Title: Time and Einstein in the 21st Century: The coolest stuff in the universe

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Abstract: At the beginning of the 20th century Einstein published three revolutionary ideas that changed forever how we view Nature. At the beginning of the 21st century Einstein\'s thinking is shaping one of the key scientific and technological wonders of contemporary life: atomic clocks, the best timekeepers ever made. Such super-accurate clocks are essential to industry, commerce, and science; they are the heart of the Global Positioning System (GPS), which guides cars, airplanes, and hikers to their destinations. Today, atomic clocks are still being improved, using Einstein\'s ideas to cool the atoms to incredibly low temperatures. Atomic gases reach temperatures less than a billionth of a degree above Absolute Zero, without solidifying. Such atoms enable clocks accurate to better than a second in 60 million years as well as both using and testing some of Einstein\'s strangest predictions. This will be a lively, multimedia presentation, including experimental demonstrations and down-to-earth explanations about some of today\'s most exciting science. <kw> Low temperature physics, atomic clock, global positioning system, laser cooling, Bose Einstein condensate, photoelectric effect, Brownian motion, special relativity </kw>

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# Time and Einstein in the 20th Century: the coolest stuff in the universe

Perimeter Institute Waterloo, 4 June 2008

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\$upport: NIST, Office of Naval Research

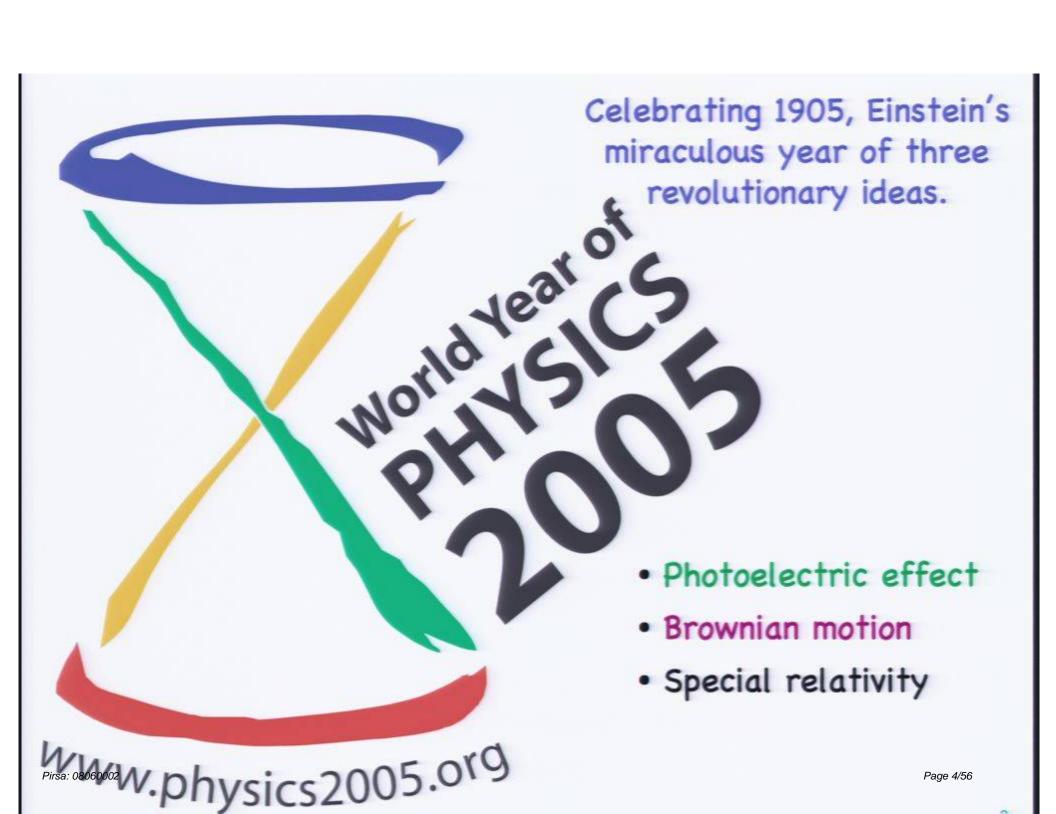






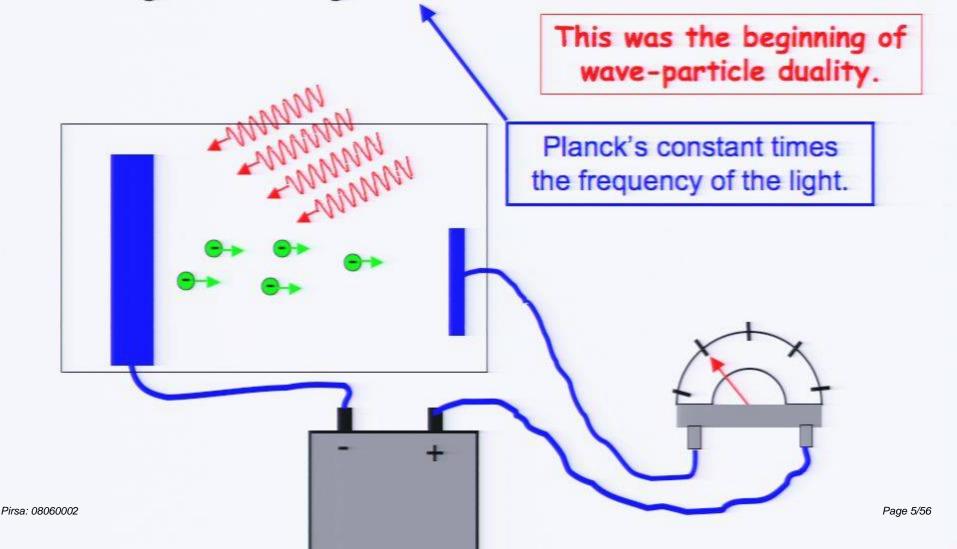
# NIST laser cooling group



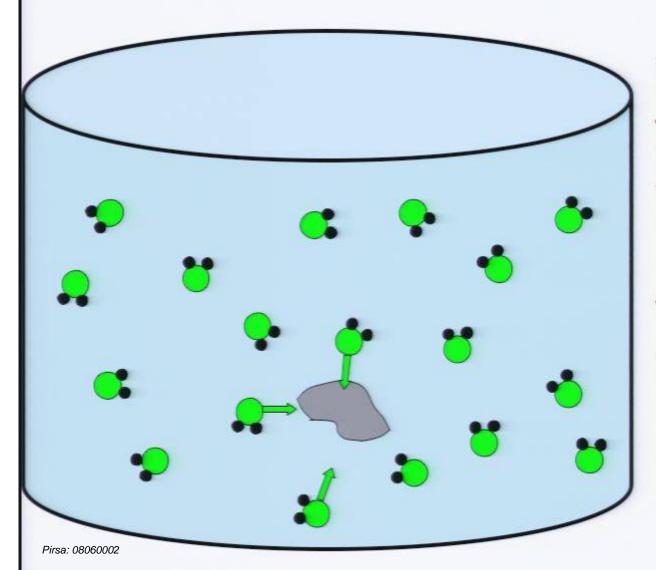


### Photoelectric Effect

Einstein explained the light-induced emission of electrons from a metal as single particles of light (now called photons) delivering their energy, hv, to the electrons in the metal.



