

Title: Quantum Black Holes and High Energy

Date: May 22, 2007 04:00 PM

URL: <http://pirsa.org/07050053>

Abstract:

Exotic states of hot and dense matter, Perimeter Institute, 2007

The role of (broken) scale invariance of QCD at RHIC

D. Kharzeev

BNL

Exotic states of hot and dense matter, Perimeter Institute, 2007

The role of (broken) scale invariance of QCD at RHIC

D. Kharzeev

BNL

Outline

Running QCD coupling and the initial conditions in hydrodynamical calculations

Based on work with E.Levin, M.Nardi;
T.Hirano, U.Heinz, Y.Nara, R.Lacey

Broken scale invariance and
bulk viscosity close to T_c

Recent work with K.Tuchin

QCD and quantum anomalies

$$\mathcal{L} = -\frac{1}{4}G_{\mu\nu}^a G_{\mu\nu}^a + \sum_f \bar{q}_f^a (i\gamma_\mu D_\mu - m_f) q_f^a;$$

Classical scale invariance is broken by quantum effects:

scale anomaly

$$\theta_\mu^\mu = \frac{\beta(g)}{2g} G^{\alpha\beta a} G_{\alpha\beta}^a + \sum_q m_q \bar{q}q$$

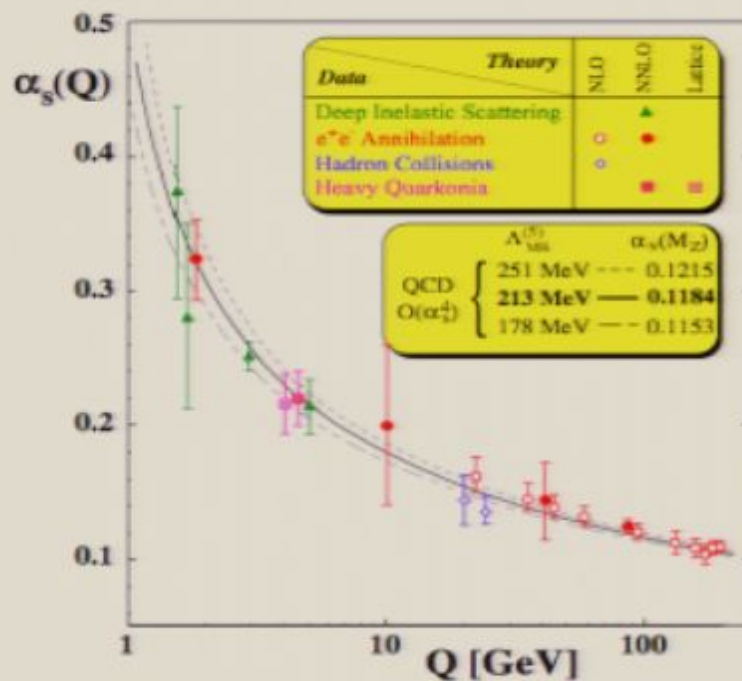
trace of the energy-momentum tensor

“beta-function”; describes the dependence of coupling on momentum

$$\mu \frac{dg(\mu)}{d\mu} = \beta(g)$$

Hadrons get masses \longleftrightarrow coupling runs with the distance

Asymptotic Freedom



At short distances,
the strong force becomes weak
(**anti-screening**) -
one can access the “asymptotically
free” regime in hard processes

and in super-dense matter
(inter-particle distances $\sim 1/T$)

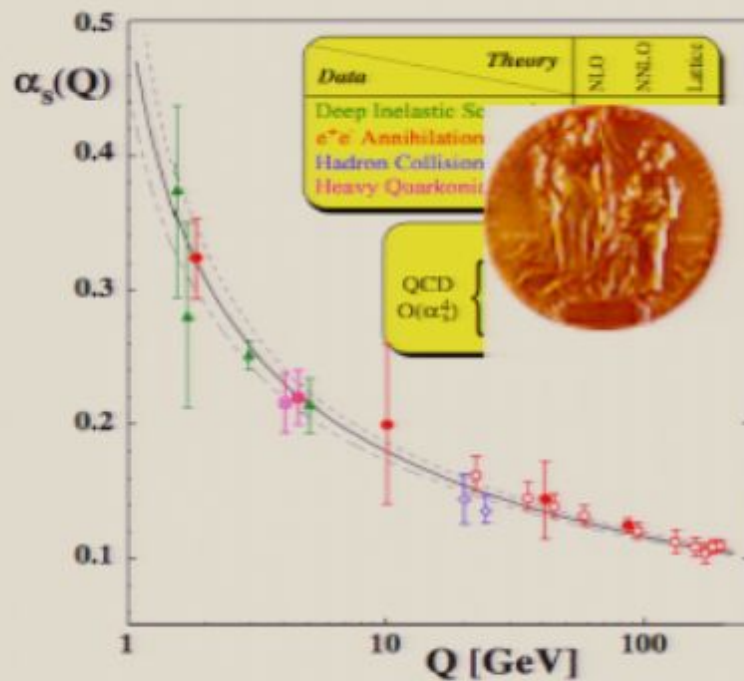
$$\alpha_s(Q) \simeq \frac{4\pi}{b \ln(Q^2/\Lambda^2)}$$

number
of colors

number
of flavors

$$b = (11N_c - 2N_f)/3$$

Asymptotic Freedom



At short distances,
the strong force becomes weak
(**anti-screening**) -
one can access the “asymptotically
free” regime in hard processes

and in super-dense matter
(inter-particle distances $\sim 1/T$)

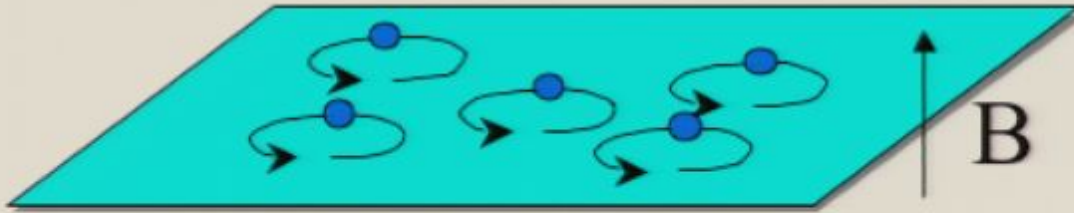
$$\alpha_s(Q) \simeq \frac{4\pi}{b \ln(Q^2/\Lambda^2)}$$

number
of colors

number
of flavors

$$b = (11N_c - 2N_f)/3$$

Asymptotic freedom and Landau levels of 2D parton gas



The effective potential: sum over 2D Landau levels

$$V_{\text{pert}}(H) = \frac{g H}{4 \pi^2} \int dp_z \sum_{n=0}^{\infty} \sum_{s_z=\pm 1} \sqrt{2 g H (n + 1/2 - s_z) + p_z^2}.$$

Paramagnetic response of the vacuum: V

$$\text{Re } V_{\text{pert}}(H) = \frac{1}{2} H^2 + (g H)^2 \frac{b}{32 \pi^2} \left(\ln \frac{g H}{\mu^2} - \frac{1}{2} \right)$$

1. The lowest level $n=0$ of radius $\sim (gH)^{-1/2}$ is **unstable!**

2. Strong fields \longleftrightarrow Short distances

QCD and the classical limit

Classical dynamics applies when the action $S = \int d^4x \mathcal{L}(x)$ is large in units of the Planck constant (Bohr-Sommerfeld quantization)

$$\frac{S_{QCD}}{\hbar} \sim \frac{1}{g^2 \hbar} \int d^4x \operatorname{tr} G^{\mu\nu}(x) G_{\mu\nu}(x) \gg 1$$

($gA \rightarrow A$)

(equivalent to setting $\hbar \rightarrow 0$)

\Rightarrow Need weak coupling and strong fields

$$D_\mu = \partial_\mu - igA_\mu^a t^a$$

$$A^2 \ll \frac{p^2}{g^2}$$

weak
field

$$A^2 \sim \frac{p^2}{g^2}$$

strong
field

Renormalization group and the effective action

RG constraints the form of the effective action:

$$\mathcal{L}_{\text{eff}} = -\frac{1}{4\bar{g}^2(t)} G^2, \quad t \equiv \ln \left(\frac{G^2}{\Lambda^4} \right)$$

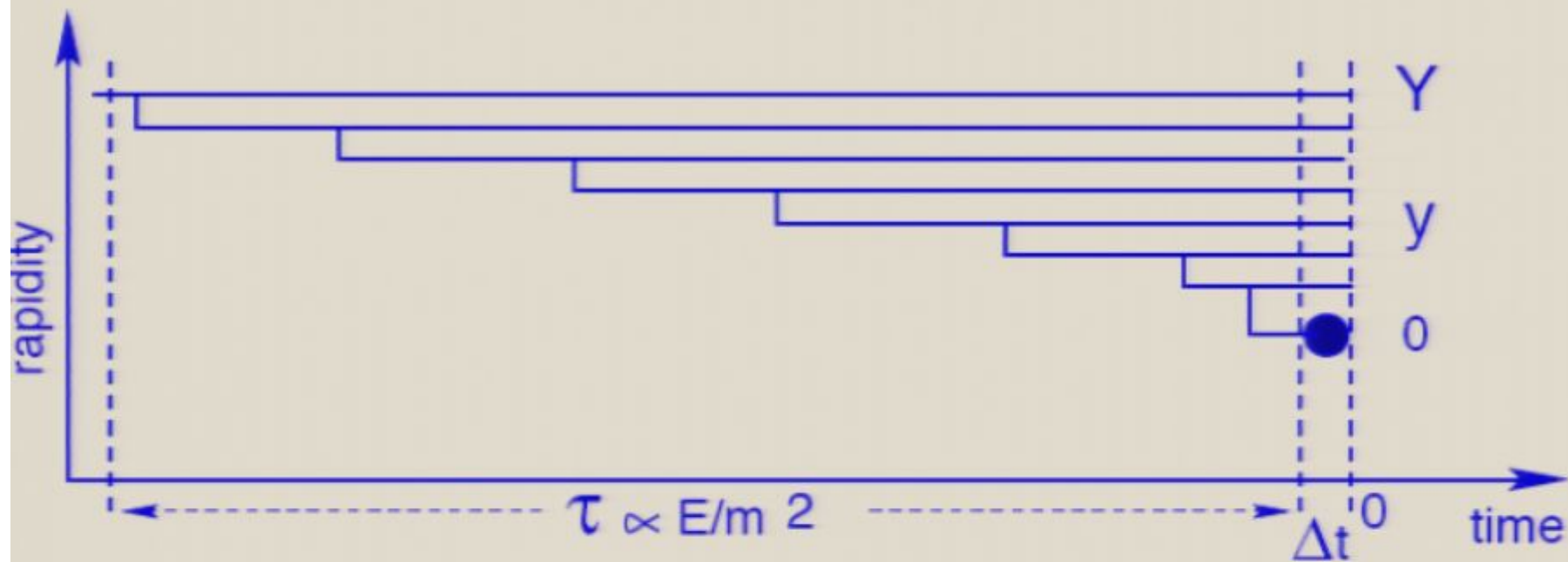
the coupling is defined through

$$t = \int_g^{\bar{g}(t)} \frac{dg}{\beta(g)}$$

At large t (strong color field),

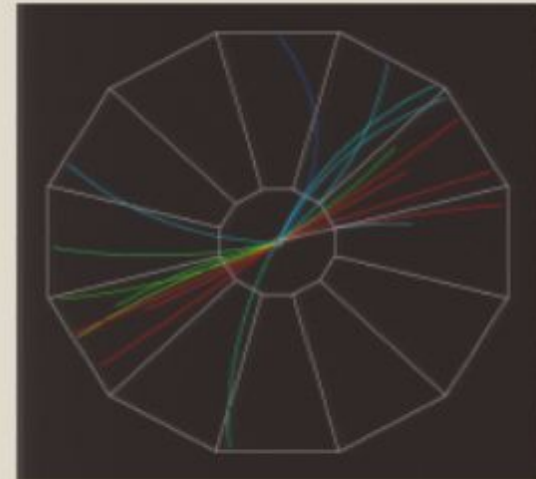
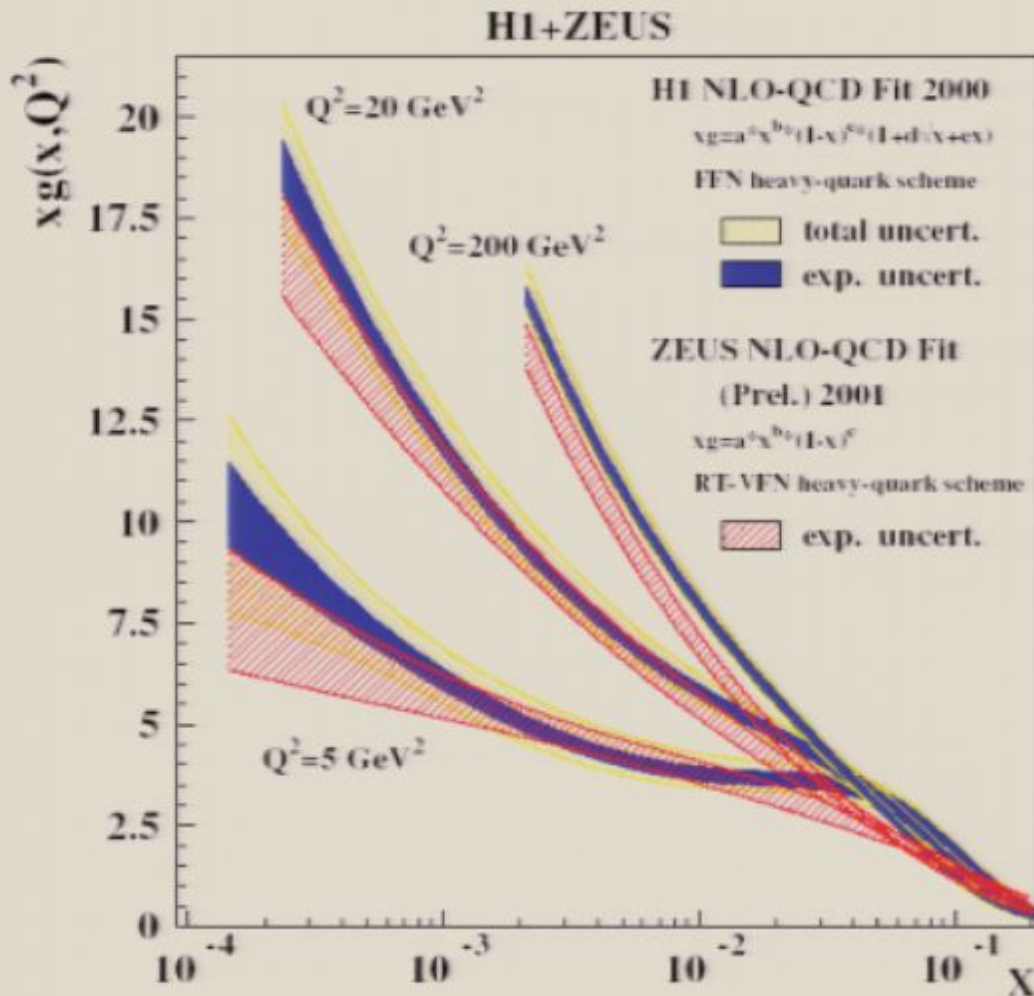
$$\frac{1}{\bar{g}^2(t)} \sim t + \dots \quad \text{and} \quad \mathcal{L}_{\text{eff}} \sim G^2 \ln \left(\frac{G^2}{\Lambda^4} \right)$$

The space-time picture of high-energy interactions in QCD



1. Fast (large y) partons live for a long time;
2. Parton splitting probability is $\sim \alpha_s y$ - not small!

Resolving the gluon cloud at small x and short distances $\sim 1/Q^2$

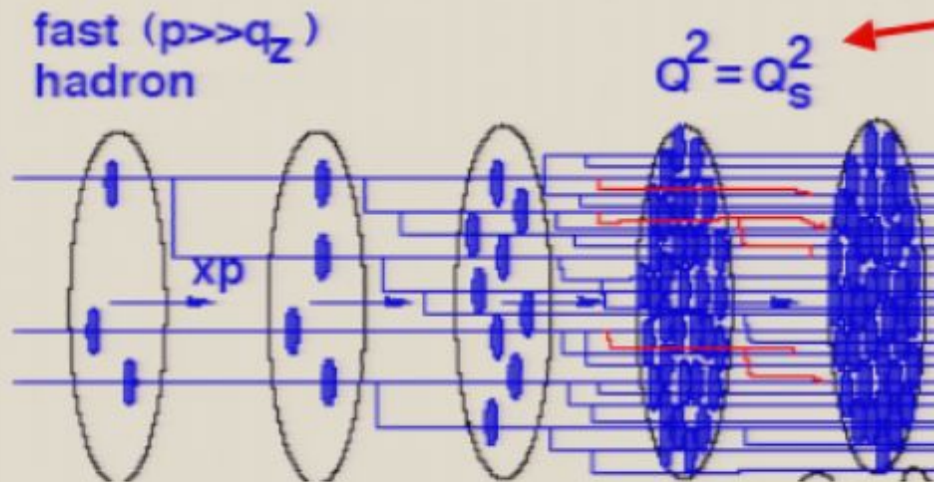
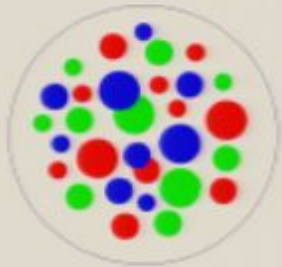


Building up strong color fields:

small x (high energy) and large A (heavy nuclei)

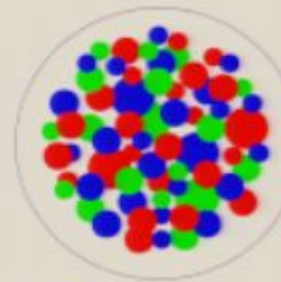
Bjorken x : the fraction of hadron's momentum carried by a parton; high energies s open access to small $x = Q^2/s$

Large x



the boundary of non-linear regime: partons of size $1/Q > 1/Q_s$ overlap

small x

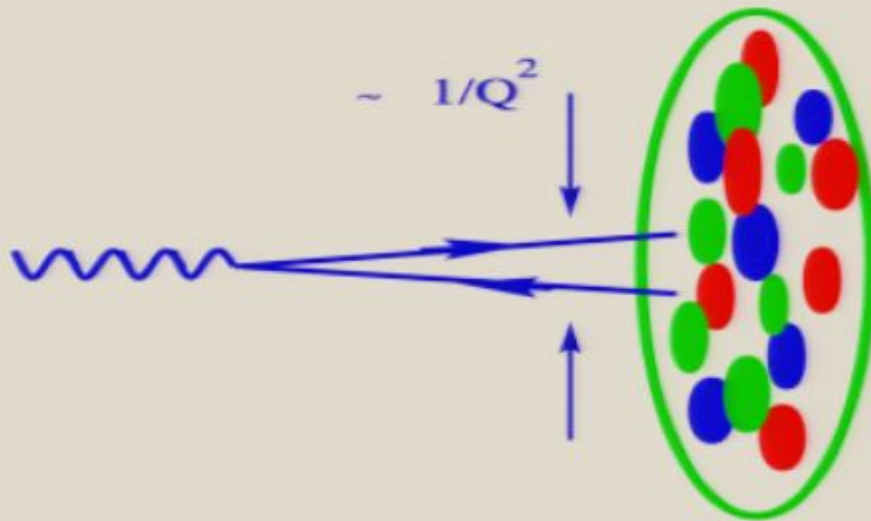


Because the probability to emit an extra gluon is $\sim \alpha_s \ln(1/x) \sim 1$, the number of gluons at small x grows; the transverse area is limited

transverse density becomes large

Strong color fields in heavy nuclei

At small Bjorken x , hard processes develop over large longitudinal distances $l_c \sim \frac{2\nu}{Q^2} = \frac{1}{mx}$

