

Title: Symmetry deformation from quantum relational observables

Date: Nov 06, 2007 11:15 AM

URL: <http://pirsa.org/07110044>

Abstract: Observables in (quantum) General Relativity can be constructed from (quantum) reference frame -- a physical observable is then a relation between a system of interest and the reference frame. A possible interpretation of DSR can be derived from the notion of deformed reference frame (cf Liberati-Sonego-Visser). We present a toy model and study an example of such quantum relational observables. We show how the intrinsic quantum nature of the reference frame naturally leads to a deformation of the symmetries, comforting DSR to be a good candidate to describe the QG semi-classical regime.

OUTLINE

- Flat semi-classical space time: **modification of symmetries** → **notion of reference frame revised**. Is this effect due to a possible quantum nature of the reference frame?
- Why a quantum reference frame is natural from the Quantum Gravity perspective?
- Quantum reference frame in a toy model.

PART I:

There is now hope to probe semi-classical QG effects in experiments: need to construct a theory and in this context provide different experimental sets-up together with predictions.

Ideally this theory should be derived from:

$$\int [d\phi_M][dg] e^{i \int \mathcal{L}_M(\phi_M, g) + \mathcal{L}_{GR}(g)},$$

$\mathcal{L}_M(\phi_M, g)$, $\mathcal{L}_{GR}(g)$, lagrangians for resp. matter and grav. dof.

This is the "conservative" point of view. Introducing a semi-classical state for flat space-time, we obtain the effective lagrangian for matter.

$$\int [d\phi_M] e^{i \int \tilde{\mathcal{L}}_M(\phi)}$$

To regularize this expression, we could use for example the spinfoam approach.

If we take a "liberal" point of view, we can implement extra-dimensions, other fundamental objects (eg strings, D-particles...)...

- In 3d, the calculation can be done exactly, one obtains Deformed Special Relativity (*Freidel, Livine, hep-th/0512113*).
- In 4d, it can't be done yet, so **one models by hand the effective theory reproducing the QG fluctuations around flat space-time:**
- \rightsquigarrow Modify flat space-time (ie **Special Relativity**) to incorporate QG effects (eg Planck scale) **at the kinematical level.**

Special Relativity is characterized by the Poincaré symmetries.

↪ If there is a new fundamental scale (M_P), Poincaré symmetries can be

- **Unmodified:** need to understand why it is so (eg discrete structure consistent with Lorentz symmetries (*causal sets, Snyder...*)).
- **Broken:** this is very much constrained at this time (*cf Stefano's talk tomorrow*).
- **Deformed:** there is still a (modified) relativity principle: symmetries are deformed to be consistent with the new scale. This is Deformed (or Doubly) Special Relativity.
- To distinguish between the different possibilities, one needs either a full derivation or some experiments.

Special Relativity is characterized by the Poincaré symmetries.

↪ If there is a new fundamental scale (M_P), Poincaré symmetries can be

- **Unmodified:** need to understand why it is so (eg discrete structure consistent with Lorentz symmetries (*causal sets, Snyder...*)).
- **Broken:** this is very much constrained at this time (cf Stefano's talk tomorrow).
- **Deformed:** there is still a (modified) relativity principle: symmetries are deformed to be consistent with the new scale. This is Deformed (or Doubly) Special Relativity.
- To distinguish between the different possibilities, one needs either a full derivation or some experiments.

Special Relativity is characterized by the Poincaré symmetries.

↪ If there is a new fundamental scale (M_P), Poincaré symmetries can be

- **Unmodified:** need to understand why it is so (eg discrete structure consistent with Lorentz symmetries (*causal sets, Snyder...*)).
- **Broken:** this is very much constrained at this time (cf Stefano's talk tomorrow).
- **Deformed:** there is still a (modified) relativity principle: symmetries are deformed to be consistent with the new scale. This is Deformed (or Doubly) Special Relativity.
- To distinguish between the different possibilities, one needs either a full derivation or some experiments.

Special Relativity is characterized by the Poincaré symmetries.

↪ If there is a new fundamental scale (M_P), Poincaré symmetries can be

- **Unmodified:** need to understand why it is so (eg discrete structure consistent with Lorentz symmetries (*causal sets, Snyder...*)).
- **Broken:** this is very much constrained at this time (cf Stefano's talk tomorrow).
- **Deformed:** there is still a (modified) relativity principle: symmetries are deformed to be consistent with the new scale. This is Deformed (or Doubly) Special Relativity.
- **To distinguish between the different possibilities, one needs either a full derivation or some experiments.**

DEFORMED SPECIAL RELATIVITY (DSR)

We modify the action of the boost on momentum, but not the Lorentz algebra:

↪ Non linear realization of the Lorentz group.

π_μ the "platonic" momentum variable carrying linear representation (therefore unbounded).

$p_\mu = F_\mu(\pi, M_P)$ the "physical" momentum bounded by M_P . F is a **non linear** invertible map.

$$\tilde{\Lambda} \cdot p = F(\Lambda \cdot F^{-1}(p)), \quad [\tilde{\Lambda}_{\mu\nu}, \tilde{\Lambda}_{\rho\sigma}] = \eta_{\mu\rho}\tilde{\Lambda}_{\nu\sigma} - \eta_{\mu\sigma}\tilde{\Lambda}_{\nu\rho} - \eta_{\nu\rho}\tilde{\Lambda}_{\mu\sigma} + \eta_{\nu\sigma}\tilde{\Lambda}_{\mu\rho}.$$

This leaves the modified mass shell invariant:

$$\pi^2 = m^2 \rightarrow E^2 = m^2 + p^2 + \sum_{n=1}^{\infty} \alpha_n(p, M_P)$$

DSR

There are various versions on how momenta add, how space-time is reconstructed, ...

$$\{x^\mu, x_\nu\} = \delta_\nu^\mu + g(p, M_P)$$

$$\{x^\mu, p_\nu\} = \delta_\nu^\mu + f(p, M_P)$$

$$p_1 \oplus p_2 = ?$$

DSR

This is not simply a rewriting of Special Relativity in a funny choice of coordinates if we choose carefully a new momenta addition for example, consistent with the deformed symmetries.