### Brownian Movement

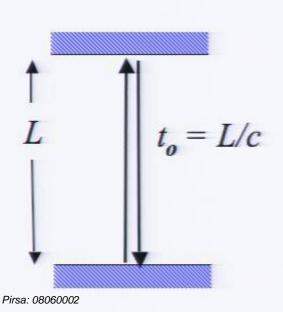


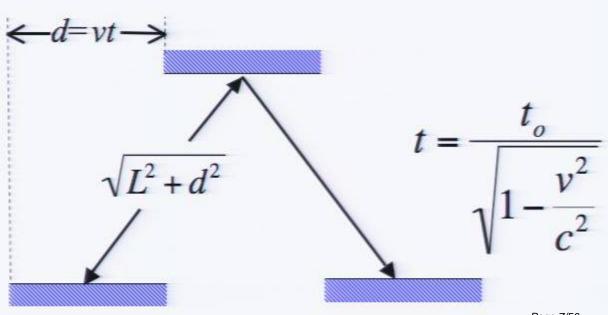
Einstein explained the jiggling of tiny particles suspended in water as coming from the random impacts of the water molecules. This was the most convincing evidence for the atomic/molecular theory of matter.

# Special Relativity

Einstein assumed that the laws of physics are unchanged when you move with constant velocity and that the speed of light in vacuum is always the same. Asking "what is time?" and concluding that "time is what a clock measures" he changed our concepts of both space and time.

Consider a clock whose "tick" is the time for light to go between two mirrors.





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moving with volocity w

at root

#### Like Einstein, we are interested in time and clocks

At NIST, making clocks, standards of time, is our business. Throughout history, clocks have continually improved.









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Harrison's H4 won the famous £20 000 Longitude Prize in the 18th century. The prize required 2 minutes accuracy on a voyage from Great Britain to the West Indies

Trials in 1762 and 1764 won the prize (39.2 s in 47 days in 1764)

Today, most quartz watches are better than H4!



# All of these clocks are imperfect.



The length of a pendulum may stretch or shrink.

Each quartz crystal is a little different from all others, and may change with heat, humidity, ...



Images Copyright Shutterstock



Even the rotation of the earth is slowed by tides and affected by storms and ocean currents.

## Atoms also "tick."



Atoms "vibrate" at specific frequencies, and all atoms of the same kind vibrate at the same frequency.

Every quartz watch is a little different from every other one. Every <sup>133</sup>Cs atom (the atom used to define time) is absolutely identical to every other one in the universe. And atoms are little affected by the environment.

Atom tickers make the best clocks.

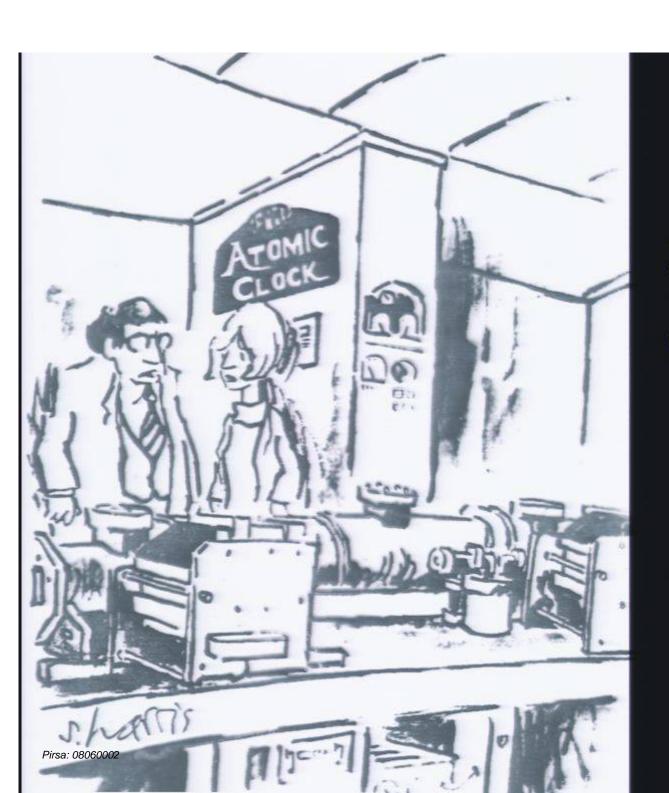
# How good are these clocks?

For less than \$100 you get a quartz watch good to better than 10-6, or 30 seconds/year

For around \$100,000 you can buy an atomic clock good to  $10^{-12}$  or 30 seconds/million years

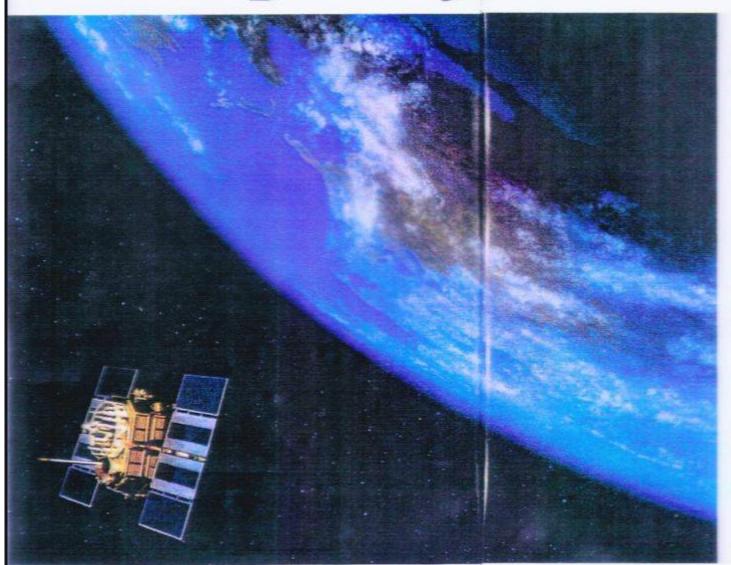
Who needs a clock that good?

