Title: Non-conformal Gauge Theory Plasma in String Theory

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Abstract:

Pirsa: 06100048



## gauge/string correspondence

Alex Buchel

(University of Western Ontario & Perimeter Institute)

<u>Based on:</u> hep-th/0311175,0405200,0406264,0408098,0506002,0507026,0507275,0509083 0510041,0605076, 0605178, 0608002, 0610145

Collaborators: J.Liu, A.Starinets, O.Aharony, A.Yarom, P.Benincasa, R.Naryshkin

Original work: D.Son, A.Starinets, · · ·

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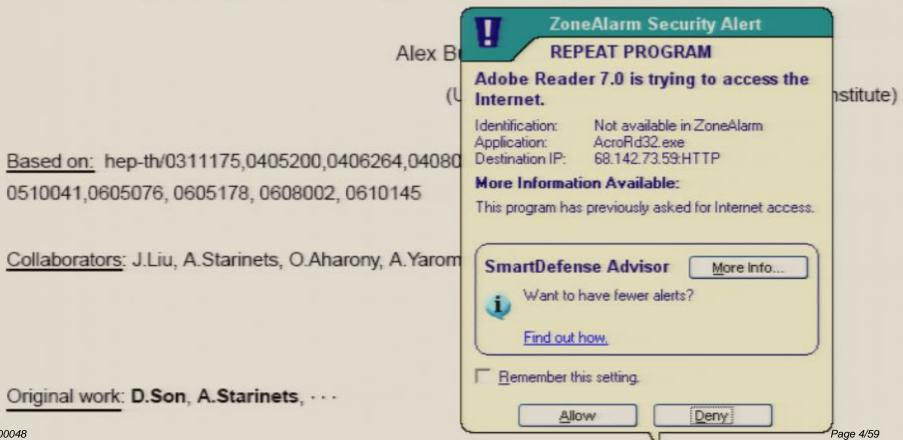
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#### Motivation:

Find a useful (experimentally testable) application of gauge theory/ string theory correspondence

Consider  $\mathcal{N}=4$  SU(N) SYM:

- ullet  $g_{YM}^2N\ll 1$  (weak effective coupling)  $\Longrightarrow$  perturbative gauge theory description
- ullet  $g_{YM}^2N\gg_1 1$  (strong effective coupling)  $\Longrightarrow$  IIB string theory on  $AdS_5 imes S^5$



Gauge theory/string theory (Maldacena correspondence)

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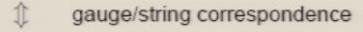


Gauge theory/string theory (Maldacena correspondence)

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- Effective description of a dynamics of a thermal system on length and time scales much longer than any relevant microscopic scale is provided by hydrodynamics
- microscopic system can be strongly coupled; in this case its hydrodynamic description (if valid) is characterized by a few "phenomenological parameters"

In this talk we discuss hydrodynamic properties of strongly coupled hot gauge theory plasma



hydrodynamics of metric fluctuations in IIB SUGRA black hole background



The latter map would allow to compute "phenomenological parameters" of the hydrodynamics from first principles

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## Outline of the talk:

- Consistencies of hydrodynamic description (gauge theory perspective)
- Consistencies of hydrodynamic description (SUGRA perspective)
- · Applications A
- renormalization of cascading gauge theories
- dynamical vs. thermodynamic instabilities of the horizon geometries
- lacksquare beyond SUGRA: lpha' corrections in  $\mathcal{F}_5$  backgrounds of string theory

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As with any duality information flows in both directions:

A: gauge theory  $\Longrightarrow$  string theory

B: string theory  $\Longrightarrow$  gauge theory

- A What hydrodynamics can teach us about string theory?
- B What are the lessons for the transport properties of hot gauge theory plasma from string theory?