DSR

There are various versions on how momenta add, how space-time is reconstructed, ...

$$\{x^\mu, x_\nu\} = \delta_\nu^\mu + g(p, M_P)$$

$$\{x^\mu, p_\nu\} = \delta_\nu^\mu + f(p, M_P)$$

$$p_1 \oplus p_2 = ?$$

EXAMPLE: "BICROSSPRODUCT" BASIS

This is the best mathematically defined and most studied example. Taking the platonic variables $\pi^A \in \mathbb{R}^5$ we have

- **Momentum:** $p_0 \equiv M_P \ln \frac{\pi_4 - \pi_0}{M_P}$, $p_i \equiv M_P \frac{\pi_i}{\pi_0 - \pi_4}$
 $\rightarrow (2M_P \sinh \frac{p_0}{2M_P})^2 - \vec{p}^2 e^{-\frac{p_0}{M_P}} = m^2$
- κ -Minkowski spacetime:
 $\{x_0, x_i\} = -\frac{1}{M_P} x_i$, $\{x_i, x_j\} = 0$
- **Poisson brackets:**
 $\{x_0, p_0\} = 1$, $\{x_i, p_j\} = -\delta_{ij}$, $\{x_0, p_i\} = -\frac{1}{M_P} p_i$
- **Momentum addition:**
 $(p_1 \oplus p_2)^0 = p_1^0 + p_2^0$, $(p_1 \oplus p_2)^i = e^{-p_2^0/M_P} p_1^i + p_2^i$

- People study this mathematically well-defined theory: **Quantum field theory on κ -Minkowski**.
- There is a number of physical concrete questions arising from the general approach.
 - What does it mean to have a non-commutative space of momenta?
 - What is the notion of "propagator"?
 - $\mathcal{P} = \mathcal{E} \otimes \mathcal{E} = \mathcal{T} \otimes \mathcal{T} = \mathcal{E} \otimes \mathcal{E}$
 - What is the physical meaning of \mathcal{E} ?
 - ...
- Can we explore this theory from a different angle? Symmetries relate reference frames: **modified symmetries means a modified notion of reference frame?**

- People study this mathematically well-defined theory: **Quantum field theory on κ -Minkowski**.
- There is a number of **physical** concrete questions arising from the general approach.
 - What does it mean to have a non-commutative sum of momenta?
 - There is the "soccer ball" problem: eg
$$|\vec{p}| < M_{Pl}c \rightarrow |\vec{p}_1 \oplus \vec{p}_2| < M_{Pl}c$$
 - What is the physical meaning of π ?
 - ...
- Can we explore this theory from a different angle? Symmetries relate reference frames: **modified symmetries means a modified notion of reference frame?**

- People study this mathematically well-defined theory: **Quantum field theory on κ -Minkowski**.
- There is a number of **physical** concrete questions arising from the general approach.
 - What does it mean to have a non-commutative sum of momenta?
 - There is the "soccer ball" problem: eg
$$|\vec{p}| < M_{Pl}c \rightarrow |\vec{p}_1 \oplus \vec{p}_2| < M_{Pl}c$$
 - What is the physical meaning of π ?
 - ...
 - Can we explore this theory from a different angle? Symmetries relate reference frames: **modified symmetries means a modified notion of reference frame?**

- People study this mathematically well-defined theory: Quantum field theory on κ -Minkowski.
- There is a number of **physical** concrete questions arising from the general approach.
 - What does it mean to have a non-commutative sum of momenta?
 - There is the "soccer ball" problem: eg
$$|\vec{p}| < M_P c \rightarrow |\vec{p}_1 \oplus \vec{p}_2| < M_P c$$
 - What is the physical meaning of π ?
 - ...
- Can we explore this theory from a different angle? Symmetries relate reference frames: modified symmetries means a modified notion of reference frame?

- People study this mathematically well-defined theory: Quantum field theory on κ -Minkowski.
- There is a number of **physical** concrete questions arising from the general approach.
 - What does it mean to have a non-commutative sum of momenta?
 - There is the "soccer ball" problem: eg
$$|\vec{p}| < M_{Pl}c \rightarrow |\vec{p}_1 \oplus \vec{p}_2| < M_{Pl}c$$
 - What is the physical meaning of π ?
 - ...
- Can we explore this theory from a different angle? Symmetries relate reference frames: modified symmetries means a modified notion of reference frame?

- People study this mathematically well-defined theory: **Quantum field theory on κ -Minkowski**.
- There is a number of **physical** concrete questions arising from the general approach.
 - What does it mean to have a non-commutative sum of momenta?
 - There is the "soccer ball" problem: eg
$$|\vec{p}| < M_{Pl}c \rightarrow |\vec{p}_1 \oplus \vec{p}_2| < M_{Pl}c$$
 - What is the physical meaning of π ?
 - ...
- Can we explore this theory from a different angle? Symmetries relate reference frames: **modified symmetries means a modified notion of reference frame?**

SYMMETRY AND FRAME

- Introduce a reference frame e^μ_α , with $\mu \equiv$ space-time coordinates, α labels vectors.
- Let π_μ be the momentum of the particle.
- Then $p_\alpha = \pi_\mu e^\mu_\alpha$ is the measurement outcome (or coordinates) of π in e . In Minkowski space-time, $e^\mu_\alpha \sim \delta^\mu_\alpha$ so $\pi \sim p$.
- Consider another reference frame, $e^\mu_\alpha \rightarrow \bar{e}^\mu_\alpha = \Lambda^\beta_\alpha e^\mu_\beta$, Λ is a Lorentz transformation.

$$p'_\alpha = \pi_\mu \bar{e}^\mu_\alpha = \Lambda^\beta_\alpha e^\mu_\beta \pi_\mu = \Lambda^\beta_\alpha p_\beta.$$

SYMMETRY AND FRAME

- Introduce a reference frame e^μ_α , with $\mu \equiv$ space-time coordinates, α labels vectors.
- Let π_μ be the momentum of the particle.
- Then $p_\alpha = \pi_\mu e^\mu_\alpha$ is the measurement outcome (or coordinates) of π in e . In Minkowski space-time, $e^\mu_\alpha \sim \delta^\mu_\alpha$ so $\pi \sim p$.
- Consider another reference frame, $e^\mu_\alpha \rightarrow \bar{e}^\mu_\alpha = \Lambda^\beta_\alpha e^\mu_\beta$, Λ is a Lorentz transformation.

$$p'_\alpha = \pi_\mu \bar{e}^\mu_\alpha = \Lambda^\beta_\alpha e^\mu_\beta \pi_\mu = \Lambda^\beta_\alpha p_\beta.$$

SYMMETRY AND FRAME

- Introduce a reference frame e^μ_α , with $\mu \equiv$ space-time coordinates, α labels vectors.
- Let π_μ be the momentum of the particle.
- Then $p_\alpha = \pi_\mu e^\mu_\alpha$ is the measurement outcome (or coordinates) of π in e . In Minkowski space-time, $e^\mu_\alpha \sim \delta^\mu_\alpha$ so $\pi \sim p$.
- Consider another reference frame, $e^\mu_\alpha \rightarrow \bar{e}^\mu_\alpha = \Lambda^\beta_\alpha e^\mu_\beta$, Λ is a Lorentz transformation.

