

Title: Intro to Supersymmetry 17

Date: Nov 13, 2007 12:30 PM

URL: <http://pirsa.org/07110013>

Abstract:

$$W = \frac{1}{2} \lambda L H^2$$

$$W = \frac{1}{2} \lambda L H^2$$

⇒ вакуум :

$$W = \frac{1}{2} \lambda L H^2$$

$\Rightarrow$   $\nu \propto \nu$  :  $H = 0$

$$W = \frac{1}{2} \lambda L H^2$$

$\Rightarrow$  вакуум :

$$H = 0$$

$$\langle H \rangle = 0$$

L

$$W = \frac{1}{2} \lambda \langle H \rangle^2$$

$\Rightarrow$  vacuum:  $H = 0$  ( $\langle H \rangle = 0$ )  
 $L$  is anything

$$W = \frac{1}{2} \lambda L H^2$$

$\Rightarrow$  vacuum:  $H = 0$  ( $\langle H \rangle = 0$ )

$L$  is anything

$\Rightarrow$  small fluctuations on the moduli space

$$L = \langle L \rangle + X$$

$$W = \frac{1}{2} \lambda L H^2$$

$$\Rightarrow \text{vacua: } H = 0 \quad (\langle H \rangle = 0)$$

$L$  is anything

$\Rightarrow$  small fluctuations on the moduli space

$$L = \langle L \rangle + x$$

$$H = \langle H \rangle + y$$



$\Rightarrow$   $X$  is a noiseless xsf

$$W(x, y)$$

surfs:  $A^L$

$$S = \frac{k}{2\pi} (S(A^L) - S(A^R))$$

$$dS = dA^L$$

$$dS =$$

$\Rightarrow X$  is a measureless xsp (West does not have a term  $\infty x^{21}$ )

$w(x, y)$

$A = \dots$

$S = \frac{1}{m} (S(A) - S(B))$

$k = \dots$

$k = \dots$

...  
 $a = e$   
 $b = 0$   
 $k = \dots$

$\Rightarrow X$  is a massless xsp (West does not have a term  $\propto X^2$ )

$w_{ij}(x,y)$

$\Rightarrow y$

$S = \frac{k}{2\pi} (\dots)$

$k = k'$

$k_3$

$A'$

$T$

$a, b$

fruit

$\Rightarrow X$  is a massless xsp (West does not have a term  $\propto X^2$ )

$w(x, y)$

$\Rightarrow y$  massive

$S = \frac{k}{2\pi} (\dots)$

$k = \dots$

$k_2$

frutt

$\Rightarrow X$  is a massless xsp (rest does not have a term  $\propto X^2$ )

$w_{ij}(x, y)$

$\Rightarrow y$  massive  $m_0^2 = |\lambda \langle L \rangle|^2$

$\Rightarrow X$  is a massless  $\chi$ SR (Wess does not have a term  $\propto X^2$ )

$$W_{10}(x, y)$$

$\Rightarrow y$  massive  $m_y^2 = |\lambda \langle L \rangle|^2$

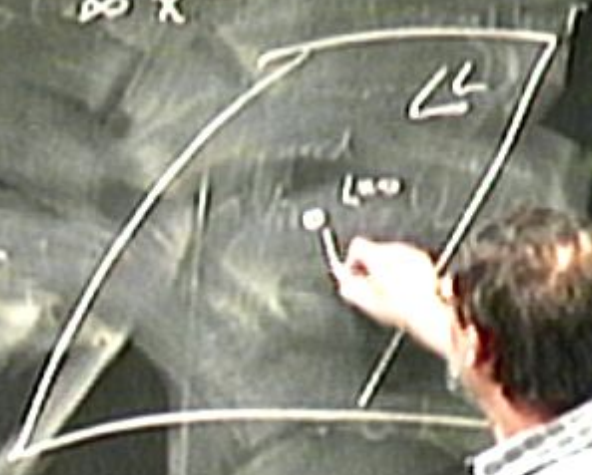
Interesting point is when  $\langle L \rangle = 0$

$\Rightarrow X$  is a massless xsp (Wess: does not have a term  $\propto X^2$ )

$w(x, y)$

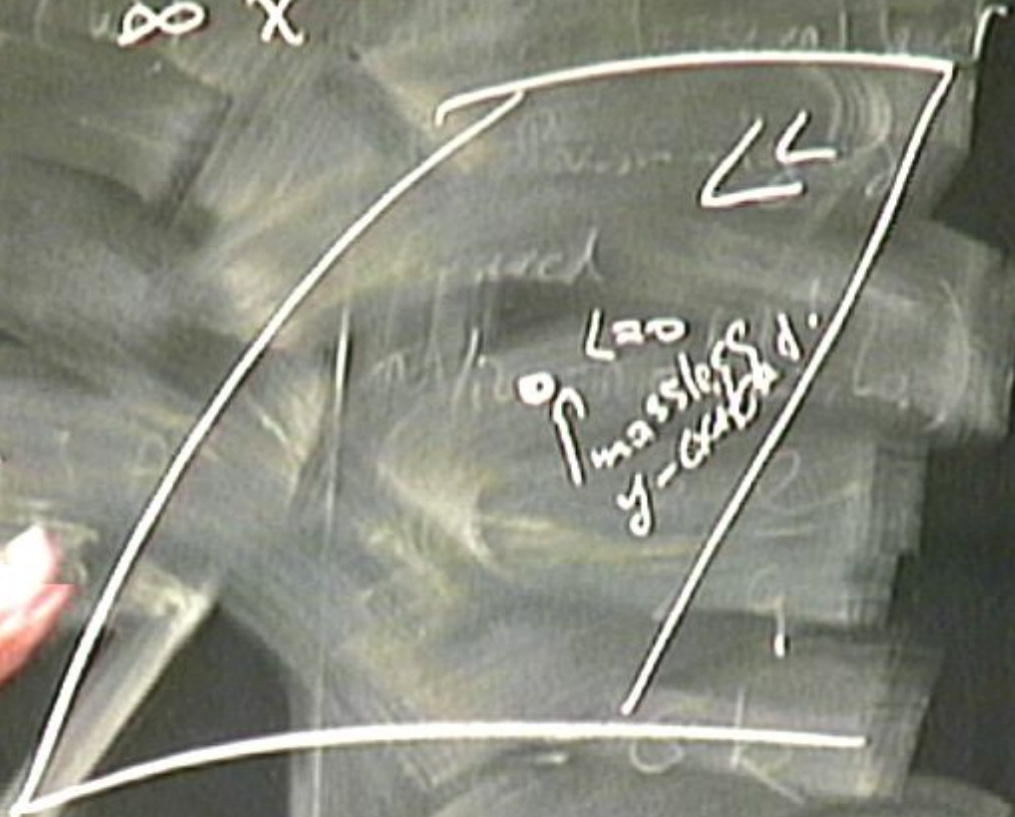
$\Rightarrow y$  massive  $m_0^2 = |\lambda \langle L \rangle|^2$

Interesting point is when  $\langle L \rangle = 0$



Wess does not have a term  
 $\infty \times \mathbb{R}^2$

$\rightarrow \langle L \rangle / 2$





$\Rightarrow X$  is a massless xsp (Wess does not have a term  $\propto X^2$ )

$$W(x, y)$$

$\Rightarrow y$  massive  $m_y^2 = \|\lambda \langle L \rangle\|^2$

Interesting point is when  $\langle L \rangle \rightarrow 0$

$\langle L \rangle$  massless y-coupling

⇒ Geometry of a moduli space.



