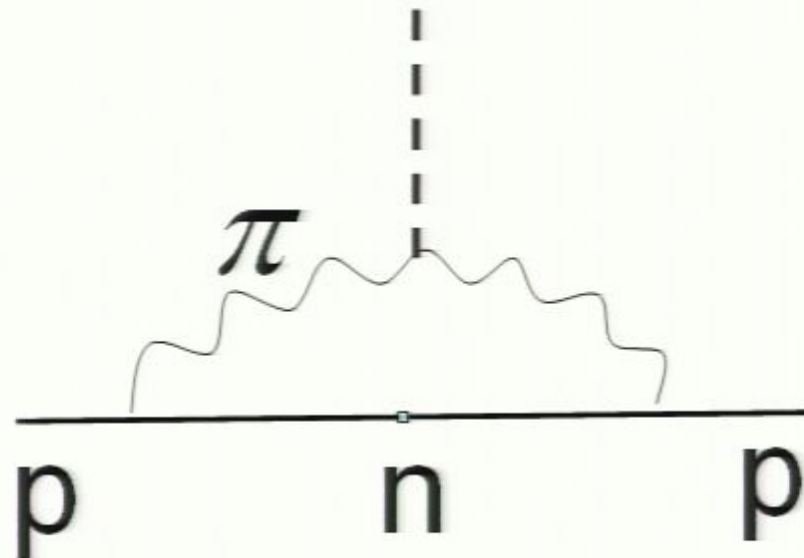
## Nucleon and meson masses

$$M = M_0 + am_q$$

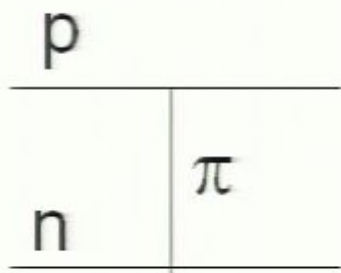
**QCD calculations:** lattice, chiral perturbation theory, cloudy bag model, Dyson-Schwinger and Faddeev equations, semiempirical.

**Nuclear calculations:** meson exchange theory of strong interaction. Nucleon mass in kinetic

Nuclear magnetic moments depends on  $\pi$ -meson mass  $m_\pi$



Nucleon magnetic moment



Spin-spin interaction between valence and core nucleons

# Nucleon magnetic moment

$$\mu = \mu_0 (1 + am_\pi + \dots) = \mu_0 (1 + b\sqrt{m_q} + \dots)$$

## Nucleon and meson masses

$$M = M_0 + am_q$$

**QCD calculations:** lattice, chiral perturbation theory, cloudy bag model, Dyson-Schwinger and Faddeev equations, semiempirical.

**Nuclear calculations:** meson exchange theory of strong interaction. Nucleon mass in kinetic

---

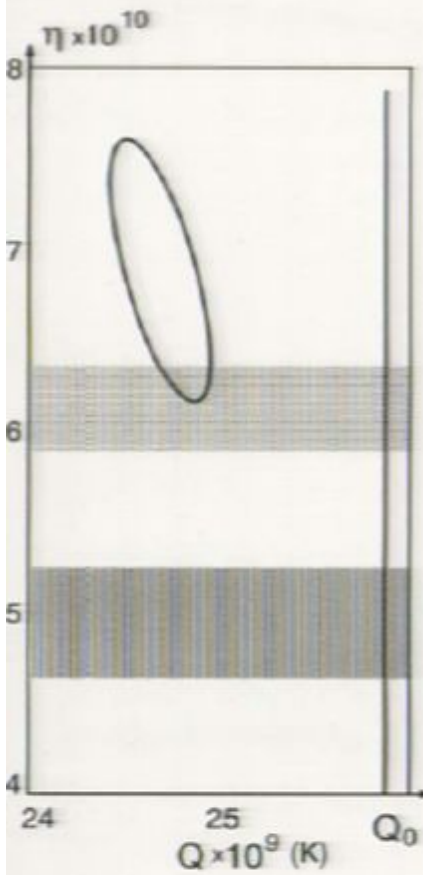
# Big Bang nucleosynthesis: dependence on quark mass

- Flambaum, Shuryak 2002
- Flambaum, Shuryak 2003
- Dmitriev, Flambaum 2003
- Dmitriev, Flambaum, Webb 2004
- Coc, Nunes, Olive, Uzan, Vangioni 2007
- Dent, Stern, Wetterich 2007
- Flambaum, Wiringa 2007
- Berengut, Dmitriev, Flambaum 2008



# Big Bang Nucleosynthesis

(Dmitriev, Flambaum, Webb)



Productions of D,  $^4\text{He}$ ,  $^7\text{Li}$  are exponentially sensitive to deuteron binding energy  $E_d$

$$\sim e^{-\frac{E_d}{T_f}}$$

-  $\eta$  from cosmic microwave background fluctuations ( $\eta$  - barion to photon ratio).

-  $\eta$  from BBN for present value of  $Q$  ( $Q = |E_d|$ )

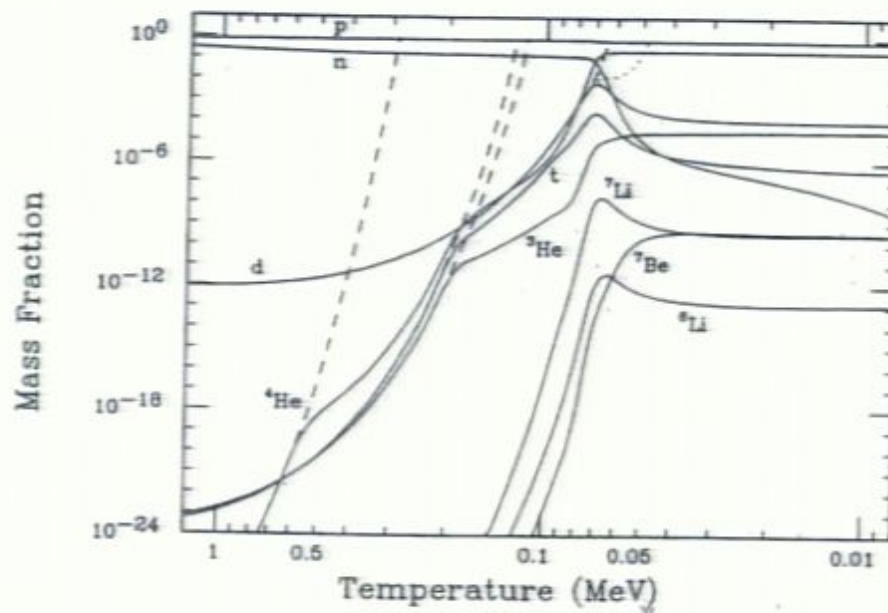


FIG. 2.—Evolution of light-element abundances with temperature, for a baryon-to-photon ratio  $\eta_{10} = 3.16$ . The dashed curves give the NSE curves of  ${}^4\text{He}$ ,  $t$ ,  ${}^3\text{He}$ , and  $d$ , respectively. The dotted curve is explained in the text.

# Deuterium bottleneck

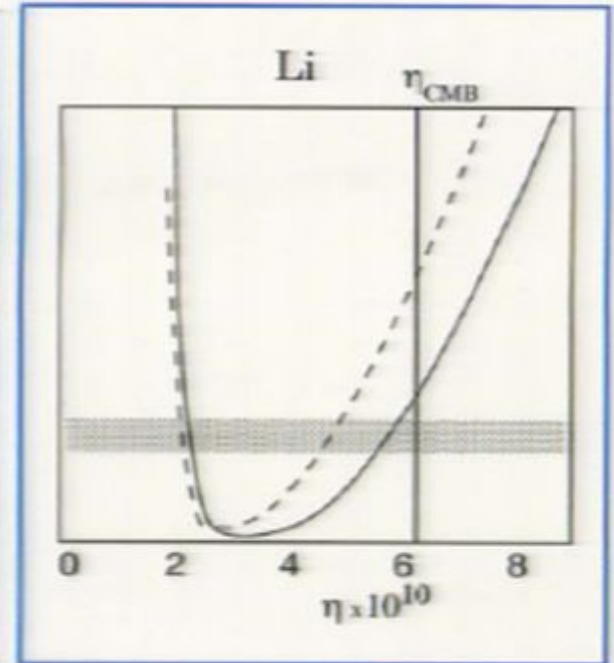
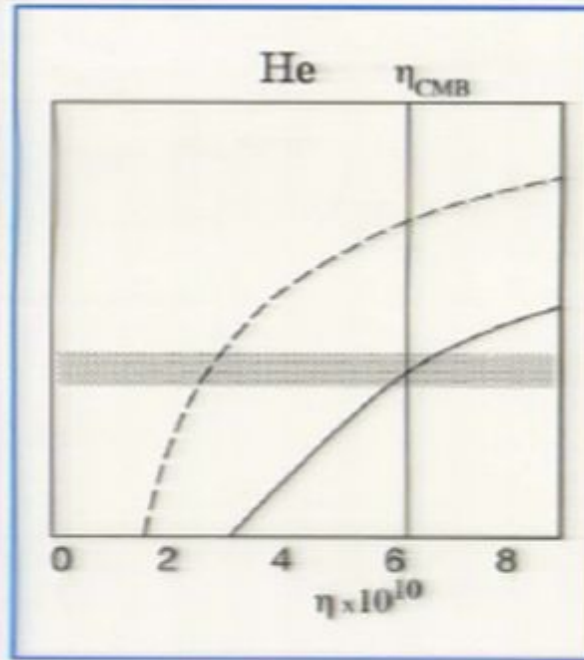
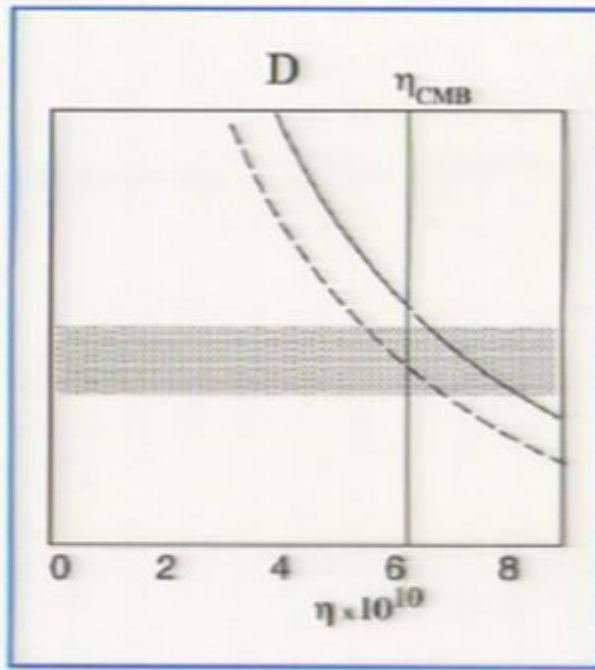
At temperature  $T < 0.3 \text{ MeV}$  all abundances follow deutron abundance

(no other nuclei produced if there are no deuterons)

Reaction  $\gamma d \rightarrow n p$ , exponentially small number of energetic photons,  $e^{-(E_d/T)}$

Exponential sensitivity to deutron binding energy  $E_d$ ,  $E_d = 2 \text{ MeV}$ ,

Freezeout temperature  $T_f = 30 \text{ KeV}$



Comparison with observations gives

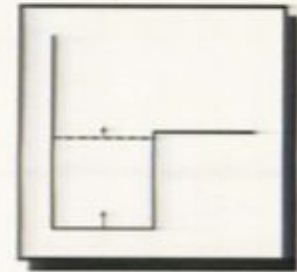
$$\frac{\delta E_d}{E_d} = -0.019 \pm 0.005$$

This also leads to agreement

$$\eta(BBN) \approx \eta(CMB)$$

**Flambaum, Shuryak:** Deuteron Binding Energy is very sensitive to variation of *strange* quark mass (4 factors of enhancement):

1. Deuteron is a shallow bound level.

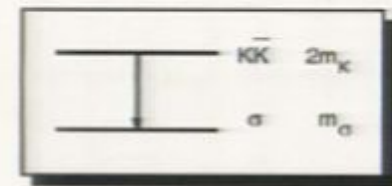


Virtual level in  $n+p \rightarrow d+\gamma$  is even more sensitive to the variation of the potential.