$$p'_\alpha = \pi_\mu \bar{e}^\mu_\alpha = \Lambda^\beta_\alpha e^\mu_\beta \pi_\mu = \Lambda^\beta_\alpha p_\beta.$$

SYMMETRY AND FRAME

- Introduce a reference frame e^μ_α , with $\mu \equiv$ space-time coordinates, α labels vectors.
- Let π_μ be the momentum of the particle.
- Then $p_\alpha = \pi_\mu e^\mu_\alpha$ is the measurement outcome (or coordinates) of π in e . In Minkowski space-time, $e^\mu_\alpha \sim \delta^\mu_\alpha$ so $\pi \sim p$.
- Consider another reference frame, $e^\mu_\alpha \rightarrow \bar{e}^\mu_\alpha = \Lambda^\beta_\alpha e^\mu_\beta$, Λ is a Lorentz transformation.

$$p'_\alpha = \pi_\mu \bar{e}^\mu_\alpha = \Lambda^\beta_\alpha e^\mu_\beta \pi_\mu = \Lambda^\beta_\alpha p_\beta.$$

SYMMETRY AND FRAME

To get a deformation of symmetry, one needs a nonlinear function of reference frame and the system ("effective reference frame"). (Liberati, Sonego,

Visser, grcq/0410113)

$$e^\mu{}_\alpha \rightarrow E^\mu{}_\alpha = \mathcal{U}(e^\mu{}_\alpha, \pi_\mu, M_P).$$

For example:

$$\mathcal{U}(e^\mu{}_\alpha, \pi_\mu, M_P) = \frac{1}{\sqrt{1 - \frac{\pi_\mu e^\mu{}_\alpha}{M_P}}} e^\mu{}_\alpha, \quad p_\alpha = \frac{1}{\sqrt{1 - \frac{\pi_\mu e^\mu{}_\alpha}{M_P}}} e^\mu{}_\alpha \pi_\mu$$

The fundamental variable is the frame $e^\mu{}_\alpha$. We have then a non-linear transformation of the effective frame:

$$p_\alpha = E^\mu{}_\alpha \pi_\mu \xrightarrow{\Lambda} p'_\alpha = \mathcal{U}(\Lambda \cdot \mathcal{U}^{-1}(E)) \pi_\mu.$$

Such frame will be associated to a **Finsler metric** (cf my talk on Saturday) or Rainbow metric.

EXAMPLE

- In a couple of papers, Liberati et al introduced a stochastic component to the tetrad induced by QG effects. (gr-qc/0511031, gr-qc/0607024)
- A particle will probe the fluctuations of the tetrad up to the scale L $E^\mu{}_\alpha(L) = \langle e^\mu{}_\alpha \rangle_L$. This scale can be naturally identified with the particle's de Broglie length $L = \hbar/p_0$.
- In this case, the effective reference frame becomes **momentum dependent** $E^\mu{}_\alpha(p_0)$.
- A similar result is obtained in the renormalization group approach (Girelli, Liberati, Percacci, Rahmede, gr-qc/0607030).
- It is however not clear how this extends in more general situations: multi-particles case...
- **The key question to answer is how the reference frame becomes system independent.**

EXAMPLE

- In a couple of papers, Liberati et al introduced a stochastic component to the tetrad induced by QG effects. (gr-qc/0511031, gr-qc/0607024)
- A particle will probe the fluctuations of the tetrad up to the scale L . $E^\mu{}_\alpha(L) = \langle e^\mu{}_\alpha \rangle_L$. This scale can be naturally identified with the particle's de Broglie length $L = \hbar/p_0$.
- In this case, the effective reference frame becomes **momentum dependent** $E^\mu{}_\alpha(p_0)$.
- A similar result is obtained in the renormalization group approach (Girelli, Liberati, Percacci, Rahmede, gr-qc/0607030).
- It is however not clear how this extends in more general situations: multi-particles case...
- **The key question to answer is how the reference frame becomes system independent.**

EXAMPLE

- In a couple of papers, Liberati et al introduced a stochastic component to the tetrad induced by QG effects. (gr-qc/0511031, gr-qc/0607024)
- A particle will probe the fluctuations of the tetrad up to the scale L . $E^\mu{}_\alpha(L) = \langle e^\mu{}_\alpha \rangle_L$. This scale can be naturally identified with the particle's de Broglie length $L = \hbar/p_0$.
- In this case, the effective reference frame becomes **momentum dependent** $E^\mu{}_\alpha(p_0)$.
- A similar result is obtained in the renormalization group approach (Girelli, Liberati, Percacci, Rahmede, gr-qc/0607030).
- It is however not clear how this extends in more general situations: multi-particles case...
- **The key question to answer is how the reference frame becomes system independent.**

- When constructing the effective theory by hand, we had the choice: the symmetry is either **broken** (**preferred frame**) or **untouched** (**usual frame?**) or **deformed** (**modified frame**).
- Can we find a general argument from the QG perspective that will pinpoint one of these different possibilities?
- In QG, we need to talk about Quantum Reference Frame (QRF).
 - ~> **Why is that?**
 - ~> **I will study QRF in a toy model that will generate some modification of the symmetry.**

PART II:

Why Quantum Reference Frame in Quantum Gravity?

WHY A RF?

- Gravity can be described as a constrained system (Dirac, Lectures on Quantum Mechanics)
- First class constraints \mathcal{C}_i can be seen as encoding symmetries (in the General Relativity case, the diffeomorphisms).
- Observables \mathcal{O} are phase space functions that commute with constraints.

$$\{\mathcal{O}, \mathcal{C}_i\} = 0$$

It is often hard to construct in general the set of observables, but using **physical** reference frame helps in general (cf Rovelli's book, Dittrich's papers, Giddings-Hartle-Marolf's paper).

WHY A RF?

- Gravity can be described as a constrained system (Dirac, Lectures on Quantum Mechanics)
- First class constraints \mathcal{C}_i can be seen as encoding symmetries (in the General Relativity case, the diffeomorphisms).
- Observables \mathcal{O} are phase space functions that commute with constraints.

$$\{\mathcal{O}, \mathcal{C}_i\} = 0$$

It is often hard to construct in general the set of observables, but using **physical** reference frame helps in general (cf Rovelli's book, Dittrich's papers, Giddings-Hartle-Marolf's paper).