⇒ Geometry of a moduli space.

Nakahara

⇒ Geometry of a moduli space.

Nakajima

Moduli space is  $\mathbb{C}^1$

⇒ Geometry of a moduli space.

Nakahara

Moduli space is  $\mathbb{C}^1$

⇒ Classically

$L, H$

$K$

⇒ Geometry of a moduli space.

Nakahara

Moduli space is  $\mathbb{C}^A$

→ Classically

$L, H$

⇒

Configuration space is  
2 complex d

K

⇒ Geometry of a moduli space.

Nakahara

Moduli space is  $\mathbb{C}^1$

→ Classically

$L, H$

⇒

Configuration space is  
2 complex dim.

$K =$

⇒ Geometry of a moduli space.

Nakahara

Moduli space is  $\mathbb{C}^1$

→ Classically

$L, H$

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Configuration space is  
2 complex dim.

$$K = L\bar{L} + H\bar{H}$$



⇒ Geometry of a moduli space.

Nakahara

Moduli space is  $\mathbb{C}^1$

⇒ Classically  $L, H$  ⇒

Configuration space is  
2 complex dim.

$$\textcircled{K} = L\bar{L} + H\bar{H}$$

$\{L, H\}$

⇒ If ~~any~~ complex manifold is Kähler

# Geometry of a moduli space.

Nakahara

Moduli space is  $\mathbb{C}^1$

Classically

$L, H$

$\Rightarrow$

Configuration space is  
2 complex dim.

$\{L, H\}$

$$\textcircled{K} = L\bar{L} + H\bar{H}$$

$\Rightarrow$  If ~~any~~ complex manifold is Kähler

$K(z_i, \bar{z}_i) \rightarrow$  a scalar function on a complex manifold

$$g_{ij} = \delta_{ij} K$$

$$\int g_{ij} = g_{ij} = 0$$

|||

$\mathbb{R}$

$$g_{ij} = \alpha_j \bar{\alpha}_i k$$

$$k = L\bar{L} + H\bar{H}$$

$$\int g_{ij} = g_{ij} = 0$$

$\mathbb{R}$

$$g_{ij} = \alpha_j \alpha_i K$$

$$\sqrt{g_{ij}} = g_{ij} = 0$$

$$K$$

$$K = LL^T + HH^T$$

$$g_{ij} = \begin{pmatrix} \alpha_i^2 & \alpha_i \alpha_j \\ \alpha_i \alpha_j & \alpha_j^2 \end{pmatrix}$$

$$\begin{matrix} \alpha_i \\ \alpha_j \end{matrix}$$

⇒ What is the metric on the model space?

$$(w(x, y)) \circlearrowleft \Delta$$

$$\rightarrow K = K$$

$$w_0 = | \langle \cdot, \cdot \rangle |^2$$

Induced by point is shown

$$e = dx$$

$$ds^2$$

$$g(x, y) = \dots$$

⇒ What is the metric on the moduli space?

$$(x, y) \in \mathbb{C}^n$$

$$\rightarrow K = K_{\text{moduli}}$$

moduli space.

→ What is the metric on the moduli space?

$$(x, y) \in \mathbb{C}^n$$

$$\rightarrow K = K_{\text{moduli}}$$

moduli space.

$$H=0, \bar{H}=0.$$



→ What is the metric on the moduli space?

$$\mathbb{C}^A$$

→

$$K_{\text{moduli}} = K$$

$$K$$

$$= \mathbb{Z} \bar{\mathbb{Z}}$$

moduli space

$$H=0, \bar{H}=0.$$

→ What is the metric on the moduli space?

$\mathbb{C}^n$

→  $K_{\text{moduli}} = K$

$= \mathbb{C}^n$

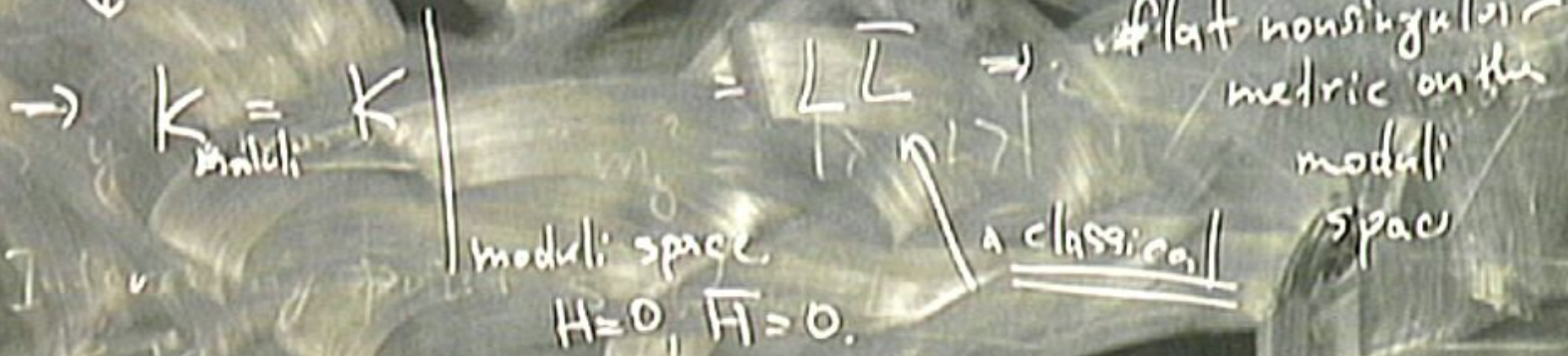
flat nonsingular metric on the moduli space

moduli space

$H=0, \bar{H}=0.$

→ What is the metric on the moduli space?

$$\mathbb{C}^n$$



$$g_{ij} = \eta_{ij} K$$

$$\sqrt{g_{ij} = g_{ij} - 0}$$

||

$$K = \begin{pmatrix} \bar{L}\bar{L} + \bar{H}\bar{H} & & \\ & 1 & \\ & & 0 & \dots & 0 \\ & & & & & 1 \end{pmatrix}$$

$$ds^2 = g_{ij} dx^i dx^j = \eta_{ij} K dx^i dx^j$$

$\Rightarrow$  Kahler potential will receive quantum corrections.