 $A \Longrightarrow B$ 

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- Consistencies of hydrodynamic description (gauge theory perspective)
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Applications A

- renormalization of cascading gauge theories
- dynamical vs. thermodynamic instabilities of the horizon geometries
- **b**eyond SUGRA:  $\alpha'$  corrections in  $\mathcal{F}_5$  backgrounds of string theory

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- Applications B
- lacktriangledown small shear viscosity of  $\mathcal{N}=4$  plasma and why this could by of relevance to RHIC physics
- bulk viscosity of "realistic plasma"
- Jet quenching in strongly coupled plasma
- Universality of shear viscosity with chemical potential
  - Conclusions, future directions, open problems

amy

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### Hydrodynamics (gauge theory perspective)

hydro mode computation produces  $\begin{array}{lll} \text{shear (sh.1)} & < T_{xy,xy}>_{R,A} + \text{Kubo formula} & \eta \\ \text{shear (sh.2)} & < T_{xz,xz}>_{R} + \text{pole} & D = \frac{\eta}{Ts} \\ & & \\ & & \\ \text{sound (sw.1)} & < T_{00}>, < T_{ii}> & v_{s} \\ \text{sound (sw.2)} & < T_{00,00}>_{R} + \text{pole} & v_{s}, \Gamma \\ \end{array}$ 

- $\blacksquare$  (sh.1) and (sh.2) produces  $\eta$  must be consistent
- lacktriangleq (sw.1) and (sw.2) produces  $v_s$  must be consistent, also

Pirsa: 06100048  $\Gamma=rac{4}{3}rac{\eta}{Ts}\left[1+rac{3\xi}{4\eta}
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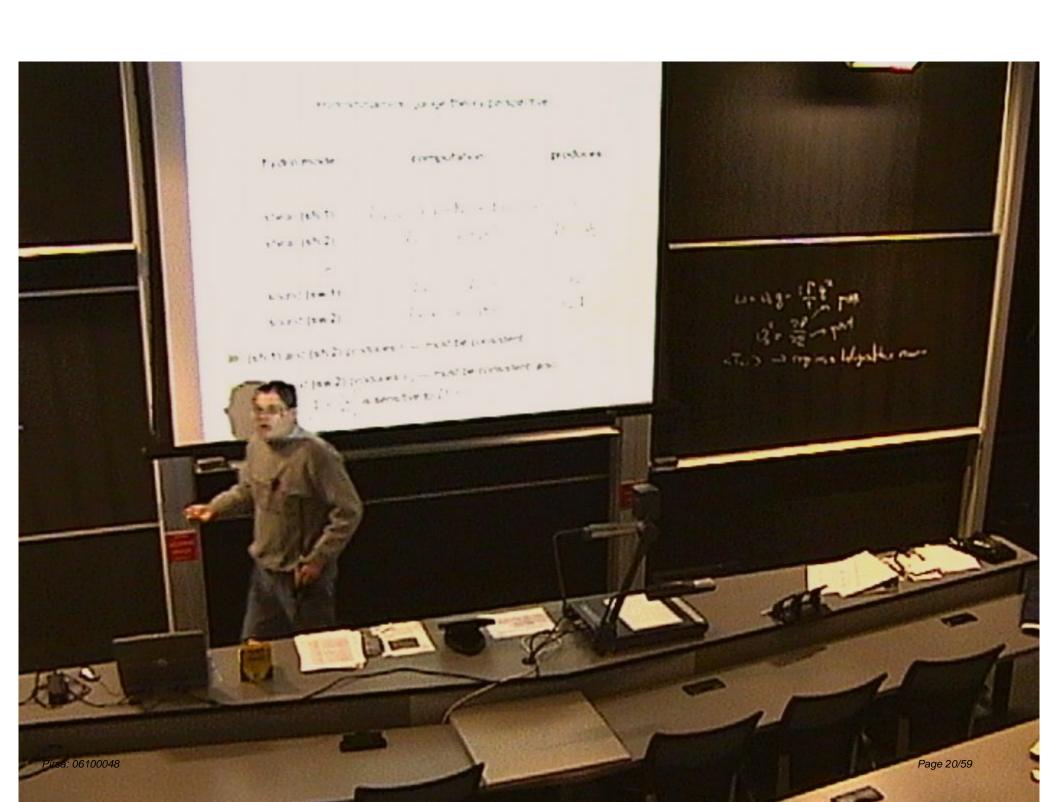
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### Hydrodynamics of supergravity fluctuations

$$T = 0$$

gauge theory string theory 
$$SU(N) \operatorname{SYM} \iff \text{N-units of 5-form flux}$$
  $g_{YM}^2 \iff g_s$ 

