

Title: Early Time Dynamics in Heavy Ion Collisions and AdS/CFT Correspondence

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Abstract:

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Yuri Kovchegov
The Ohio State University

based on work done with Anastasios
Taliotis, arXiv:0705.1234 [hep-ph]

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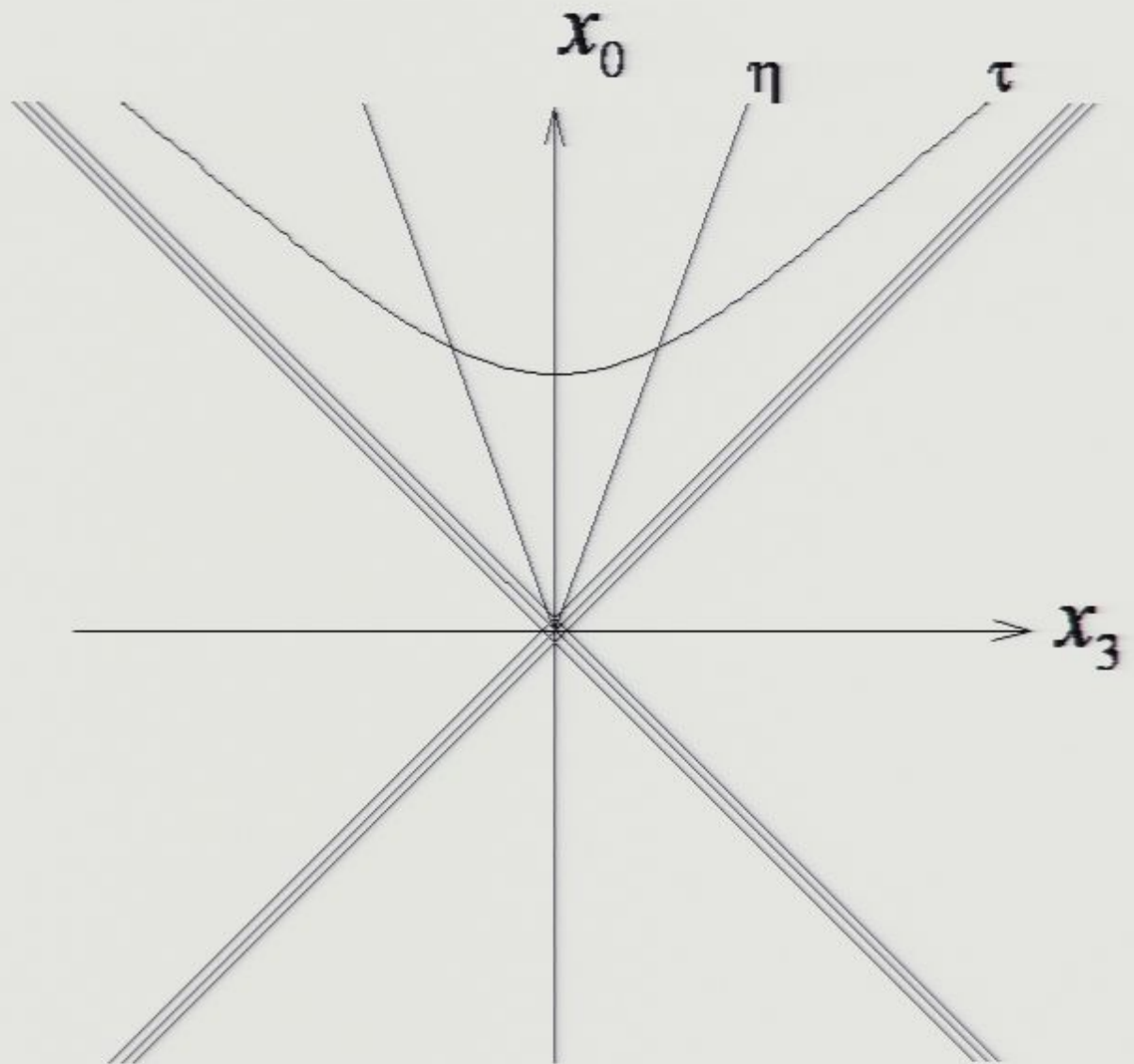
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 - Analyzed **early-time** dynamics and showed that energy density goes to a constant at early times.
 - Have therefore shown that isotropization (and hopefully thermalization) takes place in strong coupling dynamics.
 - Derived a simple formula for isotropization time and used it for heavy ion collisions at RHIC to obtain 0.3 fm/c, in agreement with hydrodynamic simulations.

Notations

We'll be using the following notations:

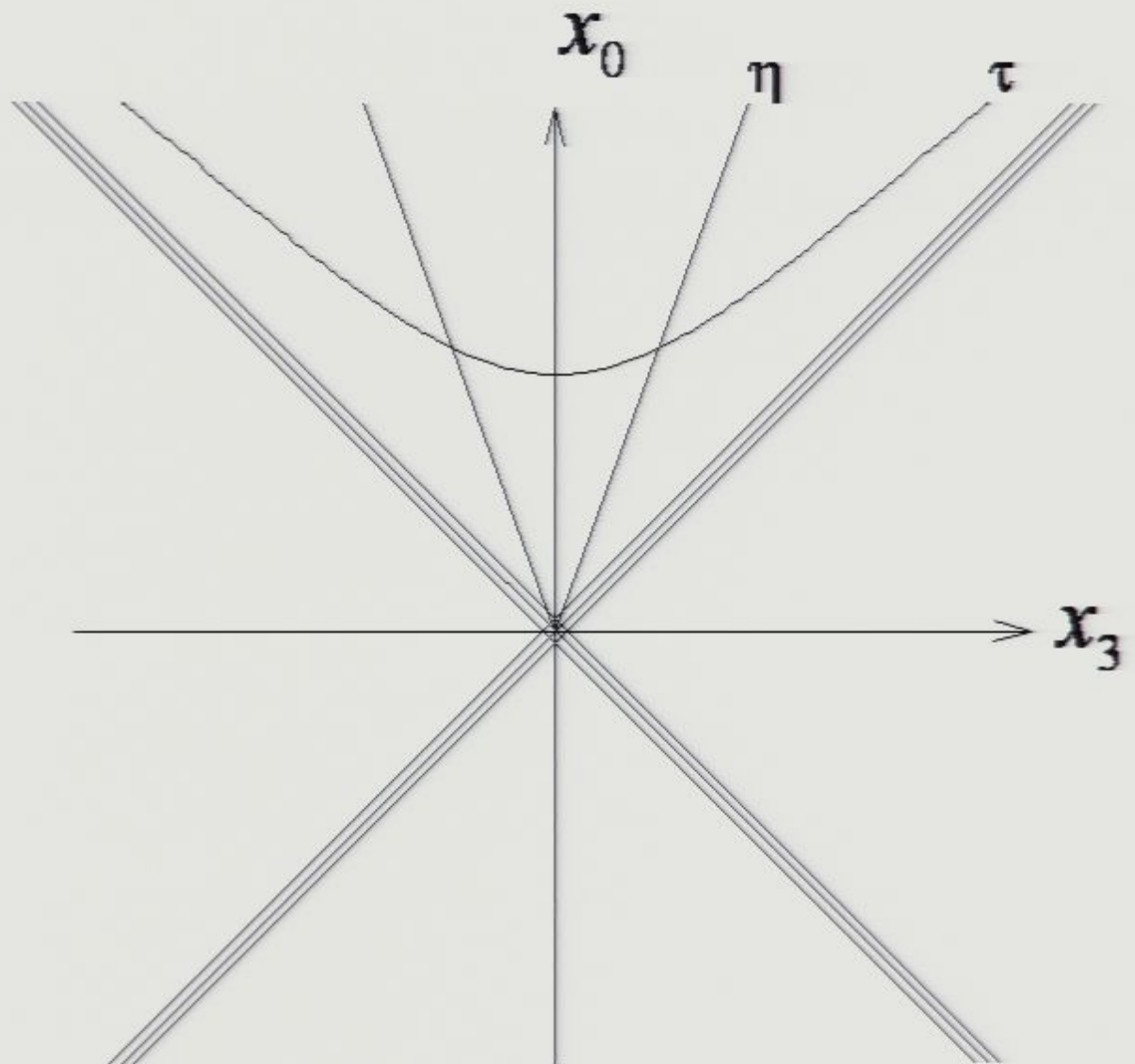


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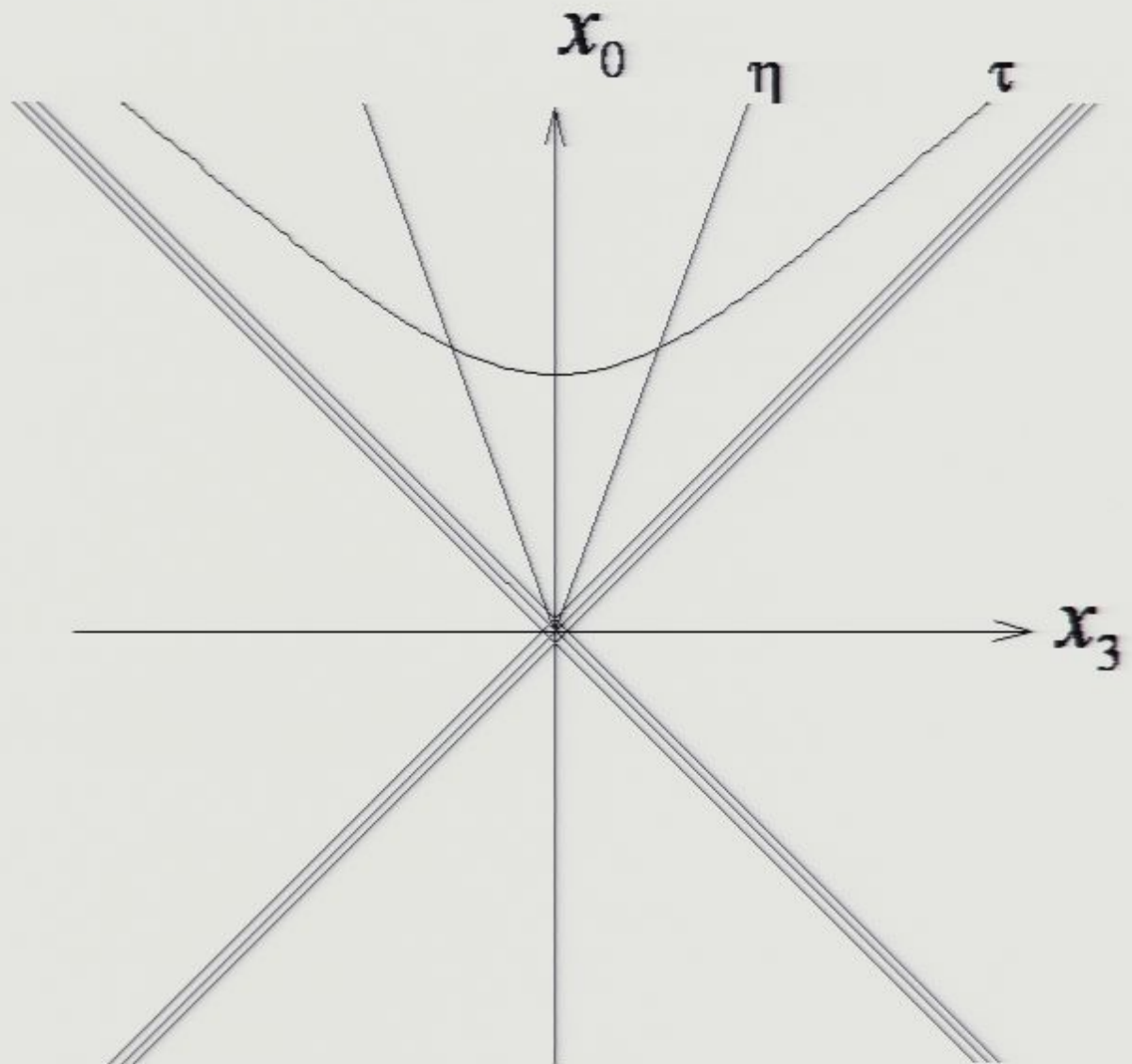
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proper time