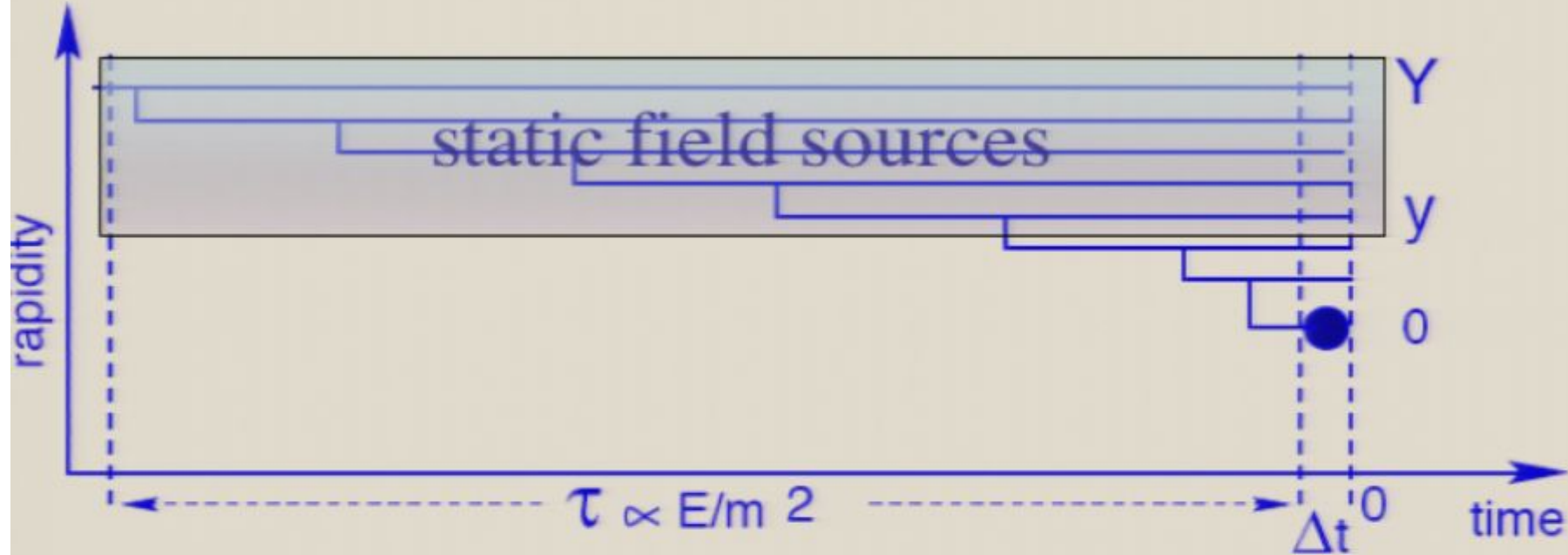


Density of partons in the transverse plane as a **new dimensionful parameter** Q_s (“saturation scale”)

Gribov, Levin, Ryskin

All partons contribute coherently \Rightarrow at sufficiently small x and/or large A strong fields, **weak coupling!**

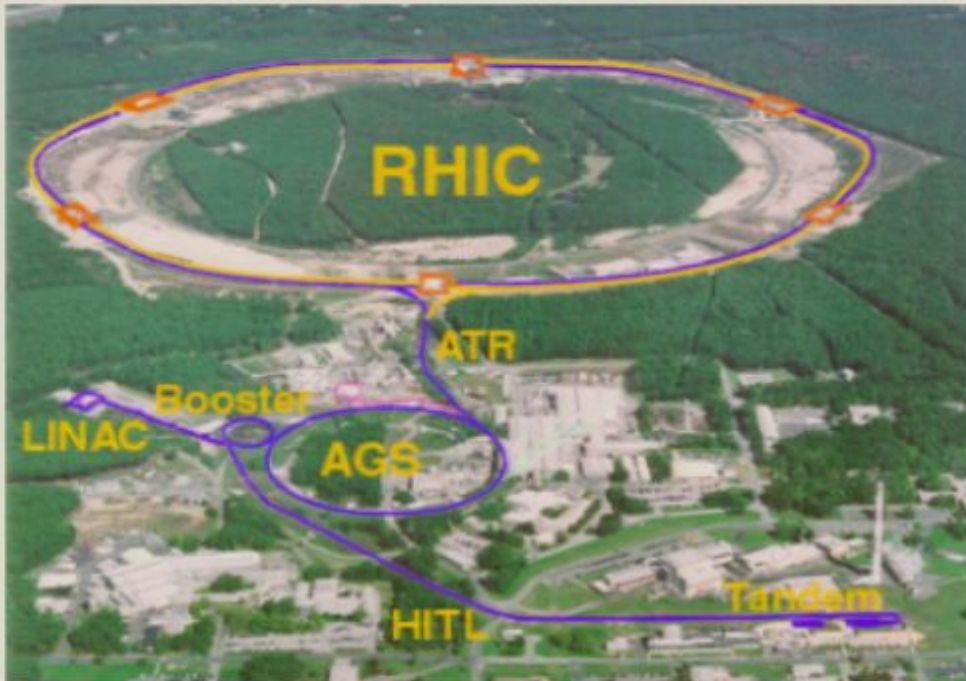
The origin of classical background field



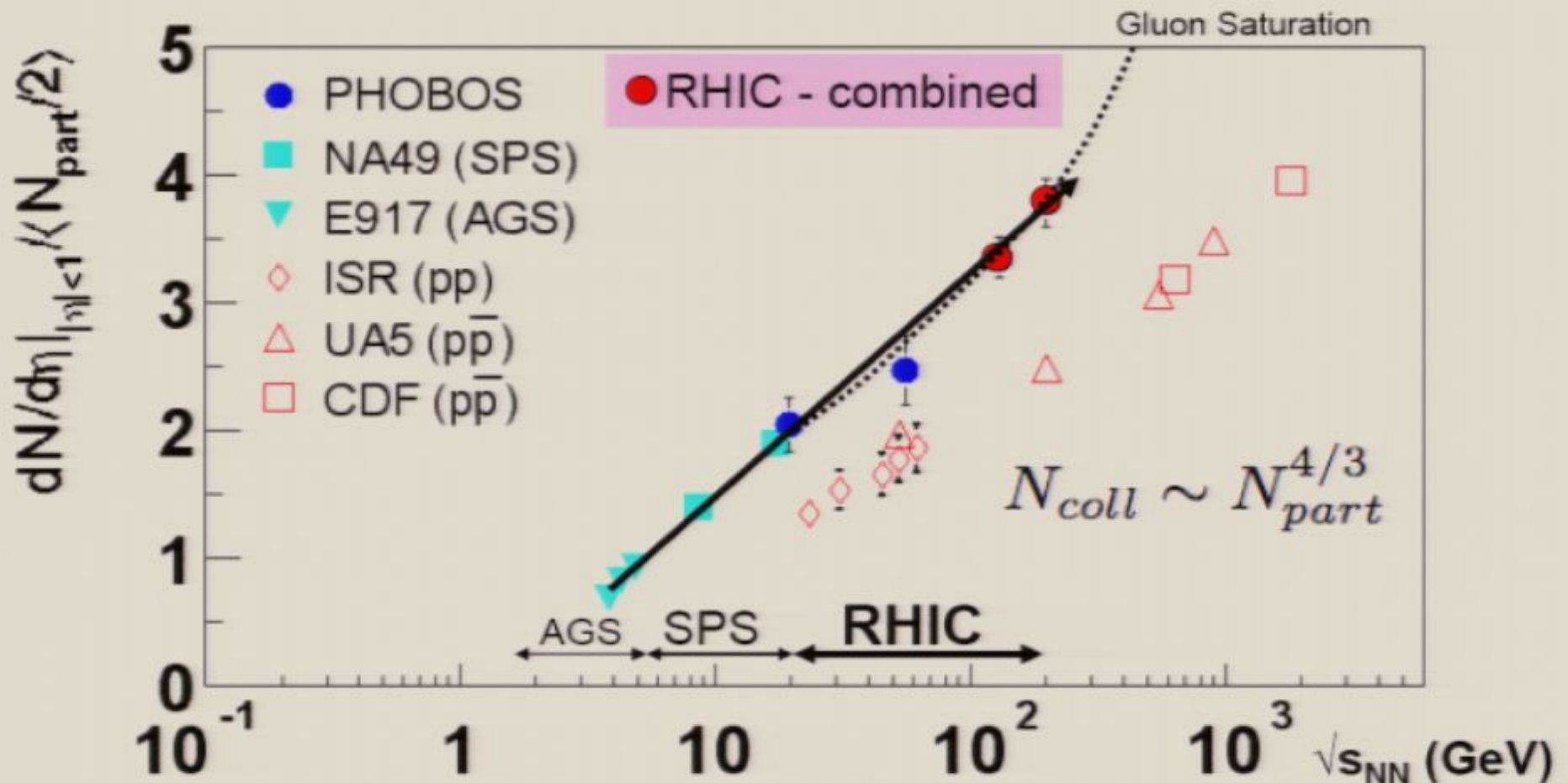
Gluons with large rapidity and large occupation number act as a background field for the production of slower gluons

“Color Glass Condensate”

Semi-classical QCD: experimental tests



Hadron multiplicities: the effect of parton coherence

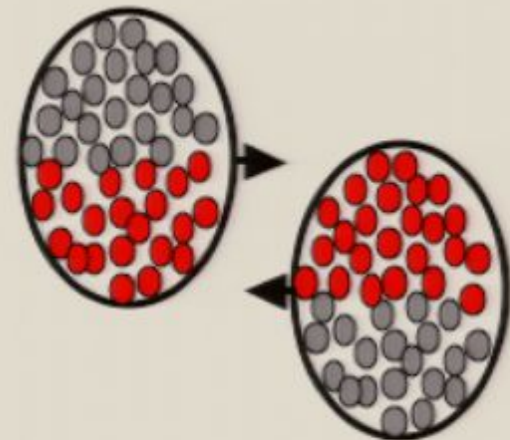


Semi-classical QCD and total multiplicities in heavy ion collisions

Expect very simple dependence of multiplicity on atomic number A / N_{part} :

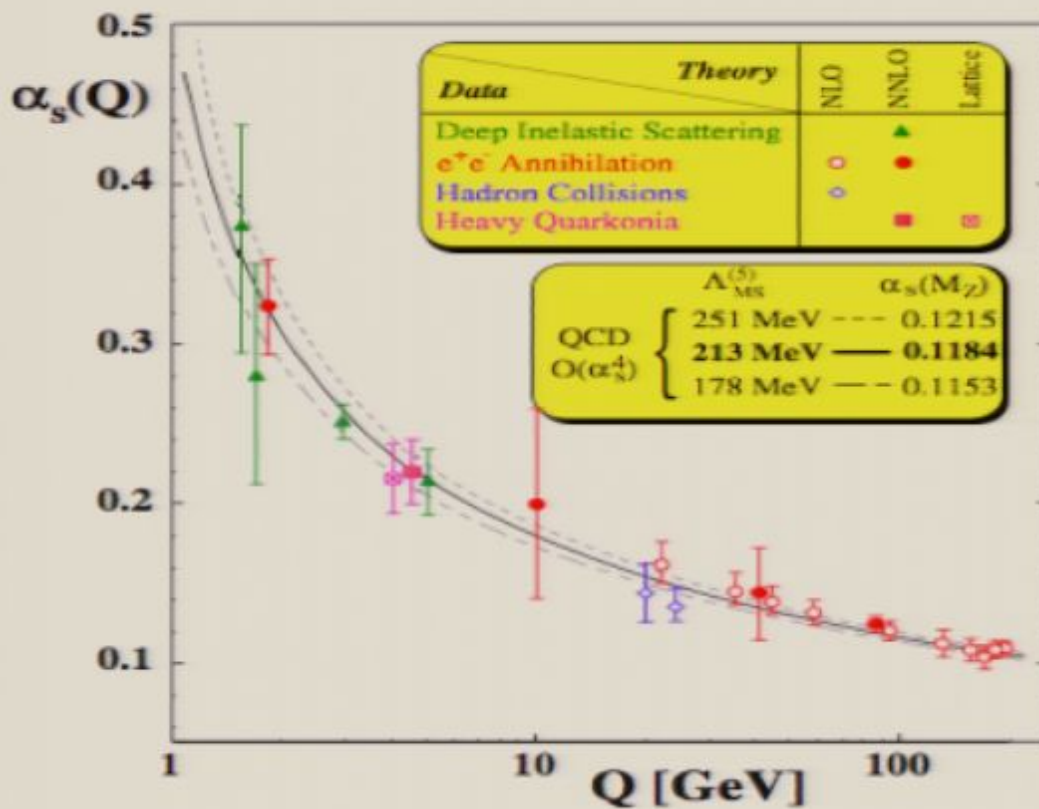
$$n \sim \frac{S_A Q_s^2}{\alpha_s(Q_s^2)} \sim N_{part} \ln N_{part}$$

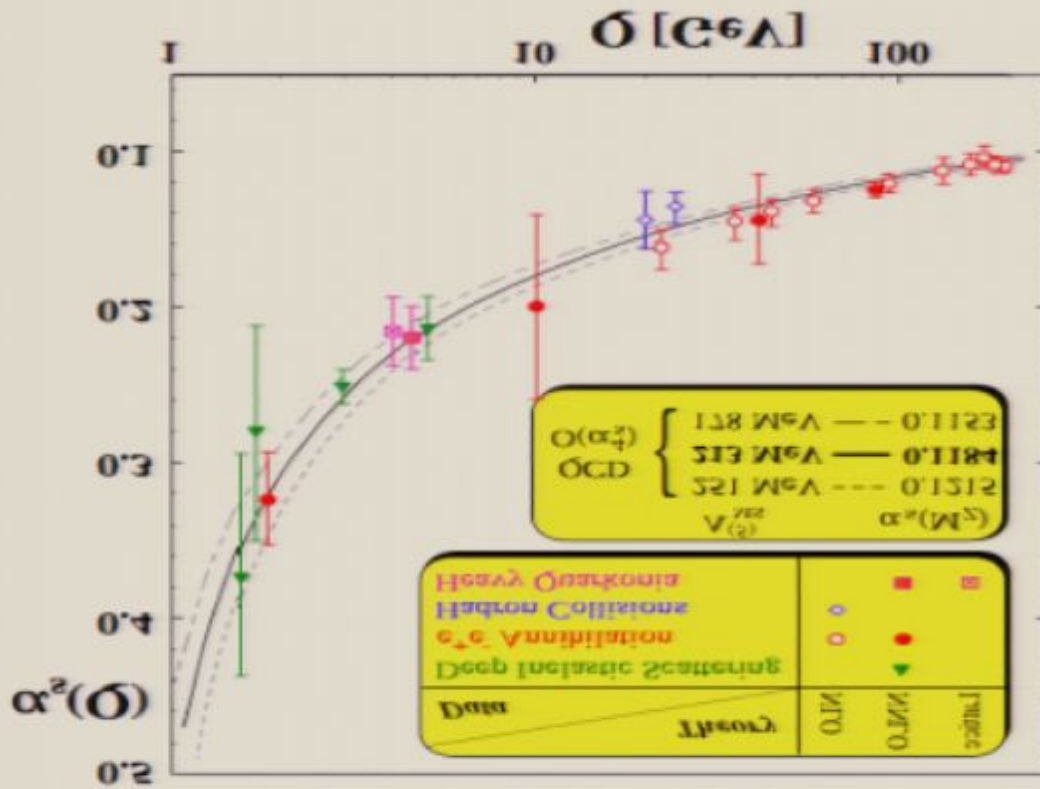
N_{part} :



Classical QCD in action

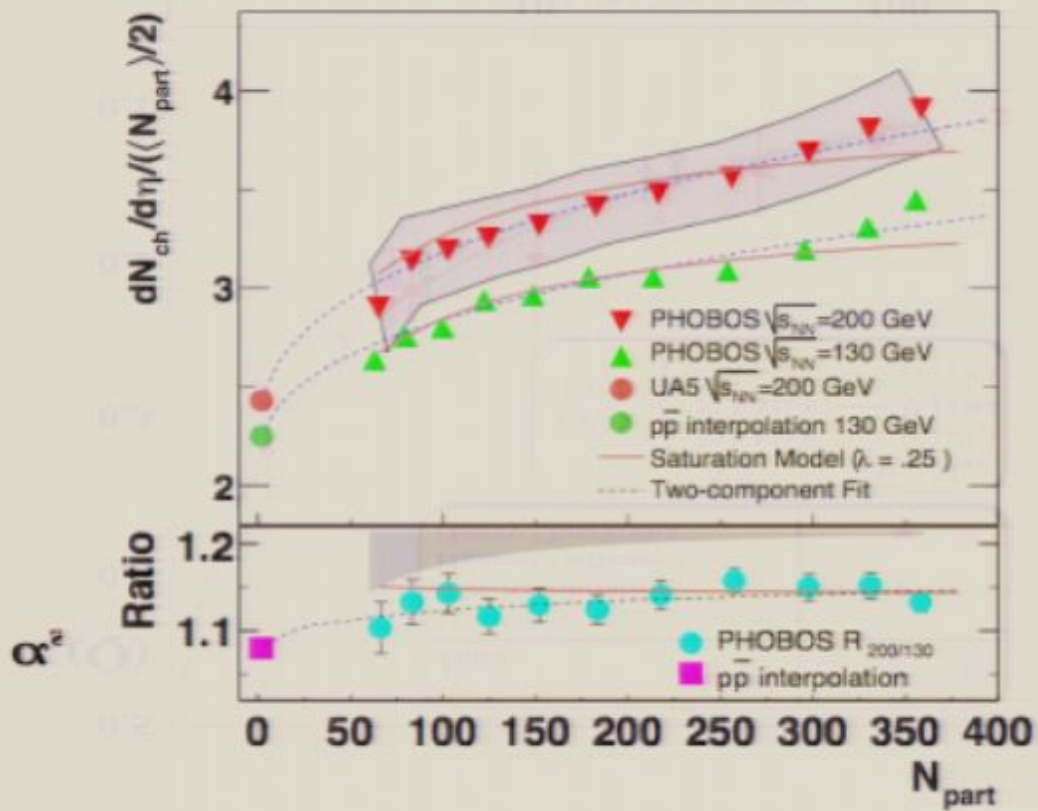
$$\frac{1}{N_{part}} \frac{dN}{d\eta} \sim \frac{1}{\alpha_s(Q)}$$