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\* ,  $y \Leftrightarrow |m_y|^2 = |\lambda L|^2$

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due to cubic interactions

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due to cubic interactions

$\Rightarrow$  there will be a wave-function renormalization



$\Rightarrow$  Kähler potential will receive quantum corrections

$$y \Leftrightarrow |m_y|^2 = |\lambda L|^2$$

due to cubic interactions

$\Rightarrow$  there will be a wave-function renormalization

$$\Rightarrow K = L\bar{L} \rightarrow K_{1-loop} =$$

$\Rightarrow$  Kahler potential will receive quantum corrections

$$* \quad y \Leftrightarrow |m_y|^2 = |\lambda L|^2$$

due to cubic interactions

$\Rightarrow$  there will be a wave-function renormalization

$$\Rightarrow K = L\bar{L} \rightarrow K_{1\text{-loop}} = Z_Y L\bar{L}$$

$$W = \frac{1}{2} \lambda \Delta H$$

$$K_{\Delta-loop} = \begin{pmatrix} 1 & + & 0 \\ 1 & & \uparrow \end{pmatrix}$$

small fluctuations

$$L = \langle L \rangle + \lambda$$

$$H = \langle H \rangle + \mu$$

$$\omega = \frac{1}{2} \lambda \langle H \rangle$$

$$K_{\Delta\text{-loop}} = \left( 1 + \frac{|\lambda|^2}{L} \right) \langle H \rangle$$

→ small fluctuations:  $\omega \approx \frac{1}{2} \lambda \langle H \rangle$

$$L = \langle L \rangle + \psi$$

$$H = \langle H \rangle + \psi$$

$\omega = \frac{1}{2} \lambda \langle H \rangle$

$$K_{\Delta\text{-loop}} = \left( 1 + |\lambda|^2 \ln \frac{\Lambda}{|m_H|} \right) \mathcal{L} \bar{\mathcal{L}}$$

$\uparrow$   
 $\mathcal{L}$

$$L = \langle L \rangle + X$$

$$H = \langle H \rangle + Y$$

$$K_{\Delta\text{-loop}} = \left( 1 + \underbrace{|\lambda|^2 \ln \frac{\Lambda}{|m_H|}}_{L} \right) \Delta \bar{L}$$

small fluctuations  
 $L = \langle L \rangle + \delta L$   
 $H = \langle H \rangle + \delta H$

$$K_{\Delta\text{-loop}} = \left( 1 + \underbrace{|\lambda|^2 \ln \frac{\Lambda}{|m_H|}}_{Z_y} \right) \Delta \bar{L}$$

$\omega = \frac{1}{2} \lambda \langle H \rangle$   
 $L = \langle L \rangle + \psi$   
 $H = \langle H \rangle + \phi$

$$K_{\Delta\text{-loop}} = \left( 1 + |\lambda|^2 \ln \frac{\Lambda}{|m_H|} \right) L\bar{L}$$

$Z_y$

$$\left( 1 + \lambda \bar{\lambda} \ln \frac{\Lambda^2}{L\bar{L}} \right) L\bar{L}$$



$$K_{\Delta\text{-loop}} = \left( 1 + |\lambda|^2 \ln \frac{\Lambda}{|m_H|} \right) L\bar{L}$$

$Z_y$   $|m_H| \propto |\Lambda|$

$$\approx \left( 1 + \lambda \bar{\lambda} \ln \frac{\Lambda^2}{L\bar{L}} \right) L\bar{L}$$

$$K_{\Delta\text{-loop}} = \left( 1 + |\lambda|^2 \ln \frac{\Lambda}{|m_H|} \right) \mathcal{L} \bar{\mathcal{L}}$$

↑

$Z_y$

$$\approx \left( 1 + \lambda \bar{\lambda} \ln \frac{\Lambda^2}{\mathcal{L} \bar{\mathcal{L}}} \right) \mathcal{L} \bar{\mathcal{L}}$$

$$\approx -|\lambda|^2 \left( \ln \mathcal{L} \bar{\mathcal{L}} \right) \mathcal{L} \bar{\mathcal{L}}$$

$$K_{\Delta\text{-loop}} = \left( 1 + |\lambda|^2 \ln \frac{\Lambda}{|m_H|} \right) LL$$

$Z_y$

$|m_H| \gg |L|$

$$\approx \left( 1 + \lambda \bar{\lambda} \ln \frac{\Lambda^2}{LL} \right) LL$$

$$\approx -|\lambda|^2 \left( \ln LL \right) LL$$

(dominant contribution for small  $|L|$ )

14

$$g_{LI} = \frac{2k_{s-100p}}{2L2L}$$

Involvement

ds

$$g_{\vec{k}} = \frac{\partial^2 \mathcal{H}^{\text{loop}}}{\partial L \partial \vec{k}} = -|\lambda|^2 \ln \frac{L}{\Lambda^2} + \text{const}$$

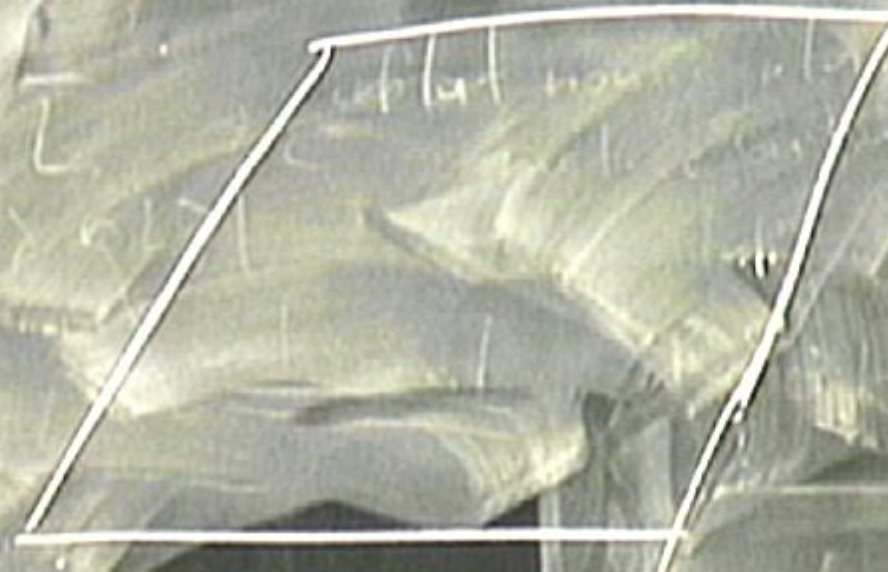
$$g_{LL} = \frac{\partial^2 K^{s-100p}}{\partial L \partial L} = -|\lambda|^2 \ln \frac{LL}{\lambda^2} + \text{const}$$

