

Title: Aspects of Hagedorn Holography

Date: Mar 08, 2008 11:00 AM

URL: <http://pirsa.org/08030050>

Abstract:

J. L. F. BARBON, C. FUERTES, E. R.

ON LITTLE HAGEDORN

HOLOGRAPHY

OVERCOME PREJUDICE - ENLIGHTENMENT

CONFIRM PREJUDICE - "PLEASURE"

LST

SEIBERG

AHARONY, BERKOOZ, KUTASOV, SEIBERG
MALACENA, MINWALLA, STROHINGER

MANY MORE...

$$S = E \frac{d}{d+1}$$



$$S = E$$

$$\frac{d}{d+1}$$

HAGEDORN

95%0

*

$$S = \beta E$$

$$T_H = T_{MAX}$$

STRINGS = COMPOSITES

*

$$C_v < 0$$

MORE THAN PREJUDICE

BOX OF SIZE $\in R$, $l_s = 1$

$$S_g \sim (E_g R)^{9/10}, S_s \sim E_s, S_b \sim \frac{1}{g^2} (g^2 E_b)^{1/2}$$

$$x_g = \frac{E_g}{E}, x_s = \frac{E_s}{E}, x_b = \frac{E_b}{E}$$

E GIVEN.

$$S(x_g, x_s, x_b) =$$

$$S_g(E) x_g^{9/10} + S_s(E) x_s + S_b(E) x_b^{1/2}$$

MAXIMIZE

$$x_g + x_s + x_b = 1$$

GENERERICALLY ONE CAN NEGLECT
ONE COMPONENT.

i.e.

$$S_g(E_{gb}) \cong S_b(E_{gb})$$

$$E_{gb} \sim \frac{1}{g_s^2} (g_s^2 R^9)^{3/17}$$

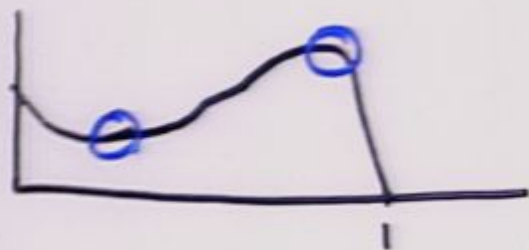
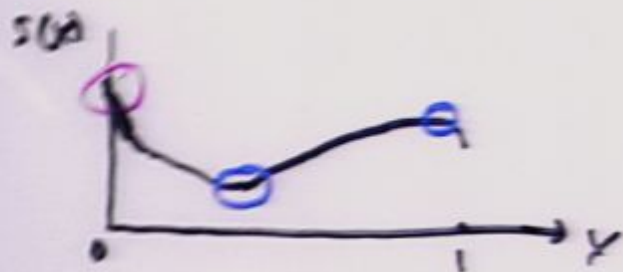
$$E_{gs} \sim R^9$$

$$E_{bs} \sim \frac{1}{g_s^2}$$

TRIPLE POINT

$$E_{\text{Triple}} \sim R^9 = \frac{1}{g_s^2}$$

$$g_s^2 R^9 = 1$$



BH
DOMINATES

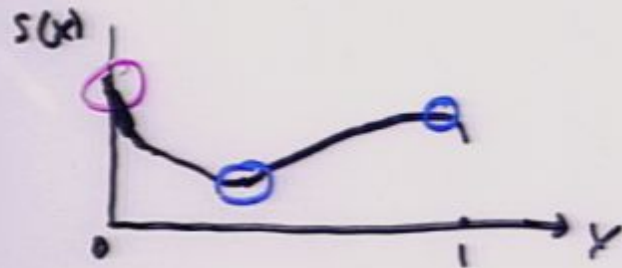
FOR $E \sim E_{pl}$

$$R_s = (g_s^2 E_{pl})^{1/2} \sim R \left(\frac{g_s^2}{R^2} \right)^{1/2} \ll R$$

EVENTUALLY R_s BECOMES $O(R)$

ADD WILL COME - - - ALL.

NE



BH
DOMINATES

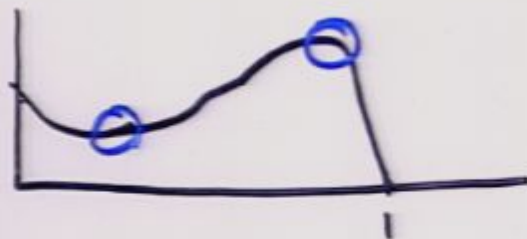
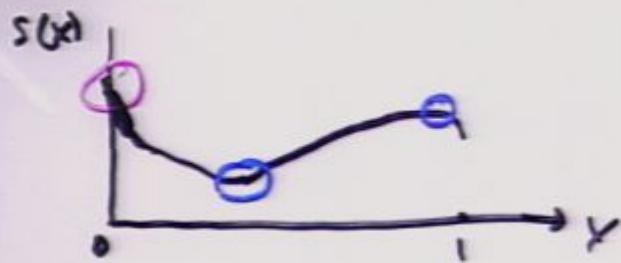
FOR $E \sim E_{gb}$

$$R_s = (g_s^2 E_{gb})^{1/2} \sim R \left(\frac{g_s^2}{R^2} \right)^{1/2} \ll R$$

EVENTUALLY R_s BECOMES $O(R)$

ADD WITH ...