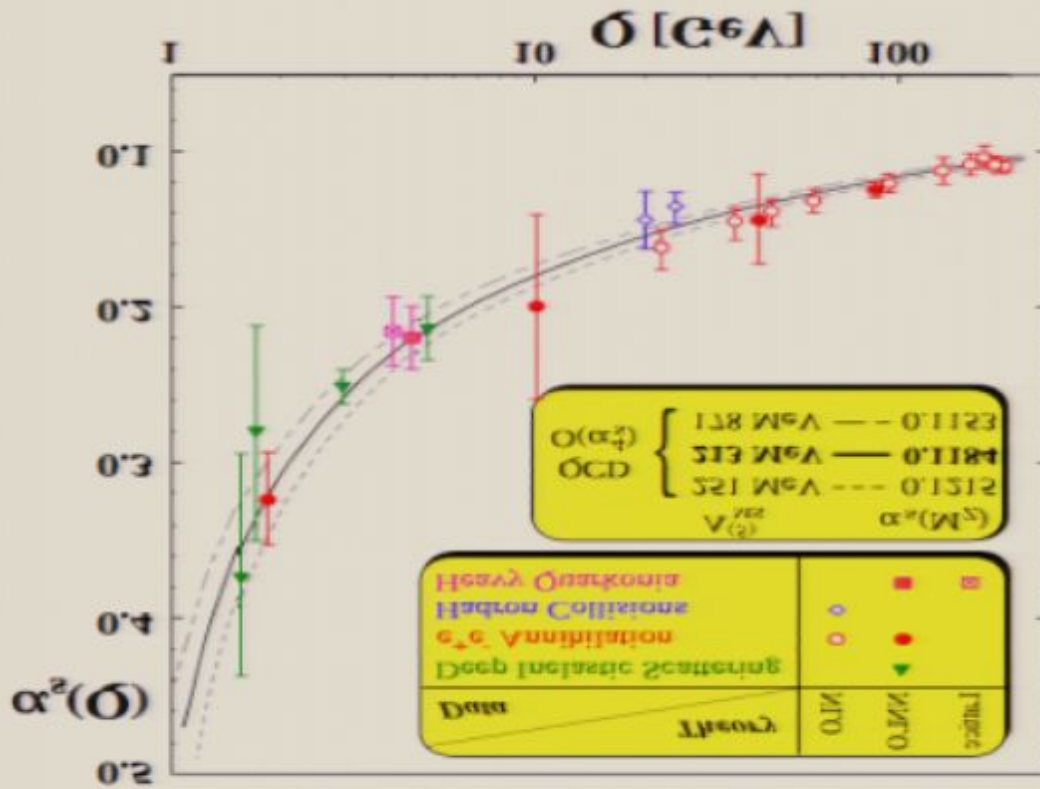




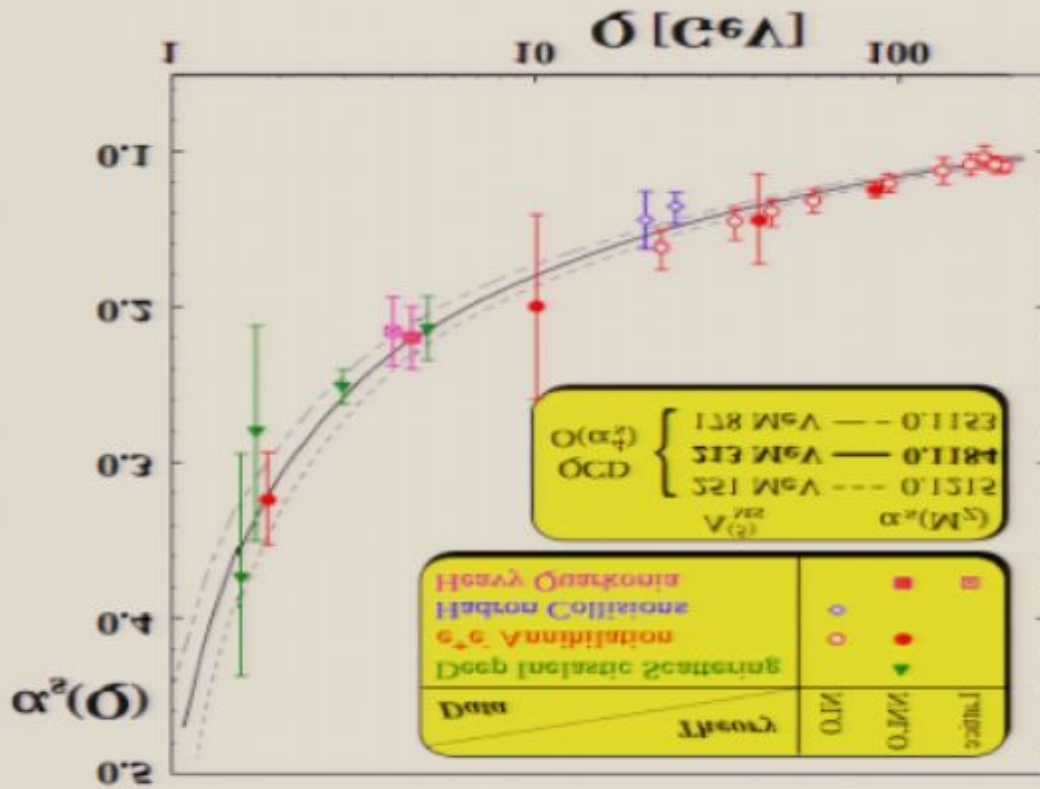
$$\frac{1}{N_{part}} \frac{dN}{d\eta} \sim \frac{1}{\alpha_s(Q)}$$

$$\frac{1}{N_{part}} \frac{dN}{d\eta} \sim \frac{1}{\alpha_s(Q)}$$



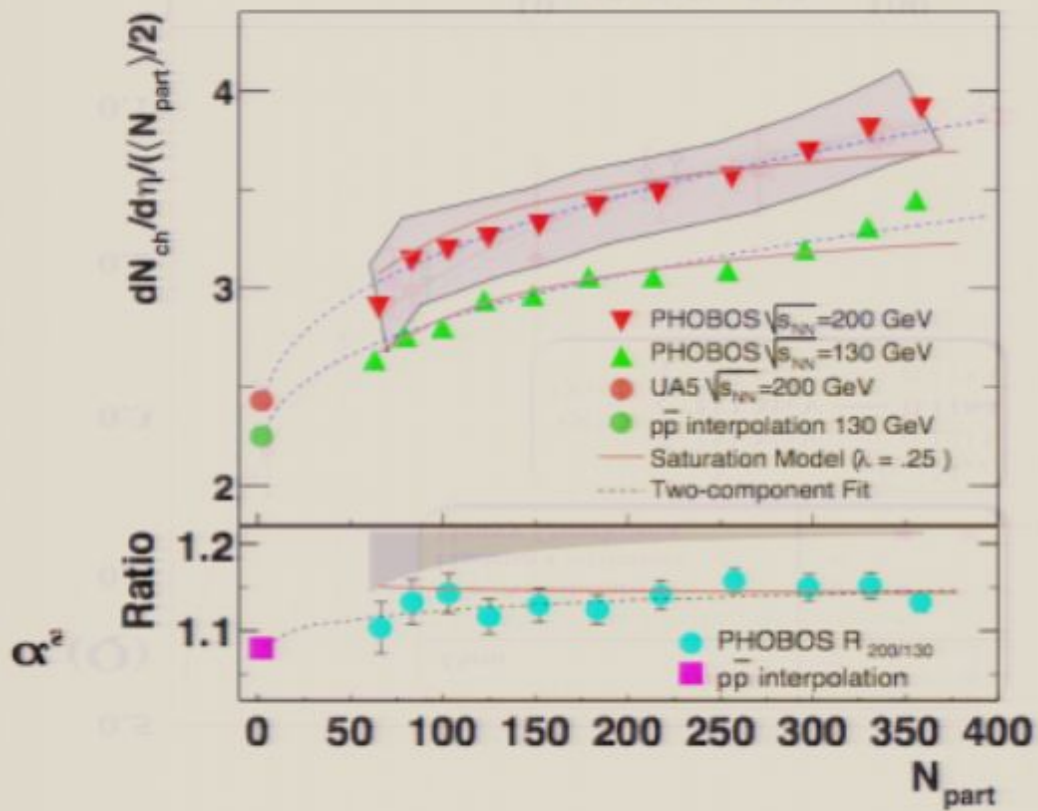


$$\frac{1}{N_{part}} \frac{dN}{d\eta} \sim \frac{1}{\alpha_s(Q)}$$



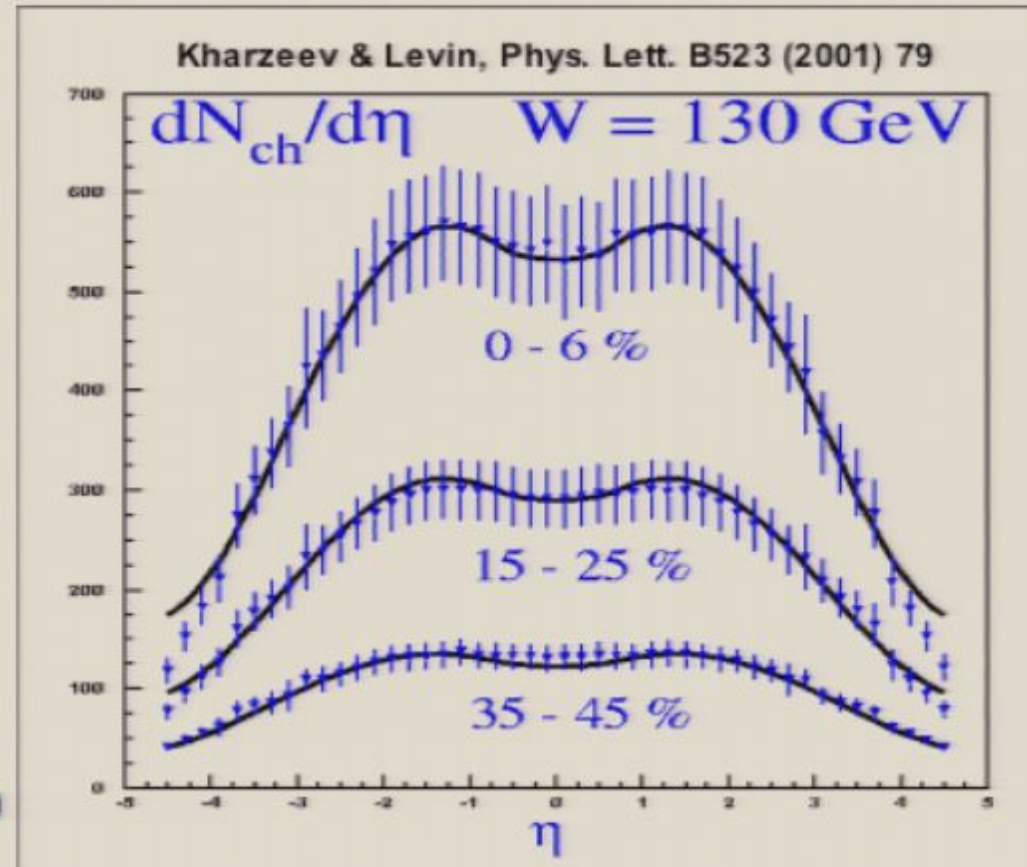
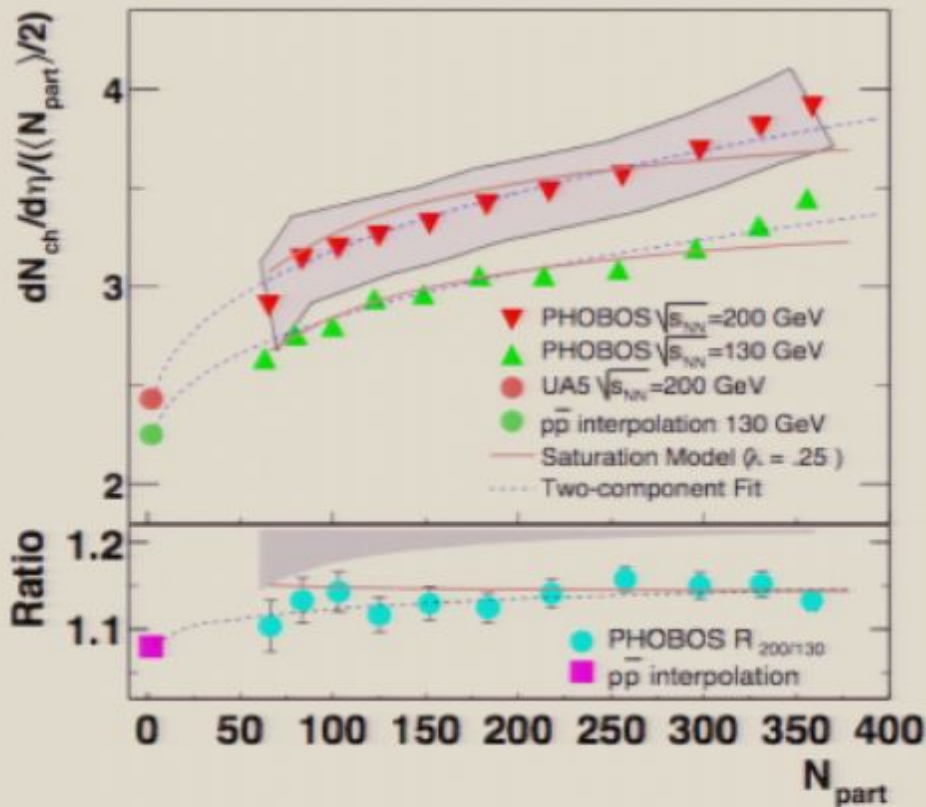
$$\frac{1}{N_{part}} \frac{dN}{d\eta} \sim \frac{1}{\alpha_s(Q)}$$

$$\frac{1}{N_{part}} \frac{dN}{d\eta} \sim \frac{1}{\alpha_s(Q)}$$



Classical QCD dynamics in action

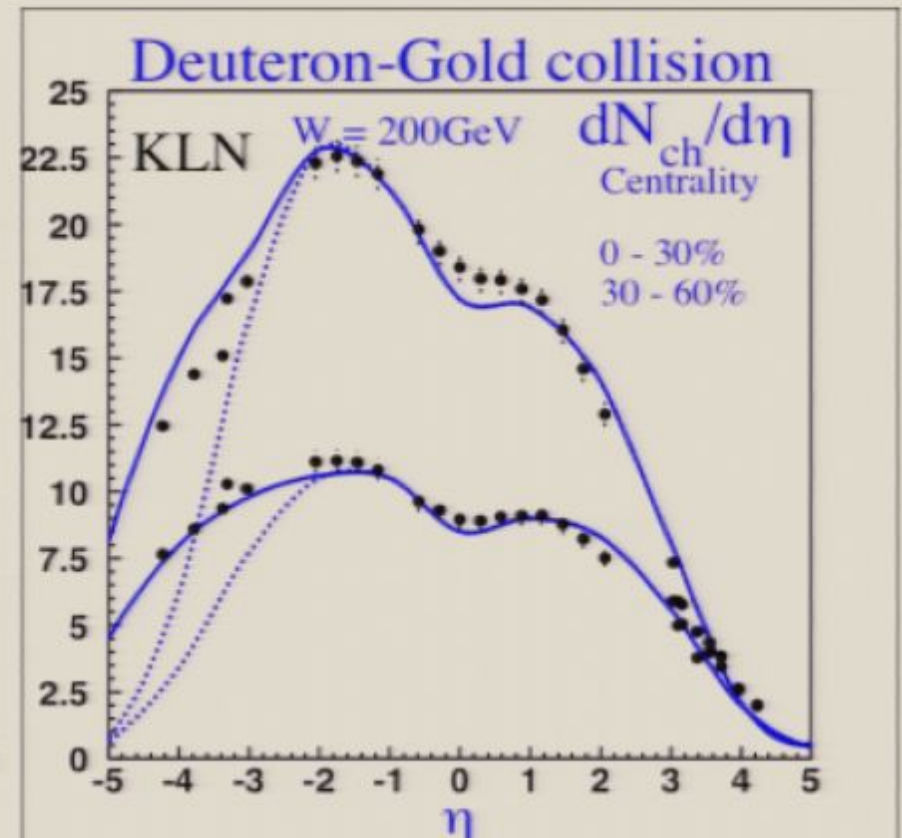
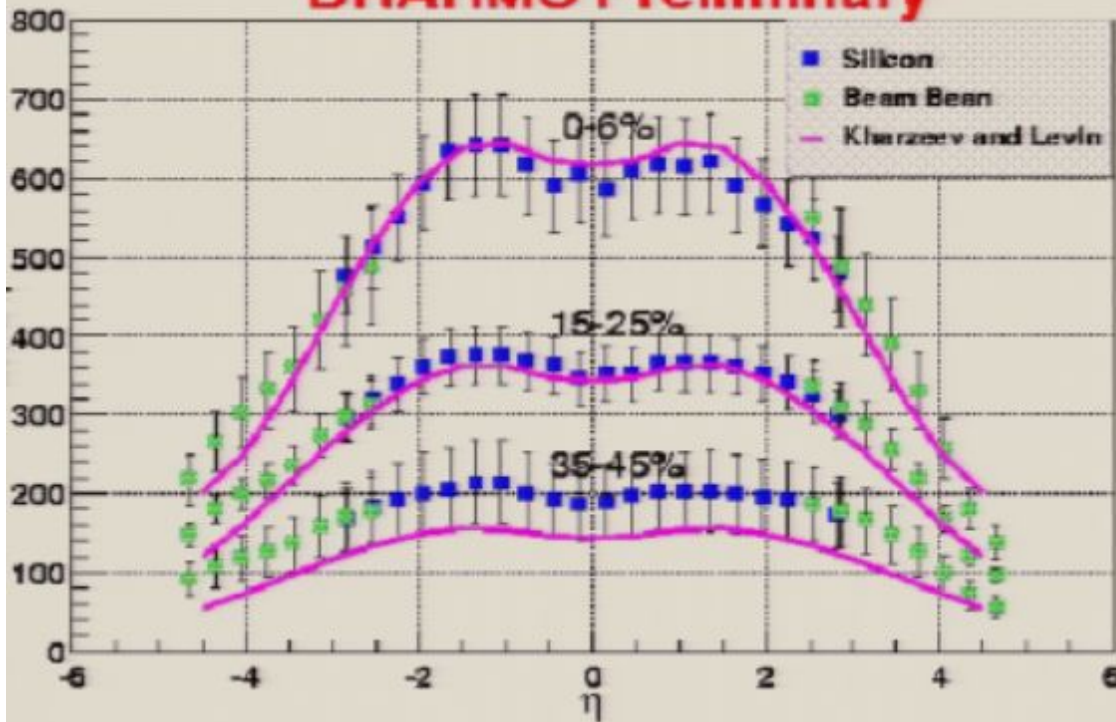
The data on hadron multiplicities in Au-Au and d-Au collisions support the quasi-classical picture