- $\blacksquare$  we study the theory in the 't Hooft (planar limit),  $N\to\infty$  ,  $g_{YM}^2\to 0$  with  $Ng_{YM}^2$  kept fixed
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$$T \neq 0$$

gauge theory at temperature 
$$\iff$$
 black brane in  $AdS_5 \times S^5$  at Hawking temperature  $T_H = T$ 

lacktriangleright Both finite and T=0 gauge/gravity correspondence can be extended to non-conformal gauge theories (by turning on fluxes) and to quiver gauge theories (by starting with branes on conical singularities)

We would end up computing correlation functions of gauge theory operators on SUGRA side



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- We will be interested in correlation functions of the stress-energy tensor  $T_{\mu\nu} \Rightarrow$  so the relevant SUGRA mode is the 5D  $\Delta$ etric fluctuations  $\delta g_{\mu\nu}$
- lacktriangle Complication: in cases with reduced SUSY and broken conformal invariance  $\delta g_{\mu\nu}$  fluctuations mix with "matter" fluctuations
- one has to worry about the issue of gauge (reparametrization) invariance —not all fluctuations are physical
  - Suppose that Z represents gauge invariant fluctuation —analog of Bardeen potentials in cosmology —for a retarded correlation function
- Z is an incoming wave at the horizon
- near the boundary

$$Z = \mathcal{A} \, r^{-\triangle_{-}} \, (1 + \cdots) \quad + \quad \mathcal{B} \, r^{-\triangle_{+}} \, (1 + \cdots)$$

$$\uparrow \qquad \uparrow$$

It is straightforward to evaluate using the general prescription above [P.Kovtun, A.Starinets, hep-th/0506184+other people]

$$<\mathcal{OO}>_R$$
  $\sim$   $\frac{\mathcal{B}}{\mathcal{A}}$  + contact terms

so to extract the poles of a 2-point correlation function one has to find the spectrum of black brane quasinormal frequencies

**Definition:** Z is a quasinormal mode if it is:

- (i) an incoming wave at the horizon;
- (ii) satisfy a Dirichlet condition at the boundary

#### Some important points:

the poles of the retarded correlation functions can be extracted without renormalizing the theory: boundary counterterms (on top of the standard Gibbons-Hawking counterterm) can modify only contact terms of correlators (which are renormalization prescription dependent anyway)

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lacktriangle Recall that for the sound wave mode  $v_s$  can be computed from

(sw.1) 
$$v_s^2 = \frac{\partial P}{\partial \epsilon}$$
 [1-point  $< T_{\mu\nu} >$  ]  $\Longleftrightarrow$  needs renormalization

(sw.2) From the sound pole of  $< T_{00} T_{00} > \implies$  using quasinormal mode approach <u>without</u> holographic renormalization

Consistency of (sw.1) and (sw.2) provides a highly nontrivial check on holographic renormalization of the theory

for the shear mode certain correlation functions do not have a pole (because they do not couple to effergy or momentum fluctuations) ⇔ corresponding (=transverse) metric fluctuations δg<sub>μν</sub> do not couple to SUGRA matter and their dynamics is that of the minimally coupled scalar in black brane geometry

(sh.1) 
$$\eta = \lim_{\omega \to 0} \frac{1}{2\omega i} \bigg[ G^A(\omega,0) - G^R(\omega,0) \bigg]$$

Sound wave mode has a pole, couples to energy and momentum fluctuations corresponding graviton fluctuations <u>do</u> couple to SUGRA matter

(sw.2) 
$$\omega = v_s q - \frac{i}{2} \frac{4}{3} \frac{\eta}{Ts} \left[ 1 + \frac{3\xi}{4\eta} q^2 \right]$$

Suppose we study conformal theory ( $\mathcal{N}=4$ )
Conformal invariance of  $\mathcal{N}=4$  is not broken by finite 't Hooft coupling  $\Rightarrow v_s=\frac{1}{\sqrt{3}},\,\xi=0$  even including  $\alpha'$  corrections but dispersion relation (sw.2)

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is sensitive to  $\alpha'$  corrections involving 5-form flux (5d SUGRA "matter")!