Taylor

$$e = \frac{1}{L}$$

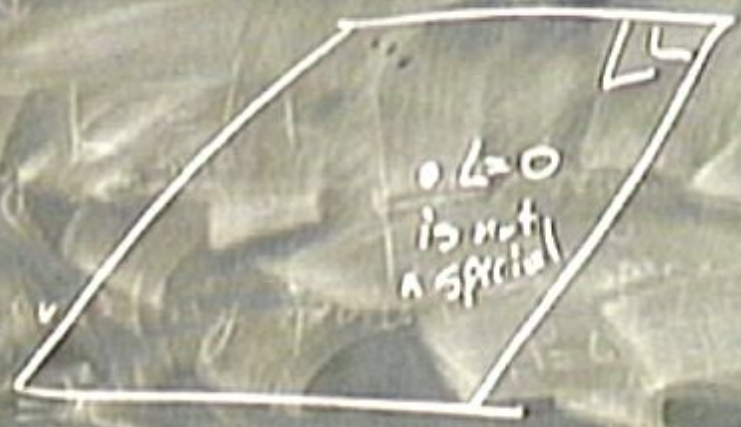
ds

$$g_{LL} = \frac{\partial^2 H_{1-loop}}{\partial L \partial \bar{L}} = -|\lambda|^2 \ln \frac{L\bar{L}}{\Lambda^2} + \text{const}$$

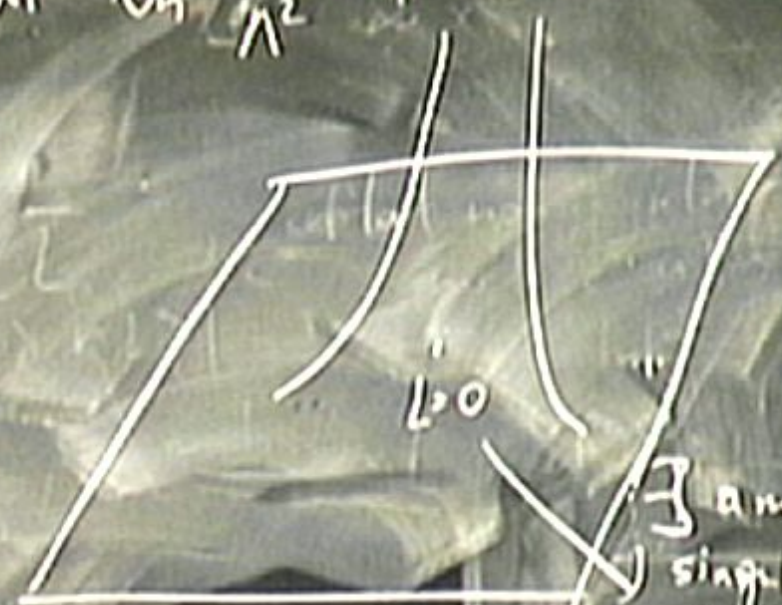


↑  
classical

$$g_{LI} = \frac{\partial^2 K}{\partial L \partial L} = -|\lambda|^2 \ln \frac{4L}{\Lambda^2} + \text{const}$$



↑  
Classical



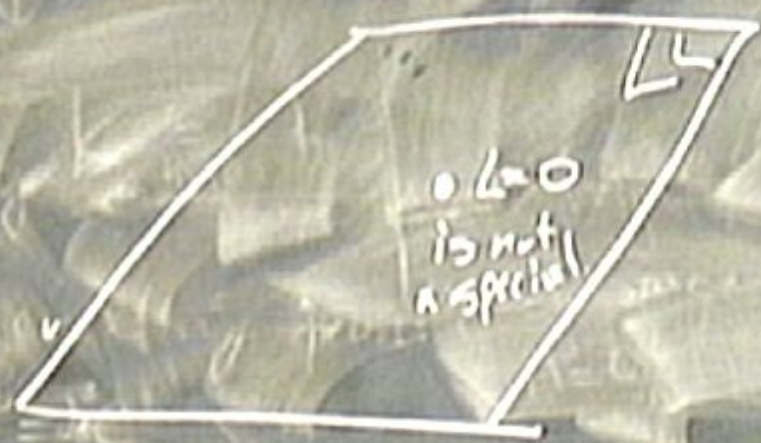
QM

∃ a metric singularity in the bulk;

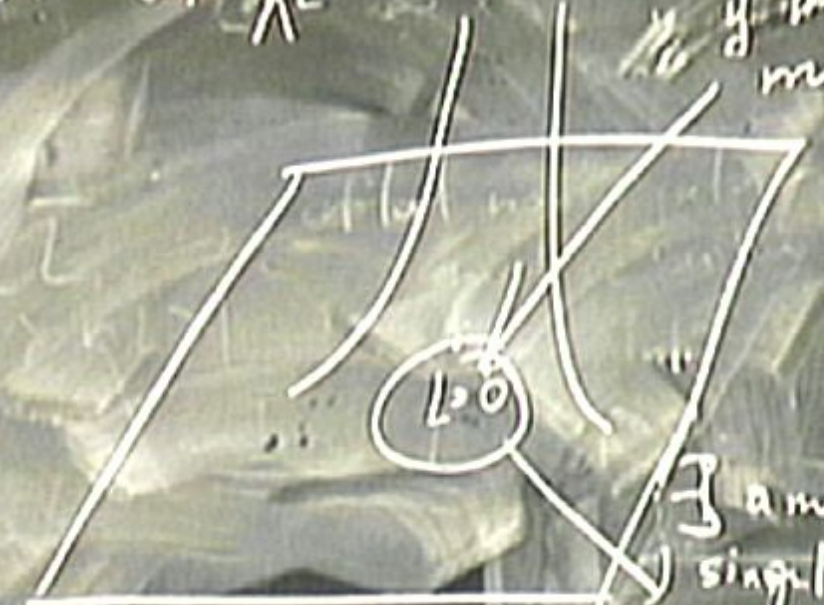


$$g_{LE} = \frac{\delta^2 H^{L=loop}}{\delta L \delta \bar{L}} = -|\lambda|^2 \ln \frac{L\bar{L}}{\Lambda^2} + \text{const}$$