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My goodness, it's 12:15:0936420175. Time for lunch

# Relax. Help Is Only 10,000 Miles Away.





Introducing Lincoln Continental's Exclusive Personal Security Package With RESCU.

Ten-thousand miles above the certh, satellites orbit of early through space, waiting for the rewer of a Lincoln Continental who major, sometry require interpretey assistance. With the mask of a beaton, the available RESCU System (Remote Emergency Starffler Cellular Unit) own global posteroming.

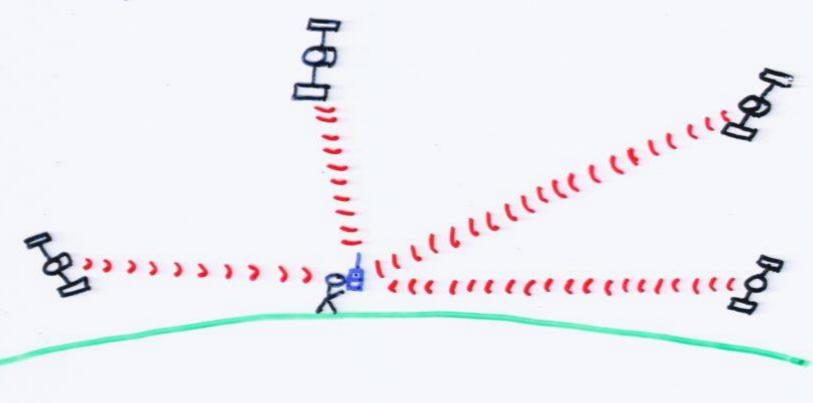
passion is their related we share collecte prove to the Carpolic Security.

Response Carpolic with tropical measurements and their recommendation of the province for their feature declarates for province for the passion. A collecte that where you of the province for and allows you in time up to \$0 miles over after a possion." Further provides Carpolic with go to accordible lengths to being you the obtained to be completed with go to accordible lengths to being you the obtained to be completed at the collection.

# LINCOLN
What A Lunury Car Should Be

Today, super - accurate atomic clocks about a constellation of 24 earth satellites ensure reliable navigation of ships at sea.

plus military and civilian land and air transport.



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### Einstein and GPS

When Einstein first formulated the theory of General Relativity it was considered so complicated and esoteric that the press claimed only three people in the world understood it.

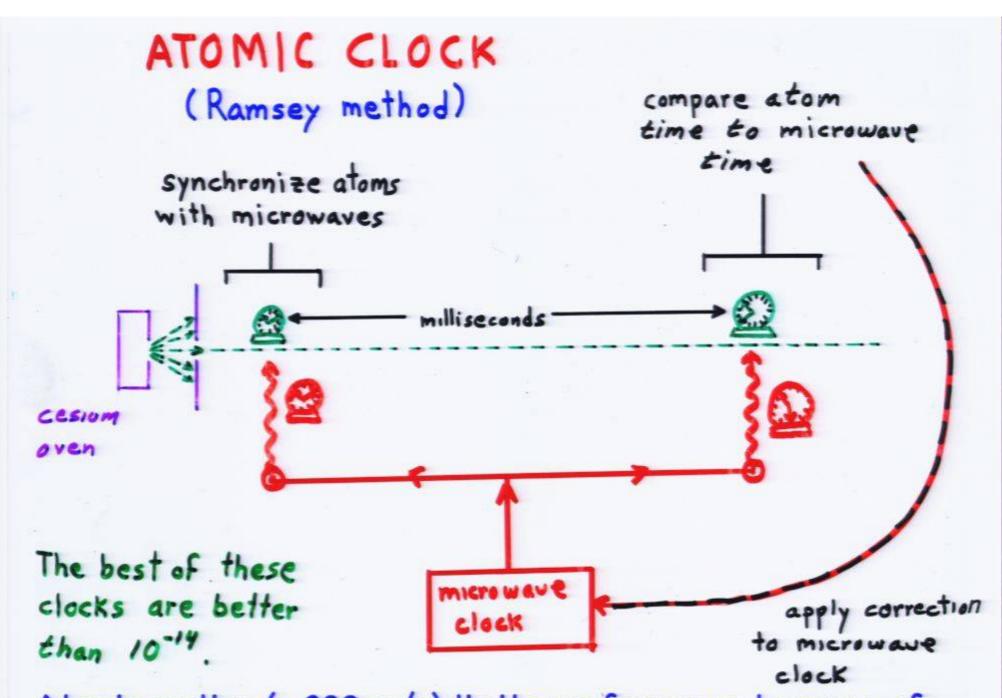
Today, without an understanding of the results of General Relativity, the Global Positioning System would be off by kilometers.

### Einstein and GPS

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Today, without an understanding of the results of General Relativity, the Global Positioning System would be off by kilometers.

General Relativity allows you to find your car in a crowded parking lot.



Atomic motion (~200 m/s) limits performance because performance perfor

#### Doppler Effect



clock, broadcasting "ticks" at frequency



sees ticks at
f (1+ 1/6)

p
speed of light

whether the receiver moves or the transmitter moves, there is a Doppler shift.

This effect allows police to measure the speed of your car.

radar transmitter

radar receiversees a higher
frequency

## Fractional Doppler Shift

$$\frac{\Delta f}{f} = \frac{v}{c}$$
Speed of light

$$V_{\rm thermal} \cong 300 \ {\rm m/s}$$
  
 $c = 3 \times 10^8 \ {\rm m/s}$   
So,  $\frac{\Delta f}{f} = 1 \times 10^{-6}$  A disaster!

Various tricks reduce this a lot, but do not eliminate it.

#### Time dilation shift

Einstein's theory of relativity tells us that moving clocks run slow (recall the twin "paradox")

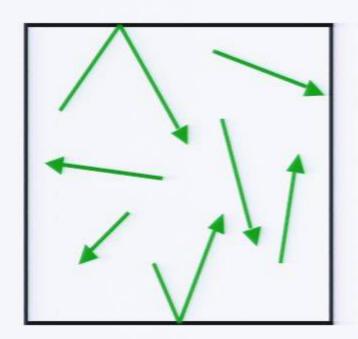
$$\frac{\Delta f}{f} = -\frac{1}{2} \frac{v^2}{c^2} \approx 10^{-12}$$

There are no tricks to get rid of this shift.