2. Strong compensation between  $\sigma$ -meson and  $\omega$ -meson exchange in potential (Walecka model):  $4\pi r V = -g_s^2 e^{-m_\sigma r} + g_v^2 e^{-m_\omega r}$

3.  $\sigma = \frac{1}{\sqrt{3}}(u\bar{u} + d\bar{d} + s\bar{s})$ ,  $m_\sigma \approx \frac{2}{3}m_s + 2\Lambda_{QCD}$

4. Repulsion of  $\sigma$  from  $K\bar{K}$  threshold



$$\text{Total } \frac{\delta E_d}{E_d} \approx -17 \frac{\delta m_s}{m_s} \quad \text{and} \quad \frac{\delta(m_s/\Lambda_{QCD})}{m_s/\Lambda_{QCD}} = (+1.1 \pm 0.3) \times 10^{-3}$$

# New BBN result

- Dent, Stern, Wetterich 2007; Berengut, Dmitriev, Flambaum 2008: dependence of BBN on energies of  ${}^2,3\text{H}$ ,  ${}^3,4\text{He}$ ,  ${}^6,7\text{Li}$ ,  ${}^7,8\text{Be}$
- Flambaum, Wiringa 2007 : dependence of binding energies of  ${}^2,3\text{H}$ ,  ${}^3,4\text{He}$ ,  ${}^6,7\text{Li}$ ,  ${}^7,8\text{Be}$  on nucleon and meson masses,
- Flambaum, Holl, Jaikumar, Roberts, Write, Maris 2006: dependence of nucleon and meson masses on light quark mass  $m_q$ .

# Big Bang Nucleosynthesis:

## Dependence on $m_q / \Lambda_{\text{QCD}}$

- $^2\text{H}$   $1+7.7x=1.07(15)$   $x=0.009(19)$
- $^4\text{He}$   $1-0.95x=1.005(36)$   $x=-0.005(38)$
- $^7\text{Li}$   $1-50x=0.33(11)$   $x=0.013(02)$

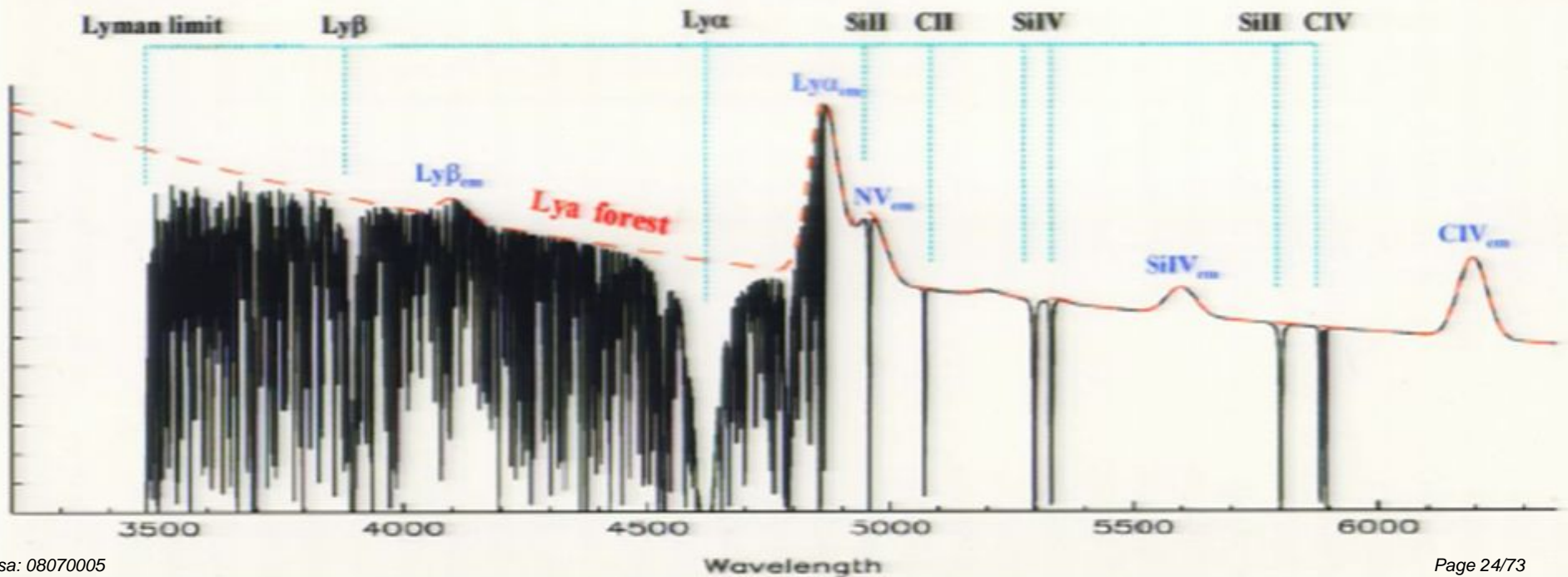
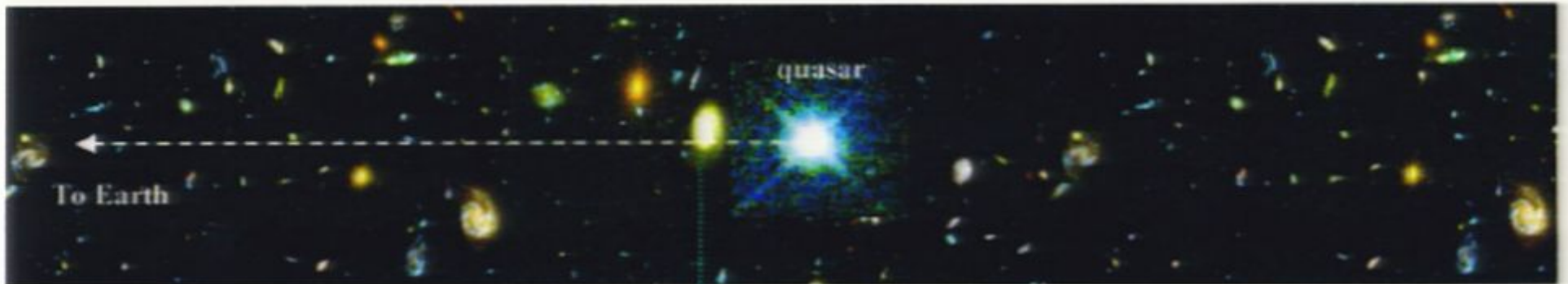
Final result

$$x = \Delta X_q / X_q = 0.013 (02), \quad X_q = m_q / \Lambda_{\text{QCD}}$$

Dominated by  $^7\text{Li}$  abundance (3 times difference), consistent with  $^2\text{H}, ^4\text{He}$

Nonlinear effects:  $x = \Delta X_q / X_q = ?$  Berengut

## 4.2 Astrophysical constraints: Quasars - probing the universe back to much earlier times





# Variation of fine structure constant $\alpha$

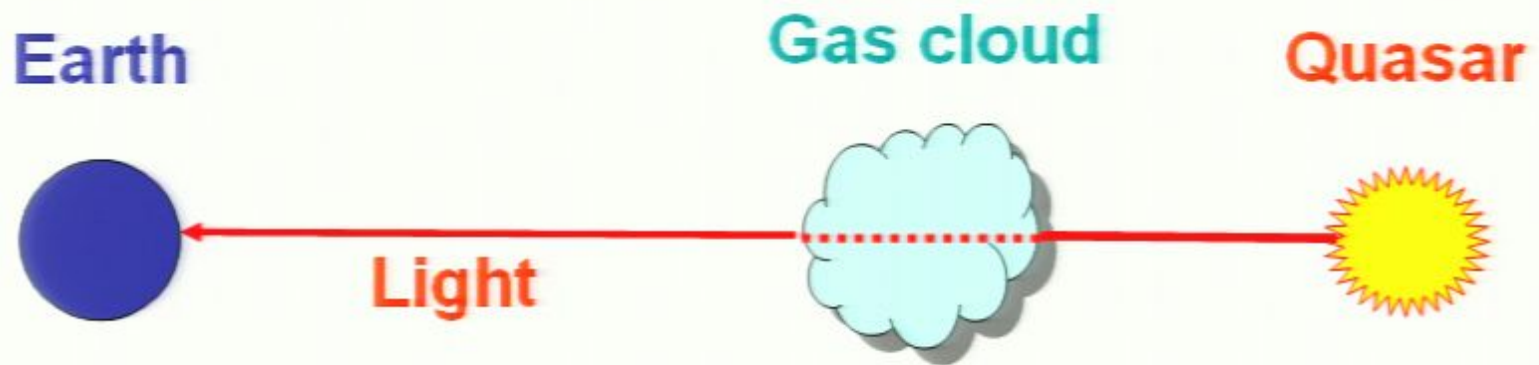
## Many-Multiplet Method

Relativistic correction to electron energy  $E_n$ :