if the comp  
massless



↑  
classical

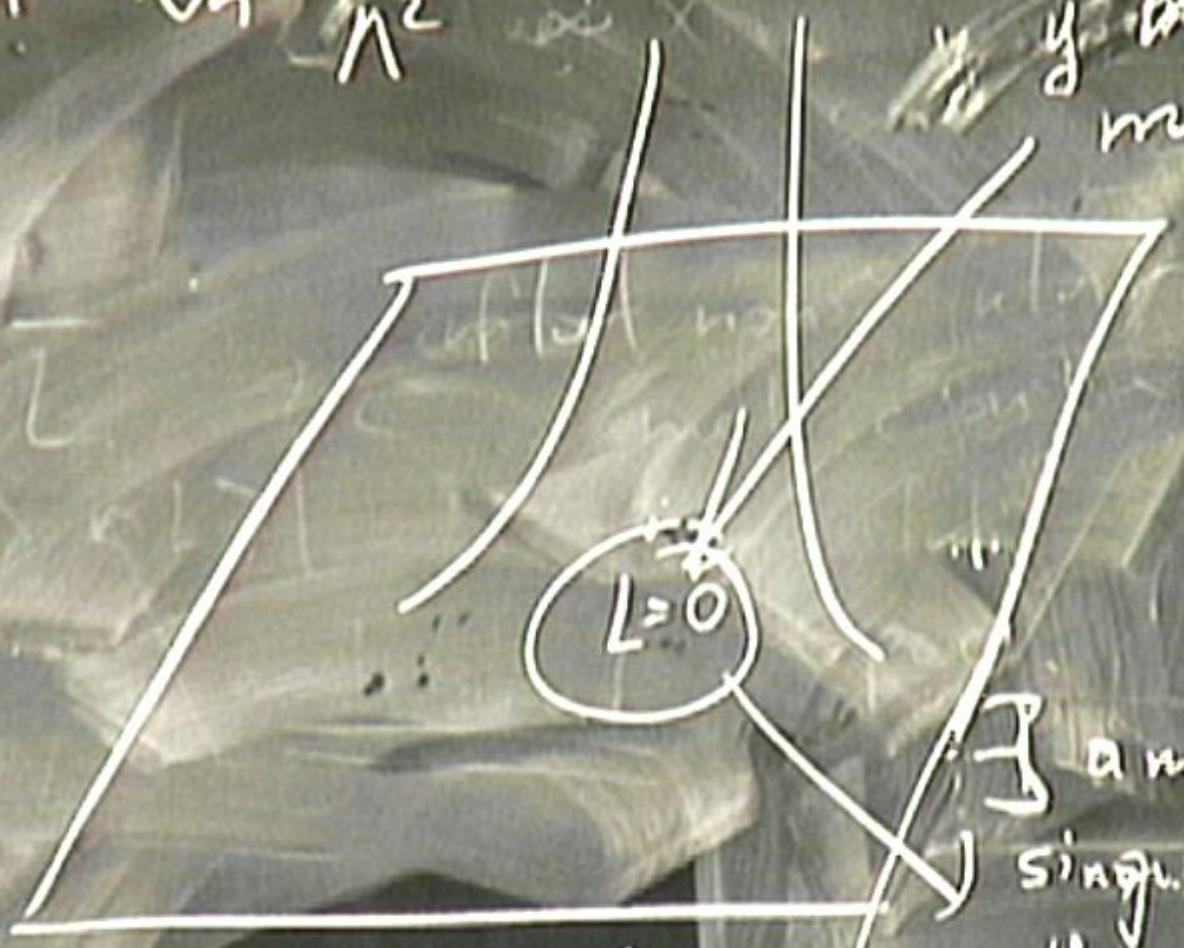


QM

∃ a metric  
singularity  
on the world;

$$-|\lambda|^2 \ln \frac{L\bar{L}}{\Lambda^2} + \text{const}$$

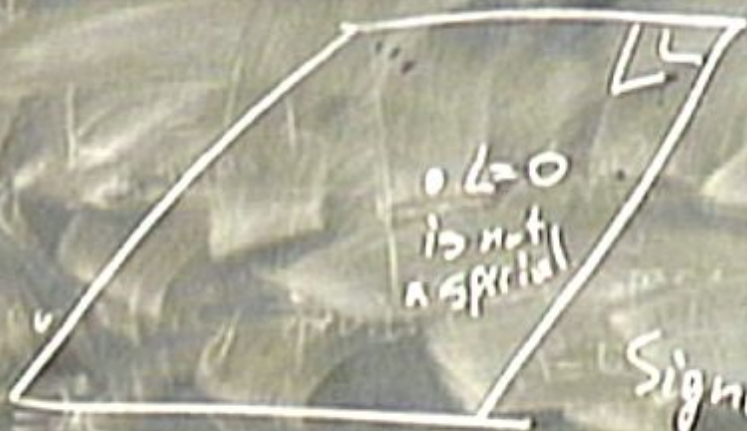
if become massless



∃ a metric singularity in the neck;

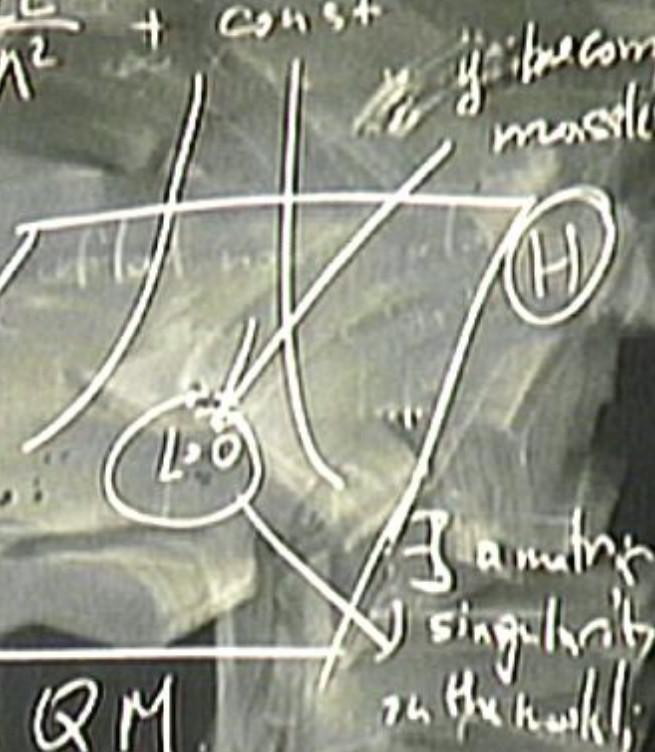
QM.