WHY A RF?

- Gravity can be described as a constrained system (Dirac, Lectures on Quantum Mechanics)
- First class constraints \mathcal{C}_i can be seen as encoding symmetries (in the General Relativity case, the diffeomorphisms).
- Observables \mathcal{O} are phase space functions that commute with constraints.

$$\{\mathcal{O}, \mathcal{C}_i\} = 0$$

It is often hard to construct in general the set of observables, but using **physical** reference frame helps in general (cf Rovelli's book, Dittrich's papers, Giddings-Hartle-Marolf's paper).

WHY A RF?

- Gravity can be described as a constrained system (Dirac, Lectures on Quantum Mechanics)
- First class constraints \mathcal{C}_i can be seen as encoding symmetries (in the General Relativity case, the diffeomorphisms).
- Observables \mathcal{O} are phase space functions that commute with constraints.

$$\{\mathcal{O}, \mathcal{C}_i\} = 0$$

It is often hard to construct in general the set of observables, but using **physical** reference frame helps in general (cf Rovelli's book, Dittrich's papers, Giddings-Hartle-Marolf's paper).

WHY A RF?

- Gravity can be described as a constrained system (Dirac, Lectures on Quantum Mechanics)
- First class constraints \mathcal{C}_i can be seen as encoding symmetries (in the General Relativity case, the diffeomorphisms).
- Observables \mathcal{O} are phase space functions that commute with constraints.

$$\{\mathcal{O}, \mathcal{C}_i\} = 0$$

It is often hard to construct in general the set of observables, but using **physical** reference frame helps in general (cf Rovelli's book, Dittrich's papers, Giddings-Hartle-Marolf's paper).

WHY A RF?

- Gravity can be described as a constrained system (Dirac, Lectures on Quantum Mechanics)
- First class constraints \mathcal{C}_i can be seen as encoding symmetries (in the General Relativity case, the diffeomorphisms).
- Observables \mathcal{O} are phase space functions that commute with constraints.

$$\{\mathcal{O}, \mathcal{C}_i\} = 0$$

It is often hard to construct in general the set of observables, but using **physical** reference frame helps in general (cf Rovelli's book, Dittrich's papers, Giddings-Hartle-Marolf's paper).

WHY A RF?

- Gravity can be described as a constrained system (Dirac, Lectures on Quantum Mechanics)
- First class constraints \mathcal{C}_i can be seen as encoding symmetries (in the General Relativity case, the diffeomorphisms).
- Observables \mathcal{O} are phase space functions that commute with constraints.

$$\{\mathcal{O}, \mathcal{C}_i\} = 0$$

It is often hard to construct in general the set of observables, but using **physical** reference frame helps in general (cf Rovelli's book, Dittrich's papers, Giddings-Hartle-Marolf's paper).

WHY A RF?

- Gravity can be described as a constrained system (Dirac, Lectures on Quantum Mechanics)
- First class constraints \mathcal{C}_i can be seen as encoding symmetries (in the General Relativity case, the diffeomorphisms).
- Observables \mathcal{O} are phase space functions that commute with constraints.

$$\{\mathcal{O}, \mathcal{C}_i\} = 0$$

It is often hard to construct in general the set of observables, but using **physical** reference frame helps in general (cf Rovelli's book, Dittrich's papers, Giddings-Hartle-Marolf's paper).

EXAMPLE 1

Relativistic particle:

$$\left(\begin{array}{l} S = m \int \sqrt{\eta_{\mu\nu} \dot{x}^\mu \dot{x}^\nu} d\tau \\ p_\nu = m \frac{\eta_{\mu\nu} \dot{x}^\mu}{\sqrt{\eta_{\mu\nu} \dot{x}^\mu \dot{x}^\nu}} \\ (\delta_{\mu\nu} - \frac{\dot{x}_\mu \dot{x}_\nu}{\eta_{\alpha\beta} \dot{x}^\alpha \dot{x}^\beta}) \ddot{x}^\mu = 0 \\ Hd\tau = p_\mu \dot{x}^\mu - \mathcal{L} = 0 \end{array} \right) \text{Legendre transform} \longleftrightarrow \left(\begin{array}{l} S = \int (\dot{x}^\mu p_\mu - \lambda(p^2 - m^2)) d\tau \\ \dot{x}^\mu = 2\lambda p^\mu \\ \dot{p}^\mu = 0 \\ p^2 = m^2 \end{array} \right)$$

- λ is a Lagrange multiplier encoding the **mass-shell constraint**.
- This constraint encodes the **time reparametrization symmetry**.
- Observables \mathcal{O} are functions on phase space such that $\{\mathcal{O}, p^2 - m^2\} = 0$.
 $J_{\mu\nu}, p_\mu$ commute since p^2 is Poincaré Casimir, but $\{x^\mu, p^2 - m^2\} = p^\mu \neq 0$.

To define observable positions, **we introduce a clock** $\mathcal{F} = x^0$, a "time reference frame". Consider

$$X^\mu = x^\mu + \frac{p^\mu}{p^0}(x^0 - T) \Rightarrow \{X^\mu, p^2 - m^2\} = 0$$

so X^μ is observable. It is a **relational** coordinate:

When $x^0 = T$ then $X^\mu = x^\mu \iff$ **when the clock indicates T , the particle is at the position x^μ .**

Other choice of relational coordinates, ie other choice of clock $\mathcal{F} = x^\mu p_\mu$:

$$\tilde{X}^\mu = x^\mu + \frac{p^\mu}{p^2}(x^\nu p_\nu - T)$$

These relational coordinates become non-commutative: $\{\tilde{X}_\mu, \tilde{X}_\nu\} = \frac{1}{p^2} J_{\mu\nu}$

When we have (first class) constraints or symmetries, relational quantities provide a natural set of observables.

To construct a relational quantity, we introduce a reference frame and define a relation between the system of interest and the reference frame.