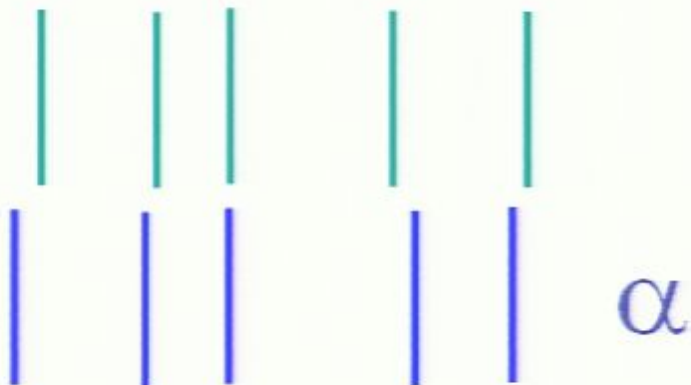
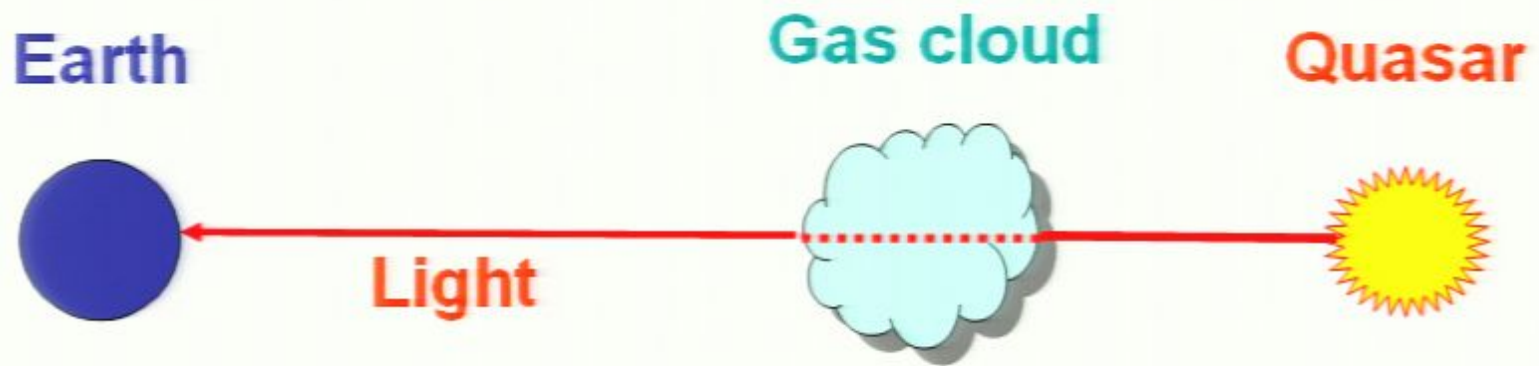
$$\Delta_n = \frac{E_n}{\nu} (Z\alpha)^2 \left[ \frac{1}{j + 1/2} - C(Z, j, l) \right] \quad C \approx 0.6$$

1. Increases with nuclear charge  $Z$ .
2. Changes sign for higher angular momentum  $j$ .

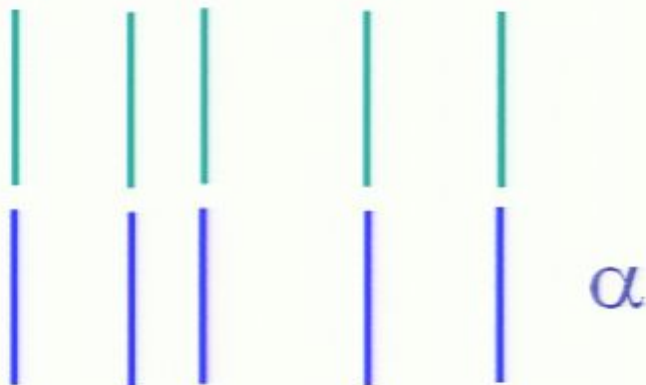
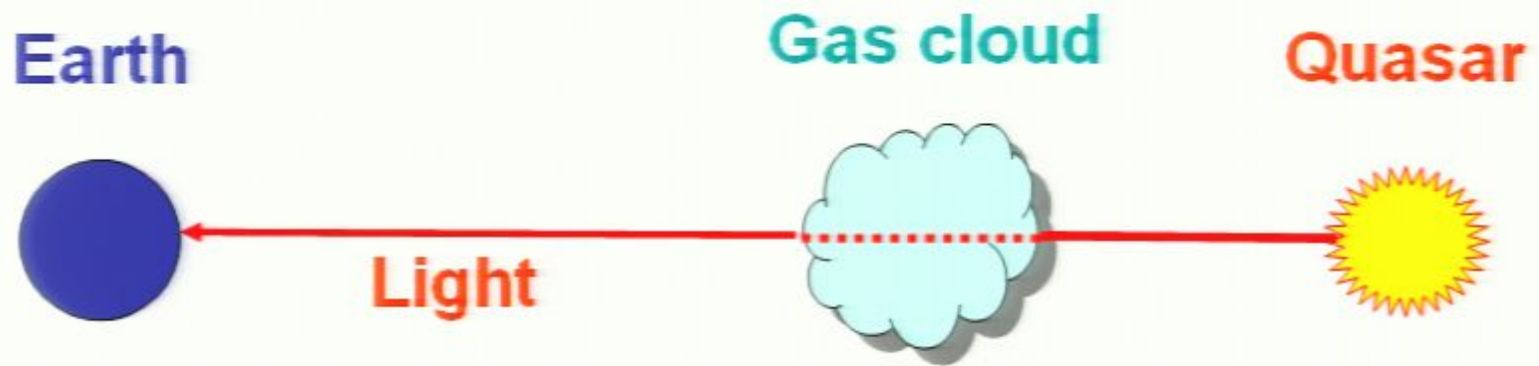
# Quasar absorption spectra



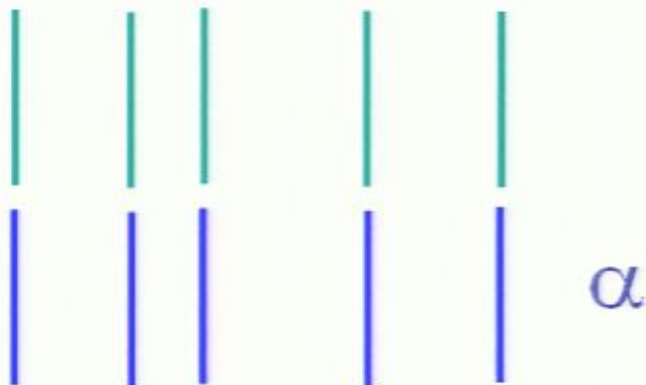
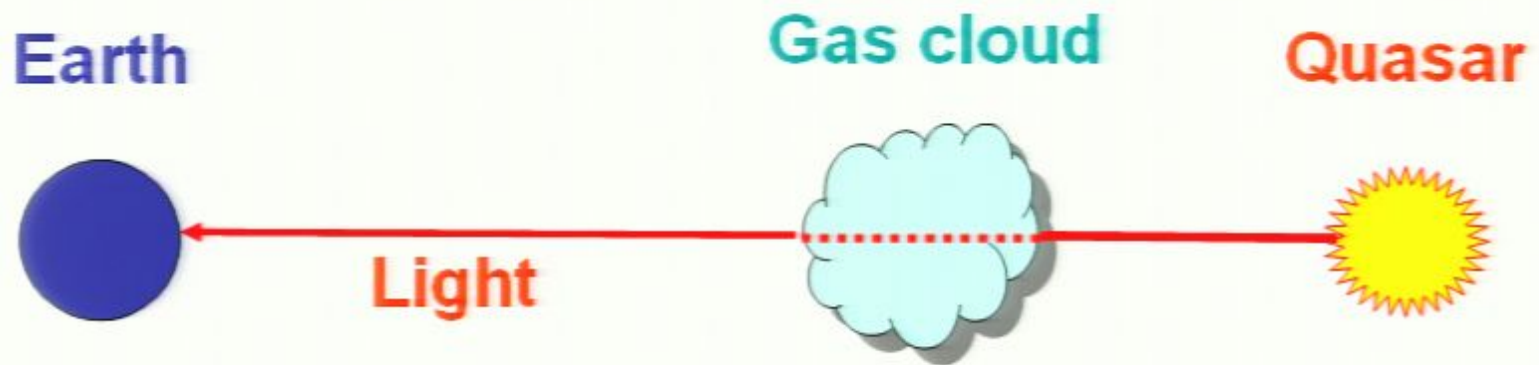
# Quasar absorption spectra



# Quasar absorption spectra



# Quasar absorption spectra



One needs to know  $E(\alpha^2)$  for each line to do the fitting

Use atomic calculations to find  $\omega(\alpha)$ .

For  $\alpha$  close to  $\alpha_0$   $\omega = \omega_0 + q(\alpha^2/\alpha_0^2 - 1)$

$q$  is found by varying  $\alpha$  in computer codes:

$$q = d\omega/dx = [\omega(0.1) - \omega(-0.1)]/0.2, \quad x = \alpha^2/\alpha_0^2 - 1$$

$\alpha = e^2/hc = 0$  corresponds to non-relativistic limit (infinite  $c$ ).

Methods were used for many important problems:

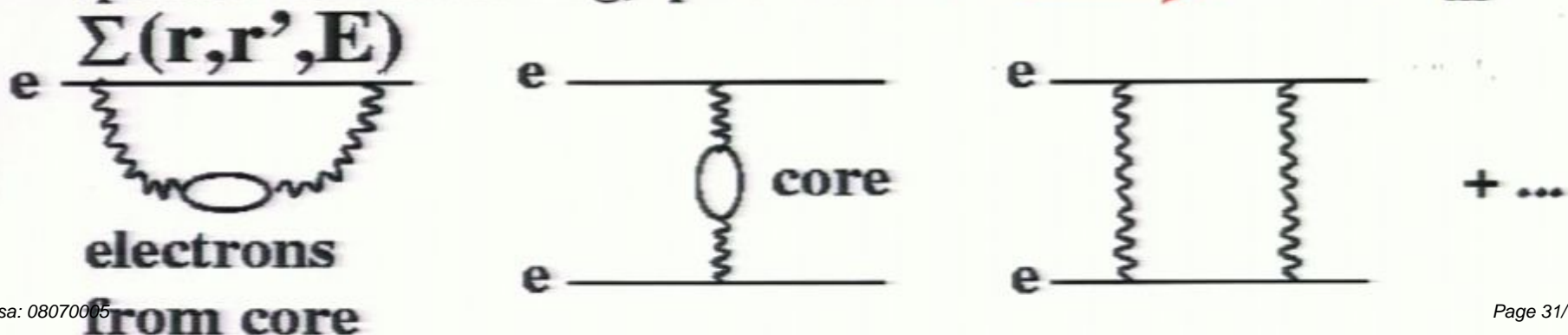
- Test of Standard Model using Parity Violation in Cs, Tl, Pb, Bi
- Predicting spectrum of **Fr (accuracy 0.1%)**, etc.

*Probing the variability of  $\alpha$  with QSO absorption lines*

To find dependence of atomic transition frequencies on  $\alpha$  we have performed calculations of atomic transition frequencies for different values of  $\alpha$ .