$$g_{LE} = \frac{\partial^2 K^{s-loop}}{\partial L \partial \bar{L}} = -|\lambda|^2 \ln \frac{L\bar{L}}{\Lambda^2} + \text{const}$$

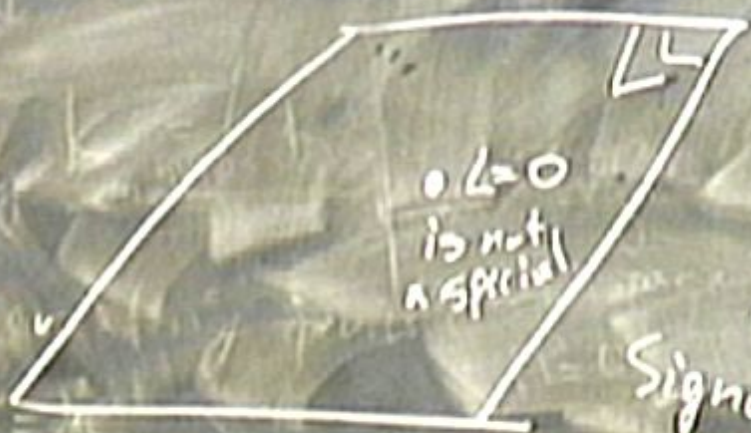


↑  
classical

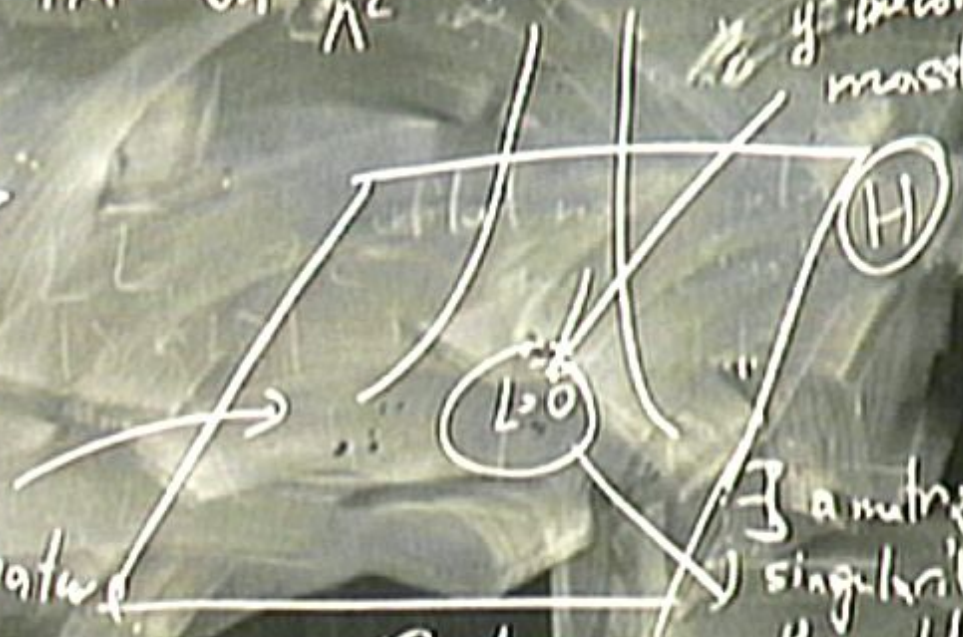
Signature  
that our effective QM  
QFT breaks down



$$g_{LE} = \frac{\partial^2 K^{L=loop}}{\partial L \partial L} = -|\lambda|^2 \ln \frac{L}{\Lambda^2} + \text{const}$$



↑  
classical



if become massless

Signature  
that our effective QM  
QFT breaks down

∫ a metric singularity in the world;

'Integration out'

$W_{mic}$



'Integration out'

$$W_{mic} = \frac{1}{2} m H^2 + \lambda L^2 H$$

'Integration out'

$$W_{\text{mic}} = \frac{1}{2} m H^2 + \lambda L^2 H$$

H is massive.

L is massless.

'Integration out'

$$W_{\text{mic}} = \frac{1}{2} m H^2 + \lambda L^2 H$$

H is massive.

L is massless.

$\Rightarrow$  non renorm

$$W_{\text{eff}} = \frac{1}{2} m H^2 + \lambda L^2 H$$

$\uparrow$

non the minimum  
effective action



'Integration out'

$$W_{\text{mic}} = \frac{1}{2} m H^2 + \lambda L^2 H$$

H is massive.

L is massless.

$\Rightarrow$  non renorm

$$W_{\text{eff}} = \frac{1}{2} m H^2 + \lambda L^2 H$$

$\uparrow$

non the minimum  
effective action

$$E \ll |m|$$

'Integration out'

$$W_{\text{mic}} = \frac{1}{2} m H^2 + \lambda L^2 H \quad \Rightarrow \quad W_{\text{eff}} = \frac{1}{2} m H^2 + \lambda L^2 H$$

H is massive.

L is massless.

non renorm

↑  
non the minimum  
effective action

→ effective action should not contain  $E \ll |m|$

$$\Rightarrow F \ll |h|$$

$\Rightarrow$  we can integrate out  $H$  - field



Solve EOM for  $H$

$\Rightarrow F \ll |m|$

$\Rightarrow$  we can integrate out  $H$  - field



Solve EOM for  $H$  and substitute it  
into the superpotential

$\Rightarrow F \ll (m)$

$\Rightarrow$  we can integrate out  $H$  - field



Solve EOM for  $H$  and substitute it  
into the superpotential

$m \rightarrow$

$\Rightarrow F \ll |m|$

$\Rightarrow$  we can integrate out  $H$  - field



Solve EOM for  $H$  and substitute it  
into the superpotential

$m \rightarrow \infty$

$$\Rightarrow F \ll |m|$$

$\Rightarrow$  we can integrate out  $H$  - field



Solve EOM for  $H$  and substitute it  
into the Superpotential

$$m \rightarrow \infty$$

$$\left( F \ll |m| \right)$$

$$\Rightarrow F \ll |m|$$

$\Rightarrow$  we can integrate out  $H$  - field



Solve EOM for  $H$  and substitute it into the superpotential

$$|m| \rightarrow \infty$$

$$\left( F \ll |m| \right)$$



of rotation out"

$$e = \frac{1}{2} m H^2 + \lambda L^2 H$$

non renorm  $\Rightarrow$

W eff

H is massive.