$$\tau = \sqrt{x_0^2 - x_3^2}$$

and rapidity

$$\eta = \frac{1}{2} \ln \frac{x_0 + x_3}{x_0 - x_3}$$



Most General Boost Invariant Energy-Momentum Tensor

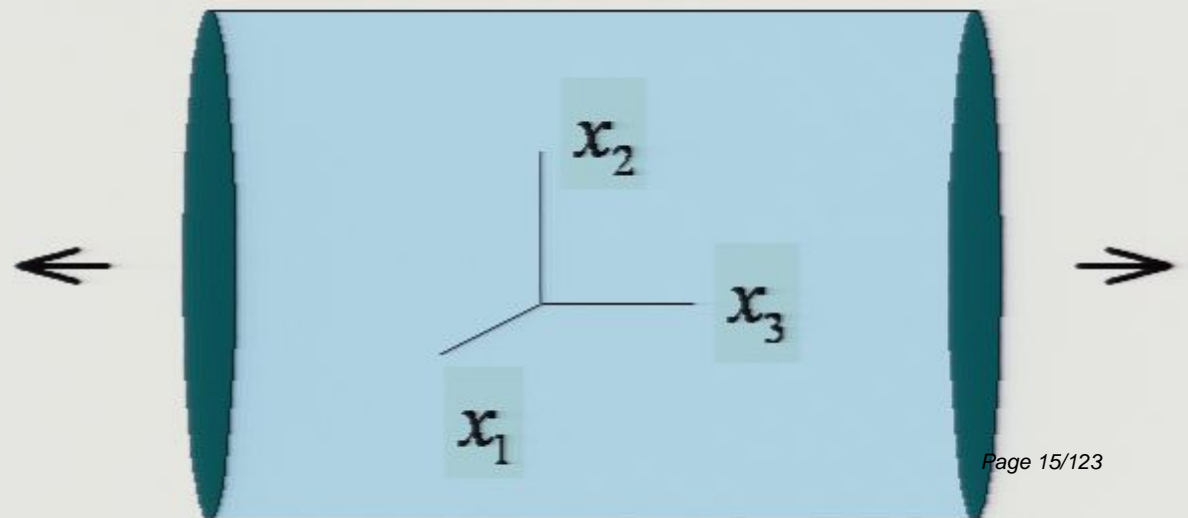
The most general boost-invariant energy-momentum tensor for a high energy collision of two very large nuclei is (at $x_3 = 0$)

$$T^{\mu\nu} = \begin{pmatrix} \varepsilon(\tau) & 0 & 0 & 0 \\ 0 & p(\tau) & 0 & 0 \\ 0 & 0 & p(\tau) & 0 \\ 0 & 0 & 0 & p_3(\tau) \end{pmatrix} \begin{matrix} x_0 \\ x_1 \\ x_2 \\ x_3 \end{matrix}$$

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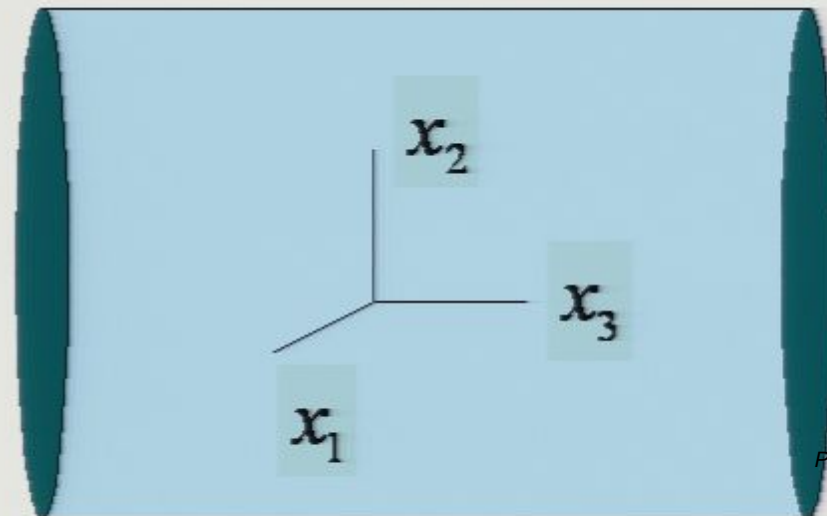
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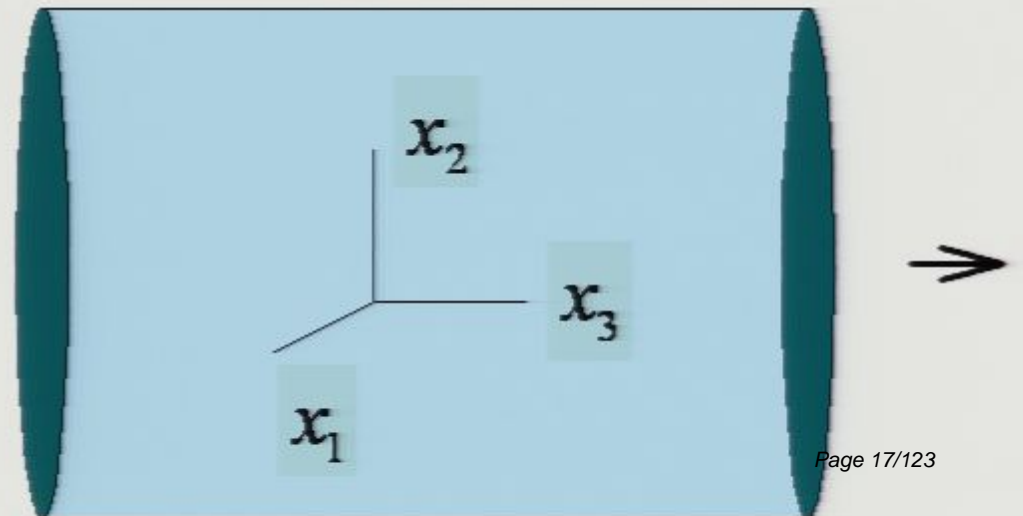
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There are 3 extreme limits.

Limit I: “Free Streaming”

Free streaming is characterized by the following “2d” energy-momentum tensor:

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$$\frac{d\varepsilon}{d\tau} = -\frac{\varepsilon}{\tau}$$

and

$$\varepsilon \sim \frac{1}{\tau}$$

- The total energy $E \sim \varepsilon \tau$ is conserved, as expected for non-interacting particles.

Limit II: Bjorken Hydrodynamics

In the case of ideal hydrodynamics, the energy-momentum tensor is symmetric in all three spatial directions (**isotropization**):

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Using the ideal gas equation of state, $\varepsilon = 3p$, yields

$$\varepsilon \sim \frac{1}{\tau^{4/3}} \quad \text{Bjorken, '83}$$

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➤ The total energy $E \sim \varepsilon \tau$ is **not** conserved, while the total entropy S is conserved.

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If $p_3 > 0$ then, as $\frac{d\varepsilon}{d\tau} = -\frac{\varepsilon + p_3}{\tau}$, one gets $\varepsilon \sim \frac{1}{\tau^{1+\Delta}}$.

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Deviations from the $\varepsilon \sim \frac{1}{\tau}$ scaling of energy density,

like $\varepsilon \sim \frac{1}{\tau^{1+\Delta}}$, $\Delta > 0$ are due to longitudinal pressure

p_3 , which does work $p_3 dV$ in the longitudinal direction

modifying the energy density scaling with tau.