1. Zero Approximation – Relativistic Hartree-Fock method: energies, wave functions, Green's functions

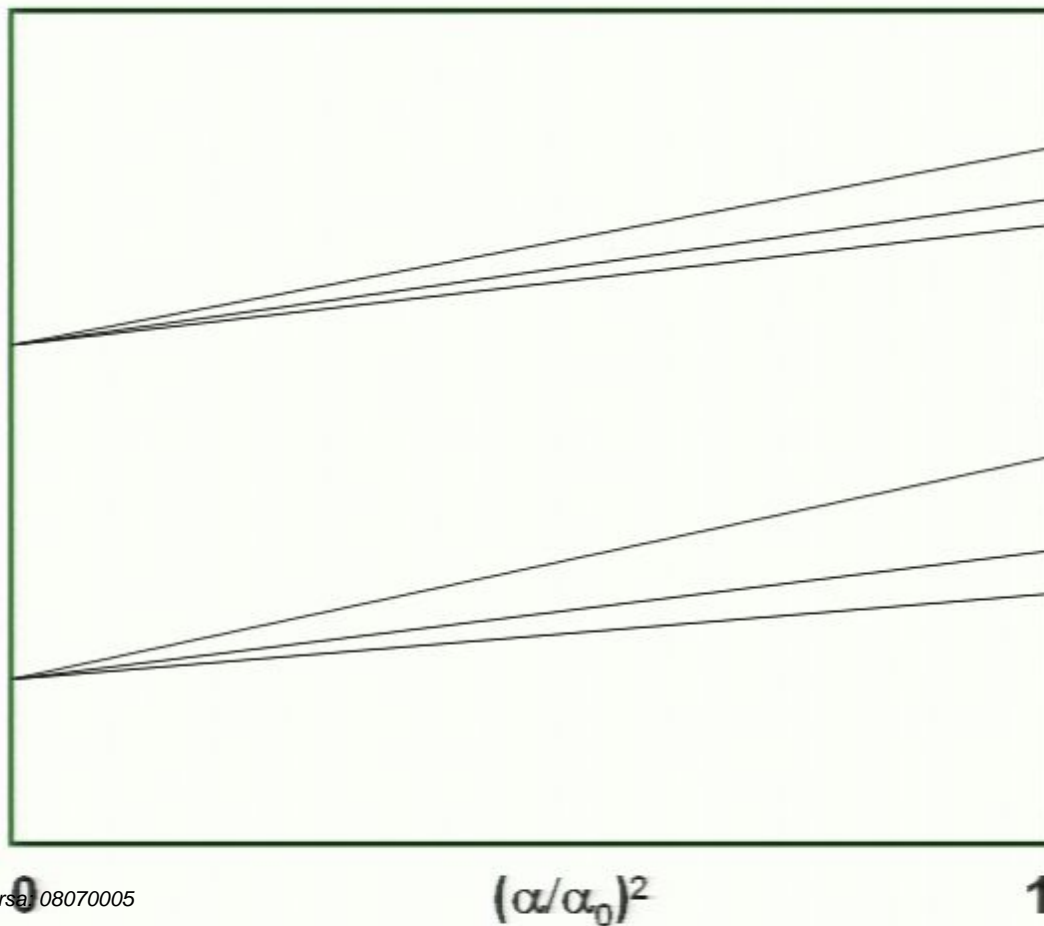
2. Many-body perturbation theory to calculate effective Hamiltonian for valence electrons including self-energy operator and screening; perturbation  $\longrightarrow V = H - H_{\text{HF}}$



3. Diagonalization of the effective Hamiltonian

# Relativistic shifts-triplets

Energies of “normal” fine structure  
triplets as functions of  $\alpha^2$

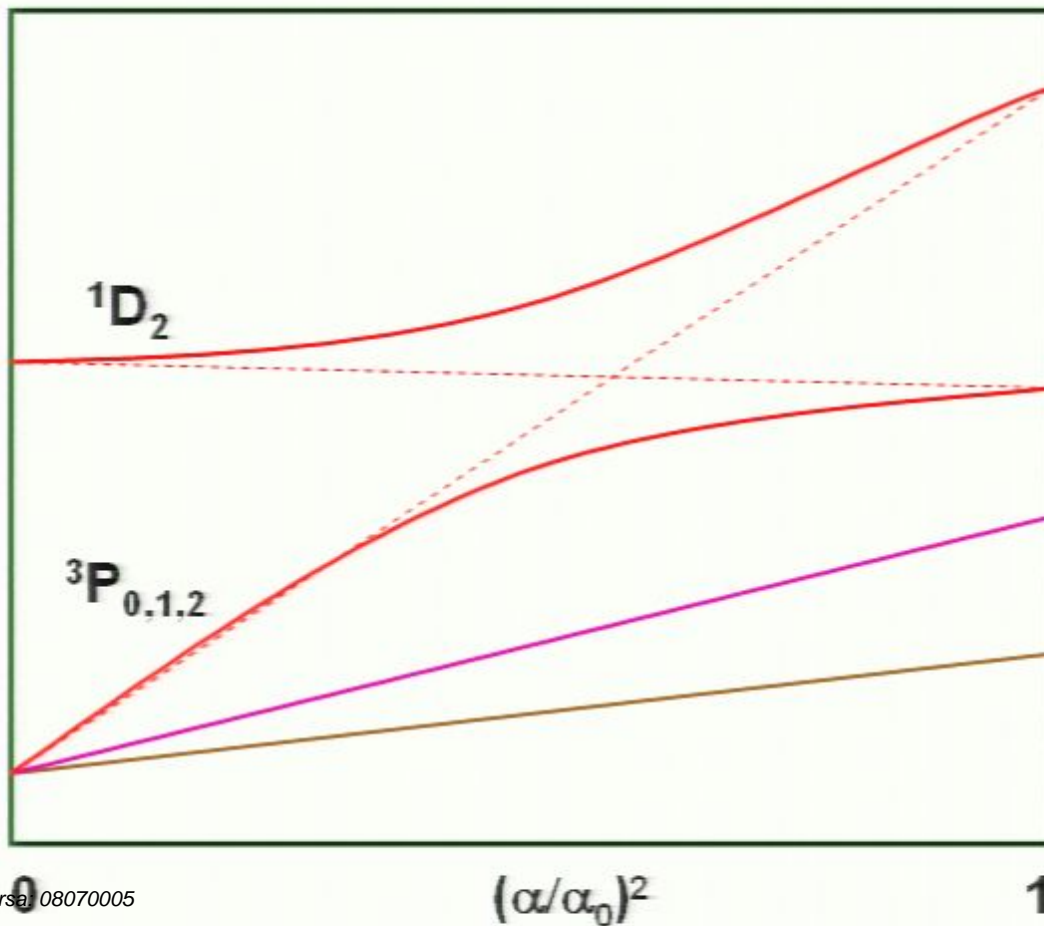


$$\Delta E = A(Z\alpha)^2$$



# Fine structure anomalies and level crossing

Energies of strongly interacting states  
as functions of  $\alpha^2$



~~$\Delta E = A(Z\alpha)^2$~~

# Results of calculations (in $\text{cm}^{-1}$ )

## Anchor lines

Atom	$\omega_0$	$q$
Mg I	35051.217	86
Mg II	35760.848	211
Mg II	35669.298	120
Si II	55309.3365	520
Si II	65500.4492	50
Al II	59851.924	270
Al III	53916.540	464
Al III	53682.880	216
Ni II	58493.071	-20

Also, many transitions in Mn II, Ti II, Si IV, C II, C IV, N V, O I, Ca I, Ca II, Ge II, O II, Pb II

## Negative shifters

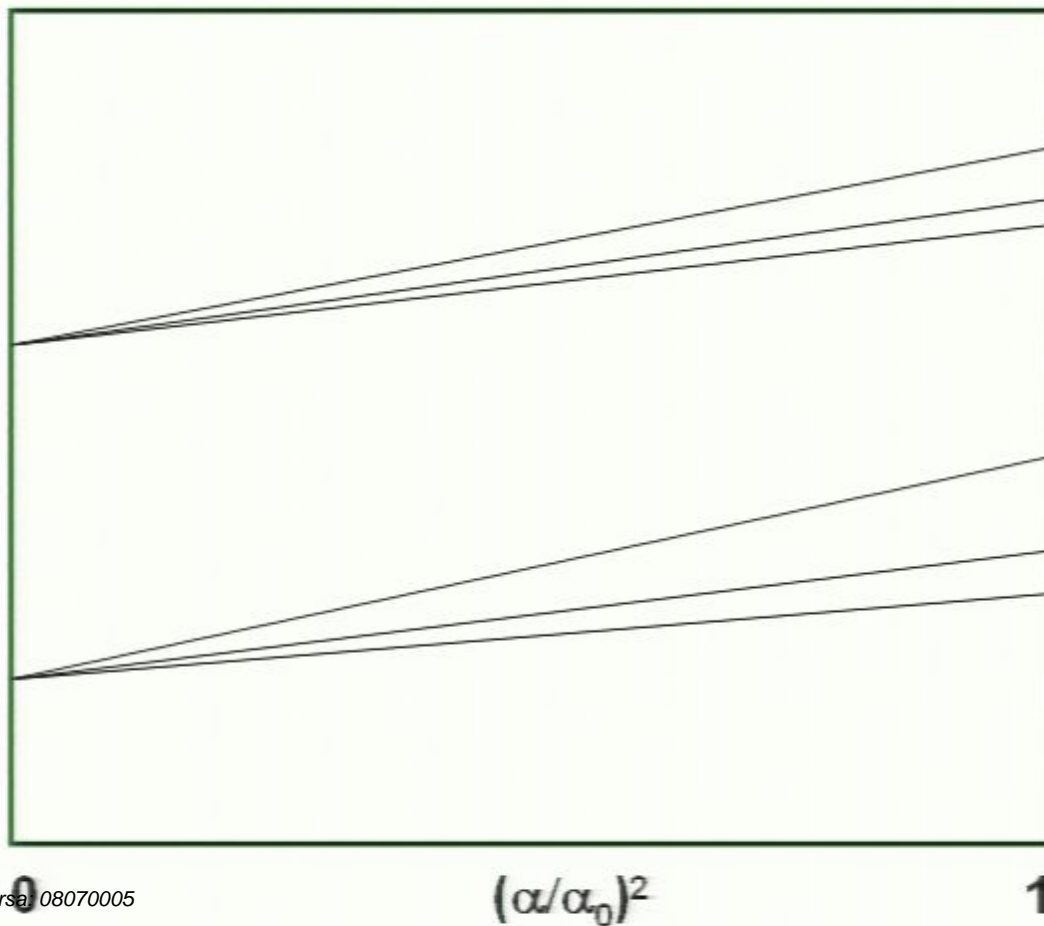
Atom	$\omega_0$	$q$
Ni II	57420.013	-1400
Ni II	57080.373	-700
Cr II	48632.055	-1110
Cr II	48491.053	-1280
Cr II	48398.862	-1360
Fe II	62171.625	-1300

## Positive shifters

Atom	$\omega_0$	$q$
Fe II	62065.528	1100
Fe II	42658.2404	1210
Fe II	42114.8329	1590
Fe II	41968.0642	1460
Fe II	38660.0494	1490
Fe II	38458.9871	1330
Zn II	49355.002	2490
Zn II	48841.077	1584

