

Title: Computing the Massless Spectrum of Non-Kahler Compactifications

Date: Dec 07, 2007 09:00 AM

URL: <http://pirsa.org/07120038>

Abstract: While Calabi-Yau compactifications of string theory are mathematically elegant, they typically result in many massless scalars in the low-energy, four-dimensional theory. Thus, it is interesting to consider non-Kahler compactifications in the hopes of deriving more phenomenologically interesting models. These models have received little attention in the heterotic theory owing to their mathematical complexity, however in recent work we have found a potential way to derive interesting features of such compactifications using gauged linear sigma models.

COMPUTING THE MASSLESS SPECTRUM OF NON-KÄHLER COMPACTIFICATIONS

Joshua Lapan
Harvard University

based on work with A. Adams and M. Ernebjerg: [hep-th/0611084](https://arxiv.org/abs/hep-th/0611084)



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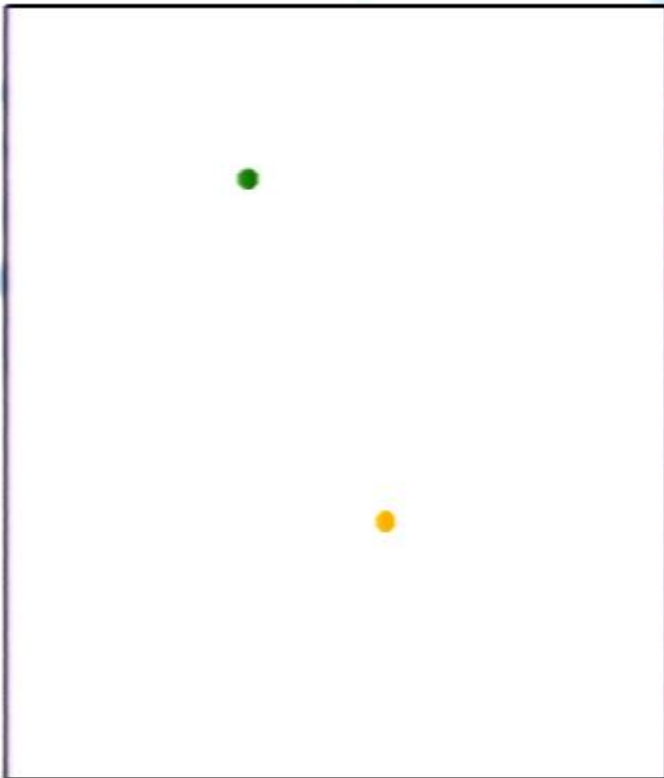


Compactifications:

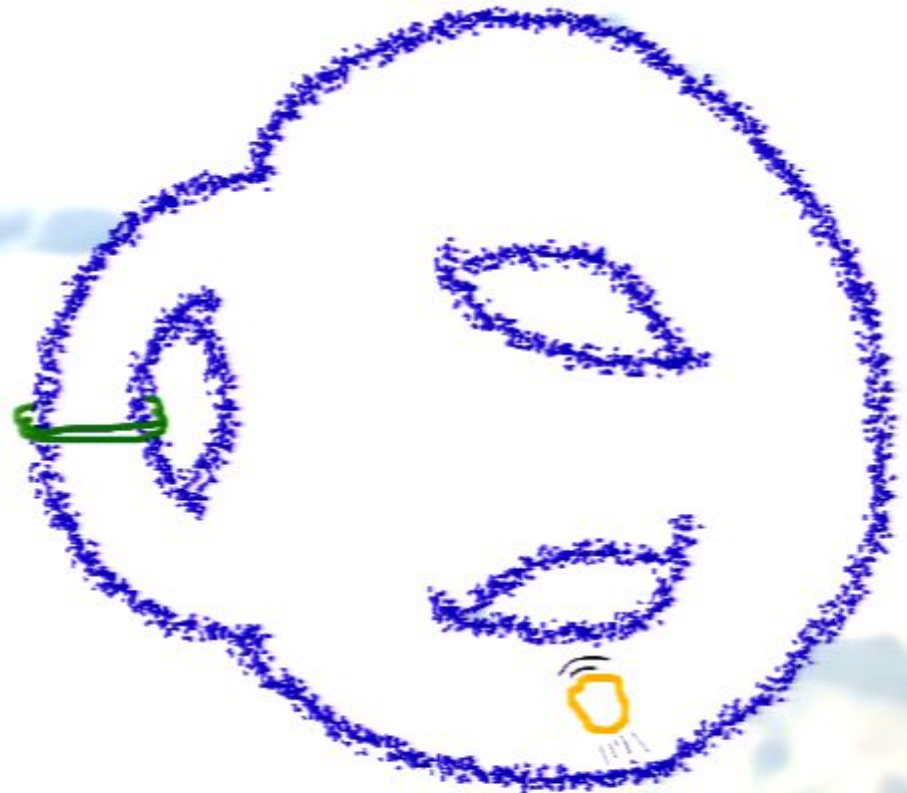
$\mathbb{R}^{3,1}$

\times

K



\times



CALABI-YAU

Pros:

- Mathematically
“Simple”
- Qualitative Physics
↔ Topology

CALABI-YAU

Pros:

- Mathematically “Simple”
- Qualitative Physics
 \rightsquigarrow Topology

Cons:

- Many Massless Scalars!
- = Many Moduli = “Continuous” lack of Predictivity

HETEROTIC STRING THEORY:

- “Natural” Appearance of Chiral, Charged Fermions
- *Exact* Worldsheet CFT (no RR-Fluxes)

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- “Natural” Appearance of Chiral, Charged Fermions
- *Exact* Worldsheet CFT (no RR-Fluxes)

NON-KÄHLER

- Potentially fewer massless scalars/moduli (e.g. Radial mode is fixed)

WORKSHEET SIGMA MODEL

$$S \sim \int_{\Sigma} d^2\sigma \left\{ \frac{1}{\alpha'} (G_{MN} + iB_{MN}) \partial X^M \bar{\partial} X^N + \psi^{\hat{M}} \left(\eta_{\hat{M}\hat{N}} \partial + \partial X^P \omega_{P\hat{M}\hat{N}}^{(-)} \right) \psi^{\hat{N}} \right. \\ \left. + \gamma^A (\delta^{AB} \bar{\partial} - i \bar{\partial} X^M A_M^{AB}) \gamma^B + \frac{1}{2} F_{\hat{M}\hat{N}}^{AB} \psi^{\hat{M}} \psi^{\hat{N}} \gamma^A \gamma^B \right\}$$

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$$X^M: \Sigma \rightarrow M = \mathbb{R}^{3,1} \times K$$

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$V =$ Spacetime Gauge Bundle

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$(H_{NS})_{MNP} = 3\partial_{[M} B_{NP]} + \dots$

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CALABI-YAU:

- EOMs

$$\square\Phi = \square_4\Phi + \square_{CY}\Phi = 0$$

- Hodge / de Rham / Dolbeault

$$\mathcal{H}_{\square}^* \cong H_d^* \cong H_{\bar{\partial}}^*$$

- Large radius limit

$$\beta_{\mu\nu}^G \sim R_{\mu\nu} + H_{\mu\rho\sigma}H^{\rho\sigma}{}_{\nu} + \dots$$

Invariance:

$$G_{\mu\nu} \rightarrow \lambda G_{\mu\nu} \quad B_{\mu\nu} \rightarrow \lambda B_{\mu\nu}$$

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NON-KÄHLER:

- EOMs

$$0 \neq \square\Phi ; \quad \cancel{\mathcal{H}_{\square}^*}$$

- de Rham / Dolbeault

$$H_d^* \not\cong H_{\bar{\partial}}^*$$

- No large radius limit

Bianchi Identity:

$$dH = \alpha'(\text{tr}\mathcal{R}^2 - \text{Tr}F^2)$$

$$\rightarrow \lambda dH = \alpha'(\text{tr}\mathcal{R}^2 - \text{Tr}F^2)$$

STANDARD CALABI-YAU MODEL

SUSY Variations

(1) $K = \text{complex}$ and $c_1(TK) = 0 \longrightarrow K = CY_3$;

(2) $i(\bar{\partial} - \partial)J = H_{NS} \longrightarrow H_{NS} = 0$

(3) $F = HYM$

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$$\langle A_M \rangle = \left(\begin{array}{c|c} \omega_M(K) & 0 \\ \hline 0 & 0 \end{array} \right) \Rightarrow \text{tr}\mathcal{R}^2 = \text{Tr}F^2$$

NON-CALABI-YAU MODEL

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FU-YAU GEOMETRY

based on [hep-th/0604063](#)

Manifold:

$$S_{(1)}^1 \times S_{(2)}^1 = T^2 \longrightarrow K$$
$$\downarrow \pi$$
$$CY_2 = S$$

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$$F^{(2,0)} = F^{(0,2)}$$
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$$\int_S \text{Bianchi: } \int_S \sum_l \omega_l^2 = Ch_2(E) - Ch_2(S)$$

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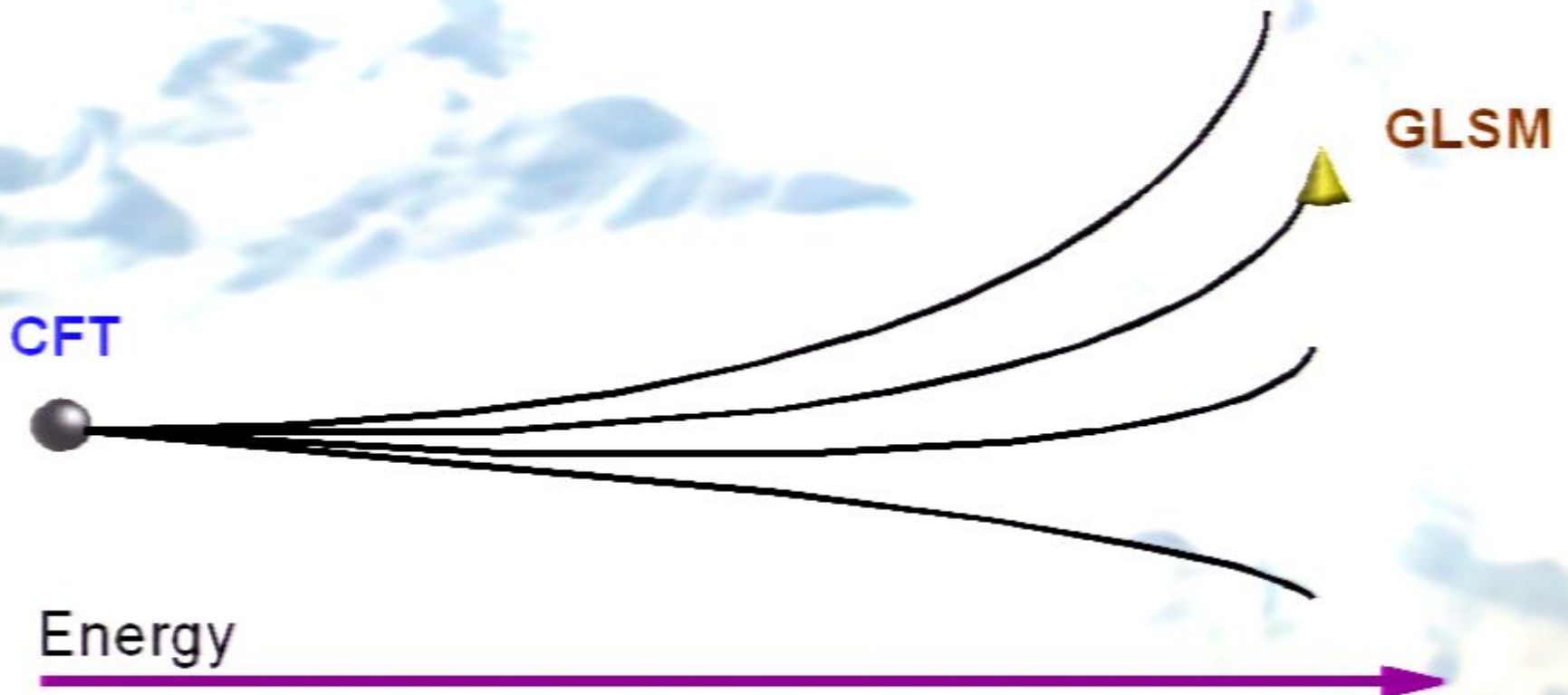