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Limit III: Color Glass at Early Times

In CGC at very early times $\varepsilon \sim \log^2 \frac{1}{\tau}$, $\tau Q_s \ll 1$ (Lappi, '06)

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AdS/CFT Approach

Start with the metric in Fefferman-Graham coordinates in AdS_5 space

$$ds^2 = \frac{1}{z^2} \left[-A(\tau, z) d\tau^2 + \tau^2 B(\tau, z) d\eta^2 + C(\tau, z) dx_{\perp}^2 + dz^2 \right]$$

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$$\tilde{g}_{\mu\nu}(x, z) = \tilde{g}_{\mu\nu}^{(0)}(x) + z^2 \tilde{g}_{\mu\nu}^{(2)}(x) + z^4 \tilde{g}_{\mu\nu}^{(4)}(x) + \dots$$

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$$\langle T_{\mu\nu} \rangle = \frac{N_c^2}{2\pi^2} \tilde{g}_{\mu\nu}^{(4)}(x)$$

Iterative Solution

General solution of Einstein equations is not known and is hard to obtain. One first assumes a specific form for energy density

$$\mathcal{E} = \mathcal{E}(\tau)$$

and then solves Einstein equations perturbatively order-by-order in z :

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The solution in AdS space (if found) determines which function of proper time is allowed for energy density.

At the order z^4 it gives the following familiar conditions:

$$\partial_\mu \langle T^{\mu\nu} \rangle = 0 \quad \text{and} \quad \langle T^\mu_\mu \rangle = 0$$

Iterative Solution

We begin by expanding the coefficients of the metric

$$ds^2 = \frac{1}{z^2} [-A(\tau, z) d\tau^2 + \tau^2 B(\tau, z) d\eta^2 + C(\tau, z) dx_\perp^2 + dz^2]$$

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into power series in z :

$$a(\tau, z) = \sum_{n=0}^{\infty} a_n(\tau) z^{4+2n}, \quad b(\tau, z) = \sum_{n=0}^{\infty} b_n(\tau) z^{4+2n}, \quad c(\tau, z) = \sum_{n=0}^{\infty} c_n(\tau) z^{4+2n}.$$

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To illustrate their structure let me display one of them:

$$a_2(\tau) = -\frac{1}{384} \left[a_0 \tau^{\Delta-4} (4 \Delta^2 - \Delta^4) + a_0^2 \tau^{2\Delta} 8 (8 + 8 \Delta + 3 \Delta^2) \right]$$

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(only if $\Delta > -4$!)

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$$-4 \leq \Delta \leq 0.$$

The above conclusion about which term dominates at what time is safe!

Late Time Solution: Scaling

At late times the perturbative (in z) series becomes

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The metric coefficients become:

$$a(\tau, z) = a(v), \quad b(\tau, z) = b(v), \quad \text{and} \quad c(\tau, z) = c(v)$$

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$$a(\tau, z) = a(v), \quad b(\tau, z) = b(v), \quad \text{and} \quad c(\tau, z) = c(v)$$

Here $a_0 < 0$ is the normalization of the energy density

$$\epsilon(\tau) = -\frac{N_c^2}{2\pi^2} a_0 \tau^\Delta$$

Janik and Peschanski's Late Time Solution

The late time solution reads (in terms of scaling variable v , for v fixed and τ going to infinity):

$$a(v) = \frac{1}{2} \left(1 - \frac{1}{D} \right) \ln(1 + D v^4) + \frac{1}{2} \left(1 + \frac{1}{D} \right) \ln(1 - D v^4)$$

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At this point Janik and Peschanski fixed the power Δ by requiring that the curvature invariant has no singularities:

$$\mathcal{R} = R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} < \infty$$

Late Time Solution: Branch Cuts

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This is not your run of the mill singularity: this is a branch cut! This means that the metric becomes complex and multivalued for $1 - D v^4 \leq 0$! Since the metric has to be real and single-valued we conclude that the metric (and the curvature invariant) do not exist for $1 - D v^4 \leq 0$. That is unless the coefficients in front of the logarithms are integers!

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If the coefficients in front of the logarithms are integers, functions A , B and C would be single-valued and real.

Late Time Solution: Fixing the Power

Requiring the coefficients in front of the logarithms to be integers l, m, n

$$\begin{aligned}\frac{1}{2} \left(1 + \frac{1}{D} \right) &= n \\ \frac{1}{2} \left(1 + \frac{\Delta + 1}{D} \right) &= m \\ \frac{1}{2} \left(1 - \frac{\Delta + 2}{2D} \right) &= l\end{aligned}$$

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after simple algebra (!) one obtains that the only allowed power is $\Delta = -\frac{4}{3}$, giving the Bjorken hydrodynamic scaling of the energy density, reproducing the result of Janik and Peschanski

$$\epsilon(\tau) = -\frac{N_c^2}{2\pi^2} a_0 \frac{1}{\tau^{4/3}} \propto \frac{1}{\tau^{4/3}}$$

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Iterative Solution

We begin by expanding the coefficients of the metric

$$ds^2 = \frac{1}{z^2} [-A(\tau, z) d\tau^2 + \tau^2 B(\tau, z) d\eta^2 + C(\tau, z) dx_{\perp}^2 + dz^2]$$

$$A(\tau, z) = e^{a(\tau, z)}, \quad B(\tau, z) = e^{b(\tau, z)}, \quad C(\tau, z) = e^{c(\tau, z)}$$

into power series in z :

$$a(\tau, z) = \sum_{n=0}^{\infty} a_n(\tau) z^{4+2n}, \quad b(\tau, z) = \sum_{n=0}^{\infty} b_n(\tau) z^{4+2n}, \quad c(\tau, z) = \sum_{n=0}^{\infty} c_n(\tau) z^{4+2n}.$$

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Assuming power-law scaling

$$\varepsilon \sim \tau^{\Delta}$$

we iteratively obtain coefficients in the expansion

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To illustrate their structure let me display one of them:

$$a_2(\tau) = -\frac{1}{384} \left[a_0 \tau^{\Delta-4} (4 \Delta^2 - \Delta^4) + a_0^2 \tau^{2\Delta} 8 (8 + 8 \Delta + 3 \Delta^2) \right]$$

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(only if $\Delta > -4$!)

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Early Time Solution: Scaling

Let us apply the same strategy to the early-time solution: using perturbative (in z) solution at early times give the following series

$$\begin{aligned} a(\tau, z) &= \# z^4 \tau^\Delta + \# ' z^6 \tau^{\Delta-2} + \# '' z^8 \tau^{\Delta-4} + \dots \\ &= z^4 \tau^\Delta \left(\# + \# ' \frac{z^2}{\tau^2} + \# '' \frac{z^4}{\tau^4} + \dots \right) \end{aligned}$$

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such that

$$a(\tau, u) = \tau^{\Delta+4} u^4 \left(\# + \# ' u^2 + \# '' u^4 + \dots \right)$$

Early Time Solution: Ansatz

Keeping u fixed and taking $\tau \rightarrow 0$, we write the following ansatz for the metric coefficients:

$$A(\tau, u) = e^{a(\tau, u)} = e^{\tau^{\Delta+4} \alpha(u)} = 1 + \tau^{\Delta+4} \alpha(u) + o(\tau^{2\Delta+8})$$

$$B(\tau, u) = 1 + \tau^{\Delta+4} \beta(u) + o(\tau^{2\Delta+8})$$

$$C(\tau, u) = 1 + \tau^{\Delta+4} \gamma(u) + o(\tau^{2\Delta+8})$$

with α , β and γ some unknown functions of u .