# Relativistic shifts-triplets

Energies of “normal” fine structure  
triplets as functions of  $\alpha^2$



$$\Delta E = A(Z\alpha)^2$$

# Variation of fine structure constant $\alpha$

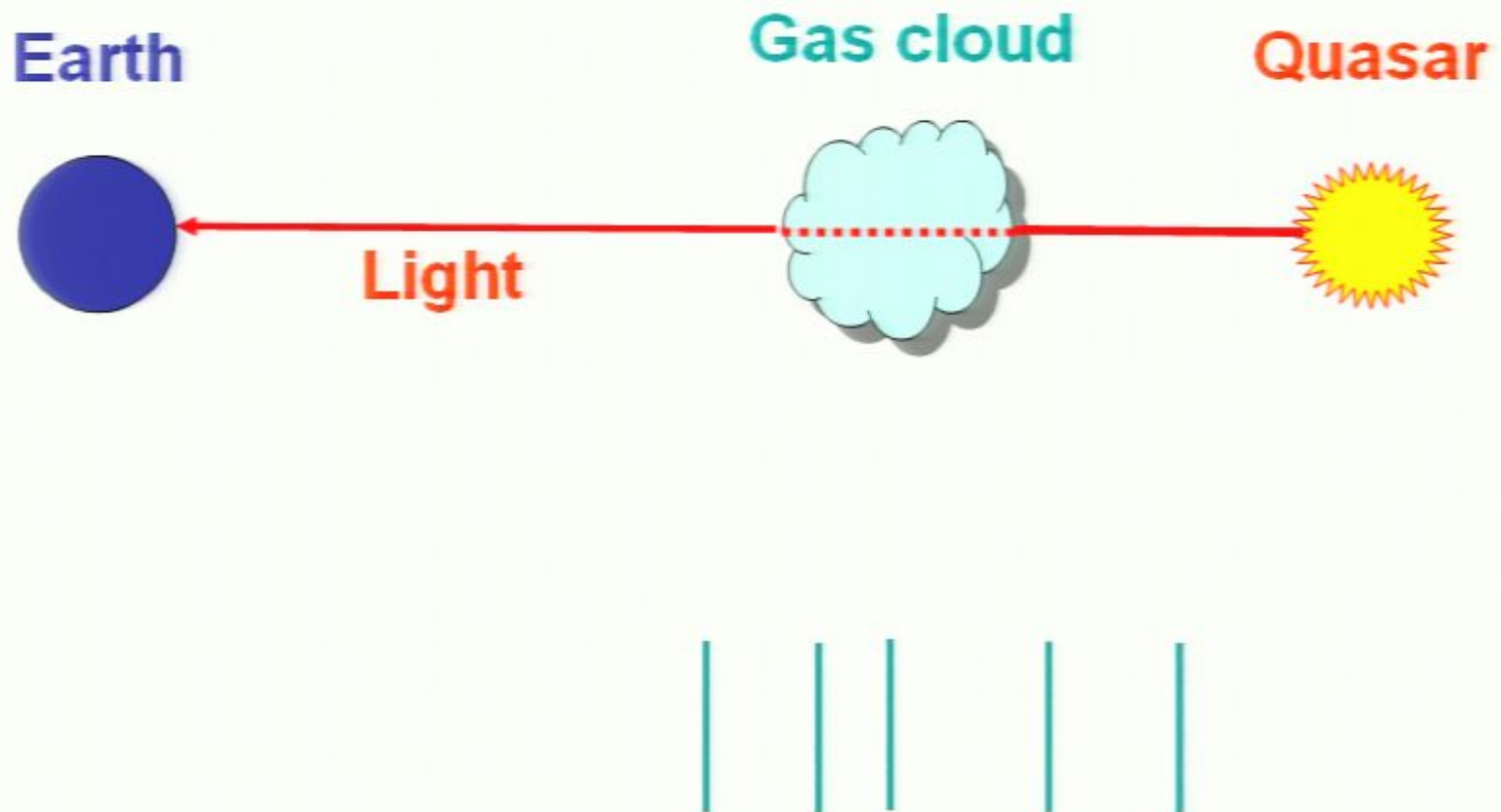
## Many-Multiplet Method

Relativistic correction to electron energy  $E_n$ :

$$\Delta_n = \frac{E_n}{\nu} (Z\alpha)^2 \left[ \frac{1}{j + 1/2} - C(Z, j, l) \right] \quad C \approx 0.6$$

1. Increases with nuclear charge  $Z$ .
2. Changes sign for higher angular momentum  $j$ .

# Quasar absorption spectra



Methods were used for many important problems:

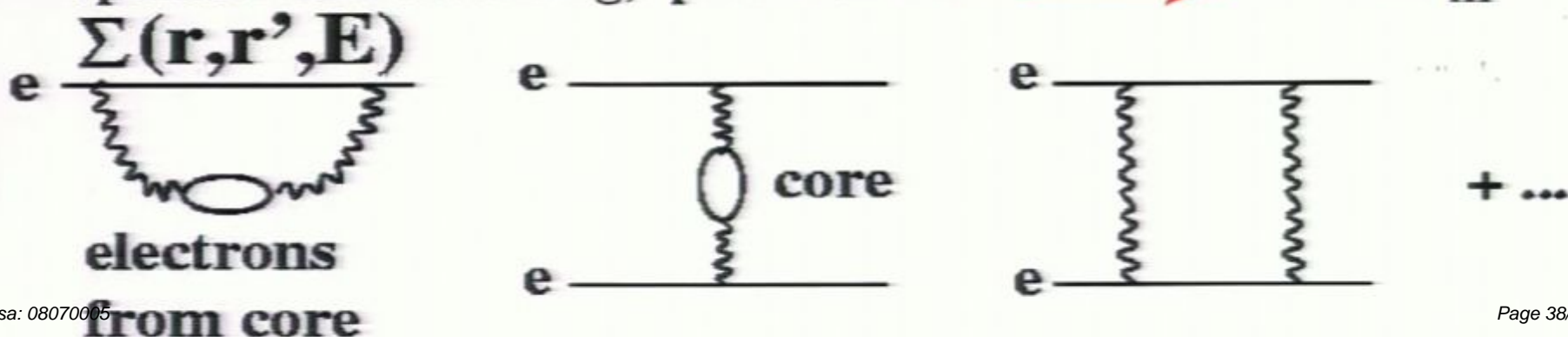
- Test of Standard Model using Parity Violation in Cs, Tl, Pb, Bi
- Predicting spectrum of **Fr (accuracy 0.1%)**, etc.

*Probing the variability of  $\alpha$  with QSO absorption lines*

To find dependence of atomic transition frequencies on  $\alpha$  we have performed calculations of atomic transition frequencies for different values of  $\alpha$ .

1. Zero Approximation – Relativistic Hartree-Fock method: energies, wave functions, Green's functions

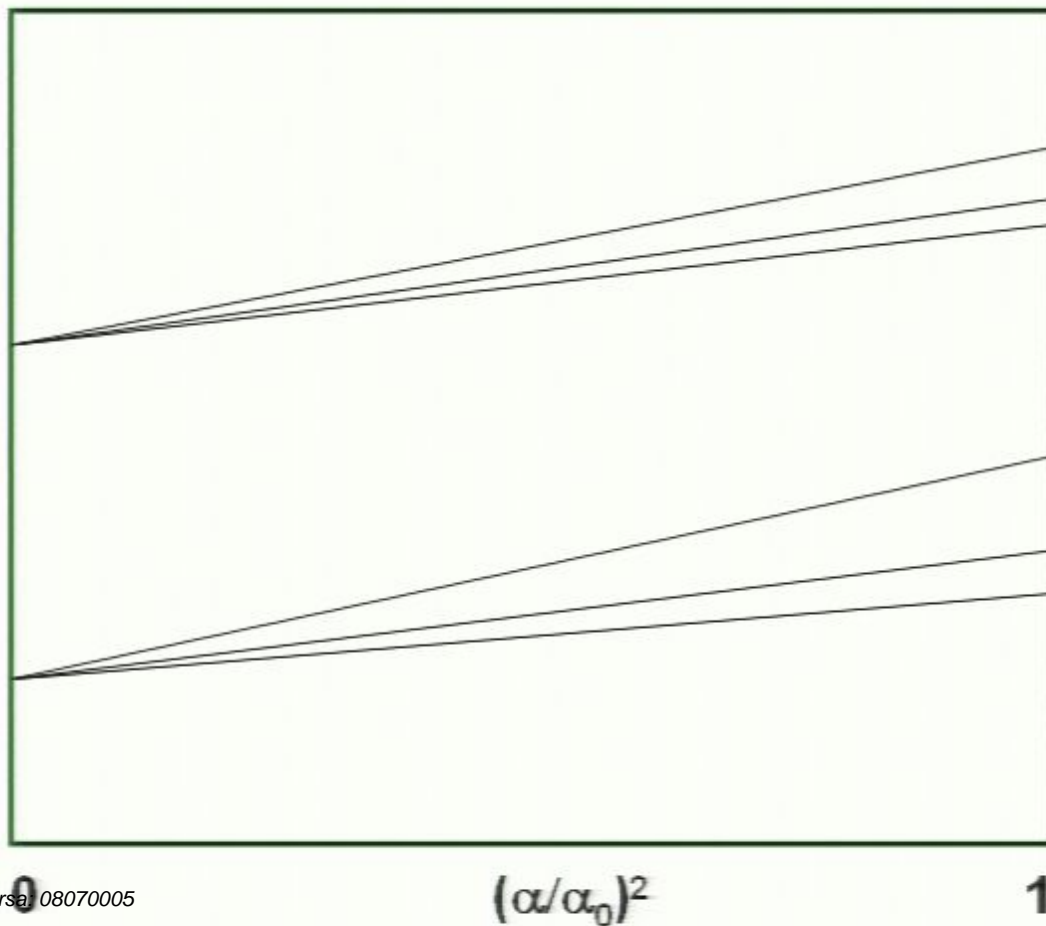
2. Many-body perturbation theory to calculate effective Hamiltonian for valence electrons including self-energy operator and screening; perturbation  $\longrightarrow V = H - H_{\text{HF}}$