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2) Fixed volume \Rightarrow pert. series may break down



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GAUGED LINEAR SIGMA MODEL



$\mathcal{N} = (0, 2)$ FIELDS

$$\bar{D}_+ = -\frac{\partial}{\partial \bar{\theta}^+} + 2i\theta^+ \partial_+$$

- **Chiral Superfields:**

$$\bar{D}_+ \Phi^i = 0 \quad \Rightarrow \quad \Phi^i = \phi^i + \sqrt{2}\theta^+ \psi_+^i - 2i\theta^+ \bar{\theta}^+ \partial_+ \phi^i$$

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- **Gauge Fields (Real Superfields):**

$$V_+ = 2\theta^+ \bar{\theta}^+ v_+, \quad \frac{1}{2}V_- = v_- - i\theta^+ \bar{\lambda}_- - i\bar{\theta}^+ \lambda_- + \theta^+ \bar{\theta}^+ D$$

$$\Upsilon = \bar{D}_+(2\partial_- V_+ + iV_-)$$

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- **Gauge Parameter (Chiral):** $\bar{D}_+ \Lambda = 0$

$$\Phi^i \rightarrow e^{-iQ^i \Lambda} \Phi^i, \quad V_+ \rightarrow V_+ + \frac{i}{2}(\Lambda - \bar{\Lambda}), \quad \text{etc.}$$

| | |
|------------|--------|
| | $U(1)$ |
| Φ^i | Q^i |
| P | $-n$ |
| Π | $-d$ |
| Γ^m | q_m |

SUPERPOTENTIAL

$$\sqrt{2}W = W_S + W_E + W_{FI}$$

$$W_S = \int d\theta^+ \Pi F(\Phi)$$

$$W_E = \int d\theta^+ P \Gamma^m E_m(\Phi)$$

$$W_{FI} = \frac{it}{4} \int d\theta^+ \Upsilon, \quad \text{where } t = r + i\theta$$

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\Rightarrow $F(\Phi)$ = Homogeneous poly. of degree d
 $E_m(\Phi)$ = Homogeneous poly. of degree $n - q_m$

$$\mathcal{L} \sim \mathcal{L}_{kin} + \mathcal{L}_{Yuk} + \underbrace{\mathcal{L}_D + \mathcal{L}_F + \mathcal{L}_E}_{-U} + 2\theta v_{+-}$$

$$\mathcal{L}_D = \frac{1}{2e^2} D^2 + D \left(\sum_i Q^i |\phi^i|^2 - n|p|^2 - r \right)$$

$$\int \mathcal{D}D \longrightarrow -\frac{e^2}{2} \left(\sum_i Q^i |\phi^i|^2 - n|p|^2 - r \right)^2$$