Classical QCD in action

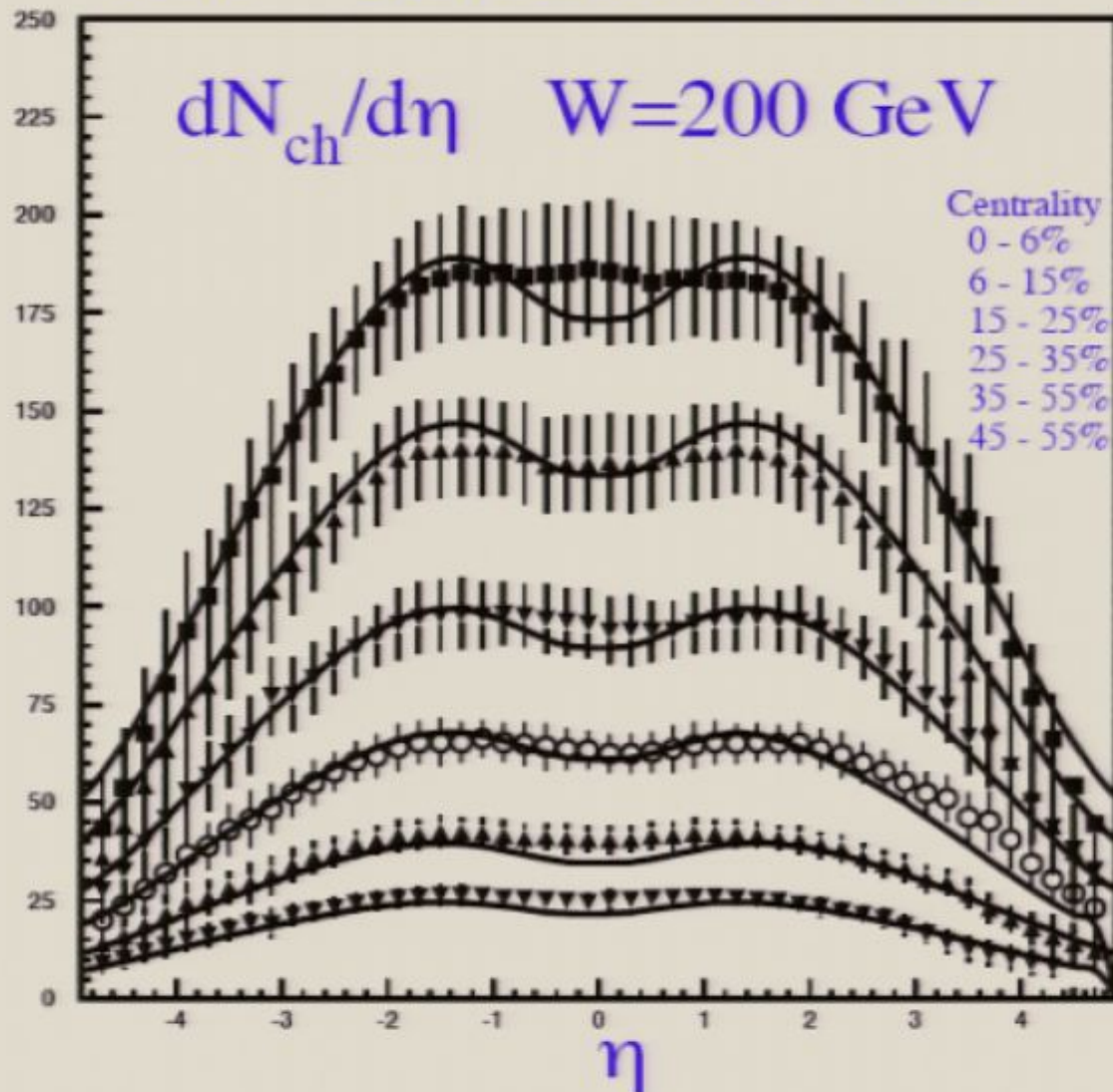
The data on hadron multiplicities in Au-Au and d-Au collisions support the semi-classical picture

BRAHMS Preliminary



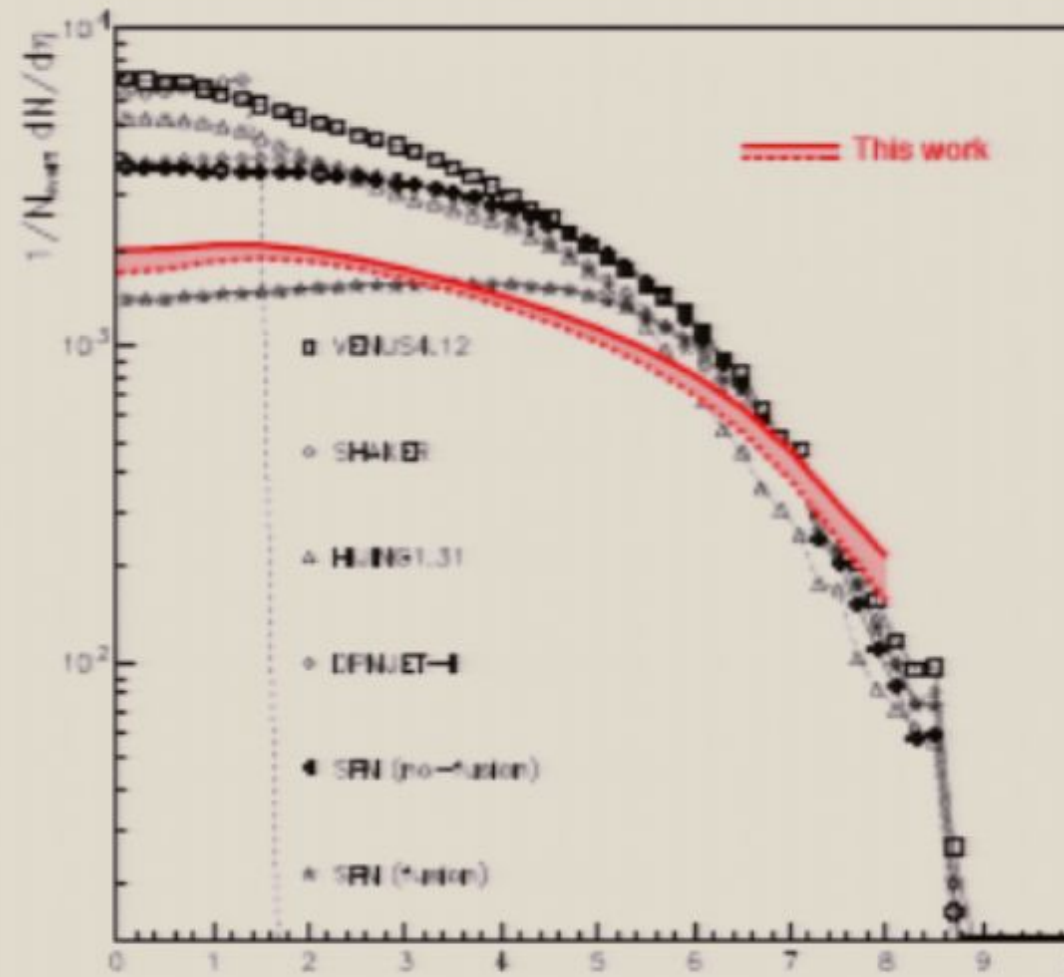
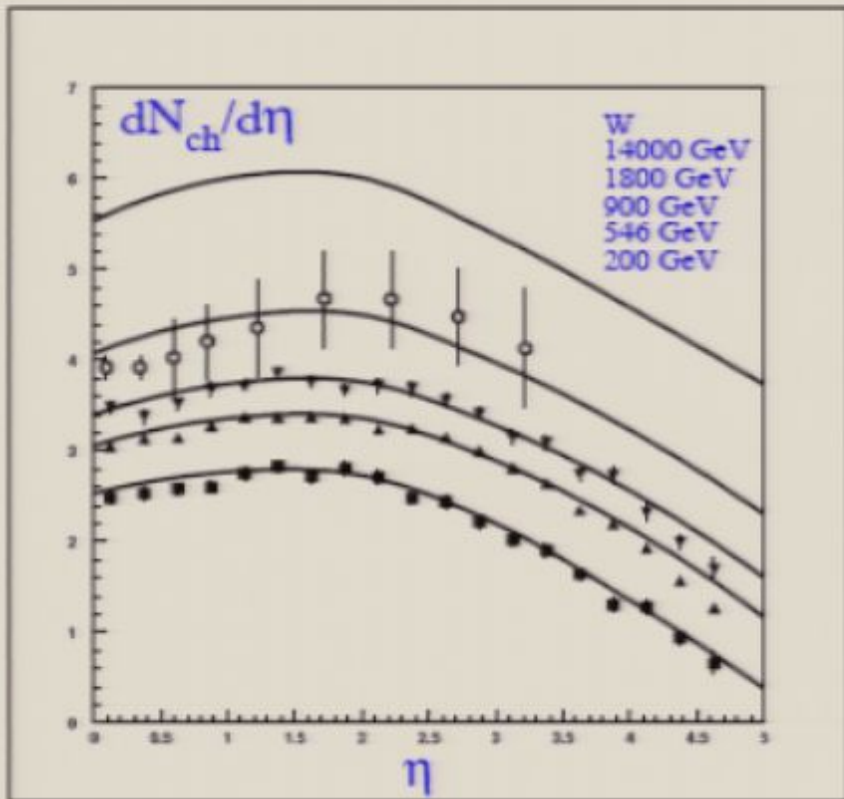
First look at the CuCu data

Cu - Cu collisions



KLN

Predictions for the LHC



KLN, hep-ph/0408050

How dense is the produced matter?

The initial energy density achieved:

$$\epsilon_{initial} \simeq \frac{\langle k_t \rangle}{\tau_0} \frac{d^2 N}{d^2 b d\eta} \simeq Q_s^2 \frac{d^2 N}{d^2 b d\eta} \simeq 18 \text{ GeV}/\text{fm}^3$$

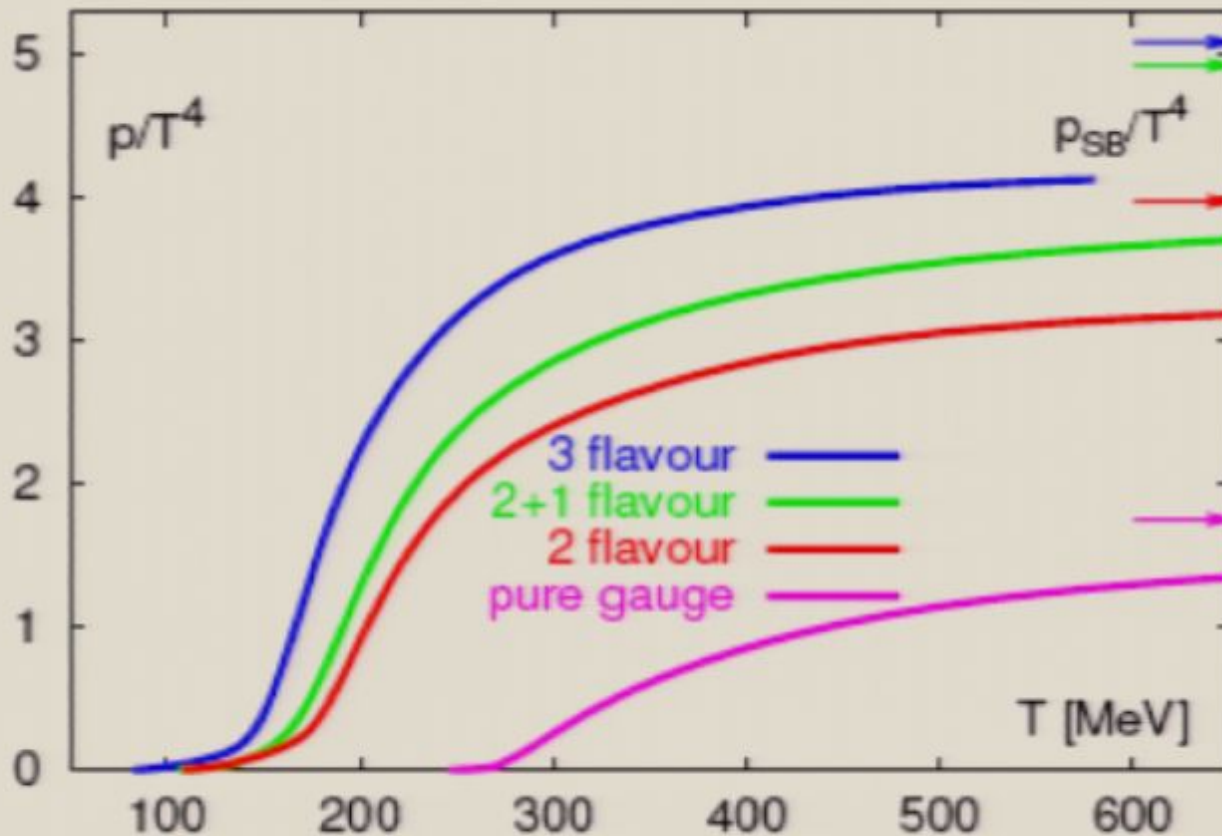
mean
transverse
momentum
of produced
gluons

gluon
formation
time

the density
of the gluons
in the transverse
plane and in
rapidity

about
100 times
nuclear
density !

What happens at such energy densities?



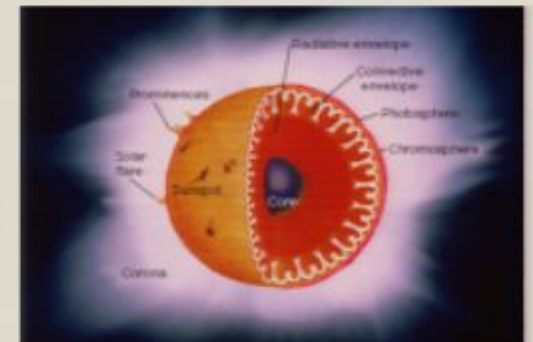
Phase transitions:

deconfinement

Chiral symmetry restoration

$U_A(1)$ restoration

Data from lattice QCD simulations F. Karsch et al



critical temperature $\sim 10^{12}$ K; cf temperature inside the Sun $\sim 10^7$ K

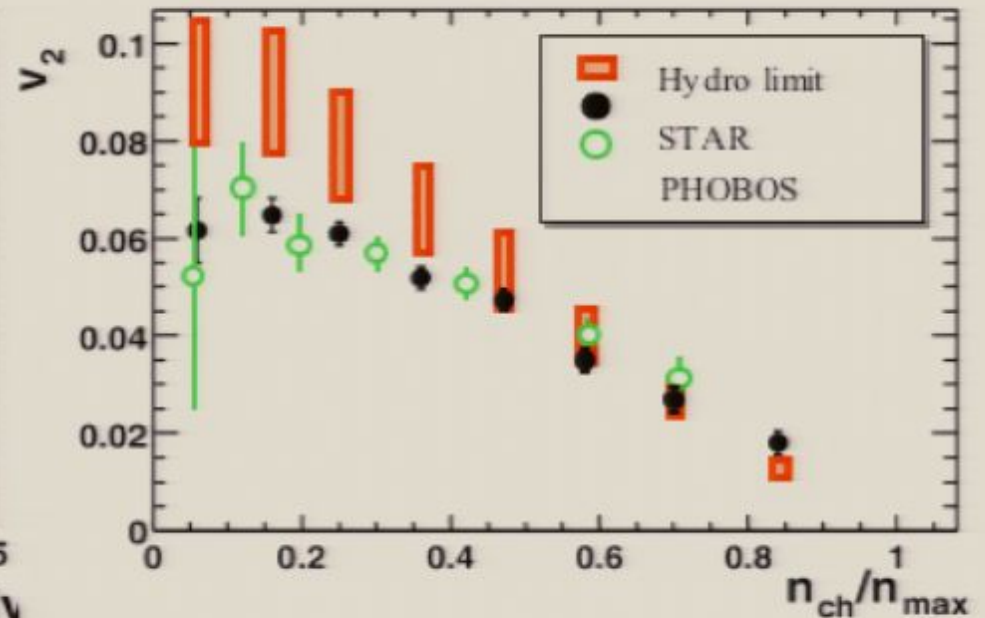
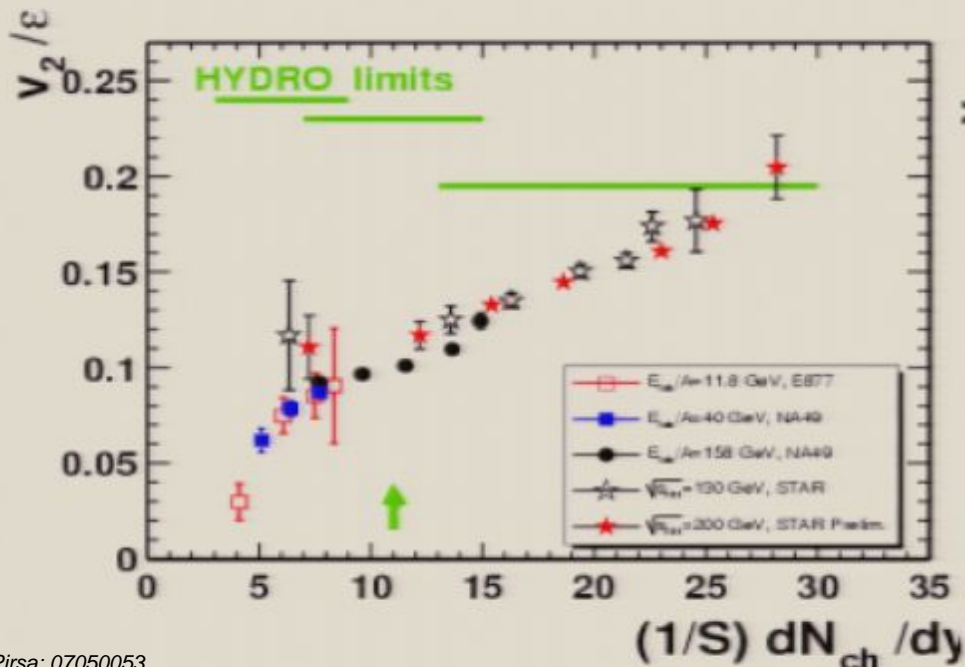
Do we see the QCD matter at RHIC ?