EXAMPLE 2

Spin systems: Consider a set of **non interacting** classical spin particles $\vec{S}^{(i)}$. We impose that the total spin is zero: we impose **global rotational symmetry**, through the constraints

$$S_x^{(tot)} = 0 \quad S_y^{(tot)} = 0, \quad S_z^{(tot)} = 0.$$

A **relative angle** is a natural observable when dealing with rotational symmetry: Consider the system $\vec{S}^{(k)} = \vec{S}$, pick one set of spins \vec{J}^1 as our reference axis, then

$$s^1 = \vec{S} \cdot \vec{J}^1 \Rightarrow \{s^1, S_a^{(tot)}\} = 0$$

To construct a 3d reference frame, we pick up another set of spins \vec{J}^2 . We can then define $\vec{J}^3 = \vec{J}^1 \wedge \vec{J}^2$. We have then the physical relational coordinates:

$$s^i = \vec{S} \cdot \vec{J}^i.$$

EXAMPLE 2

Spin systems: Consider a set of **non interacting** classical spin particles $\vec{S}^{(i)}$. We impose that the total spin is zero: we impose **global rotational symmetry**, through the constraints

$$S_x^{(tot)} = 0 \quad S_y^{(tot)} = 0, \quad S_z^{(tot)} = 0.$$

A **relative angle** is a natural observable when dealing with rotational symmetry: Consider the system $\vec{S}^{(k)} = \vec{S}$, pick one set of spins \vec{J}^1 as our reference axis, then

$$s^1 = \vec{S} \cdot \vec{J}^1 \Rightarrow \{s^1, S_a^{(tot)}\} = 0$$

To construct a 3d reference frame, we pick up another set of spins \vec{J}^2 . We can then define $\vec{J}^3 = \vec{J}^1 \wedge \vec{J}^2$. We have then the physical relational coordinates:

$$s^i = \vec{S} \cdot \vec{J}^i.$$

EXAMPLE 2

Spin systems: Consider a set of **non interacting** classical spin particles $\vec{S}^{(i)}$. We impose that the total spin is zero: we impose **global rotational symmetry**, through the constraints

$$S_x^{(tot)} = 0 \quad S_y^{(tot)} = 0, \quad S_z^{(tot)} = 0.$$

A **relative angle** is a natural observable when dealing with rotational symmetry: Consider the system $\vec{S}^{(k)} = \vec{S}$, pick one set of spins \vec{J}^1 as our reference axis, then

$$s^1 = \vec{S} \cdot \vec{J}^1 \Rightarrow \{s^1, S_a^{(tot)}\} = 0$$

To construct a 3d reference frame, we pick up another set of spins \vec{J}^2 . We can then define $\vec{J}^3 = \vec{J}^1 \wedge \vec{J}^2$. We have then the physical relational coordinates:

$$s^i = \vec{S} \cdot \vec{J}^i.$$

EXAMPLE 3

GPS coordinates: (cf Rovelli's book)

Consider a point P in (2d) space-time. We impose diffeomorphisms symmetry: P can have any arbitrary coordinates as parametrization of the manifold.

We can define physical system by considering two "satellites" with relativistic velocities V_{μ}^1, V_{μ}^2 , emitting radio signals.

Then the physical coordinates of P with respect to the satellites are:

$$s^{\alpha} = x^{\mu} V_{\mu}^{\alpha} - \sqrt{(x^{\mu} V_{\mu}^{\alpha})^2 - x^{\mu} x_{\mu}}.$$

This is the Einstein's "incidence point".

EXAMPLE 3

GPS coordinates: (cf Rovelli's book)

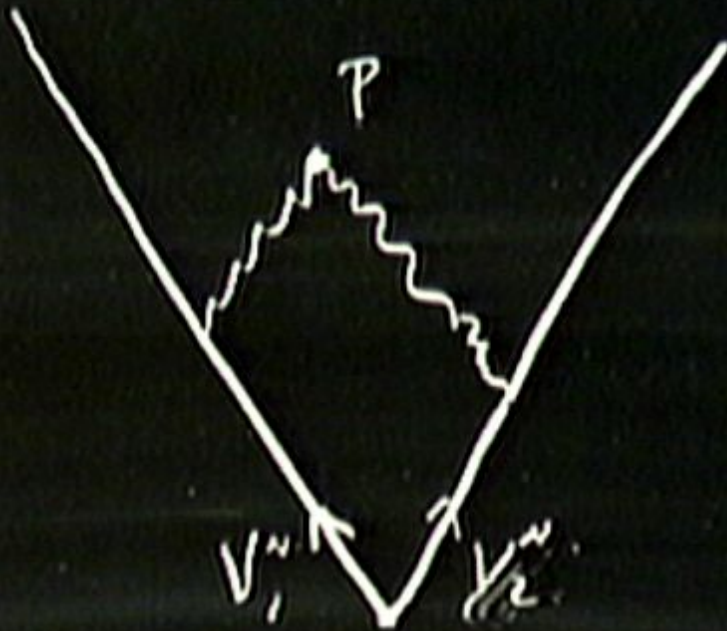
Consider a point P in (2d) space-time. We impose diffeomorphisms symmetry: P can have any arbitrary coordinates as parametrization of the manifold.

We can define physical system by considering two "satellites" with relativistic velocities V_{μ}^1, V_{μ}^2 , emitting radio signals.

Then the physical coordinates of P with respect to the satellites are:

$$s^{\alpha} = x^{\mu} V_{\mu}^{\alpha} - \sqrt{(x^{\mu} V_{\mu}^{\alpha})^2 - x^{\mu} x_{\mu}}.$$

This is the Einstein's "incidence point".



EXAMPLE 3

GPS coordinates: (cf Rovelli's book)

Consider a point P in (2d) space-time. We impose diffeomorphisms symmetry: P can have any arbitrary coordinates as parametrization of the manifold.

We can define physical system by considering two "satellites" with relativistic velocities V_{μ}^1, V_{μ}^2 , emitting radio signals.

Then the physical coordinates of P with respect to the satellites are:

$$s^{\alpha} = x^{\mu} V_{\mu}^{\alpha} - \sqrt{(x^{\mu} V_{\mu}^{\alpha})^2 - x^{\mu} x_{\mu}}.$$

This is the Einstein's "incidence point".

QUANTIZATION

One follows Dirac canonical procedure:

- Phase space \rightarrow "kinematical" Hilbert space
- $C_i \rightarrow \hat{C}_i$ (careful with ordering ambiguities...)
- Kernel of \hat{C}_i is the physical Hilbert space.
- Observables: $[\hat{O}, \hat{C}_i] = 0$. It is now natural to look at quantum relational observables, constructed out of a quantum reference frame.

EXAMPLE OF A QRF

Spin systems revisited: Consider a set of **non interacting quantum spin particles** $\vec{\xi}^{(i)}$. The quantum constraints are obtained in a straightforward manner

$$\hat{\xi}_x^{(tot)} = 0 \quad \hat{\xi}_y^{(tot)} = 0, \quad \hat{\xi}_z^{(tot)} = 0.$$

Quantum relational observables:

$$\mathfrak{s}^i = \vec{\xi} \cdot \vec{J}^i \Rightarrow \{\mathfrak{s}^i, \hat{\xi}_a^{(tot)}\} = 0$$

These are quantum relational observables and J^μ_α is a quantum reference frame.