L is massless.

on H

tive



Handwritten mathematical notes on a chalkboard, including the equation  $\frac{M}{H} = 0$  and  $27R + H = 2$ . The board is heavily obscured by large, dark, diagonal brush strokes.



$2\gamma + H_3 + \lambda L^2$   
 $O = \frac{H}{M/C} = \frac{H}{M/C}$   
 $H = H$

$$\frac{H}{M} = \frac{M}{H} \Rightarrow 0$$

$$mH + \lambda L^2$$

$$\langle H \rangle = \frac{\lambda L^2}{3}$$

$$W_{eff}(L) =$$

$$\frac{H}{Mc} = 0$$

$$mH + \lambda L^2$$

$$\langle H \rangle = - \frac{\lambda^2}{3}$$

$$W_{eff}(L) = - \frac{\lambda^2 L^4}{2m}$$

$$\begin{aligned}
 & \frac{mH}{\lambda^2} + H \\
 & \langle H \rangle = \frac{\lambda^2}{2m} * L \quad \text{it's a massless} \\
 & \underbrace{W_{eff}(L)} = \frac{\lambda^2}{2m} * L \quad \text{it's a massless} \\
 & \quad \quad \quad * E \ll m
 \end{aligned}$$

$$p(L) = \frac{\hbar^2 k^2}{2m} \quad * L \text{ is a massless}$$

$$* E \ll m$$

$\Rightarrow$  decide on Light

$$P(L) = -\frac{\hbar^2 L^2}{2m} \quad * L \text{ is a massless}$$

$$* E \ll m$$

- $\Rightarrow$  decide on Light
- $\Rightarrow$  look at symmetries



$$P(L) = -\frac{\lambda^2 L^4}{2m} \quad * L \text{ is a massless}$$

$$* E \ll m$$

$\Rightarrow$  decide on Light

$\Rightarrow$  look at symmetries

$\rightarrow$  most general

$$L) = -\frac{\lambda^2 L^2}{2m} \quad * \quad L \text{ is a massless}$$

$$* \quad E \ll m$$

- $\Rightarrow$  decide on Light
- $\Rightarrow$  look at symmetries
- $\rightarrow$  most general W (holomorphic in complex)

$$L) = -\frac{\lambda^2 L^4}{2m} \quad * L \text{ is a massless}$$

$$* E \ll m$$

- $\Rightarrow$  decide on Light
- $\Rightarrow$  look at symmetries

$\rightarrow$  most general  $W$  (holomorphic in  $cr$  plane)

→  $W_{eff} = W_{eff}(4)$

$\Rightarrow W_{\text{eff}} = W_{\text{eff}}(L)$

↑  
the origin in  $L, m, \lambda$  is not a special

$$\Rightarrow W_{\text{eff}} = W_{\text{eff}}(L) = \sum_{n_1, n_2, n_3}$$

↑  
 the origin in  
 $L, m, \lambda$  is  
 not a special

$$\Rightarrow W_{\text{eff}} = W_{\text{eff}}(L) = \sum_{n_1, n_2, n_3} \binom{n_1}{m} \binom{n_2}{\lambda} n_3$$

↑  
 the origin in  
 $L, m, \lambda$  is  
 not a special

$$\Rightarrow W_{\text{eff}} = W_{\text{eff}}(L) = \sum_{n_1, n_2, n_3} \langle n_1, n_2, n_3 |$$

↑  
the origin in  
 $L, m, \lambda$  is

$$W_{\text{min}} = \frac{1}{2} m H^2 + \lambda L H$$

not a special



$$\Rightarrow W_{\text{eff}} = W_{\text{eff}}(L) = \sum_{n_1, n_2, n_3} L^{n_1} m^{n_2} \lambda^{n_3}$$

↑  
the origin in  
 $L, m, \lambda$  is

$$W_{\text{micr}} = \frac{1}{2} m H^2 + \lambda L H$$

not a special

$$\Rightarrow W_{\text{eff}} = W_{\text{eff}}(L) = \sum_{n_1, n_2, n_3} \langle n_1, m, n_2, n_3 | C_{n_1 n_2 n_3}$$

↑  
the origin in  
 $L, m, \lambda$  is

$$W_{\text{micr}} = \frac{1}{2} m H^2 + \frac{H^2}{\lambda L} \quad \text{not a special}$$

$U(1)_H$

$U(1)_L$

$U(1)_X$

Q.9

CTT

L

0

+1

H

+1

0

X

L

$U(1)_H$  $U(1)_L$  $U(1)_R$ 

L

0

+1

H

+1

0

S

-2

+  
R

$$H(W_{\text{micro}}) = 0$$

$$L(W_{\text{micro}}) = 0$$

$$R(W_{\text{micro}}) = 2$$

	$U(1)_H$	$U(1)_L$	$U(1)_R$
L	0	+1	+1
H	+1	0	
$\bar{3}$	-2	0	
$\bar{2}$	-1	-2	

$$H(W_{\text{micro}}) = 0 \quad L(W_{\text{micro}}) = 0 \quad R(W_{\text{micro}}) = 2$$

	$U(1)_H$	$U(1)_L$	$U(1)_R$
L	0	+1	+1
H	+1	0	0
$\bar{3}$	-2	0	+2
$\bar{2}$	-1	-2	0

$$H(W_{micro}) = 0 \quad L(W_{micro}) = 0 \quad R(W_{micro}) = 2$$

⇒ construct momenta

$$V(l)_H: -2n_2 - n_3 = 0$$

$$V(l)_L: n_1 - 2n_3 = 0$$

⇒ construct molecules.

$$U(1)_H: -2n_2 - n_3 = 0$$

$$U(1)_L: n_1 - 2n_3 = 0$$

$$U(1)_R: 2n_2 + n_1 = 2$$



⇒ construct monomials.

$$U(l)_H: -2n_2 - n_3 = 0$$

$$W_{\text{eff}} \propto \frac{L^4}{\pi^2} \lambda^2$$

$$U(l)_L: n_1 - 2n_3 = 0$$

$$U(l)_R: 2n_2 + n_1 = 2$$

In superpotential, integrate out massive field



In superpotential, integrate out massive field  
classically



In superpotential integrate out massive field  
classically and the result is correct  
QM

In superpotential integrate out massive field  
classically and the result is correct  
QM