We must reduce the velocity!

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### Hot and Cold

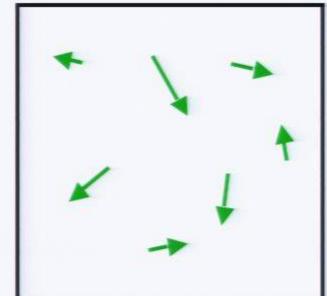


HOT:

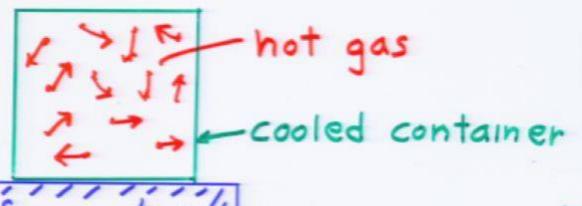
fast atoms

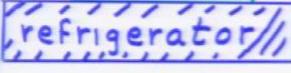
COLD:

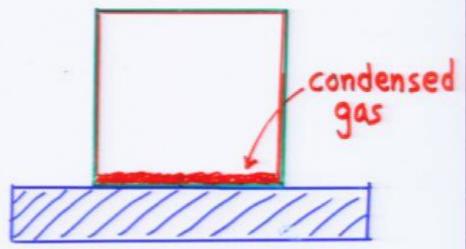
slow atoms



We cannot use conventional refrigeration to cool a gas to millikelvin temperatures:







How do we cool something without touching it?

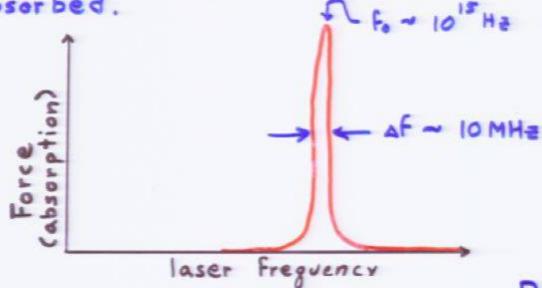




#### Resonance

Light exerts a force on atoms.

but only when absorbed; only light of specific frequencies (color) is absorbed.



Doppler Shift

2 mm

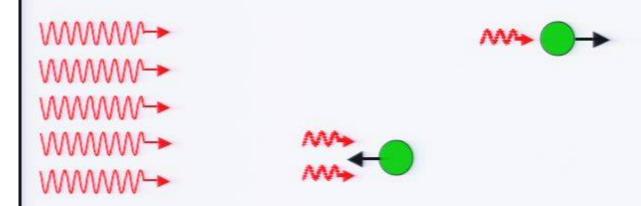
observer moving toward a light

source sees it as having Page 26/56

higher frequency (blue-)

# Laser Cooling (1975)

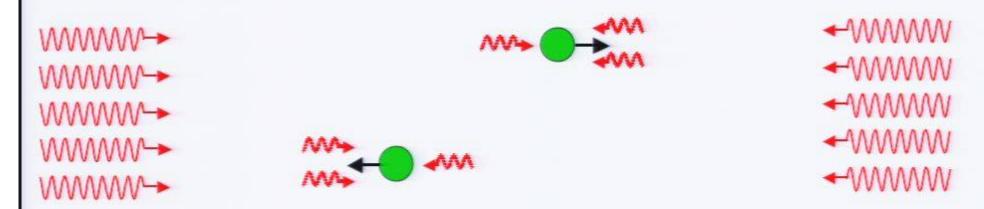
Wineland & Dehmelt and Hänsch & Schawlow



laser beam tuned below resonance

# Laser Cooling (1975)

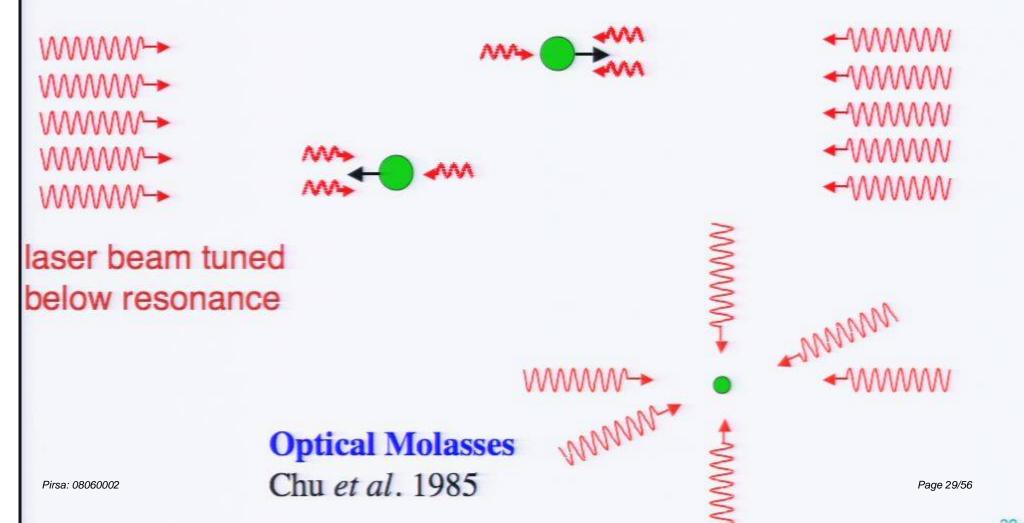
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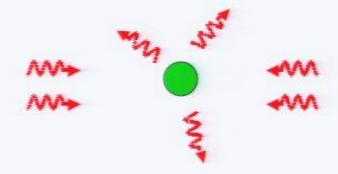
# Laser Cooling (1975)

Wineland & Dehmelt and Hänsch & Schawlow



## Laser Heating

Letokhov, Minogin, and Pavlik (1977); Wineland and Itano (1979)



Randomness of absorption and emission HEATS the atoms.