3. Diagonalization of the effective Hamiltonian

# Relativistic shifts-triplets

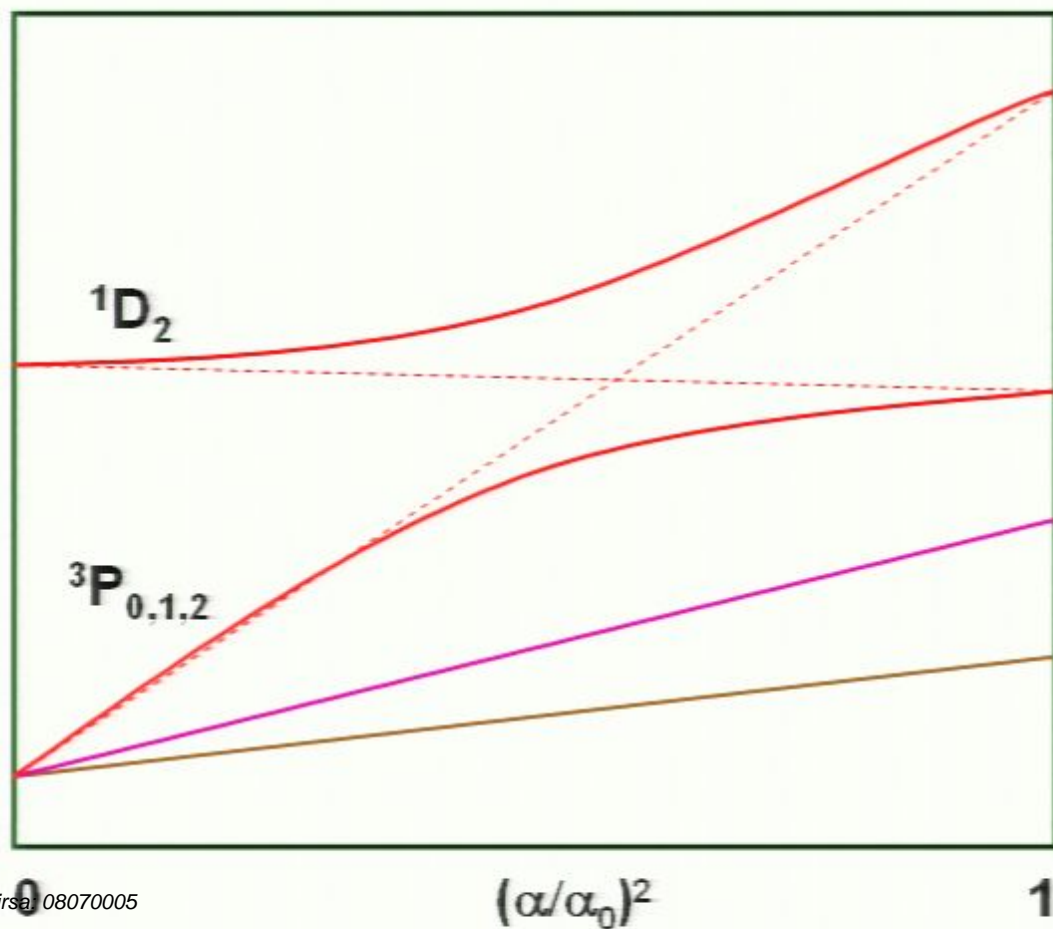
Energies of “normal” fine structure  
triplets as functions of  $\alpha^2$



$$\Delta E = A(Z\alpha)^2$$

# Fine structure anomalies and level crossing

Energies of strongly interacting states  
as functions of  $\alpha^2$



~~$\Delta E = A(Z\alpha)^2$~~



# Results of calculations (in $\text{cm}^{-1}$ )

## Anchor lines

Atom	$\omega_0$	$q$
Mg I	35051.217	86
Mg II	35760.848	211
Mg II	35669.298	120
Si II	55309.3365	520
Si II	65500.4492	50
Al II	59851.924	270
Al III	53916.540	464
Al III	53682.880	216
Ni II	58493.071	-20

Also, many transitions in Mn II, Ti II, Si IV, C II, C IV, N V, O I, Ca I, Ca II, Ge II, O II, Pb II

## Negative shifters

Atom	$\omega_0$	$q$
Ni II	57420.013	-1400
Ni II	57080.373	-700
Cr II	48632.055	-1110
Cr II	48491.053	-1280
Cr II	48398.862	-1360
Fe II	62171.625	-1300

## Positive shifters

Atom	$\omega_0$	$q$
Fe II	62065.528	1100
Fe II	42658.2404	1210
Fe II	42114.8329	1590
Fe II	41968.0642	1460
Fe II	38660.0494	1490
Fe II	38458.9871	1330
Zn II	49355.002	2490
Zn II	48841.077	1584

# Results of calculations (in $\text{cm}^{-1}$ )

## Anchor lines

Atom	$\omega_0$	$q$
Mg I	35051.217	86
Mg II	35760.848	211
Mg II	35669.298	120
Si II	55309.3365	520
Si II	65500.4492	50
Al II	59851.924	270
Al III	53916.540	464
Al III	53682.880	216
Ni II	58493.071	-20

Also, many transitions in Mn II, Ti II, Si IV, C II, C IV, N V, O I, Ca I, Ca II, Ge II, O II, Pb II

## Negative shifters

Atom	$\omega_0$	$q$
Ni II	57420.013	-1400
Ni II	57080.373	-700
Cr II	48632.055	-1110
Cr II	48491.053	-1280
Cr II	48398.862	-1360
Fe II	62171.625	-1300

## Positive shifters

Atom	$\omega_0$	$q$
Fe II	62065.528	1100
Fe II	42658.2404	1210
Fe II	42114.8329	1590
Fe II	41968.0642	1460
Fe II	38660.0494	1490
Fe II	38458.9871	1330
Zn II	49355.002	2490
Zn II	48841.077	1584

# Results of calculations (in $\text{cm}^{-1}$ )

## Anchor lines

Atom	$\omega_0$	$q$
Mg I	35051.217	86
Mg II	35760.848	211
Mg II	35669.298	120
Si II	55309.3365	520
Si II	65500.4492	50
Al II	59851.924	270
Al III	53916.540	464
Al III	53682.880	216
Ni II	58493.071	-20

Also, many transitions in Mn II, Ti II, Si IV, C II, C IV, N V, O I, Ca I, Ca II, Ge II, O II, Pb II

Different signs and magnitudes of  $q$  provides opportunity to study systematic errors!

## Negative shifters

Atom	$\omega_0$	$q$
Ni II	57420.013	-1400
Ni II	57080.373	-700
Cr II	48632.055	-1110
Cr II	48491.053	-1280
Cr II	48398.862	-1360
Fe II	62171.625	-1300

## Positive shifters

Atom	$\omega_0$	$q$
Fe II	62065.528	1100
Fe II	42658.2404	1210
Fe II	42114.8329	1590
Fe II	41968.0642	1460
Fe II	38660.0494	1490
Fe II	38458.9871	1330
Zn II	49355.002	2490
Zn II	48841.077	1584

hyperfine= $\alpha^2 g_p m_e / M_p$  atomic units

Rotation= $m_e/M_p$  atomic units

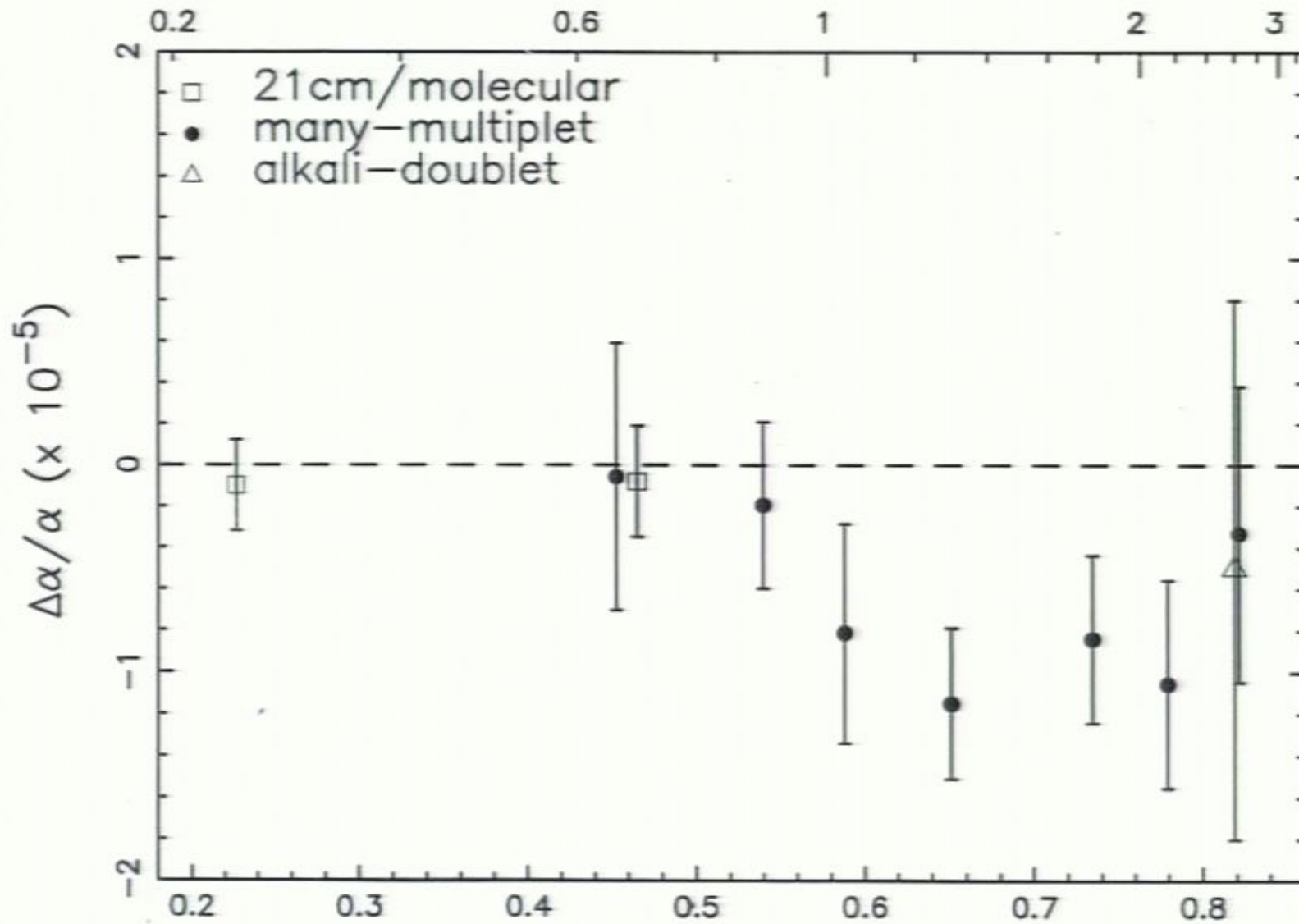
Variation in the fine structure constant?: Recent results and the future

## Radio constraints:

- Hydrogen hyperfine transition at  $\lambda_H = 21\text{cm}$ .
- Molecular rotational transitions CO, HCO<sup>+</sup>, HCN, HNC, CN, CS ...
- $\omega_H/\omega_M \propto \alpha^2 g_p$  where  $g_p$  is the proton magnetic  $g$ -factor.

$$g_p = g_p \left( \frac{m_p}{\Lambda_{QED}} \right)$$

Redshift



Fractional look-back time

Murphy et al, 2003: **Keck telescope**, 143 systems, 23 lines,  $0.2 < z < 4.2$

$$\Delta\alpha/\alpha = -0.54(0.12) \times 10^{-5}$$

• Murphy et al, 2003: **Keck telescope**, 143 systems, 23 lines,  $0.2 < z < 4.2$

$$\Delta\alpha/\alpha = -0.54(0.12) \times 10^{-5}$$

- Quast et al, 2004: **VL telescope**, 1 system, Fe II, 6 lines, 5 positive  $q$ -s, one negative  $q$ ,  $z=1.15$

$$\Delta\alpha/\alpha = -0.4(1.9)(2.7) \times 10^{-6}$$

Molaro et al 2007  $-0.12(1.8) \times 10^{-6}$ ,  $z=1.84$   $5.7(2.7) \times 10^{-6}$

- Srianand et al, 2004: **VL telescope**, 23 systems, 12 lines, Fe II, Mg I, Si II, Al II,  $0.4 < z < 2.3$

$$\Delta\alpha/\alpha = -0.06(0.06) \times 10^{-5}$$

**Murphy et al 2007**  $\Delta\alpha/\alpha = -0.64(0.36) \times 10^{-5}$

Further revision may be necessary.