In superpotential integrate out massive field  
classically and the result is correct  
QM

$$K = \overline{L}L -$$

In superpotential integrate out massive field  
classically and the result is correct  
QM

$$K = \overline{L L} + H \overline{H} \rightarrow$$

In superpotential integrate out massive field  
 classically and the result is correct  
 QM.

$$K = \bar{L} \bar{L} + H \bar{H} \quad \rightarrow \quad K_{\text{eff}} = \bar{L} \bar{L} + \frac{|\lambda|^2}{|m|^2} (\bar{L} \bar{L})^2$$



In superpotential integrate out massive field  
classically and the result is correct

QM

$$H = -\frac{\Delta^2}{m}$$

$$K = \bar{L}L + \bar{H}H$$



$$K_{\text{eff}} = \bar{L}L + \frac{|\Delta|^2}{m} (\bar{L}L)^2$$

not exact

In superpotential integrate out massive field  
 classically and the result is correct

classically  
 QM

$$H = -\frac{\Delta^2}{m}$$

$$K = \bar{L}L + H\bar{H}$$

→

$$K_{\text{eff}} = \bar{L}L + \frac{|\Delta|^2}{m} (\bar{L}L)^2$$

not exact

class  
 $\bar{L}L =$

$$= \bar{L}L$$

In superpotential integrate out massive field  
 classically and the result is correct

QM

$$H = -\frac{\Delta^2}{m}$$

$$K = L\bar{L} + H\bar{H}$$

→

$$K_{\text{eff}} = L\bar{L} + \frac{|\Delta|^2}{m} (L\bar{L})^2$$

not exact

class

$$Z = 1 + \frac{|\Delta|^2}{m} L\bar{L} = Z_L(L\bar{L})$$

संज्ञा

we ...

→

and substitute it

संरचना

$$L = -\frac{1}{4g^2} F_{\mu\nu} F^{\mu\nu}$$

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

and substitute it

int  $\rightarrow$

संरचना

$$\mathcal{L} = -\frac{1}{4g^2} F_{\mu\nu} F^{\mu\nu} \quad F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

$$A_\mu \rightarrow A_\mu + \partial_\mu \alpha$$

$\alpha \rightarrow$

⇒ vector superfield

most general superfield

$$V = V^\dagger + 0$$

⇒

$$V = C(x) + i\theta\chi(x) - i\bar{\theta}\bar{\chi}(x)$$

⇒ vector superfield

most general superfield

$$V(x, \theta, \bar{\theta}) = V + \dots + \theta^2 \dots + \bar{\theta}^2 \dots + \dots$$

$V = V^\dagger$

⇒

$$V = C(x) + i\theta\chi(x) - i\bar{\theta}\bar{\chi}(x) + \dots$$

↑  
real scalar



⇒ vector superfield

most general superfield

$$V(x, \theta, \bar{\theta}) = V + \dots = 0$$

⇒

$$V = C(x) + i\theta\chi(x) - i\bar{\theta}\bar{\chi}(x) + \frac{i}{2}\theta^2[M+iN] - \frac{i}{2}\bar{\theta}^2[M-iN]$$

↑  
real scalar

⇒ vector superfield

most general superfield

$V(x)$

$$V = V + \dots$$

⇒

$$V = C(x) + i\theta\chi(x) - i\bar{\theta}\bar{\chi}(x) + \frac{i}{2}\theta^2[M + iN] - \frac{i}{2}\bar{\theta}^2[M - iN]$$

↑  
real scalar

$$- \theta\sigma^{\mu\nu}\bar{\theta}V_{\mu\nu} + i\theta^2\bar{\theta}\left[\lambda + \frac{i}{2}\sigma^{\mu\nu}\partial_\mu\chi\right]$$

⇒ vector superfield

most general superfield

$$V(x, \theta, \bar{\theta}) = V + \dots = 0$$

$$V = V^\dagger$$

⇒

$$V = C(x) + i\theta\chi(x) - i\bar{\theta}\bar{\chi}(x) + \frac{i}{2}\theta^2[M + iN] - \frac{i}{2}\bar{\theta}^2[M - iN]$$

↑  
real scalar

$$- \theta\sigma^\mu\bar{\theta}V_\mu + i\theta^2\bar{\theta}\left[\lambda + \frac{i}{2}\sigma^\mu\partial_\mu\chi\right]$$

$$- i\bar{\theta}^2\theta\left[\bar{\lambda} + \frac{i}{2}\sigma^\mu\partial_\mu\bar{\chi}\right]$$

- look at symmetries
- most general W (holomorphic in  $\theta$ )

⇒ vector superfield

most general superfield

$$V = V + \dots$$

⇒

$$V = C(x) + i\theta\chi(x) - i\bar{\theta}\bar{\chi}(x) + \frac{i}{2}\theta^2[M + iN] - \frac{i}{2}\bar{\theta}^2[M - iN]$$

↑  
real scalar

$$- \theta\sigma^{\mu\nu}\bar{\theta}V_{\mu\nu} + i\theta^2\bar{\theta}\left[\lambda + \frac{i}{2}\sigma^{\mu\nu}\partial_{\mu}\chi\right] - i\bar{\theta}^2\theta\left[\lambda + \frac{i}{2}\sigma^{\mu\nu}\partial_{\mu}\bar{\chi}\right] + \frac{i}{2}\theta^2\bar{\theta}^2\left[\dots\right]$$

→ look at symmetries

→ most general W (holomorphic in superfield)

$C, M, N, O_m, D$

Well  $\infty$   $L^4$   $T^2$

$$+\frac{i}{2}\theta^2[M+iN] - \frac{i}{2}\theta^2[M-iN]$$

$C, M, N, O_m, D$

spin - 1/2

Weyl  $\infty$

$$+ \frac{i}{2} \theta^2 [M + iN] - \frac{i}{2} \bar{\theta}^2 [M - iN]$$

$C, M, N, \mathcal{O}_m, D, \lambda, \chi, \bar{\lambda}, \bar{\psi}$

spin - 1/2

weff  $\infty$

$$+ \frac{i}{2} \theta^2 [M + iN] - \frac{i}{2} \bar{\theta}^2 [M - iN]$$

$C, M, N, O_m, D, \lambda, \chi, \bar{\psi}, \bar{\psi}$

spin - 1/2

Dirac field  
+ 0

$$\bar{\psi}(x) + \frac{i}{2} \theta^2 [M + iN] - \frac{i}{2} \bar{\theta}^2 [M - iN]$$

$$\bar{\theta} \theta + i \theta^2 \bar{\theta} [\bar{\psi} + i \psi] \chi$$