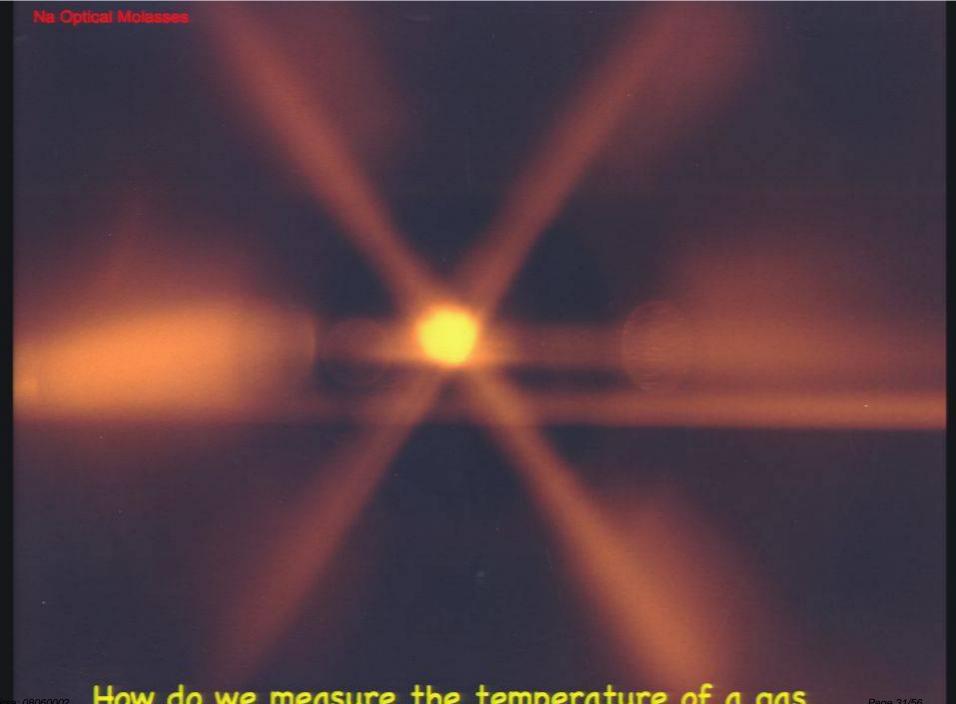
Doppler COOLING balances the heating, producing equilibrium at a temperature, TDODD

"Doppler Limit"  $k_B T_{Dopp} > \hbar \Gamma$ 

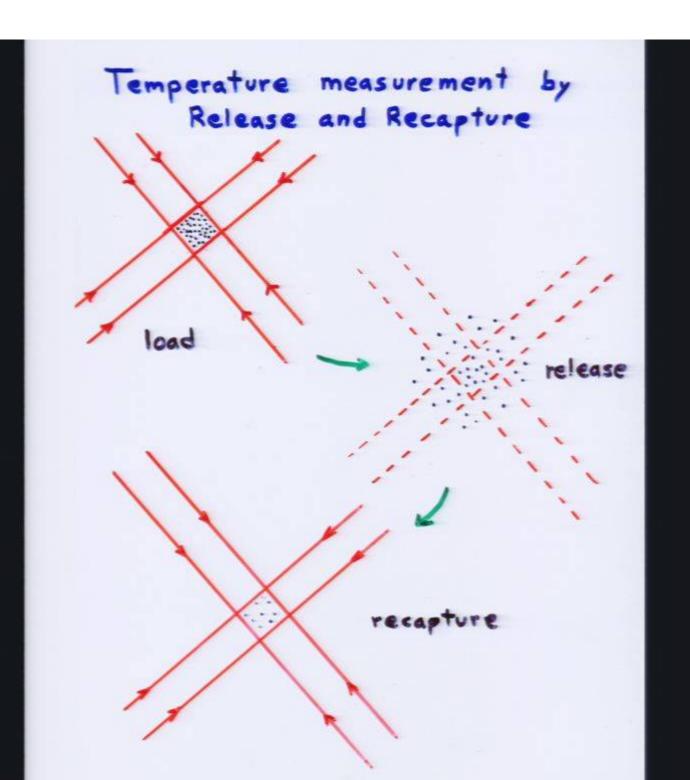
$$k_{\rm B} T_{\rm Dopp} > \hbar \Gamma$$

 $=\hbar\omega$ 

For Na,  $T_{\text{Dopp}} \geq 240 \,\mu\text{K}$ ;  $V_{\text{thermal}}$ 



How do we measure the temperature of a gas that is supposed to be as cold as 2/0 uk2



Laser-Cooling Temperatures by Release-and-Recapture

#### Bell Labs (1985):

S. Chu, L. Hollberg, J. Bjorkholm, A. Cable, Art Ashkin

T = 240 - 60 UK

NBS-Gaithersburg (1987)

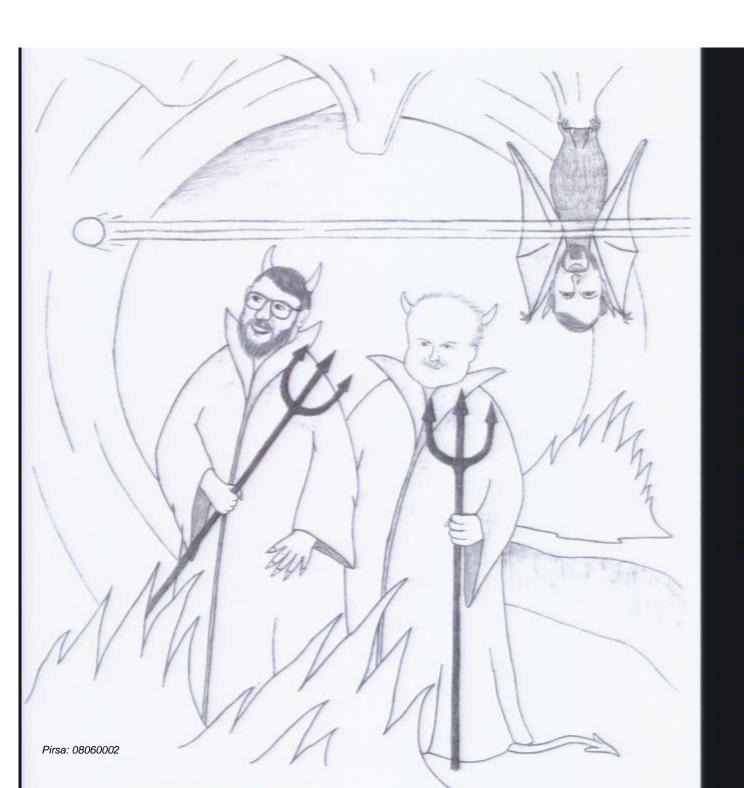
T = 240 MK

other measurements were consistent with Doppler-cooling theory....