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• Check on holographic renormalization of cascading gauge theories

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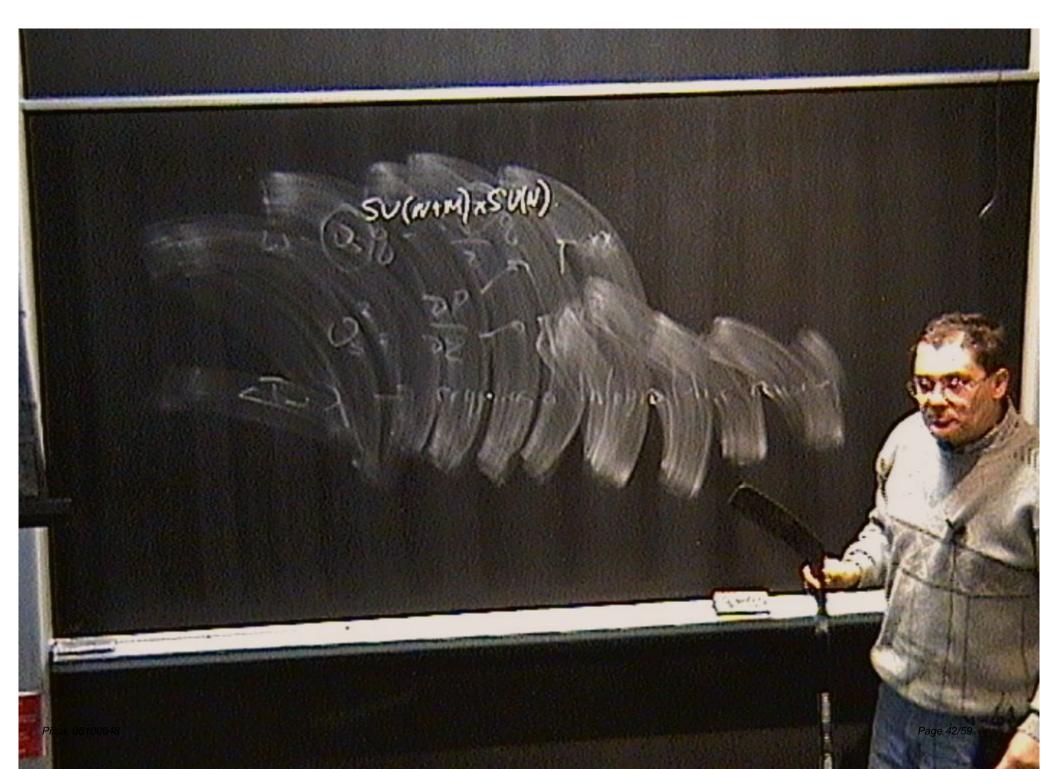
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$$N=N(E) \sim 2M^2 \ln \frac{E}{\Lambda}, \qquad E \gg \Lambda$$

where  $\Lambda$  is the strong coupling scale of the theory

- If we define cascading theory at some scale  $\mu$  with  $g_{YM}^2N(\mu)\ll 1$ , in the UV we always encounter  $g_{YM}^2N\gg 1$   $\Rightarrow$  it is not possible to renormalize the theory conventionally and one unavoidably has to use holographic renormalization
- Note that the entracted Holographic renormalization of this theory was done in hep-th/0506002, O.Aharony, A.Yarom, AB. Specifically, we computed  $< T_{\mu\nu} >$  at finite temperature and extracted

$$v_s^2 = \frac{1}{3} - \frac{2}{9 \ln \frac{T}{\Lambda}}, \quad T \gg \Lambda$$

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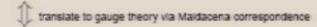
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· Gubser-Mitra conjecture

Gravitational backgrounds with translationary invariant horizon develop an instability precisely when the specific heat of a black brane geometry is negative



Finite temperature gauge theories with a negative specific heat must have a dynamical instability



trivial to identify such an instability:

Consider a sound wave mode in such gauge theory plasma

$$\omega(q) = v_s q + \mathcal{O}\left(\frac{q^2}{T}\right), \qquad v_s^2 = \frac{\partial P}{\partial \epsilon}$$

now, at zero chemical potential, f = -P

$$-\left(\frac{\partial P}{\partial T}\right)_{V} = \left(\frac{\partial f}{\partial T}\right)_{V} = -s, \quad \text{also} \quad c_{v} = \left(\frac{\partial \epsilon}{\partial T}\right)_{V}$$

SO

$$v_s^2 = \frac{\partial P}{\partial \epsilon} = \frac{\left(\frac{\partial P}{\partial T}\right)_V}{\left(\frac{\partial \epsilon}{\partial T}\right)_V} = \frac{s}{c_v}$$