Collective flow =>

Au-Au collisions at RHIC produce strongly interacting matter

shear viscosity - to - entropy ratio

hydrodynamics: QCD liquid is **more fluid than water**

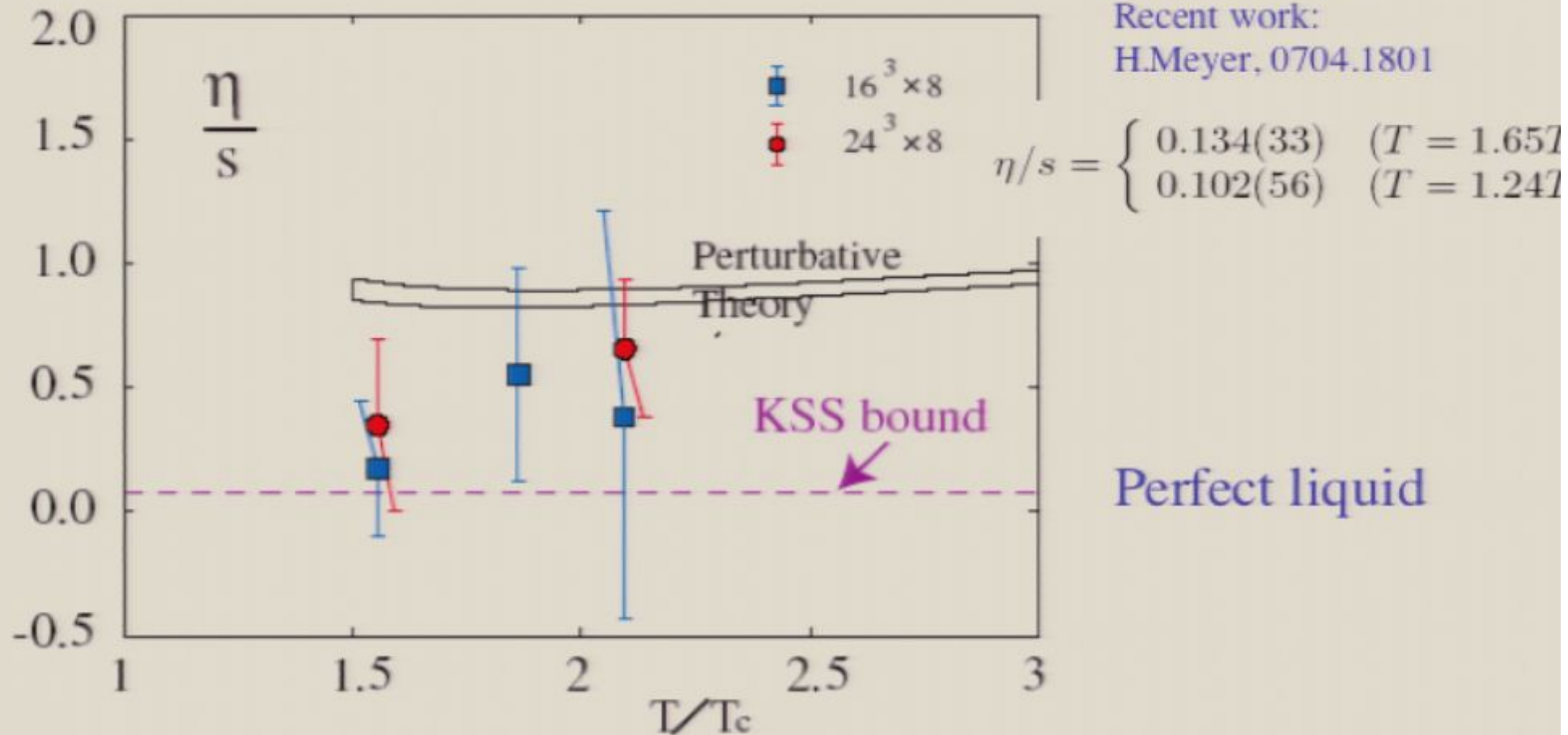


Shear viscosity of sQGP

A.Nakamura and S.Sakai,
 hep-lat/0406009;

Recent work:

H.Meyer, 0704.1801

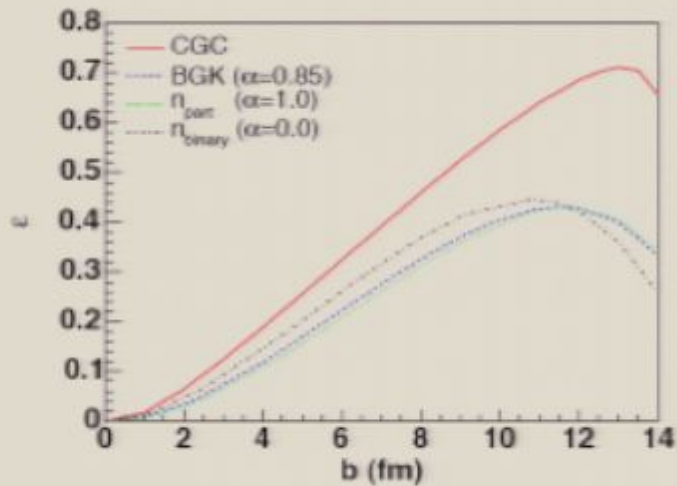


Kovtun - Son - Starinets bound:

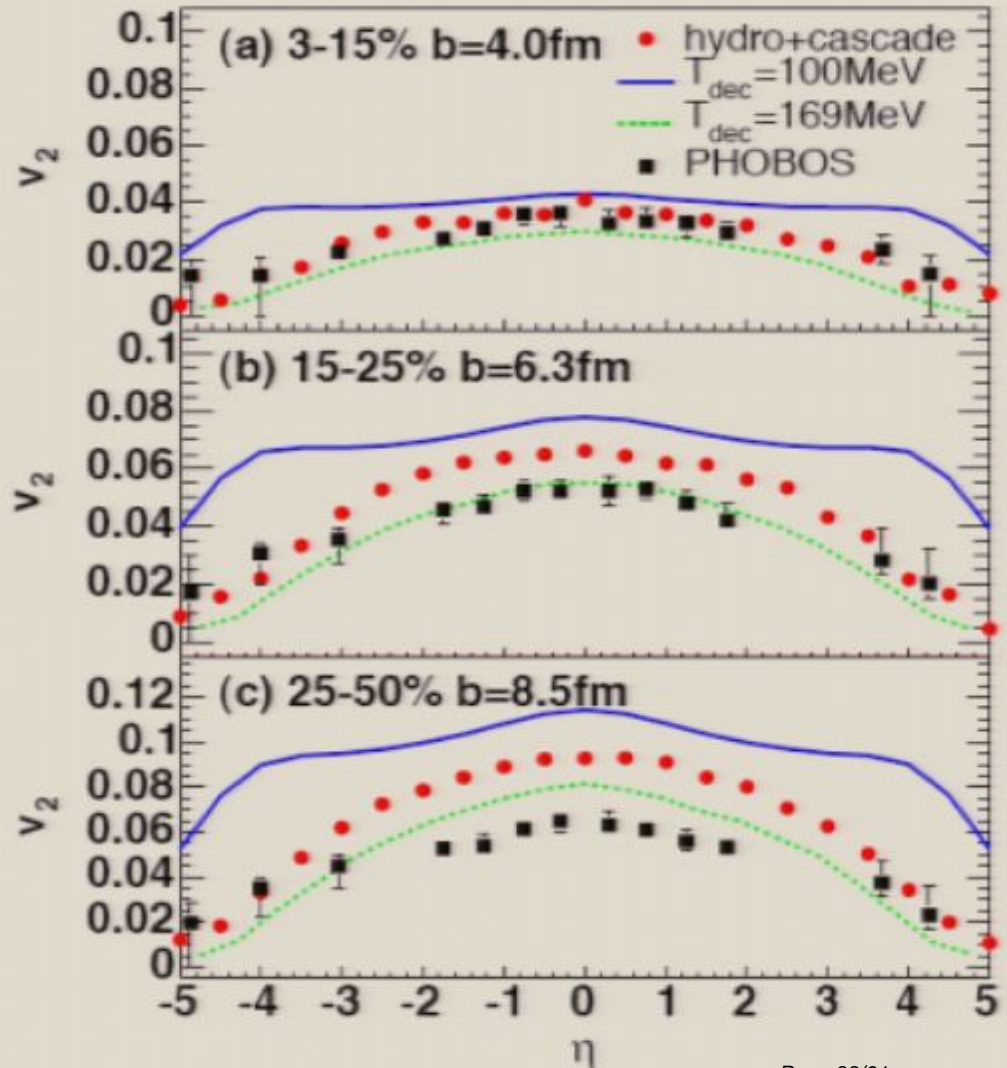
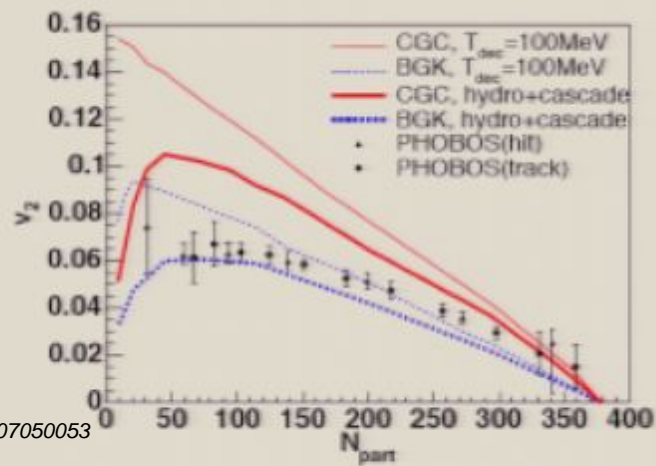
strongly coupled SUSY QCD = classical supergravity

How small really is the viscosity?

CGC initial conditions lead to larger ellipticity,



require some viscous effects:



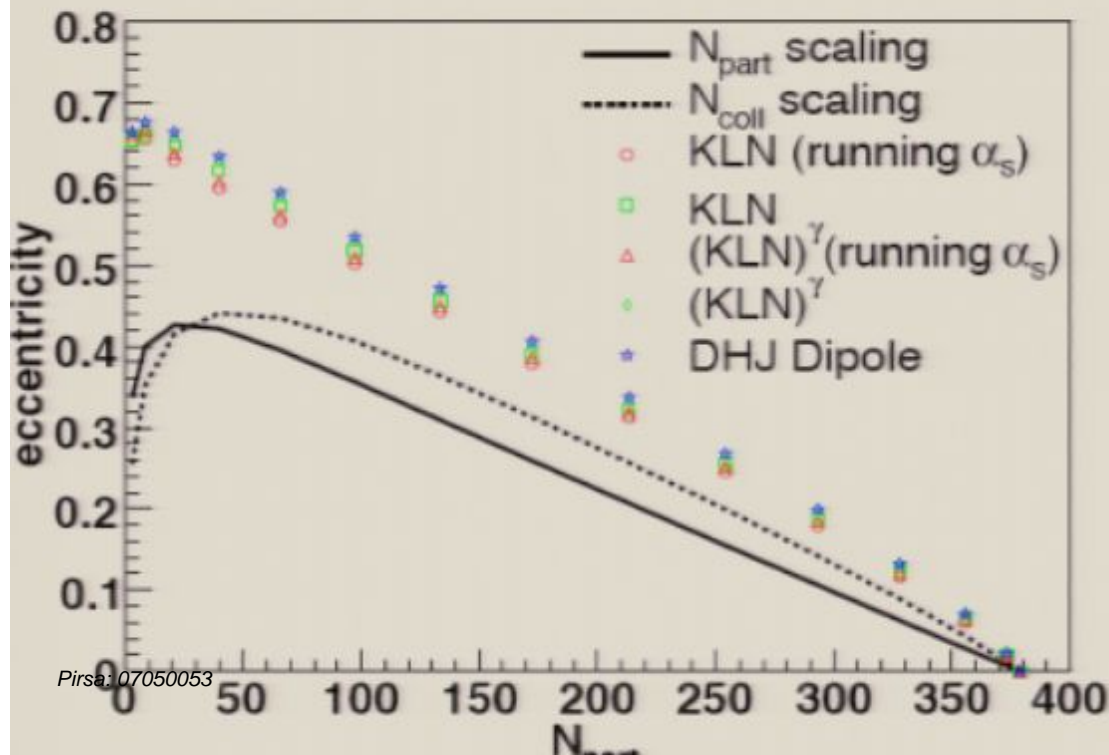
How small is really the viscosity?

KLN initial conditions lead to larger ellipticity,

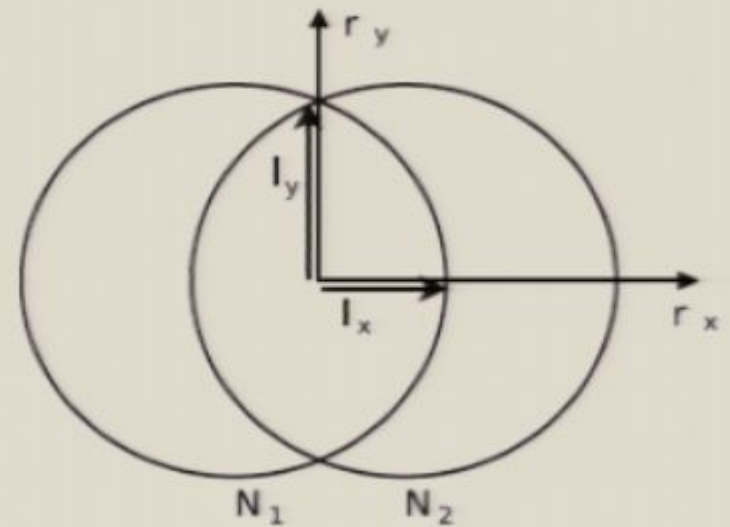
T.Hirano, U.Heinz, DK, R.Lacey, Y. Nara, hep-ph/0511046

this is not an artifact of a particular model for the gluon distribution, but a generic feature of saturation

H.Drescher, A.Dumitru, A.Hayashigaki, Y.Nara, nucl-th/0605012



Pirsa: 07050053

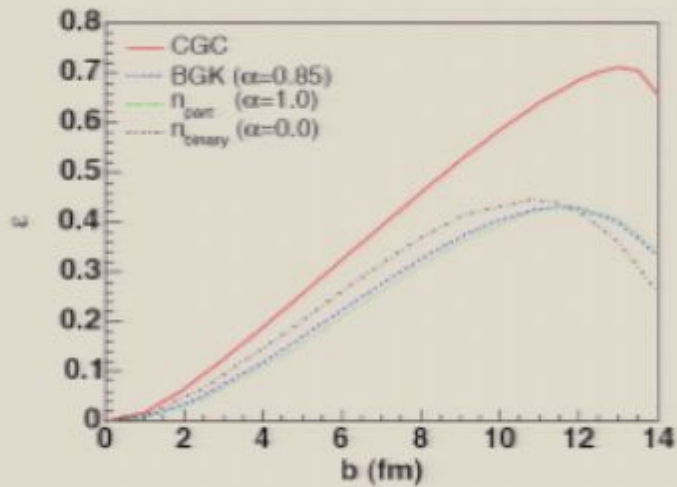


$$n \sim Q_{s,min}^2 \ln \left(\frac{Q_{s,max}^2}{Q^2} \right)$$

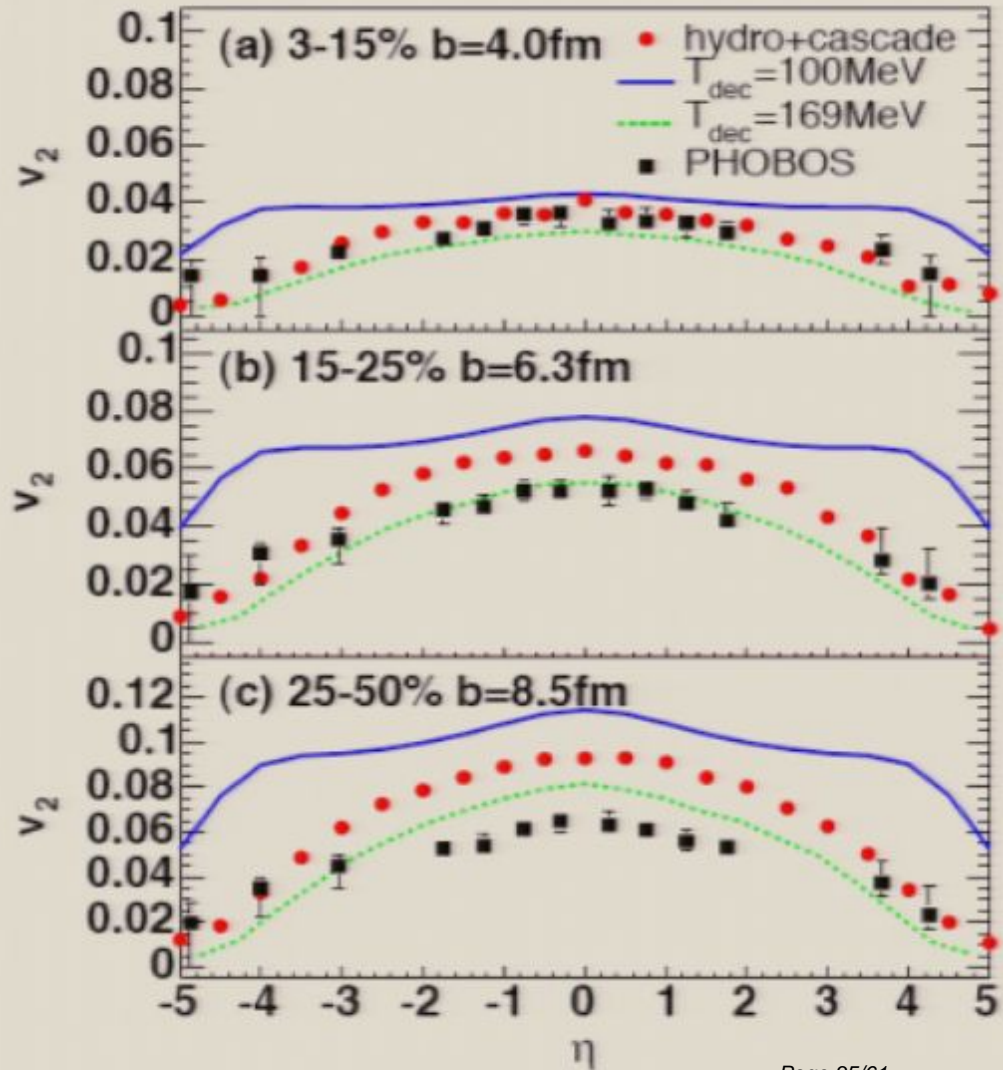
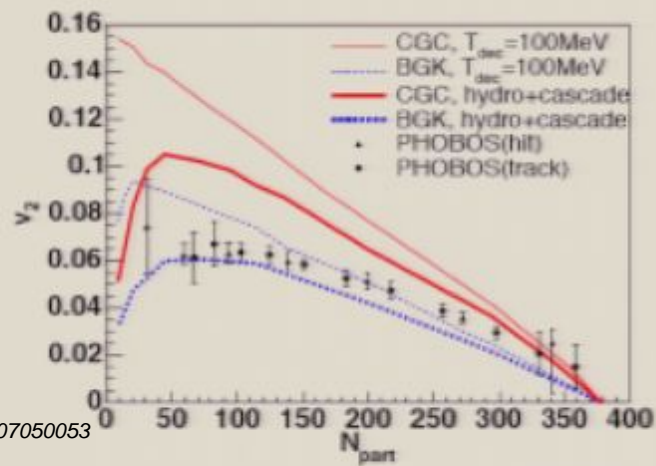
Page 34/61

How small really is the viscosity?