# Spatial variation (Steinhardt list update)

$$10^5 \Delta\alpha/\alpha$$

Murphy et al

- North hemisphere -0.66(12)
- South (close to North) -0.36(19)

Strianand et al (South) -0.06(06)??

Murphy et al (South) -0.64(36)



---

Measurements  $m_e / M_p$  or  $m_e / \Lambda_{\text{QCD}}$

- Tsanavaris, Webb, Murphy, Flambaum, Curran PRL 2005

Hyperfine H/optical , 9 quasar absorption systems with Mg, Ca, Mn, C, Si, Zn, Cr, Fe, Ni

Measured  $X = \alpha^2 g_p m_e / M_p$

$\Delta X / X = 0.6(1.0)10^{-5}$  **No variation**

# Best limit from ammonia NH<sub>3</sub>

Inversion spectrum: exponentially small “quantum tunneling” frequency  $\omega_{\text{inv}} = W \exp(-S)$

$S = (m_e / M_p)^{-0.5} f(E_{\text{vibration}} / E_{\text{atomic}})$  ,  $E_{\text{vibration}} / E_{\text{atomic}} = \text{const} (m_e / M_p)^{-0.5}$   
 $\omega_{\text{inv}}$  is exponentially sensitive to  $m_e / M_p$

Proposal of lab experiment: ND<sub>3</sub> Veldhoven et al

Flambaum, Kozlov 2007 **Enhanced effect in quasar spectra, 5 times.**

$\Delta(m_e / M_p) / (m_e / M_p) = -0.6(1.9)10^{-6}$  **No variation**  
 $z=0.68$ , 6.5 billion years ago,  $-1(3)10^{-16}$  /year

More accurate measurements Murphy, Flambaum, Henkel,  
Muller. Science 2008  $-0.74(0.47)(0.7)10^{-6}$

Levshakov, Molaro, Kozlov 2008 our Galaxy  $0.5(0.14)10^{-7}$

## Measurements $m_e / M_p$ or $m_e / \Lambda_{\text{QCD}}$

- Reinhold, Buning, Hollenstein, Ivanchik, Petitjean, Ubachs PRL 2006 , H<sub>2</sub> molecule, 2 systems

$$\Delta(m_e / M_p) / (m_e / M_p) = -2.0(0.6) 10^{-5} \quad \text{Variation}$$

**3.5  $\sigma$  !** Higher redshift,  $z=2.8$

Space-time variation? Grand Unification model?

2008 Wendt, Reimers  $< 4.9 \cdot 10^{-5}$

# Best limit from ammonia $\text{NH}_3$

Inversion spectrum: exponentially small “quantum tunneling” frequency  $\omega_{\text{inv}} = W \exp(-S)$

$S = (m_e / M_p)^{-0.5} f(E_{\text{vibration}} / E_{\text{atomic}})$  ,  $E_{\text{vibration}} / E_{\text{atomic}} = \text{const} (m_e / M_p)^{-0.5}$   
 $\omega_{\text{inv}}$  is exponentially sensitive to  $m_e / M_p$

Proposal of lab experiment:  $\text{ND}_3$  Veldhoven et al

Flambaum, Kozlov 2007 **Enhanced effect in quasar spectra, 5 times.**

$\Delta(m_e / M_p) / (m_e / M_p) = -0.6(1.9)10^{-6}$  **No variation**  
 $z=0.68$ , 6.5 billion years ago,  $-1(3)10^{-16}$  /year

More accurate measurements Murphy, Flambaum, Henkel,  
Muller. Science 2008  $-0.74(0.47)(0.7)10^{-6}$

Levshakov, Molaro, Kozlov 2008 our Galaxy  $0.5(0.14)10^{-7}$

## Measurements $m_e / M_p$ or $m_e / \Lambda_{\text{QCD}}$

- Reinhold, Buning, Hollenstein, Ivanchik, Petitjean, Ubachs PRL 2006 , H<sub>2</sub> molecule, 2 systems

$$\Delta(m_e / M_p) / (m_e / M_p) = -2.0(0.6) 10^{-5} \quad \text{Variation}$$

**3.5  $\sigma$  !** Higher redshift,  $z=2.8$

Space-time variation? Grand Unification model?

2008 Wendt, Reimers  $< 4.9 \cdot 10^{-5}$

# Oklo natural nuclear reactor

1.8 billion years ago

$n + {}^{149}\text{Sm}$  capture cross section is dominated  
by  $E_r = 0.1 \text{ eV}$  resonance

Shlyakhter; Damour, Dyson; Fujii et al

$$\Delta E_r = 1 \text{ MeV } \Delta\alpha/\alpha$$

Limits on variation of alpha

# Oklo: limits on $X_q = m_q / \Lambda_{\text{QCD}}$

Flambaum, Shuryak 2002, 2003 Dmitriev, Flambaum 2003  
Flambaum, Wiringa 2007

$$^{150}\text{Sm} \quad \Delta E_r = 10 \text{ MeV} \quad \Delta X_q / X_q - 1 \text{ MeV} \quad \Delta \alpha / \alpha$$

Limits on  $x = \Delta X_q / X_q - 0.1 \Delta \alpha / \alpha$  from

$$\text{Fujii et al} \quad |\Delta E_r| < 0.02 \text{ eV} \quad |x| < 2 \cdot 10^{-9}$$

$$\text{Petrov et al} \quad |\Delta E_r| < 0.07 \text{ eV} \quad |x| < 8 \cdot 10^{-9}$$

$$\text{Gould et al} \quad |\Delta E_r| < 0.026 \text{ eV} \quad |x| < 3 \cdot 10^{-9}, < 1.6 \cdot 10^{-18} \text{ y}^{-1}$$

There is second, non-zero solution  $x = 1.0(1) \cdot 10^{-8}$

# Atomic clocks:

Comparing rates of different clocks over long period of time can be used to study time variation of fundamental constants!

Optical transitions:  $\alpha$

Microwave transitions:  $\alpha, (m_e, m_q)/\Lambda_{\text{QCD}}$



## Calculations to link change of frequency to change of fundamental constants:

Optical transitions: atomic calculations (as for quasar absorption spectra) for many narrow lines in Al II, Ca I, Sr I, Sr II, In II, Ba II, Dy I, Yb I, Yb II, Yb III, Hg I, Hg II, Tl II, Ra II ...

$$\omega = \omega_0 + q(\alpha^2/\alpha_0^2 - 1)$$

Microwave transitions: hyperfine frequency is sensitive to  $\alpha$  (Prestage et al), nuclear magnetic moments (Karshenboim) and nuclear radii

# We performed atomic, nuclear and QCD calculations

of powers  $\kappa, \beta$  for H, D, He, Rb, Cd<sup>+</sup>, Cs, Yb<sup>+</sup>, Hg<sup>+</sup> ...

$$V = C(Ry)(m_e/M_p)\alpha^{2+\kappa} (m_q/\Lambda_{\text{QCD}})^\beta, \quad \Delta\omega/\omega = \Delta V/V$$

$$^{133}\text{Cs}: \kappa = 0.83, \beta = -0.016$$

**Cs standard is insensitive to variation of  $m_q/\Lambda_{\text{QCD}}$ !**

$$^{87}\text{Rb}: \kappa = 0.34, \beta = -0.026$$

$$^{171}\text{Yb}^+: \kappa = 1.5, \beta = -0.136$$

$$^{199}\text{Hg}^+: \kappa = 2.28, \beta = -0.169$$

$$^1\text{H}: \kappa = 0, \beta = -0.100$$

Complete Table in [arxiv:0805.0462](https://arxiv.org/abs/0805.0462)

# Results for variation of fundamental constants

Source	Clock <sub>1</sub> /Clock <sub>2</sub>	$d\alpha/dt/\alpha(10^{-16} \text{ yr}^{-1})$
Blatt <i>et al</i> , 2007	Sr(opt)/Cs(hfs)	-3.1(3.0)
Fortier <i>et al</i> 2007	Hg+(opt)/Cs(hfs)	-0.6(0.7) <sup>a</sup>
Rosenband <i>et al</i> /08	Hg+(opt)/Al+(opt)	-0.16(0.23)
Peik <i>et al</i> , 2006	Yb+(opt)/Cs(hfs)	4(7)
Bize <i>et al</i> , 2005	Rb(hfs)/Cs(hfs)	1(10) <sup>a</sup>

<sup>a</sup>assuming  $m_q/\Lambda_{\text{QCD}} = \text{Const}$

Combined results:  $d/dt \ln\alpha = -1.6(2.3) \times 10^{-17} \text{ yr}^{-1}$

$d/dt \ln(m_q/\Lambda_{\text{QCD}}) = 8(22) \times 10^{-15} \text{ yr}^{-1}$

$m_e/M_p$  or  $m_e/\Lambda_{\text{QCD}} = -1.9(4.0) \times 10^{-16} \text{ yr}^{-1}$

# We performed atomic, nuclear and QCD calculations

of powers  $\kappa, \beta$  for H, D, He, Rb, Cd<sup>+</sup>, Cs, Yb<sup>+</sup>, Hg<sup>+</sup> ...

$$V = C(Ry)(m_e/M_p)\alpha^{2+\kappa} (m_q/\Lambda_{\text{QCD}})^\beta, \quad \Delta\omega/\omega = \Delta V/V$$

$$^{133}\text{Cs}: \kappa = 0.83, \beta = -0.016$$

**Cs standard is insensitive to variation of  $m_q/\Lambda_{\text{QCD}}$ !**

$$^{87}\text{Rb}: \kappa = 0.34, \beta = -0.026$$

$$^{171}\text{Yb}^+: \kappa = 1.5, \beta = -0.136$$

$$^{199}\text{Hg}^+: \kappa = 2.28, \beta = -0.169$$

$$^1\text{H}: \kappa = 0, \beta = -0.100$$

Complete Table in [arxiv:0805.0462](https://arxiv.org/abs/0805.0462)