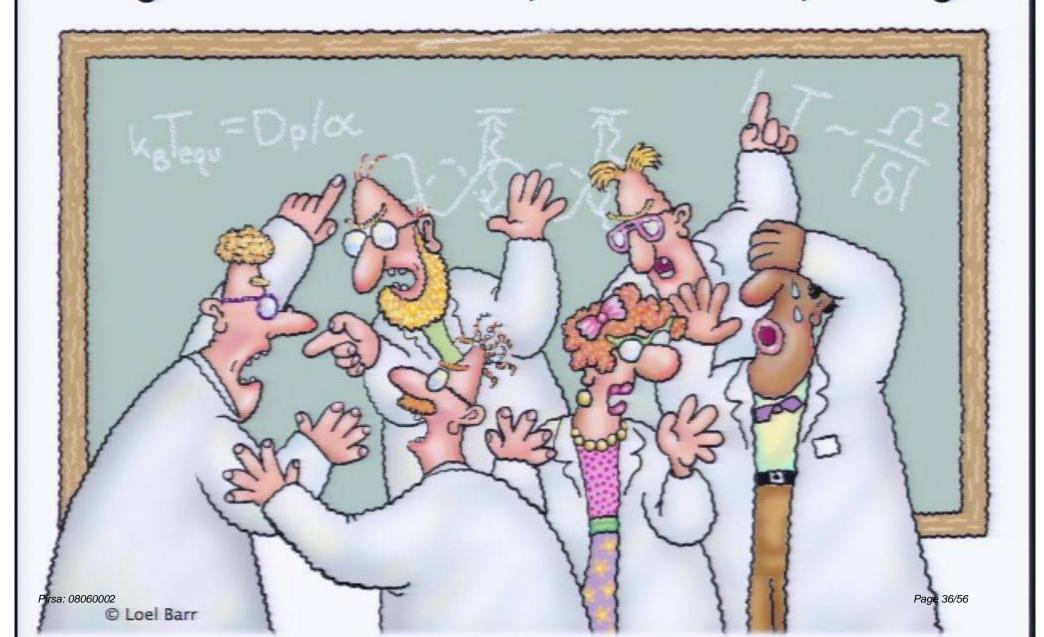
until ...

Unexpectedly, we discovered in 1988 that the temperatures could be much colder than had been predicted.

This was an astounding violation of Murphy's Law



"Look Hal, another snowball !!... I tell you, this place is slipping." Heated discussions about the nature of laser cooling ensued. Eventually a new theory emerged.



#### How cold can we get?

By 1995 we had cooled Cs atoms to

T = 700 nK! (about 200 times colder than expected)

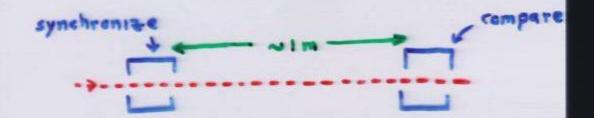
(The lowest temperature is related to the "recoil" energy)

- A hundred million times colder than liquid nitrogen
- 4 million times colder than outer space (3 K cosmic background)

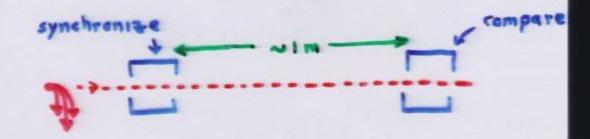
The thermal velocity of these atoms is

 $V_{thermal} < 1 cm/s$ 

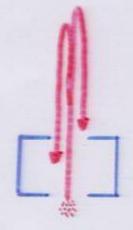
What sort of clock can we make with atoms this cold 20 37/56



Zacharias



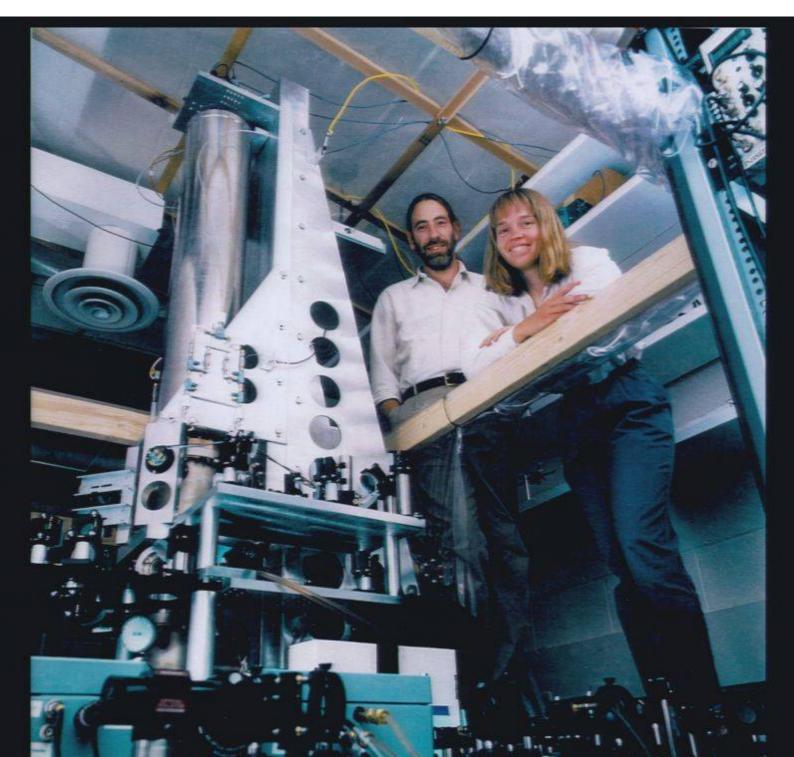
Idea of Zacharias ca 1953
"Atomic Fountain"



Synchronize on the way up

Compare on the way down

Early Fountain: Stanford 1959
"Zacharias" Fountain: Paris 199



Atomic fountain clocks are the most accurate primary frequency standards ever made.

At  $4 \times 10^{-16}$  fractional uncertainty, they are accurate to one second in 80 million years! (and getting better)

Atomic fountain clocks are the most accurate primary frequency standards ever made.

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(and getting better)

Close enough for government work!

The low temperature achieved with laser cooling has been great for clocks! It also brings us into contact with another Pirsa: Amazing prediction of Einstein.