UE



BH
DOMINATES

FOR $E \sim E_{pl}$

$$R_s = (g_s^2 E_{pl})^{1/2} \sim R \left(\frac{g_s^2}{R^2} \right)^{1/2} \ll R$$

EVENTUALLY R_s BECOMES $O(R)$

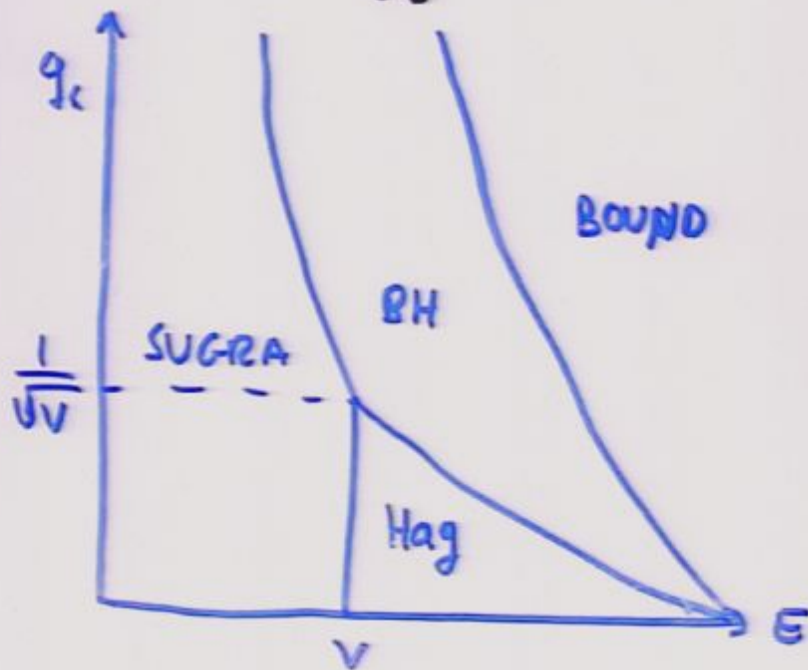
ADD WILL CRUSH THE WALL.

NEED ADS REGULATOR.

$$g_s^2 R^2 \ll 1 \quad \text{STRING PHASE ENTERS.}$$

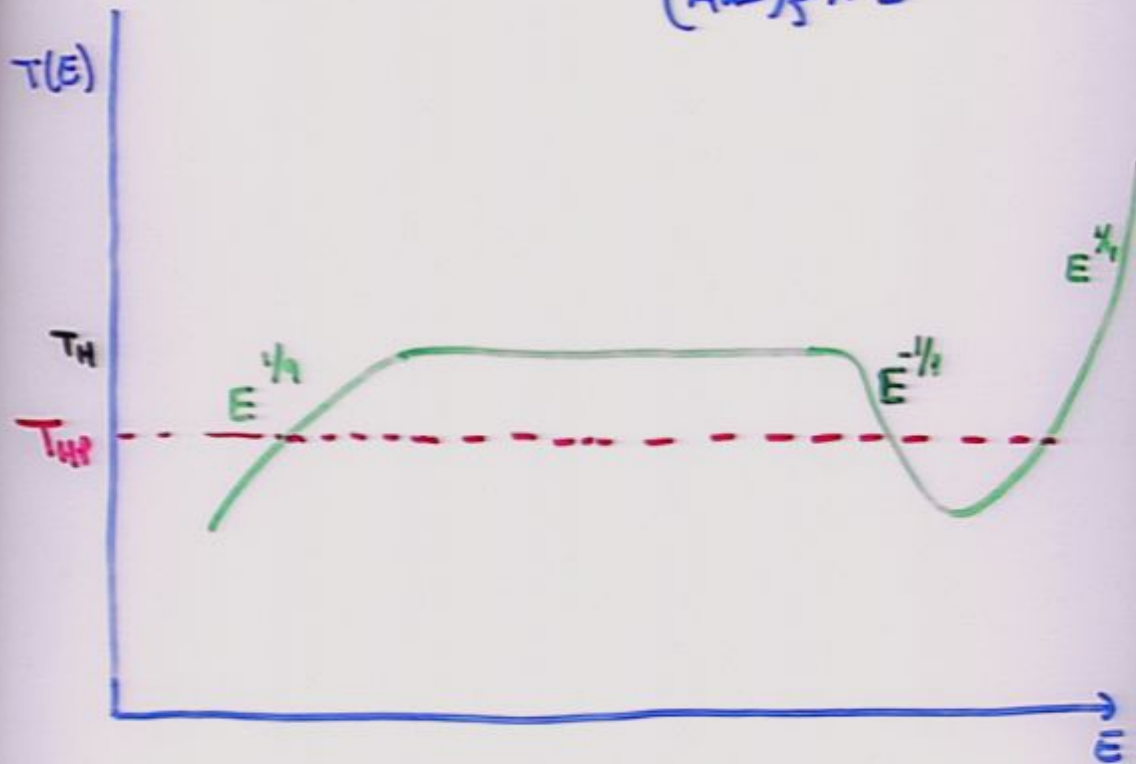
$$V_{\text{JEANS}} \approx (G_0 E)^{\frac{D-1}{D-3}}$$