- Quantum Gravity is a quantum constrained system. Observables can be constructed as relational in terms of a quantum reference frame.
- We need to understand the physical implication of having quantum reference frames.

QUESTIONS TO EXPLORE:

- What is the precision of a QRF?
- What is the robustness of a QRF?
- Do we get non-commutative coordinates?
- What is the notion of symmetry?
- How do we coarse-grain relational observables (multi-particles state)?

These questions are of interest for both the Quantum Information Theory and QG community. In particular, the QIT people would implement that in concrete experiments.

PART III:

Quantum Reference Frame in a toy model

TOY MODEL

We consider the set of quantum spin particles. The QRF axis \vec{J}^1, \vec{J}^2 are made of spin ℓ and satisfy

$$[J^\mu_\alpha, J^\nu_\beta] = \begin{cases} 0 & \text{if } \alpha \neq \beta \\ \frac{i}{\sqrt{\ell(\ell+1)}} \epsilon^{\eta\mu\nu} J^\eta_\alpha & \text{if } \alpha = \beta \end{cases}$$

We introduce also

$$\vec{J}_3 = \vec{J}_1 \wedge \vec{J}_2, \quad \mathfrak{s}_3 = \vec{\xi} \cdot \vec{J}_3$$

The quantum system $\vec{\xi}$ is a spin j . The relational coordinates satisfy the modified algebra: analog to the emergence of "non-commutativity".

$$[\mathfrak{s}_1, \mathfrak{s}_2] = i\mathfrak{s}_3$$

$$[\mathfrak{s}_2, \mathfrak{s}_3] = i\mathfrak{s}_1 + -iJ^\mu_2 J_{\mu 1} \mathfrak{s}_2 + i \frac{\mathfrak{s}_1 \mathfrak{s}_2 - (\vec{\xi} \cdot \vec{\xi}) J^\mu_2 J_{\mu 1}}{\sqrt{\ell(\ell+1)}}$$

$$[\mathfrak{s}_3, \mathfrak{s}_1] = i\mathfrak{s}_2 + -iJ^\mu_1 J_{\mu 2} \mathfrak{s}_1 + i \frac{\mathfrak{s}_1 \mathfrak{s}_2 - (\vec{\xi} \cdot \vec{\xi}) J^\mu_1 J_{\mu 2}}{\sqrt{\ell(\ell+1)}}$$

We obtain the "semi-classical relational" coordinates by taking the RF in a semi-classical state $\rho = \rho_1 \otimes \rho_2$

$$\xi_{\alpha}^{\text{semi-classical}} = \text{Tr}_{RF} (\mathfrak{s}_{\alpha} \rho),$$

taking each of the axis $a = 1, 2$ of the reference frame in the adequate coherent state $|\ell_a, \ell_a\rangle$ and $\ell_a \rightarrow \infty$.

$$\langle \ell_a, \ell_a | J_a^{\mu} | \ell_a, \ell_a \rangle = \delta_{a\mu} \frac{\ell_a}{\sqrt{\ell_a(\ell_a + 1)}} \xrightarrow{\ell_a \rightarrow \infty} \delta_{a\mu},$$

$$\xi_{\alpha}^{\text{semi-classical}} \xrightarrow{\ell_a \rightarrow \infty} \xi_{\alpha}$$

We recover the standard $su(2)$ algebra.

$$[\mathfrak{s}_i, \mathfrak{s}_j] \rightarrow [\xi_i, \xi_j] = i\epsilon_{ij}^k \xi_k$$

The semi-classical coordinates can be decomposed into a sum of $2j + 1$ orthogonal projectors P_{α}^m , $m = -j, \dots, j$ associated with distinct eigenvalues m .

MEASUREMENTS: PROJECTORS

σ state of the source particle, ρ state of reference frame. The state $\rho \otimes \sigma$ of s^α is in $j \otimes \ell \sim \bigoplus_{k=\ell-j}^{\ell+j} k$.

Spectral theorem tells us that:

$$s_\alpha = \sum_{k=\ell-j}^{\ell+j} \lambda^k \Pi_\alpha^k \text{ with eigenvalues } \lambda^k = \frac{k(k+1) - \ell(\ell+1) - j(j+1)}{2\sqrt{\ell(\ell+1)}}$$

and projector

$$\Pi_\alpha^k = \frac{1}{N^k} \prod_{k' \neq k} (\lambda^{k'} - s_\alpha),$$

where the normalization factor is

$$N^k = \prod_{k' \neq k} (\lambda^{k'} - \lambda^k).$$

Π_α^k are non-linear functions of the relational coordinates s_α (polynomials of degree $2j$).

MEASUREMENTS OF \mathfrak{s}_α

Measurement of \mathfrak{s}_α will produce the outcomes λ^k with probability

$$Pr(\lambda^k) = \text{Tr}\{\Pi_\alpha^k \rho \otimes \sigma\}$$

INDUCED MEASUREMENTS ON SYSTEM

$$Pr(\lambda^k) = \text{Tr}\{\Pi_\alpha^k \rho \otimes \sigma\} = \text{Tr}\{\Lambda_\alpha^k \sigma\}$$

with

$$\Lambda_\alpha^k = \text{Tr}_{RF}\{\Pi_\alpha^k \rho\}.$$

Λ_α^k can be seen as an **approximation** of the measure of the "semi-classical relational" coordinates.

$$\Lambda_\alpha^k = (1 - \epsilon^k) P_\alpha^k + \sum_{k' \neq k} \epsilon^{k',k} P_\alpha^{k'},$$

where $\epsilon^k, \epsilon^{k',k}$ depend on ρ and are of order ℓ^{-1} . We can then play on ρ to determine which ρ provides the optimum measurement: **this is one way to encode the precision of the QRF.**

INDUCED MEASUREMENTS ON RF: BACK-ACTION ON QRF

First note that

$$[J^\mu_\alpha, \mathfrak{s}_\alpha] = \frac{i}{\ell} \epsilon^{\mu\nu\eta} J^\nu_\alpha \xi^\eta.$$

By the uncertainty principle, a measurement of \mathfrak{s}_α will thus alter the value of J_α , and so disturb any future measurement that make use of that frame.

After measurement, we have the (normalized) state:

$$\frac{\Pi_{\alpha}^k(\rho \otimes \sigma)\Pi_{\alpha}^k}{Pr(\lambda^k)}$$

We are interested at the dynamics of the QRF induced by the measurements.