# Results for variation of fundamental constants

Source	Clock <sub>1</sub> /Clock <sub>2</sub>	$d\alpha/dt/\alpha(10^{-16} \text{ yr}^{-1})$
Blatt <i>et al</i> , 2007	Sr(opt)/Cs(hfs)	-3.1(3.0)
Fortier <i>et al</i> 2007	Hg+(opt)/Cs(hfs)	-0.6(0.7) <sup>a</sup>
Rosenband <i>et al</i> /08	Hg+(opt)/Al+(opt)	-0.16(0.23)
Peik <i>et al</i> , 2006	Yb+(opt)/Cs(hfs)	4(7)
Bize <i>et al</i> , 2005	Rb(hfs)/Cs(hfs)	1(10) <sup>a</sup>

<sup>a</sup>assuming  $m_q/\Lambda_{QCD} = \text{Const}$

Combined results:  $d/dt \ln\alpha = -1.6(2.3) \times 10^{-17} \text{ yr}^{-1}$

$d/dt \ln(m_q/\Lambda_{QCD}) = 8(22) \times 10^{-15} \text{ yr}^{-1}$

$m_e/M_p$  or  $m_e/\Lambda_{QCD} = -1.9(4.0) \times 10^{-16} \text{ yr}^{-1}$

# Dysprosium miracle

Dy:  $4f^{10}5d6s$   $E=19797.96\dots \text{cm}^{-1}$ ,  $q= 6000 \text{ cm}^{-1}$

$4f^95d^26s$   $E=19797.96\dots \text{cm}^{-1}$ ,  $q= -23000 \text{ cm}^{-1}$

Interval  $\Delta\omega = 10^{-4} \text{ cm}^{-1}$



Dzuba, Flambaum, Webb: Enhancement factor  **$K = 10^8$**  (!),  
i.e.  $\Delta\omega/\omega_0 = 10^8 \Delta\alpha/\alpha$

# Dysprosium miracle

Dy:  $4f^{10}5d6s$   $E=19797.96\dots \text{cm}^{-1}$ ,  $q= 6000 \text{ cm}^{-1}$

$4f^95d^26s$   $E=19797.96\dots \text{cm}^{-1}$ ,  $q= -23000 \text{ cm}^{-1}$

Interval  $\Delta\omega = 10^{-4} \text{ cm}^{-1}$



Dzuba, Flambaum, Webb: Enhancement factor  **$K = 10^8$**  (!),  
i.e.  $\Delta\omega/\omega_0 = 10^8 \Delta\alpha/\alpha$

Measurements (Berkeley, Los Alamos)

$$d\ln\alpha/dt = -2.7(2.6) \times 10^{-15} \text{ yr}^{-1}$$

# Dysprosium miracle

Dy:  $4f^{10}5d6s$   $E=19797.96\dots \text{cm}^{-1}$ ,  $q= 6000 \text{ cm}^{-1}$

$4f^95d^26s$   $E=19797.96\dots \text{cm}^{-1}$ ,  $q= -23000 \text{ cm}^{-1}$

Interval  $\Delta\omega = 10^{-4} \text{ cm}^{-1}$



Dzuba, Flambaum, Webb: Enhancement factor  **$K = 10^8$**  (!),  
i.e.  $\Delta\omega/\omega_0 = 10^8 \Delta\alpha/\alpha$

Measurements (Berkeley, Los Alamos)

$$d\ln\alpha/dt = -2.7(2.6) \times 10^{-15} \text{ yr}^{-1}$$

**Problem:** states are not narrow!



## More suggestions ...

Atom	State <sub>1</sub>		State <sub>2</sub>		K
Ce I	<sup>5</sup> H <sub>3</sub>	2369.068	<sup>1</sup> D <sub>2</sub>	2378.827	2000
	<sup>3</sup> H <sub>4</sub>	4762.718	<sup>3</sup> D <sub>2</sub>	4766.323	13000
Nd I	<sup>5</sup> K <sub>6</sub>	8411.900	<sup>7</sup> L <sub>5</sub>	8475.355	950
Nd I	<sup>7</sup> L <sub>5</sub>	11108.813	<sup>7</sup> K <sub>6</sub>	11109.167	10 <sup>5</sup>
Sm I	<sup>5</sup> D <sub>1</sub>	15914.55	<sup>7</sup> G <sub>2</sub>	12087.17	300
Gd II	<sup>8</sup> D <sub>11/2</sub>	4841.106	<sup>10</sup> F <sub>9/2</sub>	4852.304	1800
Tb I	<sup>6</sup> H <sub>13/2</sub>	2771.675	<sup>8</sup> G <sub>9/2</sub>	2840.170	600

# Enhancement in molecular clocks

DeMille 2004, DeMille et al 2008 – enhancement in  $\text{Cs}_2$ , **cancellation between electron excitation and vibration energies**

Flambaum 2006 Cancellations between rotational and hyperfine intervals in very narrow microwave transitions in LaS, LaO, LuS, LuO, YbF, etc.

$$\omega_0 = E_{\text{rotational}} - E_{\text{hyperfine}} = E_{\text{hyperfine}} / 100 - 1000$$

$$\Delta\omega/\omega_0 = K \Delta\alpha/\alpha \quad \text{Enhancement } \mathbf{K = 10^2 - 10^3}$$

# Cancellation between fine structure and vibrations

Flambaum, Kozlov PRL2007  **$K = 10^4 - 10^5$** ,

SiBr,  $\text{Cl}_2^+$  ... microwave transitions between narrow excited states, sensitive to  $\alpha$  and  $\mu = m_e/M_p$

$$\omega_0 = E_{\text{fine}} - E_{\text{vibrational}} = E_{\text{fine}}/K$$

$$\Delta\omega/\omega_0 = K (\Delta\alpha/\alpha - 1/4 \Delta\mu/\mu)$$

Enhancement  **$K = 10^4 - 10^5$**

$E_{\text{fine}}$  is proportional to  $Z^2\alpha^2$

$E_{\text{vibrational}} = n\omega$  is proportional to  $n\mu^{0.5}$ ,  $n=1,2,\dots$

Enhancement for all molecules along the lines  $Z(\mu,n)$

Shift 0.003 Hz for  $\Delta\alpha/\alpha=10^{-16}$ ; **width 0.01 Hz**

Compare with Cs/Rb hyperfine shift  $10^{-6}$  Hz

$\text{HfF}^+$   **$K = 10^3$**  shift 0.1 Hz

# Nuclear clocks

Suggested by Peik, Tamm 2003 Very narrow UV transition between first excited and ground state in  $^{229}\text{Th}$  nucleus Energy  $7.6(5)$  eV, width  $10^{-4}$  Hz

Flambaum 2006; He, Re 2007; Dobaczewski, Feldmayer, Flambaum, Litvinova 2008; Flambaum, Wiringa 2008; Dmitriev, Flambaum 2008

Nuclear/QCD estimate: Enhancement  $10^5$ ,

$$\Delta\omega/\omega_0 = 10^5 (0.1\Delta\alpha/\alpha + \Delta X_q/X_q)$$

$$X_q = m_q / \Lambda_{\text{QCD}}$$

Shift  $10^4$  Hz for  $\Delta\alpha/\alpha = 10^{-16}$

Compare with atomic clock shift 0.1 Hz

**Problem – to find this narrow transition using laser**

**Search: Peik et al, Lu et al, Habs et al, DeMille et al**

$^{235}\text{U}$  energy 76 eV, width  $6 \cdot 10^{-4}$  Hz

# Ultracold atomic and molecular collisions (in Bose condensate).

Cheng Chin, Flambaum PRL2006

Enhancement near Feshbach resonance.

Variation of scattering length

$$\Delta a/a = K \Delta\mu/\mu, \quad K = 10^2 - 10^{12}$$

$$\mu = m_e/M_p$$

Hart, Xu, Legere, Gibble Nature 2007

Accuracy in scattering length  $10^{-6}$

# Dependence of fundamental constants on gravitational potential or scalar potential

Projects –atomic clocks at satellites in space or close to Sun

Earth orbit is elliptic, 3% change in distance to Sun

Fortier et al – Hg<sup>+(opt)</sup>/Cs , Ashby et al -H/Cs

Flambaum, Shuryak : limits on dependence of  $\alpha$ ,  $m_e/\Lambda_{\text{QCD}}$  and  $m_q/\Lambda_{\text{QCD}}$  on gravity

$$\delta\alpha/\alpha = K_\alpha \delta(\mathbf{GM}/rc^2)$$

$$K_\alpha + 0.17K_e = -3.5(6.0) 10^{-7}$$

$$K_\alpha + 0.13 K_q = 2(17) 10^{-7}$$

New results from Dy, Sr/Cs

# Dysprosium $\delta\alpha/\alpha = K_\alpha \delta(GM/rc^2)$

Dy:  $4f^{10}5d6s$   $E=19797.96\dots \text{cm}^{-1}$ ,  $q= 6000 \text{ cm}^{-1}$

$4f^95d^26s$   $E=19797.96\dots \text{cm}^{-1}$ ,  $q= -23000 \text{ cm}^{-1}$

Interval  $\Delta\omega = 10^{-4} \text{ cm}^{-1}$



Enhancement factor  **$K = 10^8$** , i.e.  $\Delta\omega/\omega_0 = 10^8 \Delta\alpha/\alpha$

Measurements Ferrel et al 2007

$$K_\alpha = -8.7(6.6) 10^{-6}$$

$$K_e = 4.9(3.9) 10^{-6} \quad K_q = 6.6(5.2) 10^{-6}$$

# Sr(optical)/Cs comparison : S.Blatt et al 2008

New best limits

$$K_{\alpha} = 2.5(3.1) \cdot 10^{-6}$$

$$K_e = -1.1(1.7) \cdot 10^{-6}$$

$$K_q = -1.9(2.7) \cdot 10^{-6}$$



# Conclusions

- Quasar data: MM method provided sensitivity increase 100 times. Anchors, positive and negative shifters-control of systematics. Keck-variation of  $\alpha$ , VLT-?. Systematics or spatial variation.
- $m_e/M_p$  : hyperfineH/optical,  $\text{NH}_3$  – no variation,  $\text{H}_2$  - variation  $4\sigma$ . Space-time variation? Grand Unification model?
- Big Bang Nucleosynthesis: may be interpreted as a variation of  $m_q/\Lambda_{\text{QCD}}$  ?
- Oklo: sensitive to  $m_q/\Lambda_{\text{QCD}}$ , effect  $< 3 \cdot 10^{-9}$
- Atomic clocks: present time variation of  $\alpha$ ,  $m/\Lambda_{\text{QCD}}$
- Transitions between narrow close levels in atoms and molecules – huge enhancement of the **relative** effect
- $^{229}\text{Th}$  nucleus – **absolute** enhancement ( $10^5$  times larger shift)
- Dependence of fundamental constants on gravitational potential

No variation for small red shift, hints for variation at high red shift