$$E < \frac{V^{\frac{D-3}{D-1}}}{G_0}$$



$\Rightarrow \int = \frac{E}{V} \rightarrow 1 \rightarrow 0$ to TREMB

$$(AdS)_5 \times S^5$$

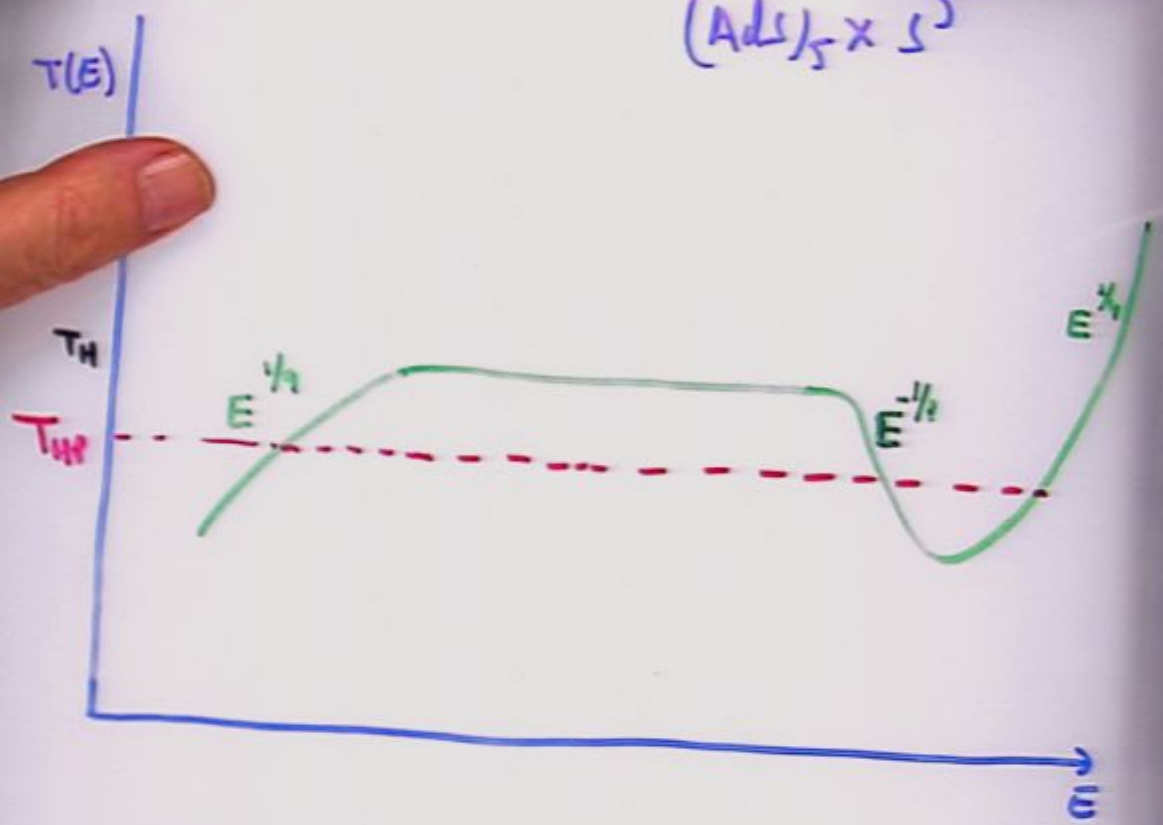


FOR LST?

NO GRAVITY

HAGEDON

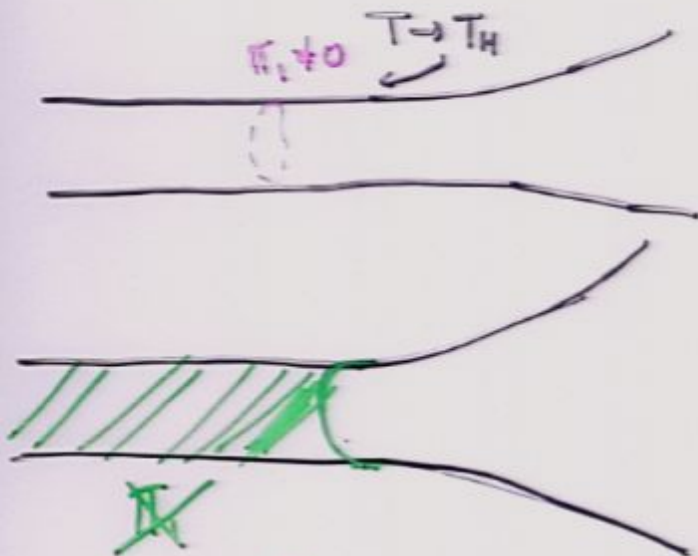
$$(Ads)_5 \times s^5$$



FOR LST ?

NO GRAVITY

HAGEDORN



EXIT
STRATEGY

~~METRIC~~
~~SPACE TIME~~

HAGEDORN
CENSORSHIP

J. BARBON, E. R.

* LST GEOMETRY

* LST THERMODYNAMICS

* HORIZON DEGENERACY - CLASSICAL

* QUANTUM RADIATION

NON RELATIVISTIC

NON HOLOGRAPHIC (AdS \neq)

* CANONICAL - PROBLEMS..

* MICROCANONICAL - LIMITS DON'T
COMMUTE.

* UU COMPLETION

NON COMMUTING LST

QUANTUM FIRST ORDER TRANSITION

HADRONIC ...

* LST THERMODYNAMICS

* HORIZON DEGENERACY - CLASSICAL

* QUANTUM RADIATION

NON RELATIVISTIC

NON HOLOGRAPHIC (AdS \neq)

* CANONICAL - PROBLEMS...

* MICROCANONICAL - LIMITS DON'T
COMMUTE.

* UU COMPLETION

NON COMMUTING LST

QUANTUM FIRST ORDER TRANSITION

HAGEORN TRANSIENT/CENSORED

LST

I.E. $g_s \rightarrow 0$ $R = \ell_s \sqrt{N}$ FIXED
OF N NS5 BRANES IN IIA.

RESULTS IN:

$$M^{45} R \phi S^3_R$$

\uparrow
N UNITS OF NS5 FLUX.

$$ds^2 = -dt^2 + dy_s^2 + dz^2 + R^2 d\Omega_3^2$$

$$e^\phi = \exp(-z/R)$$

LIMIT REMAINS

T-DUALITY, NON LOCAL

? QFT < LST < ST

BY

$$\Gamma = g_s R \exp\left(\frac{E}{R}\right)$$

AND BEING NON-EXTREMAL ONE
OBTAINS THE THERMAL BEHAVIOR

$$ds^2 = -dt^2 \left(1 - \frac{r_0^2}{r^2}\right) + dy_s^2 + \frac{R^2 dr^2}{r^2 - r_0^2} + R^2 d\phi^2$$

$$e^\phi = g_s R / \Gamma$$

$$S = \beta_H E$$

$$T \equiv T_H = \frac{1}{2\pi R} = \frac{1}{\ell_s} \frac{1}{2\pi \sqrt{N}}$$

$$\frac{\partial T}{\partial r_0} = 0 \Rightarrow \Gamma(r_0) = \beta_H E(r_0) - S(r_0) = 0$$