After a measurement, we sum over the different outcomes and trace out the system $\sigma = \sum_i s_i |\phi_i\rangle\langle\phi_i|$

$$\rho' = \sum_k \text{Tr}_S\{\Pi_{\alpha}^k(\rho \otimes \sigma)\Pi_{\alpha}^k\} = \mathcal{E}_{\alpha}(\rho) = \sum_b K_b \rho K_b^{\dagger}$$

with $a = (k, i, i')$ and the Kraus operators $K_b = \sqrt{s_j} \langle \phi_{i'} | \Pi_{\alpha}^k | \phi_i \rangle$ satisfying $\sum_b K_b^{\dagger} K_b = 1$.

EXAMPLE

If $\vec{\xi}$ is a spin $\frac{1}{2}$, assuming the QRF in coherent state, then at first order in ℓ^{-1}

$$\begin{aligned}\mathcal{E}_\alpha(\rho_\alpha) &\approx \rho_\alpha - \frac{i}{\ell} \epsilon_{\mu\nu\eta} \mathcal{J}^\mu_\alpha \langle \xi^\nu \rangle \sin \theta [J^\eta_\alpha, \rho_\alpha] \\ &= \mathcal{U}_\alpha^\dagger(J, \xi, 1/\ell)(\rho_\alpha)\end{aligned}$$

with

$$\mathcal{J}^\mu_\alpha = \text{Tr}(J^\mu_\alpha \rho), \quad \langle \xi^\nu \rangle = \text{Tr}(\xi^\nu \sigma), \quad \cos \theta = \vec{\mathcal{J}}_\alpha \cdot \langle \vec{\xi} \rangle$$

The reference frame axis undergoes **a rotation around the axis** $\vec{\mathcal{J}}_\alpha \wedge \langle \vec{\xi} \rangle$ of angle θ .

This is in the **Schrodinger picture**, we can move to the **Heisenberg picture**.

$$J^\mu_\alpha \xrightarrow{n} \mathcal{U}_\alpha^n(J^\mu_\alpha, \xi, 1/\ell),$$

after n measurements. Non-unitary effects (noise) will appear in general at higher orders.

After measurement, we have the (normalized) state:

$$\frac{\Pi_{\alpha}^k(\rho \otimes \sigma)\Pi_{\alpha}^k}{Pr(\lambda^k)}$$

We are interested at the dynamics of the QRF induced by the measurements.

After a measurement, we sum over the different outcomes and trace out the

system $\sigma = \sum_i s_i |\phi_i\rangle\langle\phi_i|$

$$\rho' = \sum_k \text{Tr}_S\{\Pi_{\alpha}^k(\rho \otimes \sigma)\Pi_{\alpha}^k\} = \mathcal{E}_{\alpha}(\rho) = \sum_b K_b \rho K_b^{\dagger}$$

with $a = (k, i, i')$ and the Kraus operators $K_b = \sqrt{s_j} \langle\phi_{i'}|\Pi_{\alpha}^k|\phi_i\rangle$ satisfying

$$\sum_b K_b^{\dagger} K_b = 1.$$

EXAMPLE

If $\vec{\xi}$ is a spin $\frac{1}{2}$, assuming the QRF in coherent state, then at first order in ℓ^{-1}

$$\begin{aligned}\mathcal{E}_\alpha(\rho_\alpha) &\approx \rho_\alpha - \frac{i}{\ell} \epsilon_{\mu\nu\eta} \mathcal{J}^\mu_\alpha \langle \xi^\nu \rangle \sin \theta [J^\eta_\alpha, \rho_\alpha] \\ &= \mathcal{U}_\alpha^\dagger(J, \xi, 1/\ell)(\rho_\alpha)\end{aligned}$$

with

$$\mathcal{J}^\mu_\alpha = \text{Tr}(J^\mu_\alpha \rho), \quad \langle \xi^\nu \rangle = \text{Tr}(\xi^\nu \sigma), \quad \cos \theta = \vec{\mathcal{J}}_\alpha \cdot \langle \vec{\xi} \rangle$$

The reference frame axis undergoes **a rotation around the axis** $\vec{\mathcal{J}}_\alpha \wedge \langle \vec{\xi} \rangle$ of angle θ .

This is in the **Schrodinger picture**, we can move to the **Heisenberg picture**.

$$J^\mu_\alpha \xrightarrow{n} \mathcal{U}_\alpha^n(J^\mu_\alpha, \xi, 1/\ell),$$

after n measurements. Non-unitary effects (noise) will appear in general at higher orders.

- The evolution can be decomposed into an invertible part and some non-invertible part (noise). When $\ell \gg n$ the number of measurements then the invertible part is dominant. This is a simplified example of the "decoupling theorem".
- The measurement of the relational quantum coordinate of a source particle induces a back-action on the RF, which is in general a non-linear function \mathcal{E}_α of the quantum relational coordinate p_α .

$$J^\mu_\alpha \rightarrow \mathcal{U}_\alpha^n(J^\mu_\alpha, \xi, 1/\ell),$$

This is to be compared with the DSR case:

$$e^\mu_\alpha \rightarrow \mathcal{U}(e^\mu_\alpha, \pi, M_P)$$

- One can then ask for which state the QRF will be the most robust, and possibly the most robust *and* the most precise as there might be a trade-off between the two.

- The evolution can be decomposed into an invertible part and some non-invertible part (noise). When $\ell \gg n$ the number of measurements then the invertible part is dominant. This is a simplified example of the "decoupling theorem".
- The measurement of the relational quantum coordinate of a source particle induces a back-action on the RF, which is in general a non-linear function \mathcal{E}_α of the quantum relational coordinate p_α .

$$J^\mu_\alpha \rightarrow \mathcal{U}_\alpha^n(J^\mu_\alpha, \xi, 1/\ell),$$

This is to be compared with the DSR case:

$$e^\mu_\alpha \rightarrow \mathcal{U}(e^\mu_\alpha, \pi, M_P)$$

- One can then ask for which state the QRF will be the most robust, and possibly the most robust *and* the most precise as there might be a trade-off between the two.

- The evolution can be decomposed into an invertible part and some non-invertible part (noise). When $\ell \gg n$ the number of measurements then the invertible part is dominant. This is a simplified example of the "decoupling theorem".
- The measurement of the relational quantum coordinate of a source particle induces a back-action on the RF, which is in general a non-linear function \mathcal{E}_α of the quantum relational coordinate p_α .

$$J^\mu_\alpha \rightarrow \mathcal{U}_\alpha^n(J^\mu_\alpha, \xi, 1/\ell),$$

This is to be compared with the DSR case:

$$e^\mu_\alpha \rightarrow \mathcal{U}(e^\mu_\alpha, \pi, M_P)$$

- One can then ask for which state the QRF will be **the most robust**, and possibly **the most robust and the most precise** as there might be a trade-off between the two.