CGC initial conditions lead to larger ellipticity,



require some viscous effects:



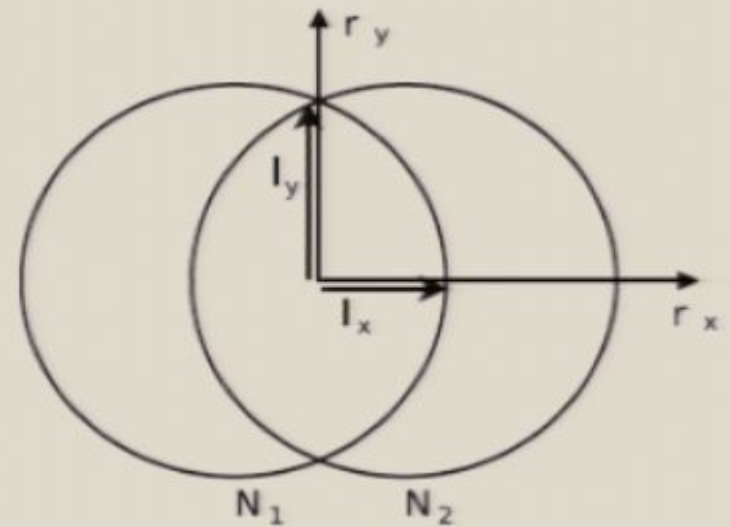
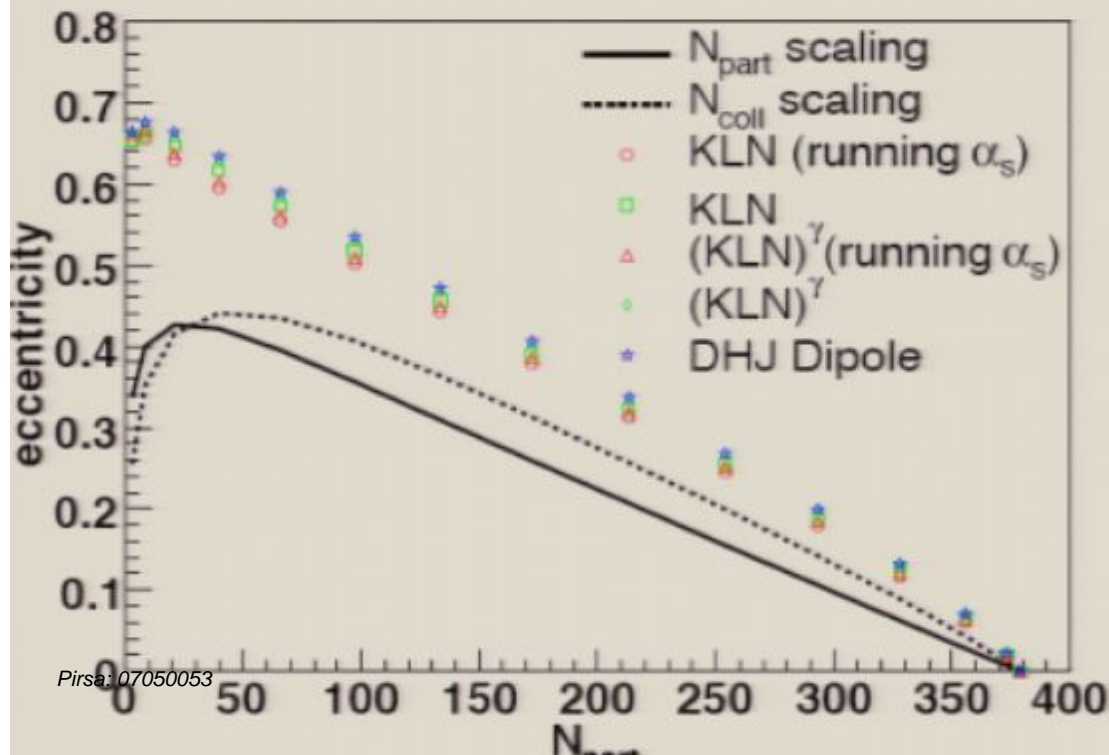
How small is really the viscosity?

KLN initial conditions lead to larger ellipticity,

T.Hirano, U.Heinz, DK, R.Lacey, Y. Nara, hep-ph/0511046

this is not an artifact of a particular model for the gluon distribution, but a generic feature of saturation

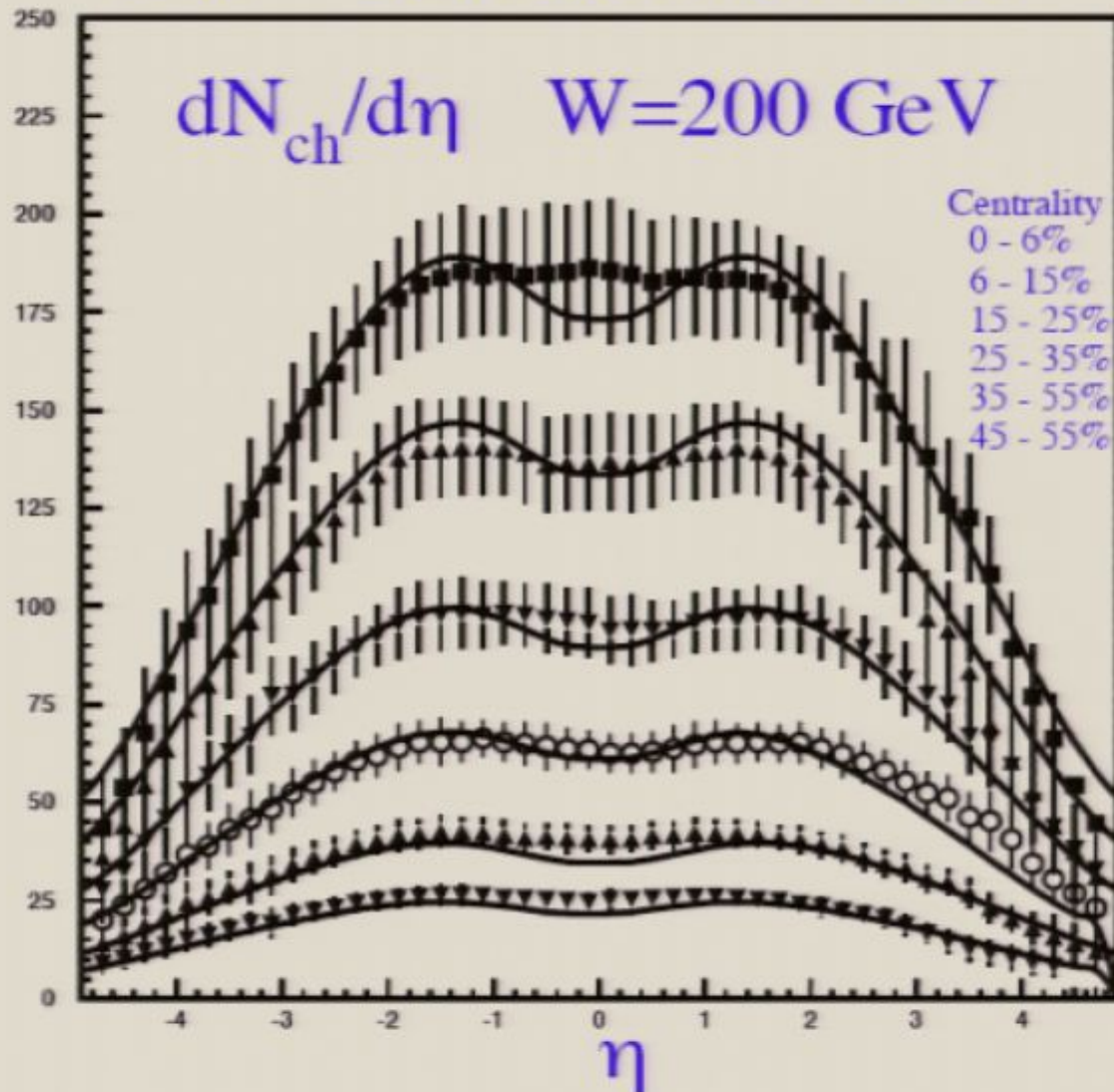
H.Drescher, A.Dumitru, A.Hayashigaki, Y.Nara, nucl-th/0605012



$$n \sim Q_{s,min}^2 \ln \left(\frac{Q_{s,max}^2}{Q^2} \right)$$

First look at the CuCu data

Cu - Cu collisions

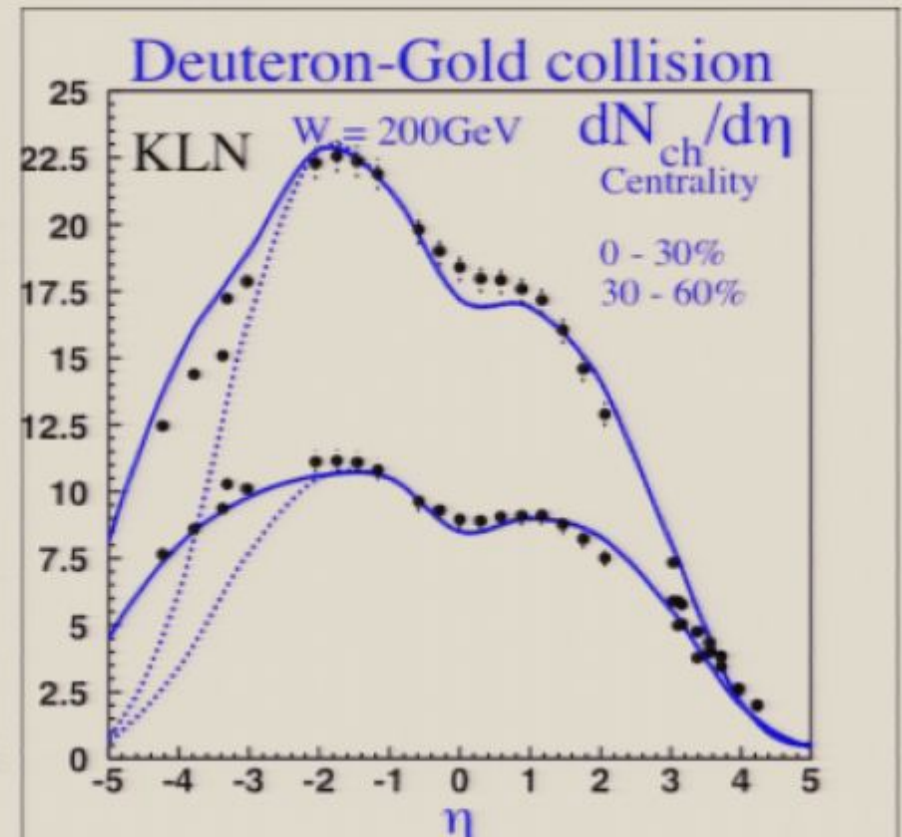
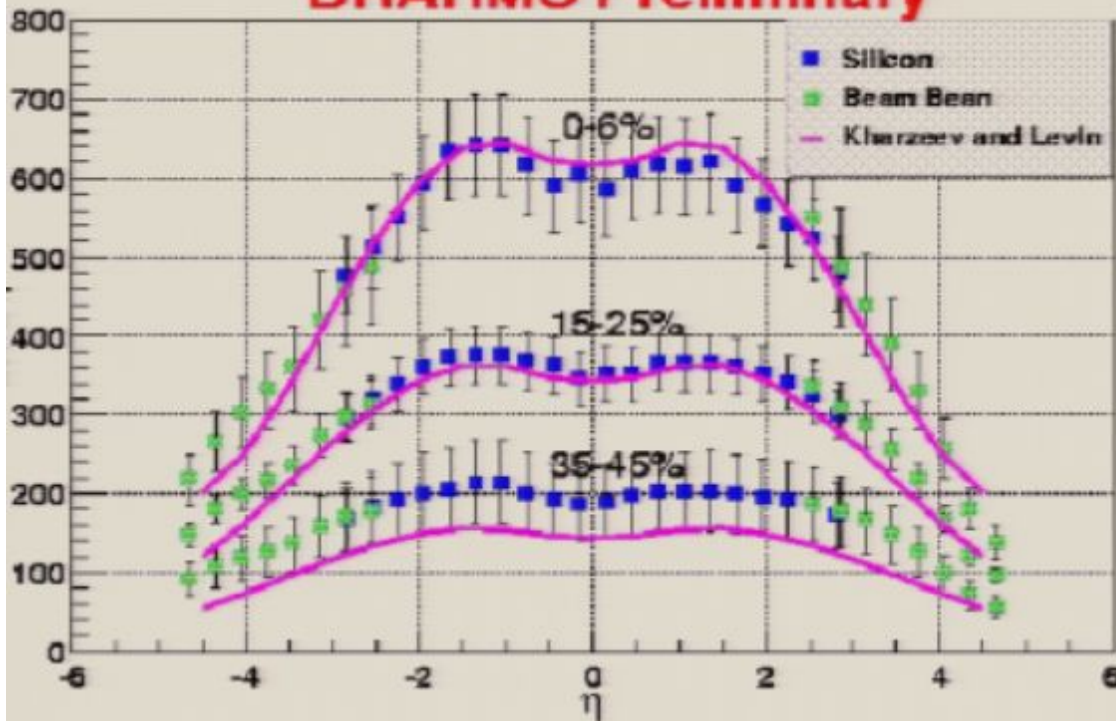


KLN

Classical QCD in action

The data on hadron multiplicities in Au-Au and d-Au collisions support the semi-classical picture

BRAHMS Preliminary



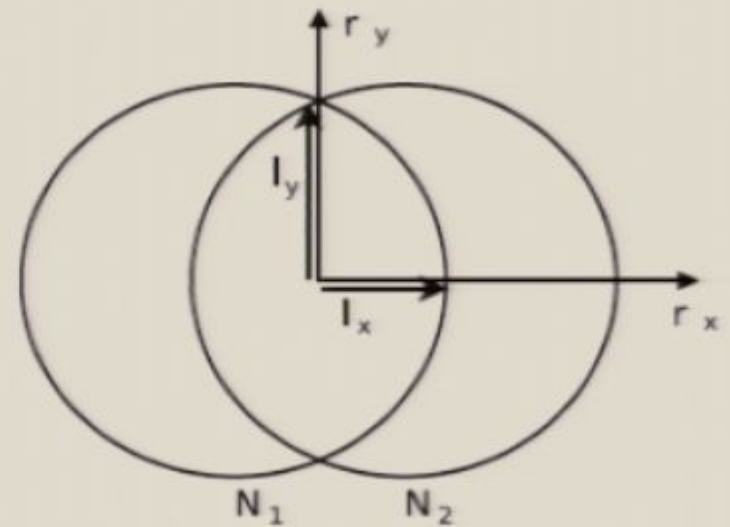
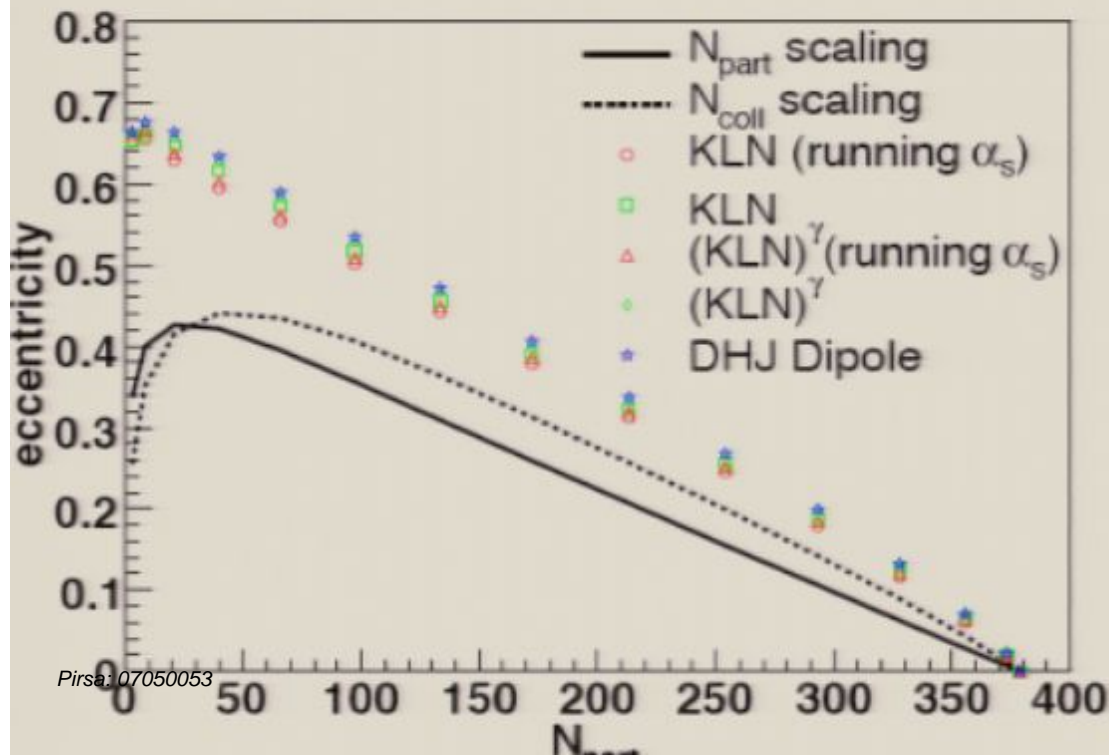
How small is really the viscosity?

KLN initial conditions lead to larger ellipticity,

T.Hirano, U.Heinz, DK, R.Lacey, Y. Nara, hep-ph/0511046

this is not an artifact of a particular model for the gluon distribution, but a generic feature of saturation

H.Drescher, A.Dumitru, A.Hayashigaki, Y.Nara, nucl-th/0605012



$$n \sim Q_{s,min}^2 \ln \left(\frac{Q_{s,max}^2}{Q^2} \right)$$

Low energy theorems and bulk viscosity

Sum rule for the spectral density:

$$2 \int_0^\infty \frac{\rho(u, \vec{0})}{u} du = - \left(4 - T \frac{\partial}{\partial T} \right) \langle \theta \rangle_T = T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v|$$

Using ansatz

we get

$$\frac{\rho(\omega, \vec{0})}{\omega} = \frac{9\zeta}{\pi} \frac{\omega_0^2}{\omega_0^2 + \omega^2} \quad \zeta = \frac{1}{9\omega_0} \left\{ T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v| \right\}$$

Low energy theorems and bulk viscosity

Sum rule for the spectral density:

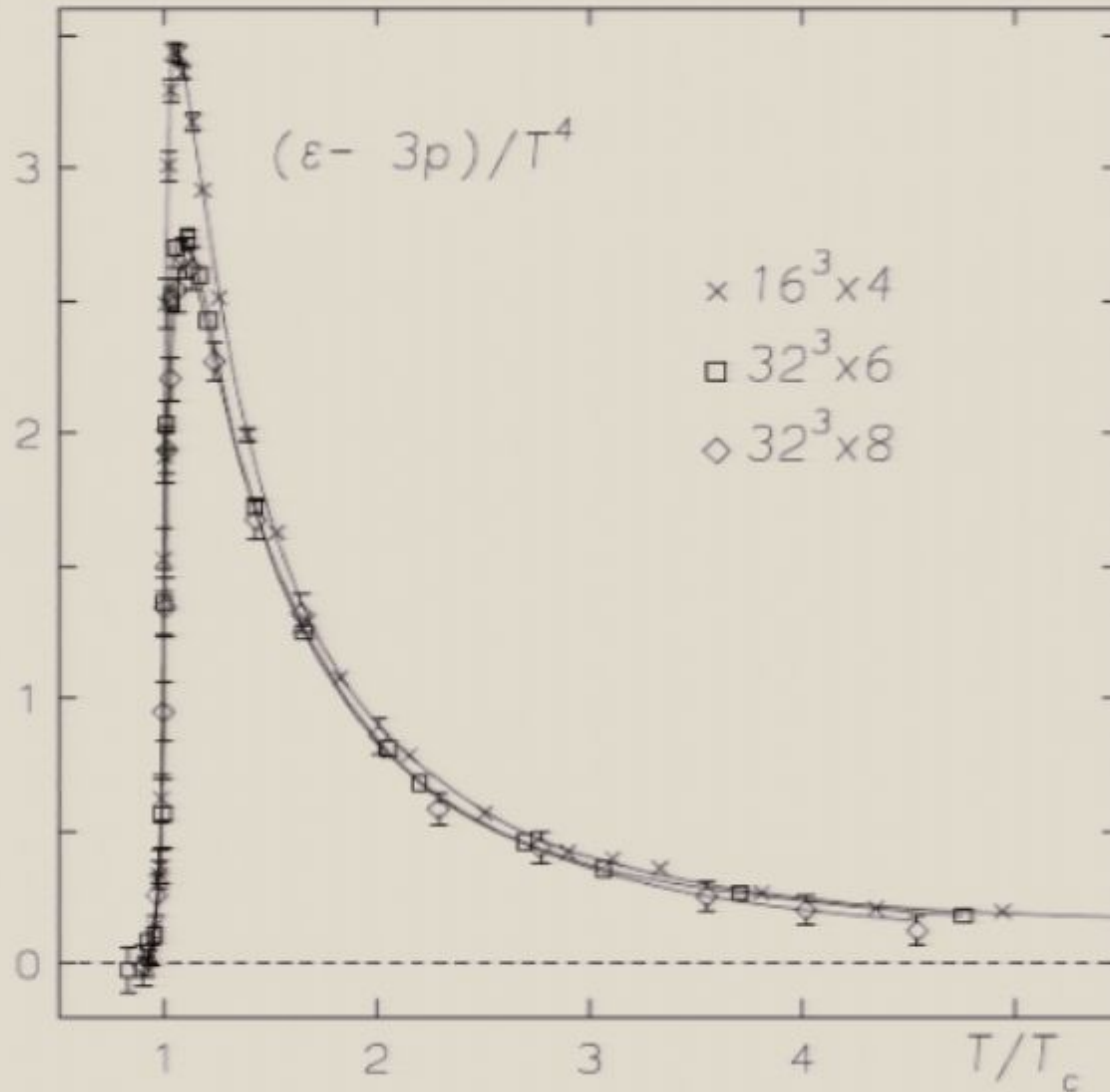
$$2 \int_0^\infty \frac{\rho(u, \vec{0})}{u} du = - \left(4 - T \frac{\partial}{\partial T} \right) \langle \theta \rangle_T = T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v|$$

Using ansatz

we get

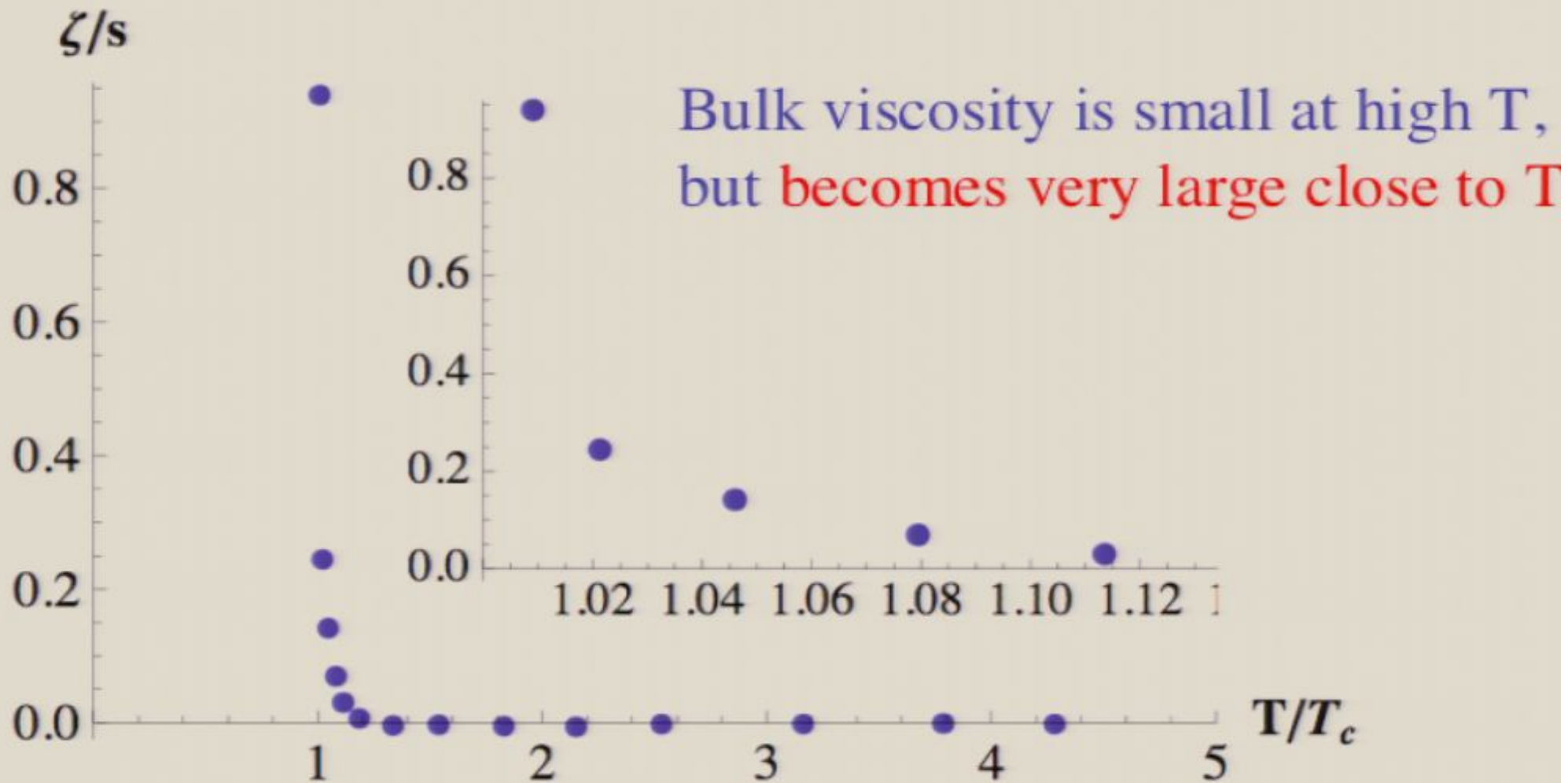
$$\frac{\rho(\omega, \vec{0})}{\omega} = \frac{9\zeta}{\pi} \frac{\omega_0^2}{\omega_0^2 + \omega^2} \quad \zeta = \frac{1}{9\omega_0} \left\{ T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v| \right\}$$

SU(3),
pure gauge

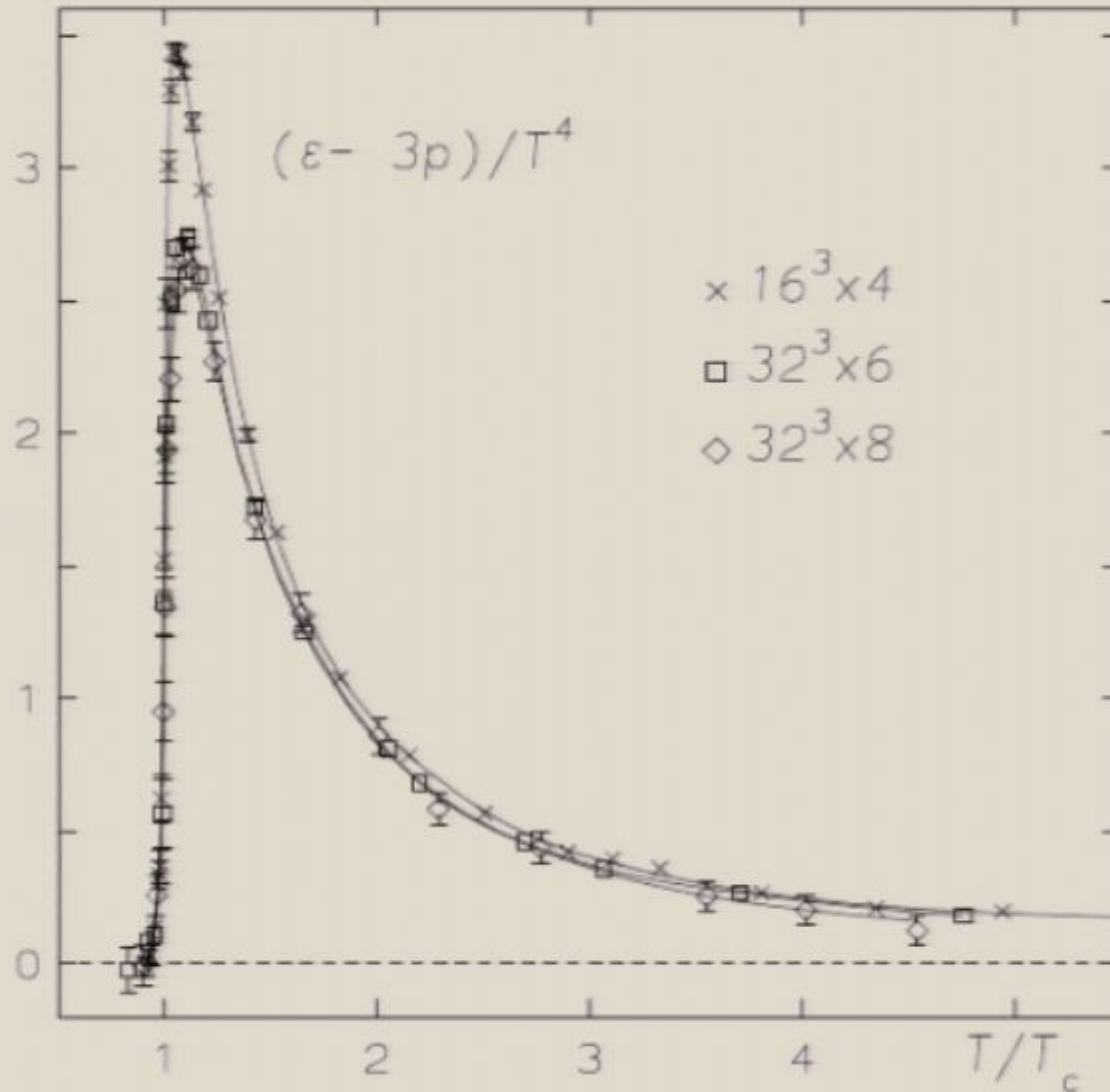


Use the lattice data from G.Boyd, J.Engels, F.Karsch, E.Laermann,
C.Legeland, M.Lutgeimer, B.Petersson, hep-lat/9602007

The result

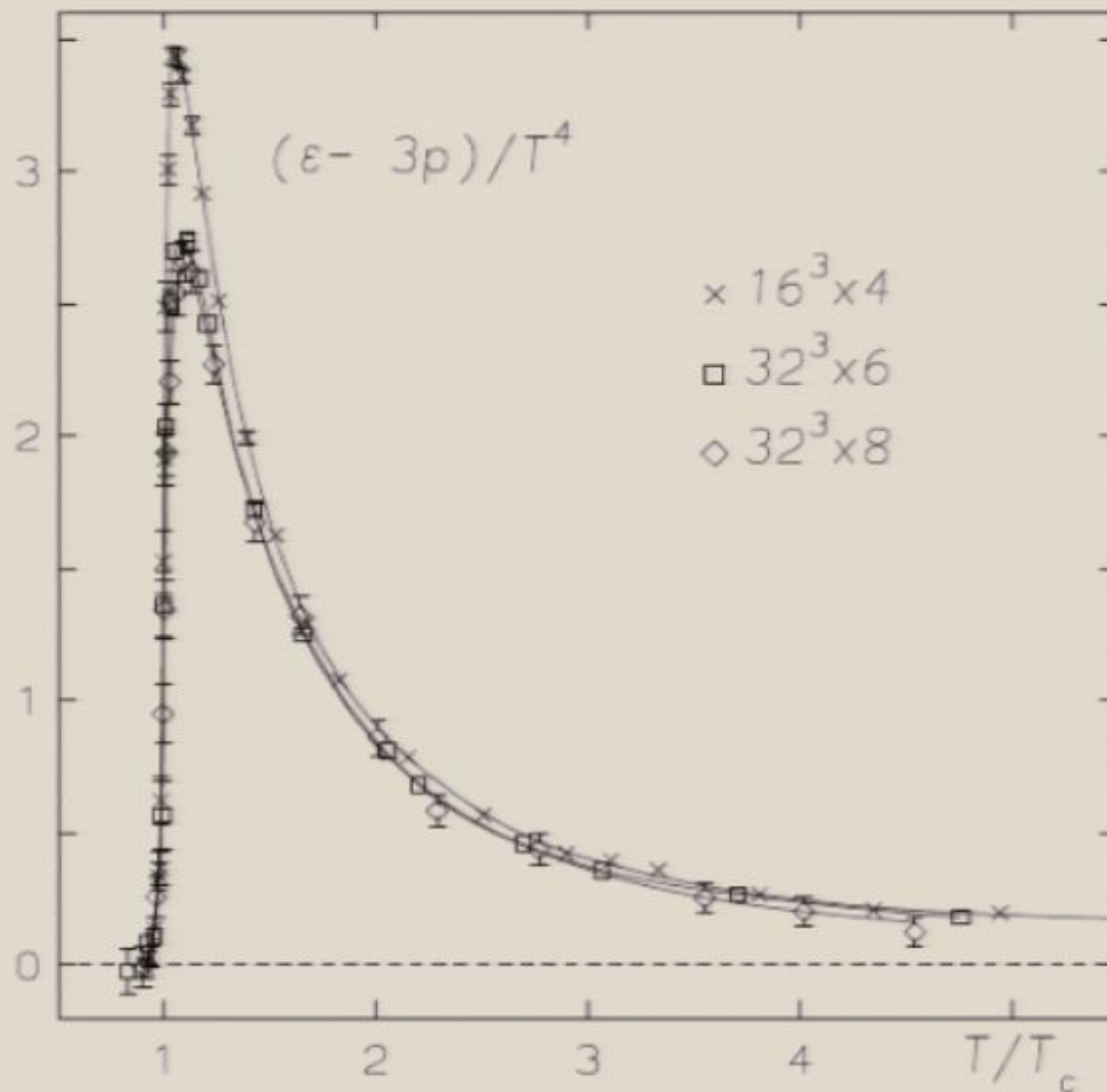


SU(3),
pure gauge



Use the lattice data from G.Boyd, J.Engels, F.Karsch, E.Laermann,
C.Legeland, M.Lutgeimer, B.Petersson, hep-lat/9602007

SU(3),
pure gauge



Use the lattice data from G.Boyd, J.Engels, F.Karsch, E.Laermann,
C.Legeland, M.Lutgeimer, B.Petersson, hep-lat/9602007

Low energy theorems and bulk viscosity

Sum rule for the spectral density:

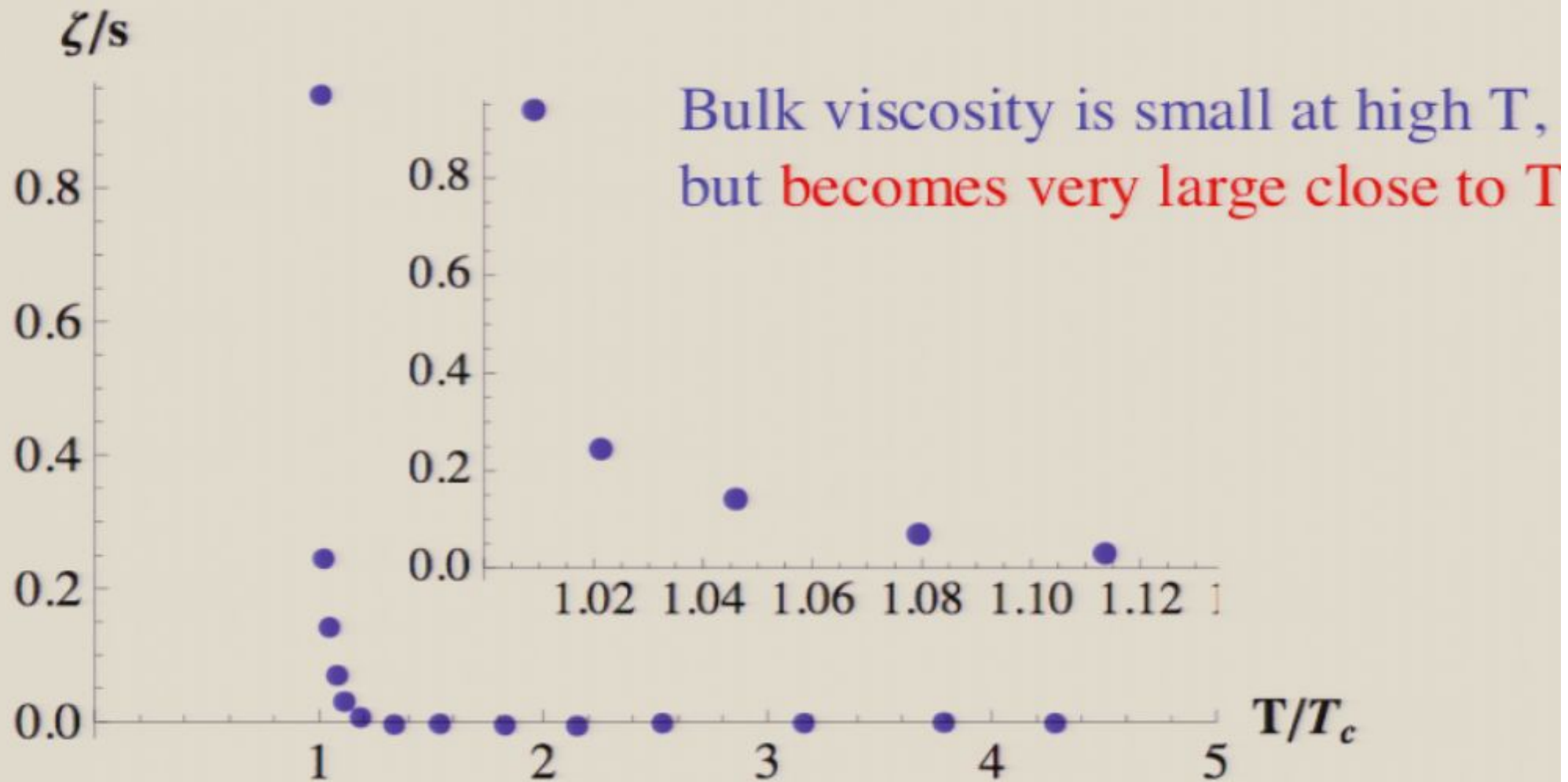
$$2 \int_0^\infty \frac{\rho(u, \vec{0})}{u} du = - \left(4 - T \frac{\partial}{\partial T} \right) \langle \theta \rangle_T = T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v|$$

Using ansatz

we get

$$\frac{\rho(\omega, \vec{0})}{\omega} = \frac{9\zeta}{\pi} \frac{\omega_0^2}{\omega_0^2 + \omega^2} \quad \zeta = \frac{1}{9\omega_0} \left\{ T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v| \right\}$$

The result



Low energy theorems and bulk viscosity

Sum rule for the spectral density:

$$2 \int_0^\infty \frac{\rho(u, \vec{0})}{u} du = - \left(4 - T \frac{\partial}{\partial T} \right) \langle \theta \rangle_T = T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v|$$

Using ansatz

we get

$$\frac{\rho(\omega, \vec{0})}{\omega} = \frac{9\zeta}{\pi} \frac{\omega_0^2}{\omega_0^2 + \omega^2} \quad \zeta = \frac{1}{9\omega_0} \left\{ T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v| \right\}$$

Can we say anything about non-perturbative effects?