$\beta_H E(r_0)$

$$\frac{\partial \Gamma}{\partial r_0} = 0$$

$$E(r_0) = \frac{V_s m_s^8 r_0^2}{(2\pi)^5 g_s^2}$$

FIXED

$$\Pi_s = \frac{V_s m_s^6 N}{(2\pi)^5} + E(r_0)$$

NON DECOUPLED METRIC, DILATON

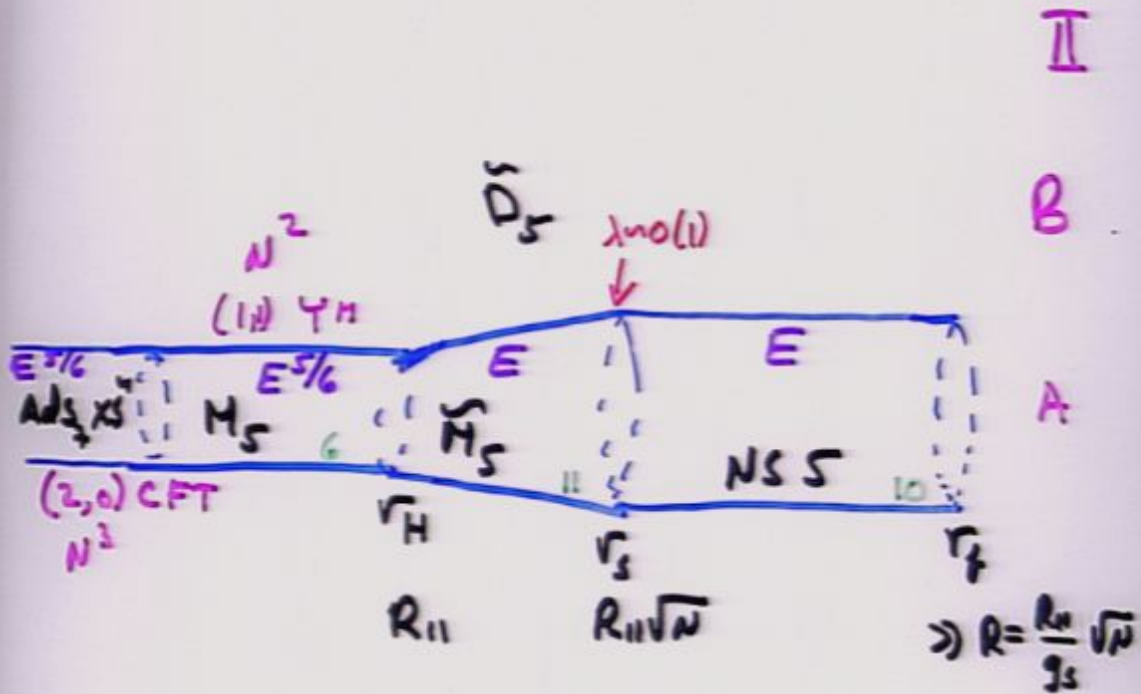
$$ds^2 = -dt^2 \left(1 - \frac{r_0^2}{r^2}\right) + dy_5^2 +$$

$$+ \left(1 + \frac{R^2}{r^2}\right) \left(\frac{d\Gamma^2}{1 - \frac{r_0^2}{r^2}} + r^2 d\Omega_3^2\right)$$

$$e^{2\phi} = g_s^2 \left(1 + \frac{R^2}{r^2}\right) \sim O(1) \text{ FOR } r = g_s R$$

$$g_s R = g_s l_s \sqrt{N} = R_{11} \sqrt{N}$$

$$R_{11} = g_s l_s$$



$$R = l_s \sqrt{N}$$

$$R_{11} = g_s l_s$$

BACK TO "FULL METRIC"

$$T(r_0) = \frac{1}{2\pi \sqrt{r_0^2 + R^2}}$$

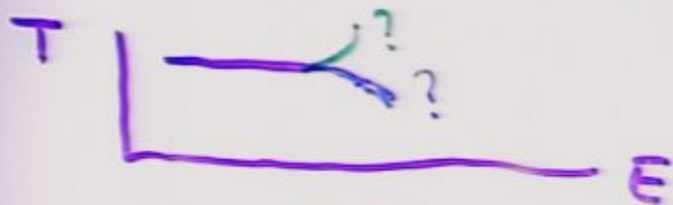
$$E(r_0) = \frac{V_S M_S^2 r_0^2}{(2\pi)^5 \eta_S^2}$$

$$\Rightarrow \frac{\partial T}{\partial E} < 0 \quad C_V < 0$$

PROBLEM.

* HOW TO REMOVE r_0 DEGENERACY

* $C_V > 0$? $C_V < 0$



CORRECTIONS:

*

l'

*

g_s

HORIZON

TUBE



HOLOGRAPHY OF RADIATION

* BH

$$(Ads)_{\substack{d \\ d+1}} \times S^d$$

$$ds^2 = \frac{r^2}{R^2} (-dt^2 + dx_{d-1}^2) + R^2 \frac{dr^2}{r^2}$$

$$(Ads)_{BH} \quad \text{HORIZON } r_0 = 2\pi R^2 T$$

LEADING TERM FOR ENTROPY $\frac{\partial r_0}{\partial T} \neq 0$

$$S_{BH} = \frac{V_{d-1} r_0^{d-1}}{4G_{d+1} R^{d-1}} = \frac{(2\pi R)^{d-1}}{G_{d+1}} V_{d-1} T^{d-1}$$

N_{eff}

HOLOGRAPHIC

HOLOGRAPHY OF RADIATION

* BH

$$(A_{dS})_{d+1} \times S^d$$

$$ds^2 = \frac{r^2}{R^2} (-dt^2 + dx_{d-1}^2) + R^2 \frac{dr^2}{r^2}$$

$$A = \Gamma_0 = 2\pi R^2 T$$

TERM FOR ENTROPY $\frac{\partial \Gamma_0}{\partial T} \neq 0$

$$= \frac{V_{d+1} \Gamma_0^{d-1}}{G_{d+1} R^{d-1}} = \frac{(2\pi R)^{d-1}}{G_{d+1}} V_{d+1} T^{d-1}$$

N_{eff}

HOLOGRAPHIC

* RADIATION

$$S_r \sim \int_{r > r_0} d\vec{V} \left(\frac{T}{\sqrt{-g_{tt}}} \right)^d$$

OPTICAL VOLUME!

$$S_r = \# T^{d+1} V_{\text{OPTICAL}}(\text{AdS}(T)) =$$

$$\sim V_{d-1} T^{d-1}$$

- * NO RADIAL EXTENSIVITY
- * NO KK S^5
- * WORKS FOR D/p BRANES

$$p < 5$$

$$p=5 \quad D5 \quad NS5$$

V_{eff}

CONSIDER

$$ds^2 = -f(r) dt^2 + \frac{dr^2}{g(r)} + \sum_i P_i^2(r) ds_i^2$$

↓
d_i dimensional
Riemannian
space

$$S_\psi = -\frac{1}{2} \int d^D x \sqrt{g} e^{-2\phi} (g^{\mu\nu} \partial_\mu \psi \partial_\nu \psi + m^2 \psi^2)$$

ψ SCALAR FROM NSNS MINIMALLY COUPLED.