Early-Time General Solution

Solving Einstein equations yields

$$A(\tau, u) = 1 + a_0 \tau^{4+\Delta} u^4 F\left(-1 - \frac{\Delta}{2}, -\frac{\Delta}{2}; 3; u^2\right),$$

$$B(\tau, u) = 1 + a_0 \tau^{4+\Delta} u^4 \left[(\Delta + 1) F\left(-1 - \frac{\Delta}{2}, -\frac{\Delta}{2}; 3; u^2\right) - \frac{\Delta(\Delta + 2)}{6} u^2 F\left(1 - \frac{\Delta}{2}, -\frac{\Delta}{2}; 4; u^2\right) \right],$$

$$C(\tau, u) = 1 + a_0 \tau^{4+\Delta} u^4 \frac{\Delta + 2}{12} \left[-6 F\left(-1 - \frac{\Delta}{2}, -\frac{\Delta}{2}; 3; u^2\right) + \Delta u^2 F\left(1 - \frac{\Delta}{2}, -\frac{\Delta}{2}; 4; u^2\right) \right].$$

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Hypergeometric functions have a branch cut for $u > 1$.

We have branch cuts again!

Allowed Powers of Proper Time

However, now hypergeometric functions are not in the exponent. The only way to avoid branch cuts is to have hypergeometric series terminate at some finite order, becoming a polynomial.

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Before we do that we note that, at early times the total energy of the produced medium is $E \sim \varepsilon \tau$.

Requiring it to be finite $E \sim \varepsilon \tau < \infty$ we conclude that for

$\varepsilon \sim \tau^{\Delta}$ the power should be $\Delta \geq -1$.

Allowed Powers of Proper Time

However, now hypergeometric functions are not in the exponent. The only way to avoid branch cuts is to have hypergeometric series terminate at some finite order, becoming a polynomial.

Before we do that we note that, at early times the total energy of the produced medium is $E \sim \varepsilon \tau$.

Requiring it to be finite $E \sim \varepsilon \tau < \infty$ we conclude that for

$\varepsilon \sim \tau^\Delta$ the power should be $\Delta \geq -1$.

Hence, at early times the physically allowed powers are:

$$\boxed{-1 \leq \Delta \leq 0}$$

Early Time Solution: Terminating the Series

Finally, we see that the hypergeometric series in the solution

$$A(\tau, u) = 1 + a_0 \tau^{4+\Delta} u^4 F\left(-1 - \frac{\Delta}{2}, -\frac{\Delta}{2}; 3; u^2\right),$$

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$$C(\tau, u) = 1 + a_0 \tau^{4+\Delta} u^4 \frac{\Delta + 2}{12} \left[-6 F\left(-1 - \frac{\Delta}{2}, -\frac{\Delta}{2}; 3; u^2\right) + \Delta u^2 F\left(1 - \frac{\Delta}{2}, -\frac{\Delta}{2}; 4; u^2\right) \right].$$

terminates only for $\Delta = 0$ in the physically allowed

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Early Time Solution

The early-time scaling of the energy density in this strongly-coupled medium is

$$\epsilon(\tau) \rightarrow \text{constant} \quad \text{as} \quad \tau \rightarrow 0.$$

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This leads to the following energy-momentum tensor, reminiscent of CGC at very early times:

$$T^{\mu\nu} = \begin{pmatrix} \epsilon(\tau) & 0 & 0 & 0 \\ 0 & \epsilon(\tau) & 0 & 0 \\ 0 & 0 & \epsilon(\tau) & 0 \\ 0 & 0 & 0 & -\epsilon(\tau) \end{pmatrix} \begin{matrix} x_0 \\ x_1 \\ x_2 \\ x_3 \end{matrix}$$

Early Time Solution: Log Ansatz

One can also look for the solution with the logarithmic ansatz (sort of like fine-tuning):

$$\epsilon(\tau) = -\frac{N_c^2}{2\pi^2} \left[a_0 \ln^\delta \left(\frac{1}{\tau} \right) + a_1 \ln^{\delta-1} \left(\frac{1}{\tau} \right) \right]$$

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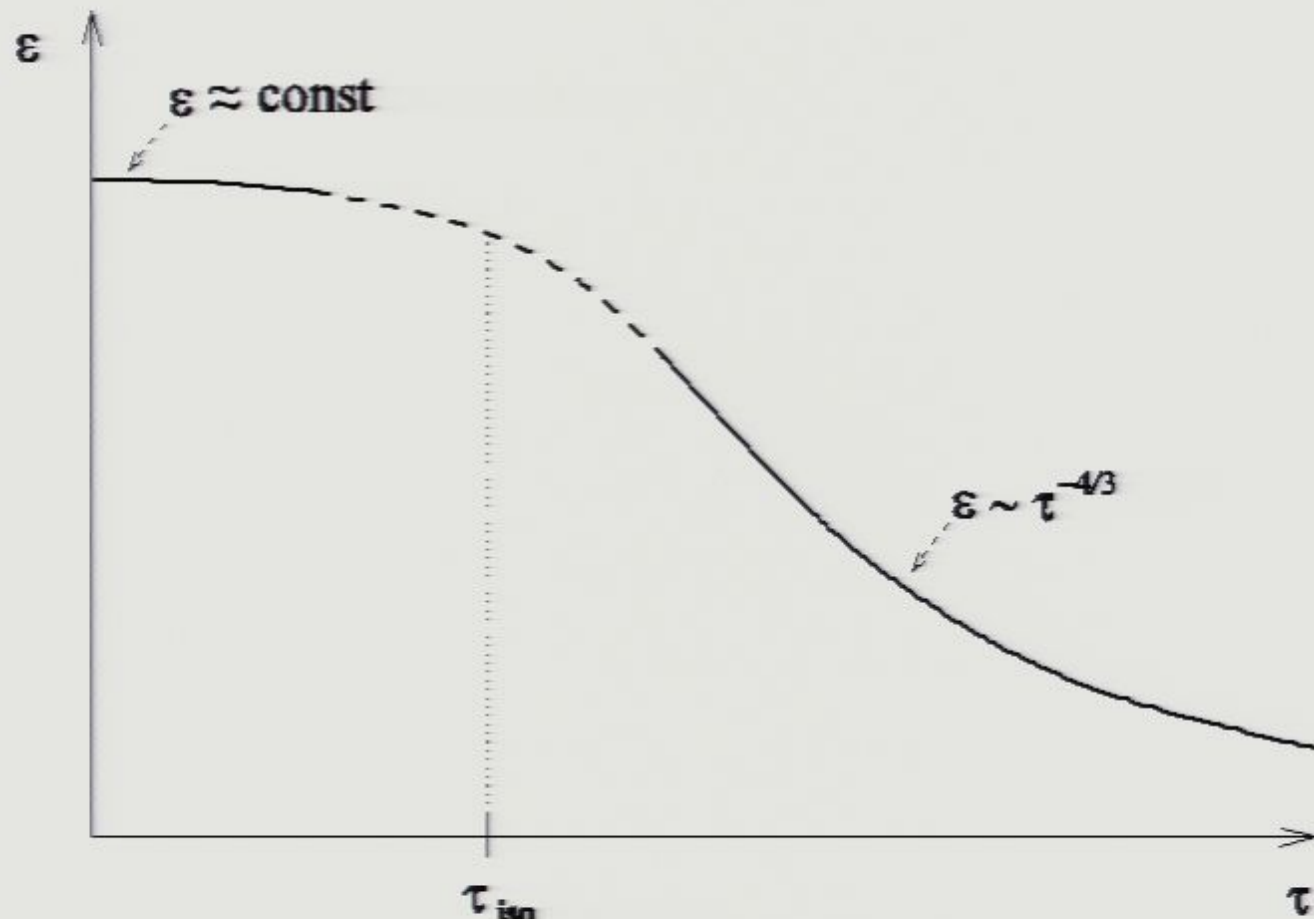
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The approach to a constant at early times could be logarithmic! (More work is needed to sort this out.)