MULTI-PARTICLES:

Remember for DSR we have eg:

$$p_0^{(tot)} = p_0^{(1)} + p_0^{(2)}, \quad p_i^{(tot)} = e^{-p_0^{(2)}/M_P} p_i^{(1)} + p_i^{(2)}.$$

Within the reference frame interpretation, $\pi^{(tot)} = \pi^{(1)} + \pi^{(2)}$ is the total intrinsic momentum. The physical momentum is naturally defined then as

$$p_\alpha^{(tot)} = \left(\pi^{(1)} + \pi^{(2)} \right)_\mu e^\mu_\alpha = p_\alpha^{(1)} + p_\alpha^{(2)}.$$

If the frame is "deformed"

$$p_\alpha^{(tot)} = \left(\pi^{(1)} + \pi^{(2)} \right)_\mu E^\mu_\alpha(e, (\pi^{(1)} + \pi^{(2)}), \gamma M_P),$$

where I added γ to solve the "soccer ball" problem.

Can we get something analog in the toy model?

The total intrinsic spin is

$$\vec{\xi}^{(tot)} = \vec{\xi}^{(1)} + \vec{\xi}^{(2)}, \quad [\xi_{\mu}^{(1)}, \xi_{\nu}^{(2)}] = 0, \quad [\xi_{\mu}^{(tot)}, \xi_{\mu}^{(k)}] = 0$$

The relational coordinate is therefore

$$\mathfrak{s}_{\alpha}^{(tot)} = \xi_{\mu}^{(tot)} J_{\alpha}^{\mu}$$

But then

$$[\mathfrak{s}_{\alpha}^{(1)}, \mathfrak{s}_{\alpha}^{(2)}] = \frac{\epsilon_{\rho}^{\mu\nu}}{\sqrt{\ell(\ell+1)}} \xi_{\mu}^{(1)} \xi_{\nu}^{(2)} J_{\alpha}^{\rho},$$

that is **the components of the physical total spin are not commuting anymore.**

Also we have

$$[\mathfrak{s}_{\alpha}^{(k)}, \mathfrak{s}_{\alpha}^{(tot)}] \neq 0 \quad k = 1, 2,$$

that is **measuring the total relational spin of two particles differs from measuring their individual relational spins and adding the outcomes.**

MULTI-PARTICLES:

Remember for DSR we have eg:

$$p_0^{(tot)} = p_0^{(1)} + p_0^{(2)}, \quad p_i^{(tot)} = e^{-p_0^{(2)}/M_P} p_i^{(1)} + p_i^{(2)}.$$

Within the reference frame interpretation, $\pi^{(tot)} = \pi^{(1)} + \pi^{(2)}$ is the total intrinsic momentum. The physical momentum is naturally defined then as

$$p_\alpha^{(tot)} = \left(\pi^{(1)} + \pi^{(2)} \right)_\mu e^\mu{}_\alpha = p_\alpha^{(1)} + p_\alpha^{(2)}.$$

If the frame is "deformed"

$$p_\alpha^{(tot)} = \left(\pi^{(1)} + \pi^{(2)} \right)_\mu E^\mu{}_\alpha(e, (\pi^{(1)} + \pi^{(2)}), \gamma M_P),$$

where I added γ to solve the "soccer ball" problem.

Can we get something analog in the toy model?

The total intrinsic spin is

$$\vec{\xi}^{(tot)} = \vec{\xi}^{(1)} + \vec{\xi}^{(2)}, \quad [\xi_{\mu}^{(1)}, \xi_{\nu}^{(2)}] = 0, \quad [\xi_{\mu}^{(tot)}, \xi_{\mu}^{(k)}] = 0$$

The relational coordinate is therefore

$$\mathfrak{s}_{\alpha}^{(tot)} = \xi_{\mu}^{(tot)} J_{\alpha}^{\mu}$$

But then

$$[\mathfrak{s}_{\alpha}^{(1)}, \mathfrak{s}_{\alpha}^{(2)}] = \frac{\epsilon_{\rho}^{\mu\nu}}{\sqrt{\ell(\ell+1)}} \xi_{\mu}^{(1)} \xi_{\nu}^{(2)} J_{\alpha}^{\rho},$$

that is **the components of the physical total spin are not commuting anymore.**

Also we have

$$[\mathfrak{s}_{\alpha}^{(k)}, \mathfrak{s}_{\alpha}^{(tot)}] \neq 0 \quad k = 1, 2,$$

that is **measuring the total relational spin of two particles differs from measuring their individual relational spins and adding the outcomes.**

ANALOG TO THE SOCCER-BALL PROBLEM?

- Measurements are defined in terms of the projector Π_{α}^k , with $k = |\ell - j|, \dots, \ell + j$. There are therefore $2\ell + 1$ projectors.
- The system $\vec{\xi}$ is defined in terms of $2j + 1$ projectors.
- If $j < \ell$, then we have enough projectors to specify a priori the information encoded in the system.
- If $j > \ell$, then we do not have enough projectors to specify completely the information encoded in the system. ℓ is comparable to the Planck scale.
- But there is not a fundamental problem a priori: I should just take a bigger reference frame such that $\ell' > j$.

CONCLUSIONS

- **Quantum Reference Frames are essential in QG to construct some observables.**
- I used a toy model to explore the consequences of dealing with a QRF: interesting structures such as **non-commutativity, symmetry deformation, unusual coarse-graining** appeared.
- These structures also appeared in the context of effective description of semi-classical flat space-time, such as Deformed Special Relativity. **The analogy suggests therefore a new argument favoring DSR as the natural QG flat semi-classical limit.**

- Physics of a QRF can also be of interest for QIT community: precision, robustness of the QRF to encode information in relational quantities.
- **New models need to be explored.** The toy-model concentrated here on spin, ie $su(2)$. For the QG flat semi-classical limit, we need to deal with Poincaré symmetries. We should aim at constructing for example a quantum reference frame using quantum fields.

CONCLUSIONS

- **Quantum Reference Frames are essential in QG to construct some observables.**
- I used a toy model to explore the consequences of dealing with a QRF: interesting structures such as **non-commutativity, symmetry deformation, unusual coarse-graining** appeared.
- These structures also appeared in the context of effective description of semi-classical flat space-time, such as Deformed Special Relativity. **The analogy suggests therefore a new argument favoring DSR as the natural QG flat semi-classical limit.**