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In 1924 Einstein predicted that if a gas of "bosons," for example:

- Na atoms
- Rb atoms

were cold enough and dense enough, something weird and wonderful would happen -

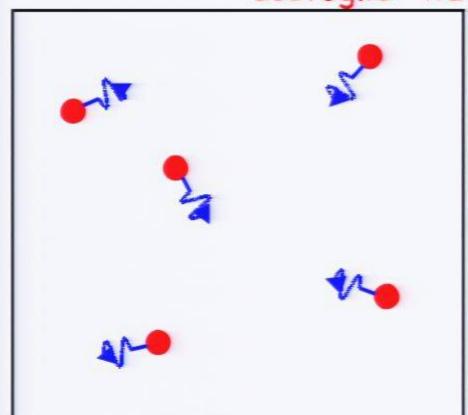
#### Bose-Einstein Condensation:

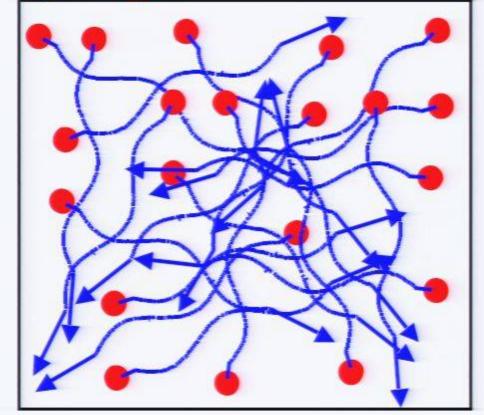
A phase transition where a large fraction of the atoms stop moving! (or, at least as much as the Heisenberg uncertainty principle allows)

#### How cold and how dense?

(N.B.: This was not the way Einstein thought about it)

Recall that all material particles have a wavelike nature, with "deBroglie" wavelength  $\lambda_{dB} = h/mv$ 





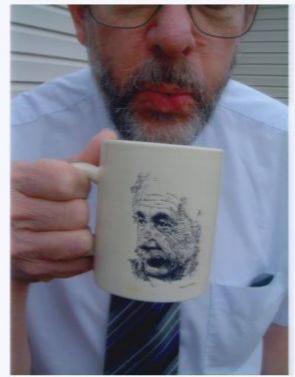
In a hot, dilute gas, λ<sub>dB</sub> is so short that it makes no difference.

In a cold, dense gas, the wavelength can become comparable to the average

# How do we get that cold?

Laser cooling is a good start, but doesn't get cold or dense enough

To get colder and denser, we use evaporative cooling



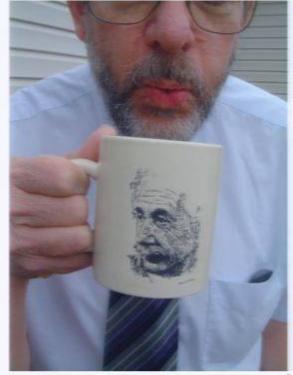
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### How do we get that cold?

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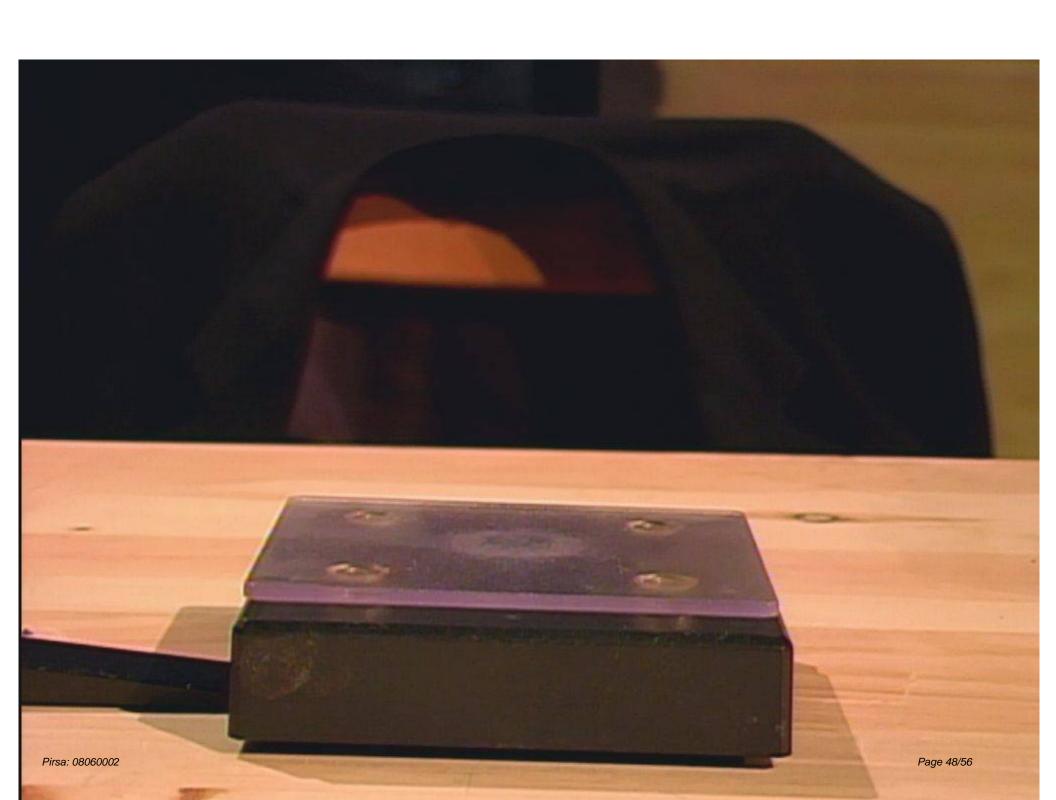
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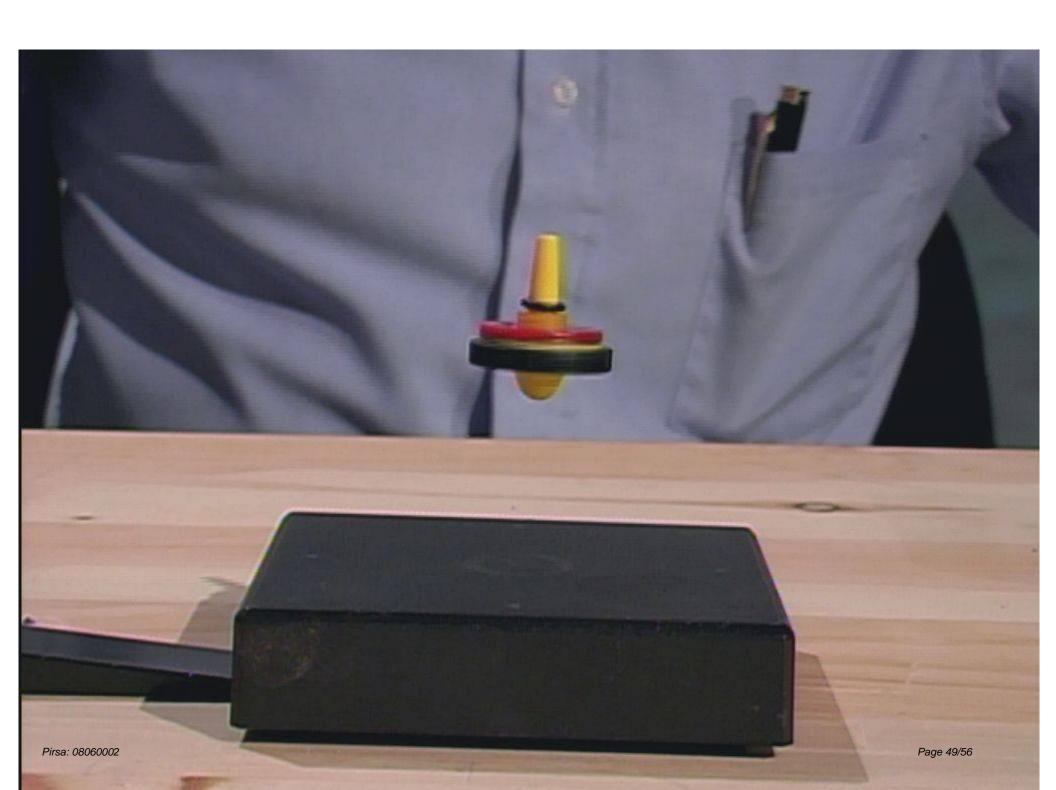
But what do we use for a container?