At zero temperature, broken scale invariance leads to a chain of low-energy theorems for the correlation functions of $\partial^\mu s_\mu = \theta^\mu_\mu$

Novikov, Shifman,
Vainshtein, Zakharov '81

(elegant geometrical interpretation - classical theory in a curved background)

Migdal, Shifman '82;
DK, Levin, Tulin '04

These theorems have been generalized to finite T:

$$G^E(0, \vec{0}) = \int d^4x \langle T\theta(x), \theta(0) \rangle = \left(T \frac{\partial}{\partial T} - 4 \right) \langle \theta \rangle_T$$

Ellis, Kapusta,
Tang '98

What is the role of broken scale invariance in the transport properties of QCD plasma?

Consider the bulk viscosity:

$$T_{ij} = P_{\text{eq}}(\epsilon) \delta_{ij} - \eta \left(\partial_i u_j + \partial_j u_i - \frac{2}{3} \delta_{ij} \partial_k u_k \right) - \zeta \delta_{ij} \nabla \cdot \mathbf{u}$$

it can be computed through

$$\zeta = \frac{1}{9} \lim_{\omega \rightarrow 0} \frac{1}{\omega} \int_0^\infty dt \int d^3r e^{i\omega t} \langle [\theta_\mu^\mu(x), \theta_\mu^\mu(0)] \rangle$$

Since $\partial^\mu s_\mu = \theta_\mu^\mu$ is a measure of deviation from conformal invariance, bulk viscosity characterizes the importance of scale anomaly

Can we say anything about non-perturbative effects?

At zero temperature, broken scale invariance leads to a chain of low-energy theorems for the correlation functions of $\partial^\mu S_\mu = \theta^\mu_\mu$

Novikov, Shifman,
Vainshtein, Zakharov '81

(elegant geometrical interpretation - classical theory in a curved background)

Migdal, Shifman '82;
DK, Levin, Tulin '04

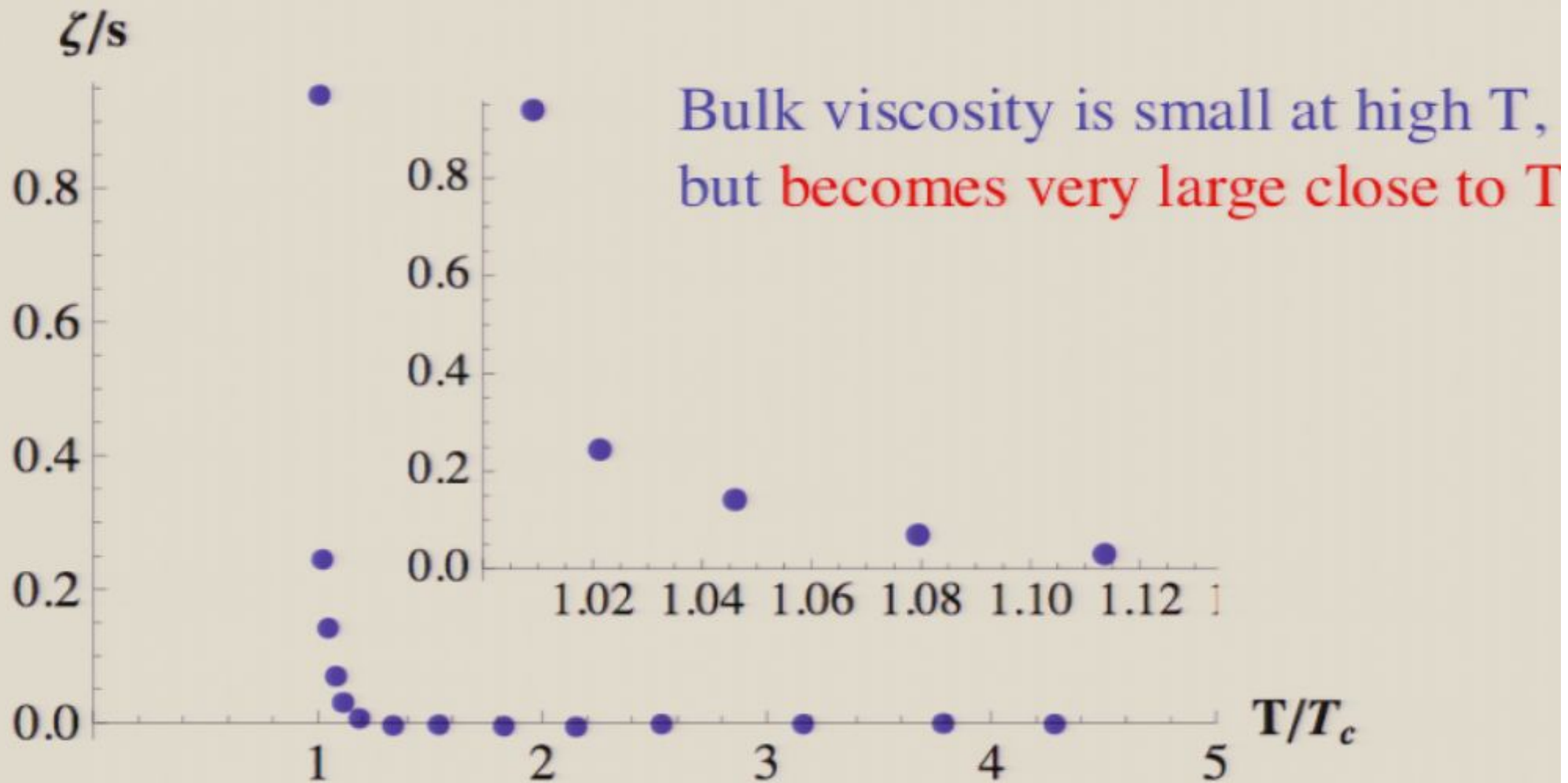
These theorems have been generalized to finite T:

$$G^E(0, \vec{0}) = \int d^4x \langle T\theta(x), \theta(0) \rangle = \left(T \frac{\partial}{\partial T} - 4 \right) \langle \theta \rangle_T$$

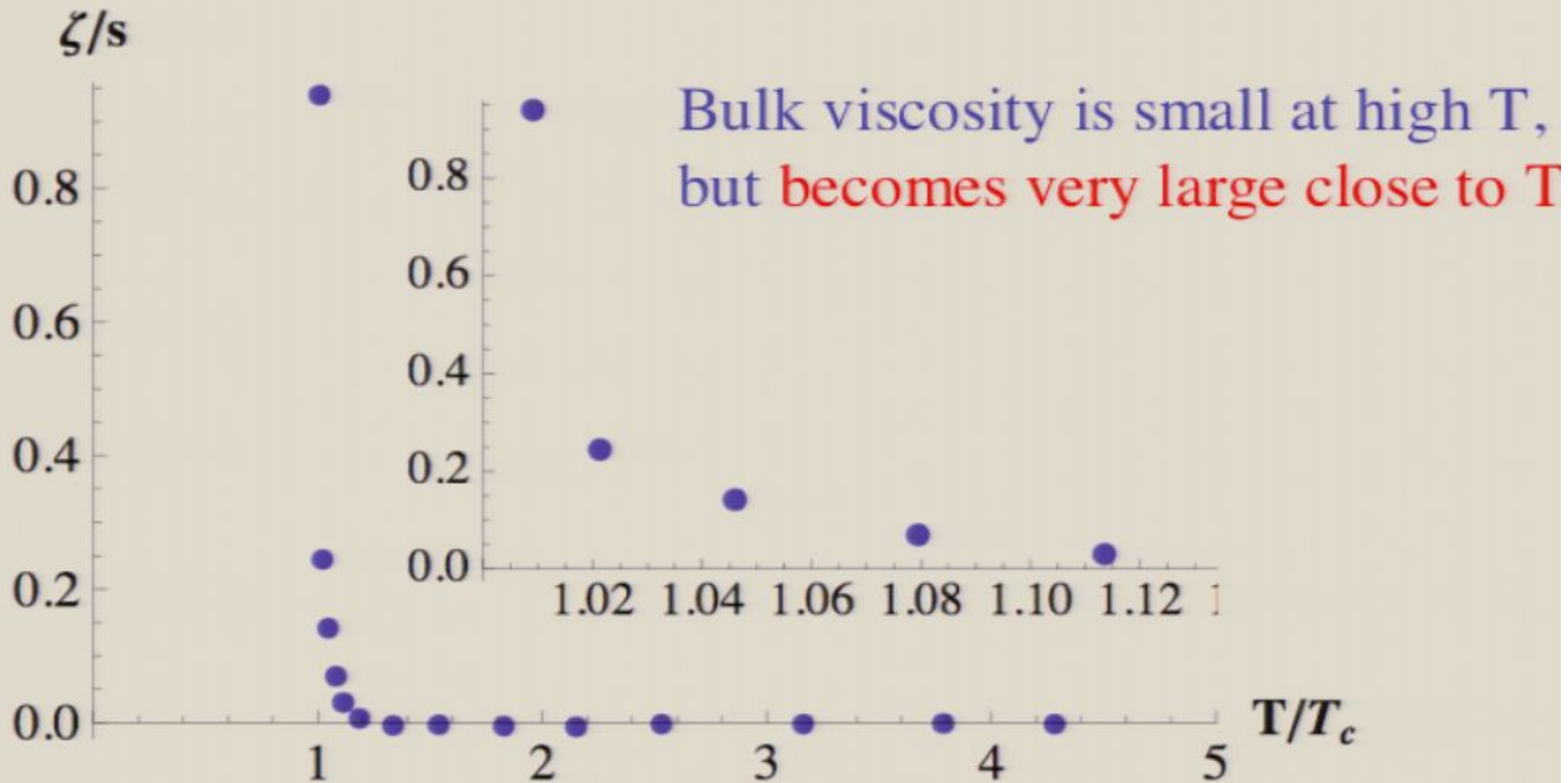
Ellis, Kapusta,
Tang '98



The result



The result



Low energy theorems and bulk viscosity

Sum rule for the spectral density:

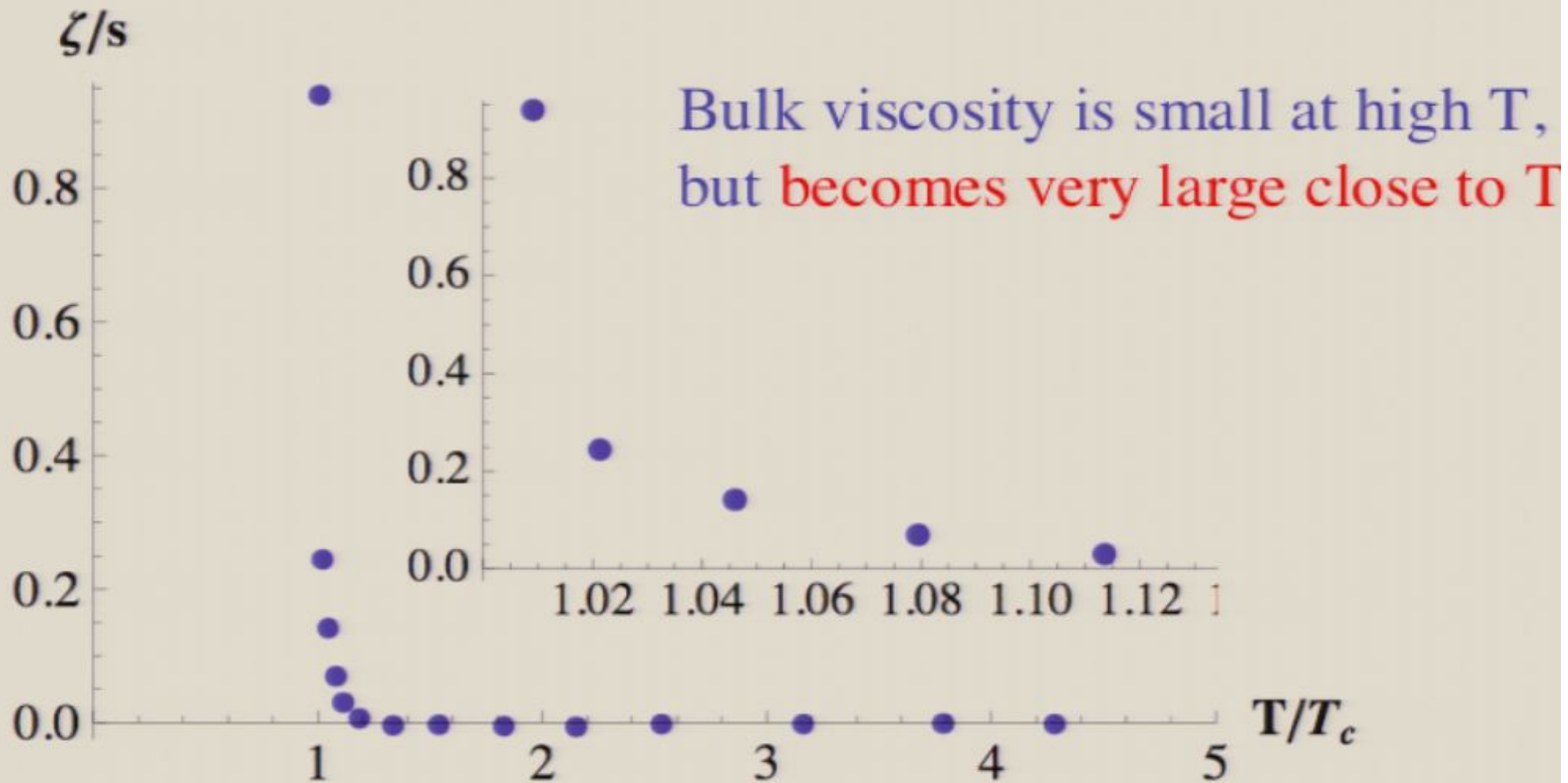
$$2 \int_0^\infty \frac{\rho(u, \vec{0})}{u} du = - \left(4 - T \frac{\partial}{\partial T} \right) \langle \theta \rangle_T = T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v|$$

Using ansatz

we get

$$\frac{\rho(\omega, \vec{0})}{\omega} = \frac{9\zeta}{\pi} \frac{\omega_0^2}{\omega_0^2 + \omega^2} \quad \zeta = \frac{1}{9\omega_0} \left\{ T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v| \right\}$$

The result



Low energy theorems and bulk viscosity

Sum rule for the spectral density:

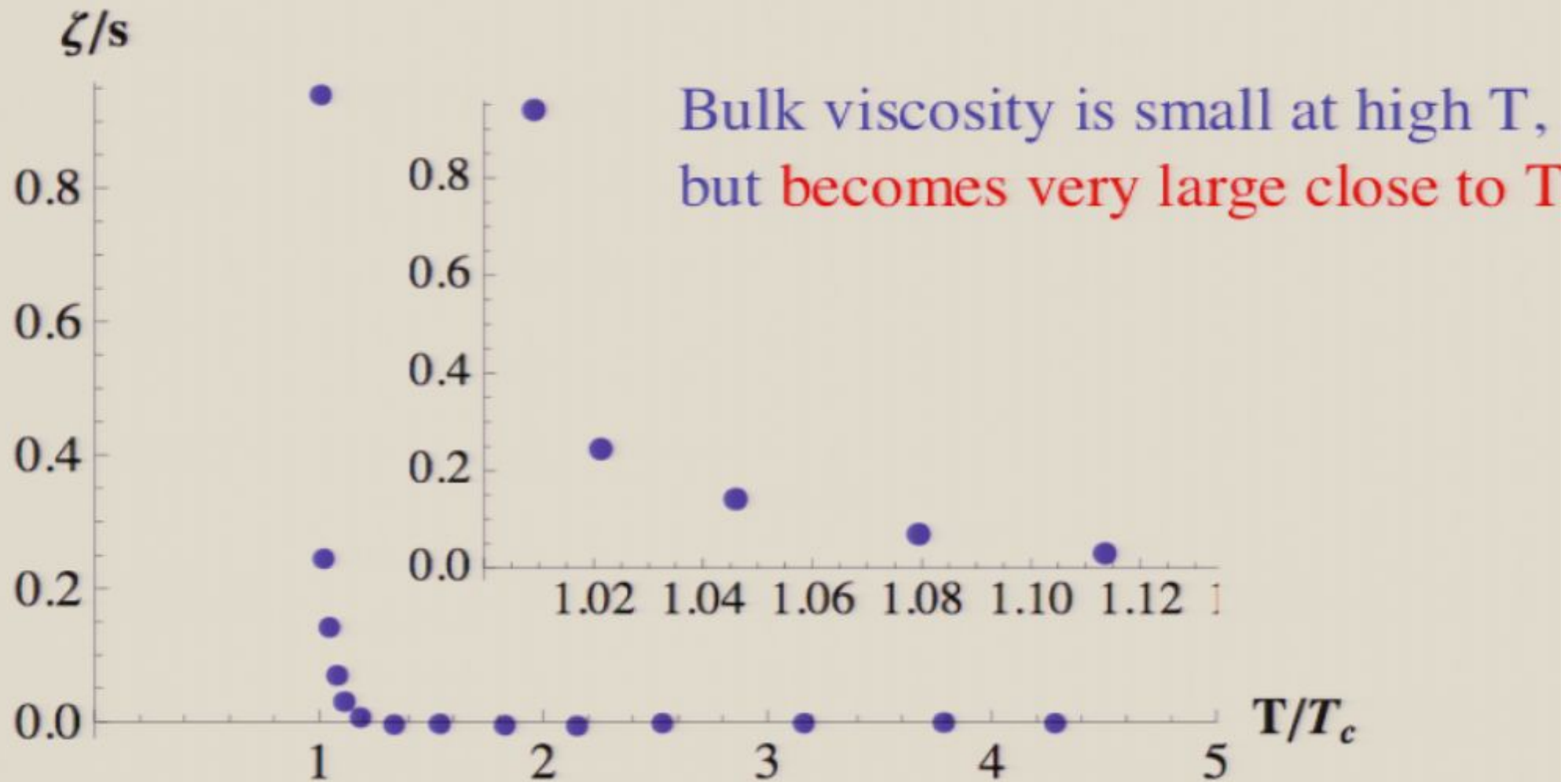
$$2 \int_0^\infty \frac{\rho(u, \vec{0})}{u} du = - \left(4 - T \frac{\partial}{\partial T} \right) \langle \theta \rangle_T = T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v|$$

Using ansatz

we get

$$\frac{\rho(\omega, \vec{0})}{\omega} = \frac{9\zeta}{\pi} \frac{\omega_0^2}{\omega_0^2 + \omega^2} \quad \zeta = \frac{1}{9\omega_0} \left\{ T^5 \frac{\partial}{\partial T} \frac{(\mathcal{E} - 3P)_{\text{LAT}}}{T^4} + 16|\epsilon_v| \right\}$$

The result



Summary

Broken scale invariance of QCD manifests itself in the bulk behavior of quark-gluon matter in at least two ways:

1. Initial conditions
2. Bulk viscosity close to T_c

Summary

Broken scale invariance of QCD manifests itself in the bulk behavior of quark-gluon matter in at least two ways:

1. Initial conditions

2. Bulk viscosity close to T_c

Incorporate scale invariance breaking in the dual description?

The result