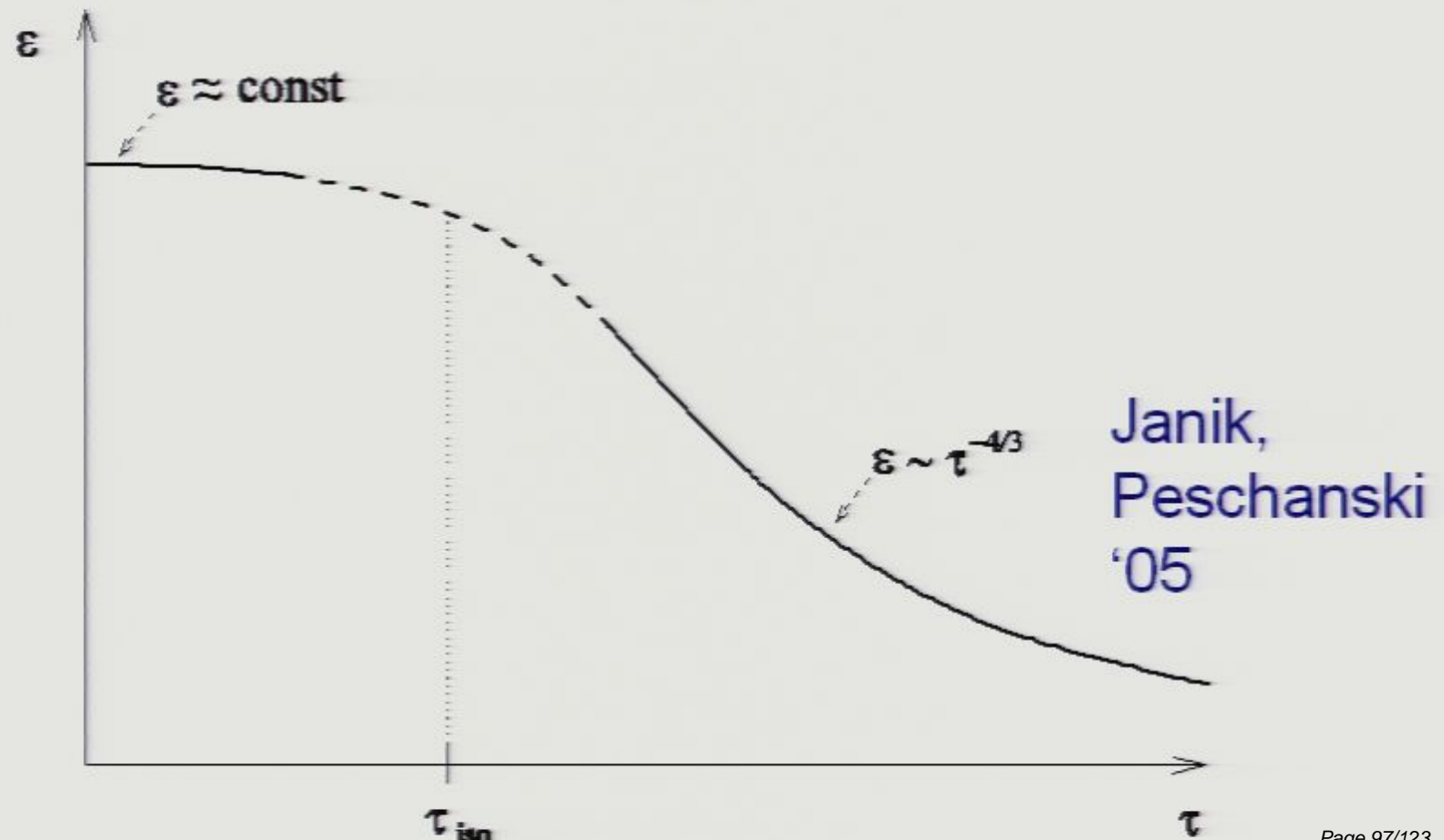
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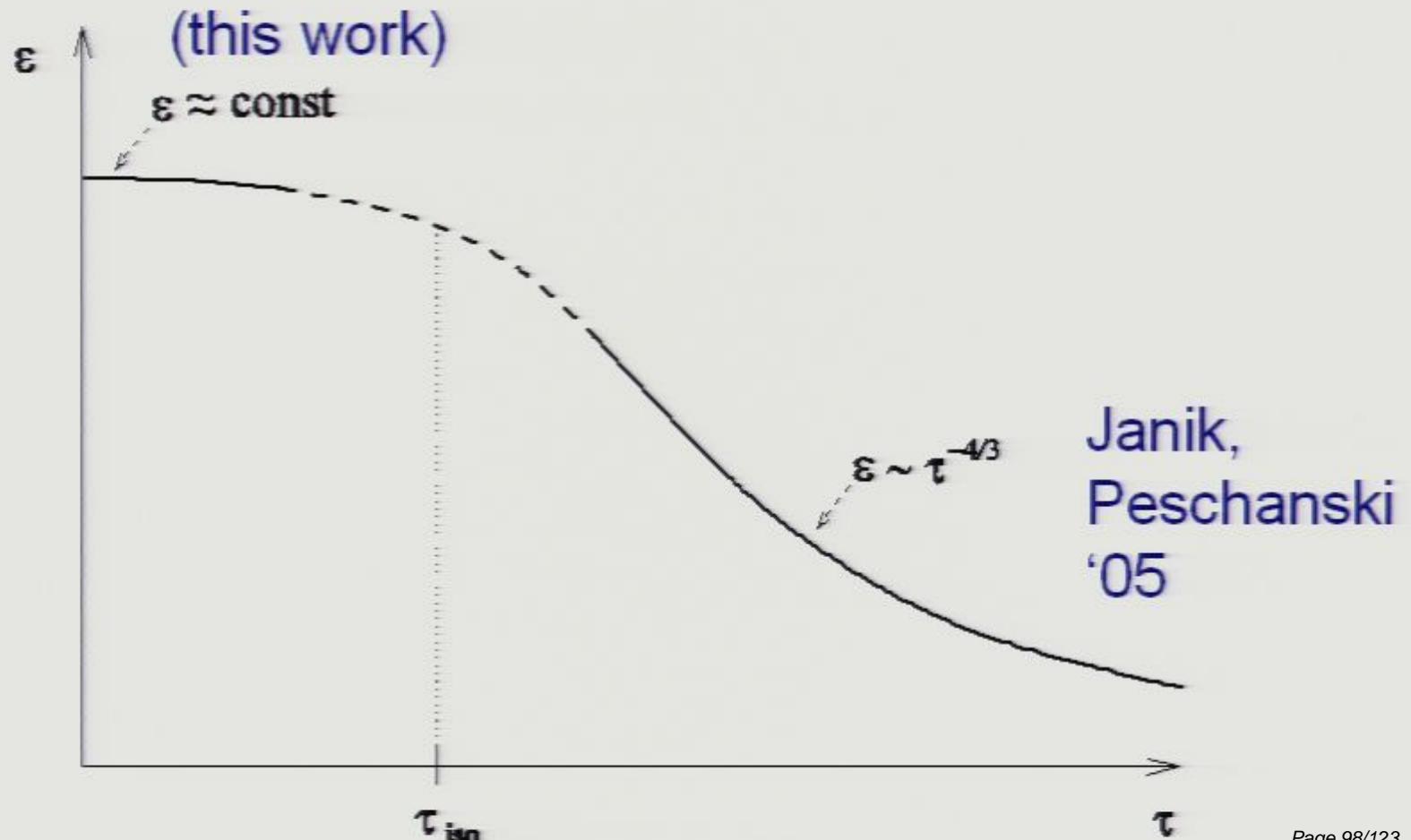
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$$a_2(\tau) = -\frac{1}{384} \left[a_0 \tau^{\Delta-4} (4 \Delta^2 - \Delta^4) + a_0^2 \tau^{2\Delta} 8 (8 + 8 \Delta + 3 \Delta^2) \right]$$