\Rightarrow

$$V_{eff} = \frac{1}{2} \partial_z^2 \log(\rho_{\pi}) + \frac{1}{4} (\partial_z \log \rho_{\pi})^2 + f(z) \left(\tilde{m}^2 + \sum_i \frac{\Delta_i}{\rho_i^2(z)} \right)$$

* (Ads)₅

$$V_{eff}(z) \sim \frac{\Lambda^2}{(z-z_{\infty})^2} \text{ "BOX"}$$



* of $f < 5$ BOX

$$\frac{r}{R_{of}} = \left(\frac{2}{5-f} \right)^{\frac{2}{5-f}} \left(\frac{R_{of}}{z_{\infty}-z} \right)^{\frac{2}{5-f}}$$

$$\sim (g_{\infty})^{\frac{1}{5-f}}$$

$$V_{\text{eff}} = \frac{1}{2} \partial_z^2 \log(\Sigma \pi) + \frac{1}{4} (\partial_z \log \Sigma \pi)^2$$

$$+ f(z) \left(\tilde{m}^2 + \sum_i \frac{\Delta_i}{f_i(z)} \right)$$

* (Ads)₅

$$V_{\text{eff}}(z) \sim \frac{\Lambda^2}{(z - z_{\infty})^2} \text{ "BOX"}$$



* of $f < 5$ BOX

$$\frac{r}{R_{\text{of}}} = \left(\frac{2}{5-f} \right)^{\frac{2}{5-f}} \left(\frac{R_{\text{of}}}{z_{\infty} - z} \right)^{\frac{2}{5-f}}$$

$$R_{\text{of}} = \text{CHARGED RADIUS OF } D_{f/2} \sim (g_s N)^{\frac{1}{5-f}}$$

*

NS5

$$ds^2 = f(z) (-dt^2 + d\vec{z}^2) + \sum_i f_i^2(z) ds_i^2$$

$$f(z) = 1 \quad \begin{array}{l} S^3 \quad f(z) = R \\ R^5 \quad f(z) = 1 \end{array}$$

$$V_{\text{eff}}(z) = \tilde{m}^2 + \vec{p}^2 + \frac{\Delta_S^2}{R^2}$$

$$\tilde{m}^2 = m^2 + \frac{1}{R^2} \rightarrow \text{DILATON SLOPE}$$

DIAGONALIZE BY $\exp(i\vec{p}\cdot\vec{z})$

$$\omega^2 = p_z^2 + \vec{p}^2 + m^2 + \frac{1}{R^2} + \frac{\Delta_S^2}{R^2}$$

*

NSS

$$ds^2 = f(z) (-dt^2 + dz^2) + \sum_i f_i^2(z) ds_i^2$$

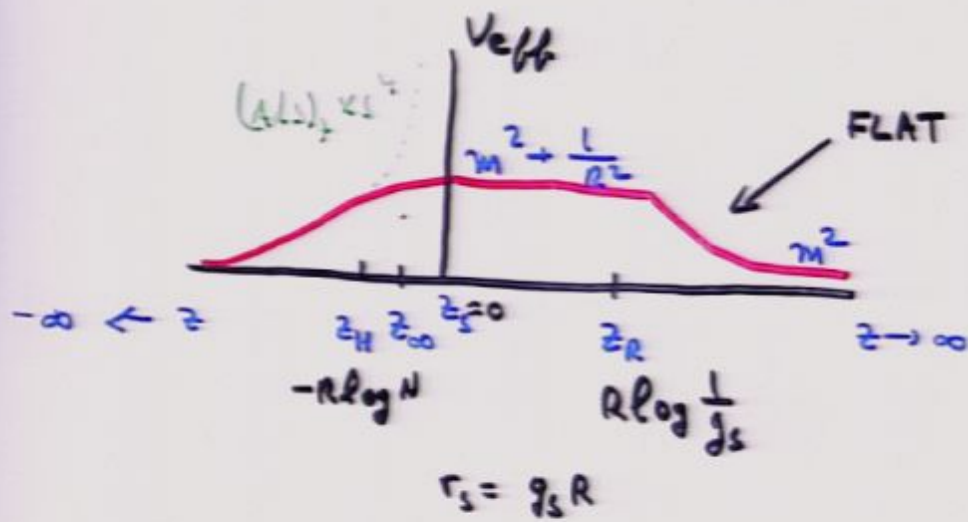
$$f(z) = 1 \quad \begin{array}{l} S^3 \quad f(z) = R \\ R^5 \quad f(z) = 1 \end{array}$$