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$$\mathcal{L}_F = |\rho|^2 - \rho F(\phi) - \bar{\rho} \overline{F(\phi)}$$

$$\xrightarrow{\int \mathcal{D}\rho} -|F(\phi)|^2$$

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$$\mathcal{L}_E = \sum_m |G_m|^2 - p G^m E_m(\phi) - \bar{p} \overline{G^m E_m(\phi)}$$

$$\int \mathcal{D}G \rightarrow -|p|^2 \sum_m |E_m(\phi)|^2$$

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$r \gg 1$ SCALAR POTENTIAL

$$U = \frac{e^2}{2} \underbrace{\left(\sum_i Q^i |\phi^i|^2 - n|p|^2 - r \right)^2}_{\mathcal{L}_D} + \underbrace{|F(\phi)|^2}_{\mathcal{L}_F} + \underbrace{|p|^2 \sum_m |E_m(\phi)|^2}_{\mathcal{L}_E}$$

Manifold:

1. $\{\phi^i\}$, $i = 1, \dots, N + 1$, parameterize \mathbb{C}^{N+1}

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Manifold:

1. $\{\phi^i\}$, $i = 1, \dots, N + 1$, parameterize \mathbb{C}^{N+1}
2. $\mathcal{L}_D = 0$ and $n, Q^i \geq 0 \Rightarrow \mathbb{C}^{N+1} \rightarrow \sim S^{2N+1}$

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3. $U(1)$ Gauge fixing $\Rightarrow S^{2N+1} \rightarrow \text{WP}_{\vec{Q}}^N$
4. $\mathcal{L}_F = 0 \Rightarrow \text{WP}_{\vec{Q}}^N \rightarrow S = \{F(\phi) = 0 | \vec{\phi} \in \text{WP}_{\vec{Q}}^N\}$

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Manifold: (cont.)

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5. Assume that $F(\phi) = E_1(\phi) = \dots = 0 \Leftrightarrow \vec{\phi} = 0$

6. $\mathcal{L}_E = 0$ (given $\mathcal{L}_F = 0$) $\Rightarrow p = 0$

$r \gg 1$ SCALAR POTENTIAL

$$U = \underbrace{\frac{e^2}{2} (\sum_i Q^i |\phi^i|^2 - n|p|^2 - r)^2}_{\mathcal{L}_D} + \underbrace{|F(\phi)|^2}_{\mathcal{L}_F} + \underbrace{|p|^2 \sum_m |E_m(\phi)|^2}_{\mathcal{L}_E}$$

Manifold: (cont.)

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$r \ll -1$ LG Phase

$U = 0$ implies that $p = \sqrt{-\frac{r}{n}}$ and $\vec{\phi} = 0$

$r \gg 1$ FERMIONS

\mathcal{L}_{Yuk} determines massless fermions

Right-Movers

Cohomology of

$$0 \rightarrow \mathcal{O}_S \rightarrow \oplus_i \mathcal{O}_S(Q^i) \rightarrow \mathcal{O}_S(d) \rightarrow 0$$

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Sections of E given by exact sequence

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$$\mathcal{L} \supset 2\theta v_{+-} = \frac{1}{2}\theta\epsilon^{\mu\nu}v_{\mu\nu} \sim B_{MN}\epsilon^{\mu\nu}\partial_\mu X^M\partial_\nu X^N$$

$\Rightarrow B \sim \theta H$, where H generates $H^2(\mathbb{W}\mathbb{P}_{\vec{Q}}^N)$

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$$\int_S \sum_i \omega_l = Ch_2(E) - Ch_2(S)$$

COMPUTING MASSLESS SPECTRUM

Right-Movers: $\mathcal{N} = (0, 2)$ SCA implies

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forming Conformal Algebra in UV

(extension of) Silverstein & Witten, [hep-th/9403054](https://arxiv.org/abs/hep-th/9403054)

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$$\underline{\text{Massless Fermions}} \equiv H_{\bar{Q}} \cap \left\{ |\mathcal{O}\rangle \mid L_0 |\mathcal{O}\rangle = 0 \right\}$$

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$$\underline{r \ll -1}$$

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\Rightarrow Compute Massless Spectrum in LG Orbifold

RESULTS & OUTLOOK

- (1) Still many massless moduli
- (2) **Puzzle**: In CY,
- (3) # generations =
- (4)
- (5)

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