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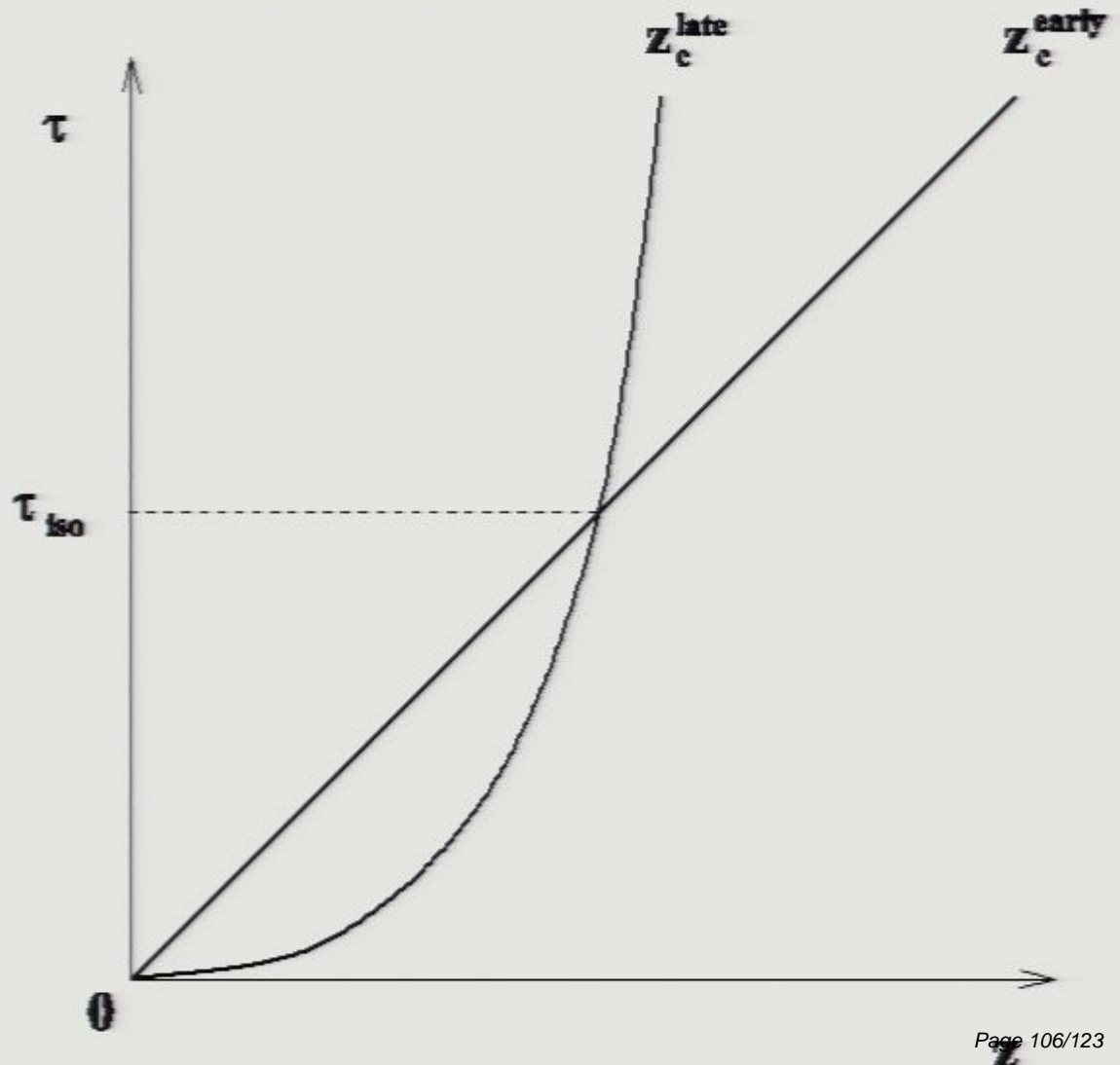
$$u = 1$$