$$V_{\text{eff}}(z) = \tilde{m}^2 + \vec{p}^2 + \frac{\Delta_S^2}{R^2}$$

$$\tilde{m}^2 = m^2 + \frac{1}{R^2} \rightarrow \text{DILATON SLOPE}$$

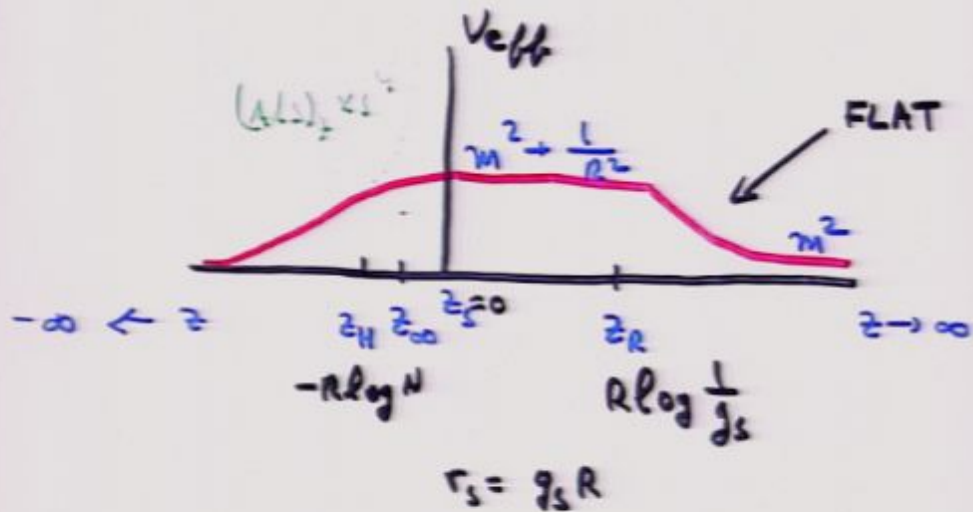
DIAGONALIZE BY $\exp(i\phi(z))$

$$\omega^2 = p_z^2 + \vec{p}^2 + m^2 + \frac{1}{R^2} + \frac{\Delta_S^2}{R^2}$$



NO HOLOGRAPHY IN RADIAL DIRECTION

ONE



NO HOLOGRAPHY IN RADIAL DIRECTION

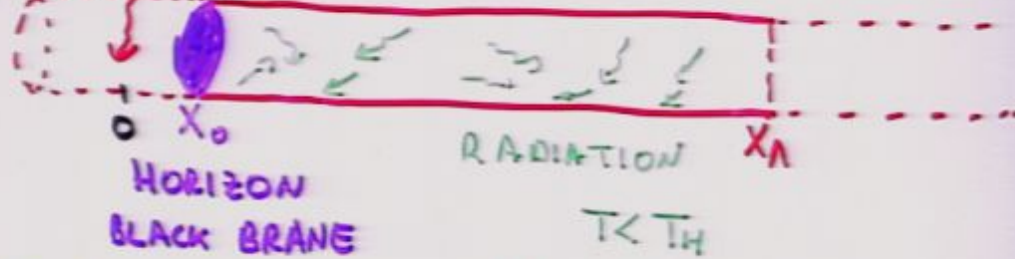
ONE NEEDS TO PROBE WITH
A CUTOFF

THERMODYNAMICS OF CUTOFF LST

"PISTON" MODEL

$$S_{IR} = N^\alpha V_S T^5$$

$\alpha = 3$
 $L = 2$



$$S = E_f(x_0) \beta_H$$

$$E_f(x_0) = E_H \exp(-2x_0)$$

$$E_H = N^\alpha V_S T_H^6$$

$$T_H = \frac{1}{2\pi R}$$

$$R = \ell_s \sqrt{N}$$

NR RADIATION

$$E_R = (x_A - x_0) R^3 V_S T^3 \exp\left(-\frac{1}{TR}\right)$$

$T < T_H$

$T > T_H$

$$E_R = -(x_A - x_0) R^3 V_S T^3$$

$$T \supset T_H \quad S = E \quad \textcircled{\frac{d}{d+1}}$$



$$\beta(\epsilon) = \frac{1}{T_H} \left(1 - \frac{1}{20} \left(\frac{q}{10} \right)^9 \frac{1+2\delta}{N^2 \epsilon} \right)$$

$$\delta = \frac{1}{2} \left(\frac{q}{10} \right)^{10} \left(\frac{\chi_\Lambda - \frac{1}{2} \log \frac{E}{E_H}}{N^2 \frac{E}{E_H}} \right)$$

$$* \quad C_U < 0 \quad \frac{1}{2} \log \frac{E}{E_H} < \chi_\Lambda \ll N^2 \frac{E}{E_H}$$

UV COMPLETION

- * SUPPLY CUTOFF + D.O.F
- * FIRST ORDER TRANSITION.
- * RADIATION WILL BE HOLOGRAPHIC

DECONSTRUCTION:

N. DOREY

ARKANI-HAMED, COHEN, KARCH, MOTL, KAMAN

DECONSTRUCTION:

N. DOREY

ARKANI-HAMED, COHEN, KARCH, MOTL, KAMAN

* THERMODYNAMICS OF $\widetilde{\text{LST}}$ II A
CLASSICAL AND QUANTUM

CLASSICAL

* $\widetilde{\text{LST}}$ B

* $N=4$ SUSY YM β DEFORMED

* HIGGS BRANCH GIVES

DECONSTRUCTED $D=6$ YM.

* DUAL GEOMETRY HAS $D5$ S

S DUALITY

* TOUGH ON SUSY YM

WORKS ON GEOMETRY $D5$ \leftrightarrow $M2$

BUT TYPE B.

T DUALITY

* CLASSICAL $\widetilde{\text{LST}}$ II A. MC HOLE
* CLASSICAL THERMO + QUANTUM

IB

$D=4$ $N=4$ $U(\hat{N})$ SUSY YM

$$W = i\kappa \left(\exp(i\bar{\beta}/2) \phi_1 \phi_2 \phi_3 - \exp(-i\bar{\beta}/2) \phi_1 \phi_3 \phi_2 \right)$$

PARAMETERS $g, \kappa, \bar{\beta}$

$$\tau = g = \frac{4\pi c}{g^2} + \frac{\theta}{2\pi}$$

* EXACTLY MARGINAL OPERATOR

* DUALITY : $SL(2, \mathbb{R})$

$$\tau_R = \tau + \frac{i\bar{\beta}}{\pi} \log \kappa$$

$$\det \begin{pmatrix} a & b \\ c & d \end{pmatrix} = 1 \quad \tau_R \rightarrow \frac{a\tau_R + b}{c\tau_R + d} ; \beta \rightarrow \frac{\beta}{c\tau_R + d} ; \kappa^2 \sin \beta \rightarrow \frac{\kappa^2 \sin \beta}{c\tau_R + d}$$