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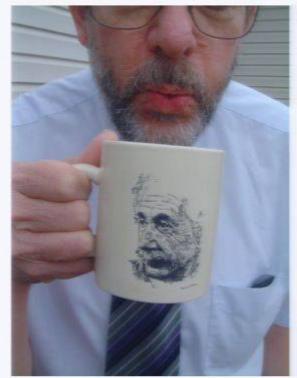


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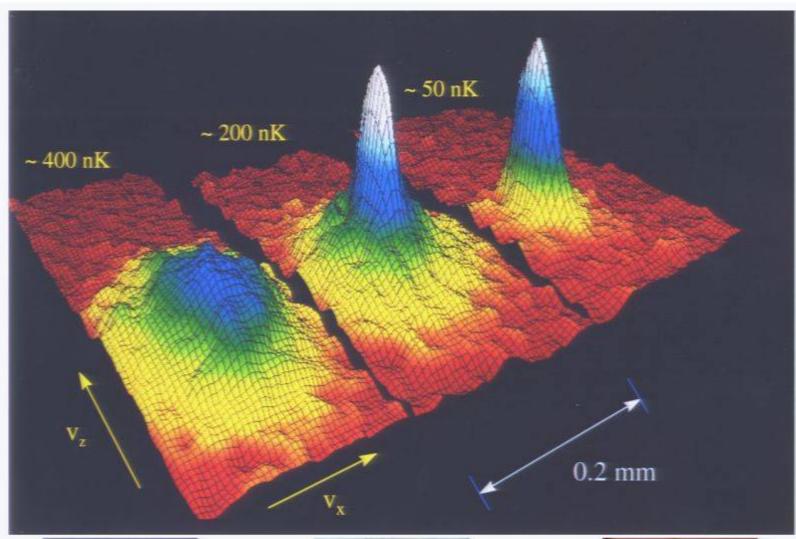


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In 1995 (70 years after Einstein's prediction) teams in Boulder, Colorado and Cambridge, Massachusetts achieved Bose-Einstein Condensation in super-cold gas (and opened a new and rapidly advancing area of research).

This feat earned those scientists the 2001 Nobel Prize for physics.

They and others have reached temperatures lower than one nanokelvin!

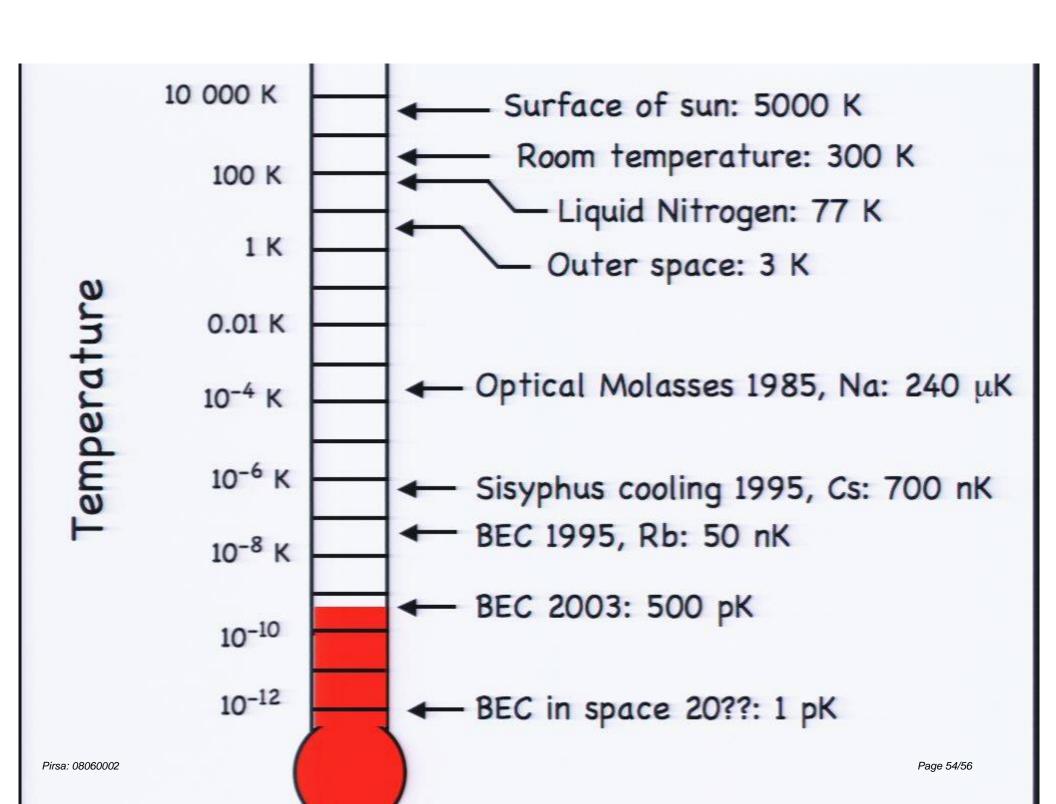








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#### What's Next?

- Better clocks
- Tests of the fundamental understanding of Nature
- Quantum Computers
- More...



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