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Isotropization Transition: Time Estimate

We plot both branch cuts in the (z, τ) plane:

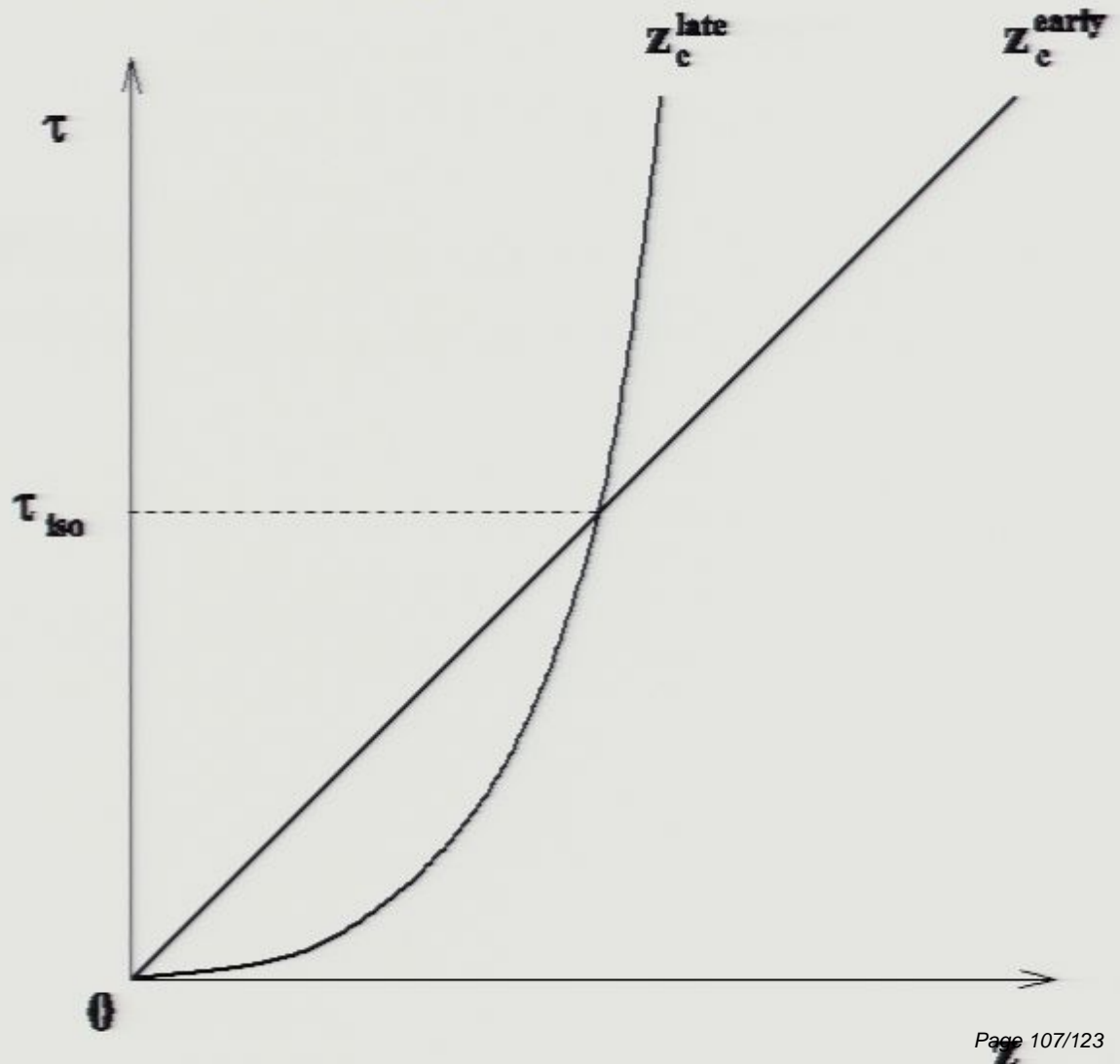


Isotropization Transition: Time Estimate

We plot both branch cuts in the (z, τ) plane:

The intercept is at the
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$$\tau_{\text{iso}} = \left(\frac{3}{-a_0} \right)^{\frac{3}{8}}$$



Isotropization Transition: Time Estimate

In terms of more physical quantities we re-write the above estimate as

$$\tau_{\text{iso}} = \left(\frac{3}{\epsilon_0} \frac{N_c^2}{2\pi^2} \right)^{\frac{3}{8}}$$

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$$\tau_{\text{iso}} \approx 0.29 \text{ fm/c}$$

Isotropization Transition Estimate: Self-Critique

An AdS/CFT skeptic would argue that our estimate

$$\tau_{\text{iso}} = \left(\frac{3}{\epsilon_0} \frac{N_c^2}{2\pi^2} \right)^{\frac{3}{8}}$$

is easy to obtain from dimensional reasoning. If one has a conformally invariant theory with $\epsilon(\tau) = \frac{\epsilon_0}{\tau^{\frac{4}{3}}}$, the only

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- Re-derived JP late-time results without requiring the curvature invariant to be finite: all we need is for the metric to exist.
- Analyzed early-time dynamics and showed that energy density goes to a constant at early times.
- Have therefore shown that isotropization (and hopefully thermalization) takes place in strong coupling dynamics.
- Derived a simple formula for isotropization time and used it for heavy ion collisions at RHIC to obtain 0.3 fm/c , in agreement with hydrodynamic simulations.

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Finally, we see that the hypergeometric series in the solution

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Iterative Solution

We begin by expanding the coefficients of the metric

$$ds^2 = \frac{1}{z^2} [-A(\tau, z) d\tau^2 + \tau^2 B(\tau, z) d\eta^2 + C(\tau, z) dx_{\perp}^2 + dz^2]$$

$$A(\tau, z) = e^{a(\tau, z)}, \quad B(\tau, z) = e^{b(\tau, z)}, \quad C(\tau, z) = e^{c(\tau, z)}$$

into power series in z :

$$a(\tau, z) = \sum_{n=0}^{\infty} a_n(\tau) z^{4+2n}, \quad b(\tau, z) = \sum_{n=0}^{\infty} b_n(\tau) z^{4+2n}, \quad c(\tau, z) = \sum_{n=0}^{\infty} c_n(\tau) z^{4+2n}.$$

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