S DUALITY INCLUDED

* FOR $\beta = \frac{2\pi}{n}$ $\beta = N\tau$

(S DUAL) $\tilde{\beta} = \frac{i2\pi^2 n}{g^2}$ $\tilde{g}^2 = \frac{16\pi^2}{g^2}$

$U(\hat{N}) \rightarrow U(N)$ BY NON COMMUTING ϕ_i
HIGGS BRANCH

WHEN

$$\langle \phi_1 \rangle = \alpha_1 \int_{U(N)} \rho_1 \rho_1^{-1}$$

$$\langle \phi_2 \rangle = \alpha_2 \int_{U(N)} \rho_2 \rho_2^{-1}$$

$$\langle \phi_3 \rangle = \alpha_3 \int_{U(N)} \rho_2^+ \rho_1^+$$

$$(\rho_1)_{ij} = \delta_{ij} w^i \quad (\rho_2)_{ij} = \delta_{i,j-1}$$

$$w^i = \exp\left(\frac{2\pi i}{n}\right)$$

* P

'S AT

* FOR $\beta = \frac{2\pi}{n}$ $\beta = N\tau$

(S DUAL) $\tilde{\beta} = \frac{i8\pi^2 n}{g^2}$ $\tilde{g}^2 = \frac{16\pi^2}{g^2}$

$U(\hat{N}) \rightarrow U(N)$ BY NON COMMUTING ϕ_i
HIGGS BRANCH

WHEN

$$\langle \phi_1 \rangle = \alpha_1 \int_{U(N)} \rho_1 \, d\mu$$

$$\langle \phi_2 \rangle = \alpha_2 \int_{U(N)} \rho_2 \, d\mu$$

$$\langle \phi_3 \rangle = \alpha_3 \int_{U(N)} \rho_2^+ \rho_1^+ \, d\mu$$

$$(\rho_1)_{ij} = \delta_{ij} w^i \quad (\rho_2)_{ij} = \delta_{i,j-1}$$

$$w^i = \exp\left(\frac{2\pi i}{n}\right)$$

*

'S AT

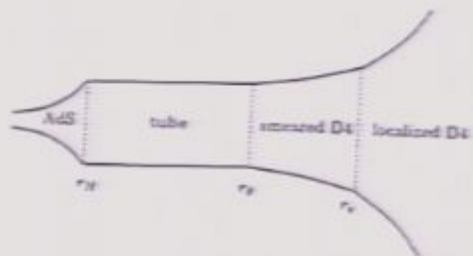


Fig. 6: Rendering of the background profile in IIA deconstruction.

We may summarize the energy hierarchies for both IIA ($\alpha = 3$) and IIB ($\alpha = 2$) cases by the relations

$$\frac{E_g}{E_H} \sim \left(\frac{n}{n_{\text{eff}}^{4-\alpha}} \right)^2, \quad \frac{E_v}{E_H} \sim \left(\frac{n}{n_{\text{eff}}^{3-\alpha}} \right)^2. \quad (6.19)$$

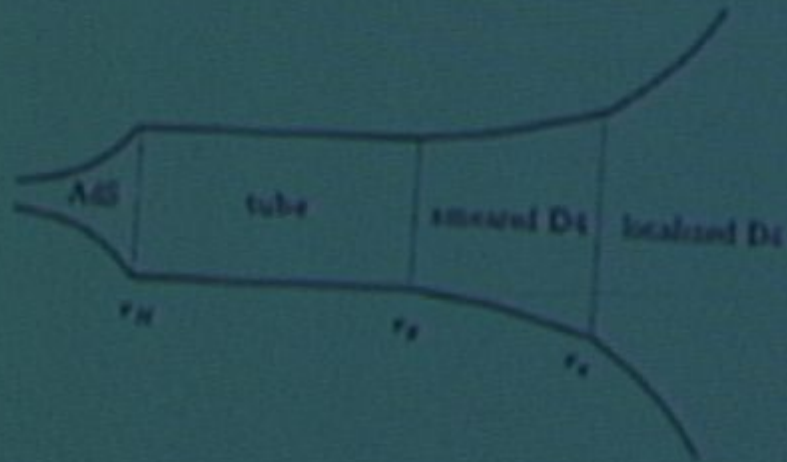


Fig. 6: Rendering of the background profile in IIA deconstruction.

We may summarize the energy hierarchies for both IIA ($\alpha = 3$) and IIB ($\alpha = 2$) cases by the relations

$$\frac{E_d}{E_H} \sim \left(\frac{n}{n_{eff}} \right)^3, \quad \frac{E_d}{E_H} \sim \left(\frac{n}{n_{eff}} \right)^2. \quad (6.19) \text{ here}$$

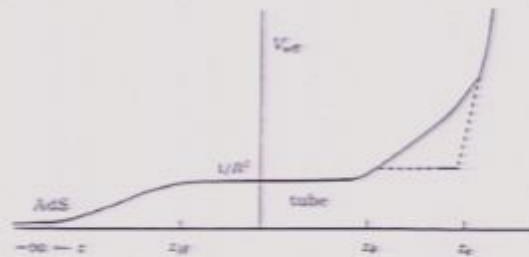


Fig. 7: Effective radiation potential in the deconstructed type IIA LST background. For $z_1 < z < z_2$ the tube plateau steps up to $(3/2R)^2$ for modes with zero momentum in the noncommutative direction, whereas we encounter an exponential wall for non-vanishing momentum modes.

$$V_{\text{eff}}(z) = \underbrace{f_c^2}_C + \underbrace{f^2(z) \vec{q}^2}_{NC \quad \vec{q} \neq 0} + \frac{\Delta_s^2}{R^2} + \left(\frac{\alpha}{2R}\right)^2$$

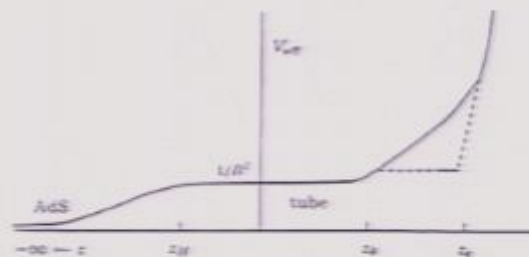


Fig. 7: Effective radiation potential in the deconstructed type IIA LST background. For $z_0 < z < z_e$ the tube plateau steps up to $(3/2R)^2$ for modes with zero momentum in the noncommutative direction, whereas we encounter an exponential wall for non-vanishing momentum modes.

$$V_{\text{eff}}(z) = \underbrace{\vec{f}_c^2}_C + \underbrace{f^2(z) \vec{q}^2}_{NC \quad q \neq 0} + \frac{\Delta_s^2}{R^2} + \left(\frac{\alpha}{2R}\right)^2$$

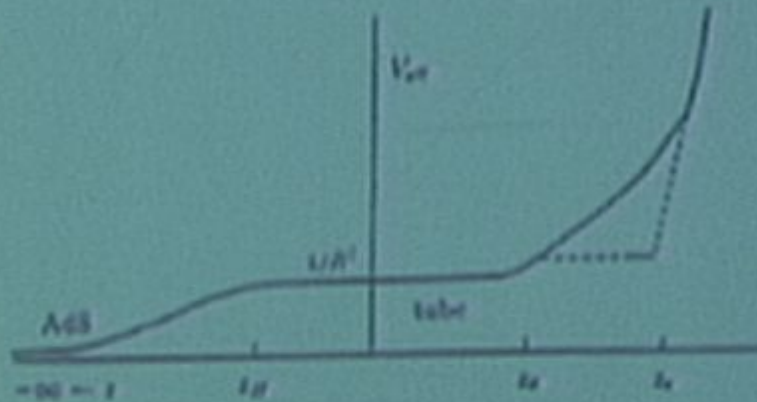


Fig. 7: Effective radiation potential in the deconstructed type IIA LST background. For $z_0 < z < z_1$ the tube plateau steps up to $(3/2R)^2$ for modes with zero momentum in the noncommutative direction, whereas we encounter an exponential wall for non-vanishing momentum modes.

35

$$V_{\text{eff}}(z) = \underbrace{\vec{p}_c^2}_C + \underbrace{f(z)^2 \vec{q}^2}_{\text{NC } \vec{q} \neq 0} + \frac{\Delta_S^2}{R^2} \left(\frac{\alpha}{2R} \right)^2$$

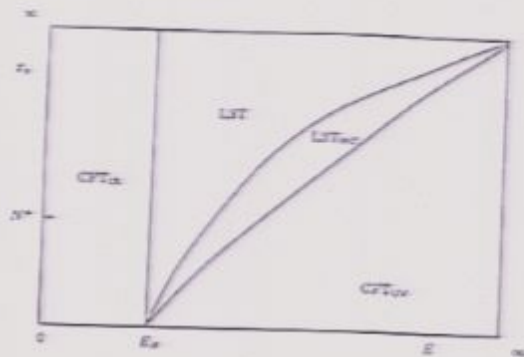


Fig. 8: Phase diagram of the deconstructed LST theory, as a function of the energy and the effective length of radiation tube, z_e , in the classical approximation. The noncommutative LST transient is supported by the ratio $n_{eff} = L/\epsilon > 1$, disappearing when it approaches unity.

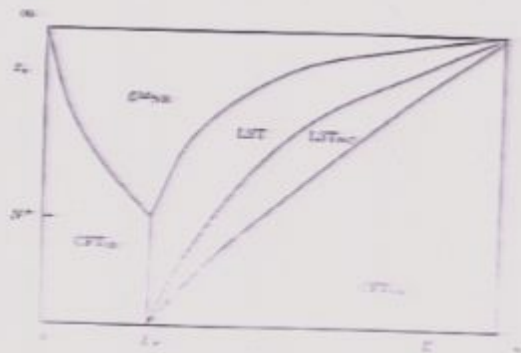


Fig. 9: The complete phase diagram of the deconstructed LST theory, as a function of the energy and the tube cutoff, z_e . Comparing with the phase diagram in the classical approximation, we see that the main effect of the radiation corrections is to open a new five-dimensional tube, $z_e > N^*$, dominated by the seven-dimensional S^2 transverse.

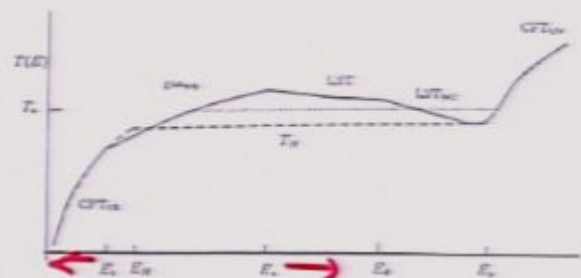


Fig. 10: The microcanonical temperature function, showing stable, unstable and metastable phases, as a function of the important thresholds in the system. The dashed line represents the classical approximation, with a Hagedorn plateau extending from E_B out to E_u . The dotted line gives the Maxwell construction of the first-order phase transition at the critical temperature T_c . In the limit when we decouple the UV fixed point, $E_c \rightarrow \infty$, the high-energy thresholds of the tube E_u and E_B diverge, whereas the low-energy threshold E_1 vanishes. In this limit, only the nonrelativistic gas remains.

FIRST ORDER PHASE

TRANSITION RADIATION ~~EPISODES~~

* "THROW THE BABY WITH BATH WATER"

$$z_E \rightarrow \infty$$

* CENSORSHIP

CONCLUSIONS:

- * LST CAN BE EMBEDDED IN UV COMPLETE SYSTEM.
 - HAGEDORN IS TRANSIENT.
 - FIRST ORDER TRANSITION
 - HAGEDORN CENSORSHIP

- * ON ITS OWN HAS IR DEPENDENCE.

- * FOLLOW V