- $\Rightarrow$   $c_v < 0 \Longleftrightarrow v_s$  is imaginary
- sound wave amplitude grows (dynamical instability)

translate back to gravity

sound wave quasinormal mode is unstable

#### Example:

# Consider $\mathcal{N}=1$ SYM from NS5 branes wrapping $S^2$ of the resolved conifold — Maldacena-Nunez model

from thermodynamics of MN black branes

$$v_s^2 = -2\left(\frac{T}{T_H} - 1\right)^2$$
, as  $\left(\frac{T}{T_H} - 1\right) \ll 1$ 

where the regime  $\left(\frac{T}{T_H}-1\right)\ll 1$  corresponds to the size of the resolved conifold  $S^2\to\infty$  (NS5 branes are almost flat);  $T_H$  is a Hagedorn temperature of flat NS5 branes  $T_H\propto\frac{1}{\sqrt{\# \text{of branes}}}$ 

Identical  $v_s^2$  can be extracted from the pole of the appropriate correlation functions of the stress energy tensor, or equivalently the dispersion relation of the sound quasinormal mode

- α' corrections in IIB SUGRA
- In hep-th/9808126 GKT constructed α' corrected nonextremal 3-brane geometry based on the following type IIB corrected SUGRA action:

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now, at zero chemical potential, f = -P

$$-\left(\frac{\partial P}{\partial T}\right)_{V} = \left(\frac{\partial f}{\partial T}\right)_{V} = -s, \quad \text{also} \quad c_{v} = \left(\frac{\partial \epsilon}{\partial T}\right)_{V}$$

SO

$$v_s^2 = \frac{\partial P}{\partial \epsilon} = \frac{\left(\frac{\partial P}{\partial T}\right)_V}{\left(\frac{\partial \epsilon}{\partial T}\right)_V} = \frac{s}{c_v}$$

- $\Rightarrow$   $c_v < 0 \Longleftrightarrow v_s$  is imaginary
- sound wave amplitude grows (dynamical instability)

translate back to gravity

sound wave quasinormal mode is unstable

#### Example:

# Consider $\mathcal{N}=1$ SYM from NS5 branes wrapping $S^2$ of the resolved conifold — Maldacena-Nunez model

from thermodynamics of MN black branes

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$$S_{IIB} = \frac{1}{16\pi G_{10}} \int d^{10}x \sqrt{-g} \left[ R - \frac{1}{2} (\partial \phi)^2 - \frac{1}{4 \cdot 5!} (F_5)^2 + \dots + \gamma e^{-\frac{3}{2}\phi} W + \dots \right]$$

where  $\phi$  is a dilaton,  $\gamma=\frac{1}{8}\xi(3)(\alpha')^3$ , and W is constructed from the Weyl tensor  $C_{mnpq}$ 

$$W \equiv C^{hmnk}C_{pmnq}C_h^{\ rsp}C_{\ rsk}^q + \frac{1}{2}C^{hkmn}C_{rqmn}C_h^{\ rsp}C_{\ rsk}^q$$

and  $\cdots$  denote other SUGRA modes and higher order  $\alpha'$  corrections

Some features of the  $\alpha'$  corrected geometry at  $T \neq 0$ 

$$lpha'=0$$
  $lpha'\neq 0$   $\phi=0$   $\phi\neq 0$ , depends on  $r$  size of  $S^5$  is constant size of  $S^5$  depends on  $r$   $S=rac{A_{horizon}}{4G_{10}}$   $S
eq rac{A_{horizon}}{4G_{10}}$  use Wald formula  $T_H\equiv T_0$   $T_H\equiv T_0(1+15\gamma)$ 

Pirsa: 06100048

Notice crucial assumption: <u>only</u> metric receives  $\alpha'$  corrections; 5-form does not <u>Claim</u>: consistency of hydrodynamics provides a highly nontrivial check on all these features & above assumption

Using Kubo formula (correlation functions without a pole) one finds, hep-th/0406264 J.Liu, A. Starinets, AB:

$$\frac{\eta}{s} = \frac{1}{4\pi} \left( 1 + 135\gamma \right) + \mathcal{O}(\gamma^2)$$

Alternatively, one can study dispersion relation for the shear quasinormal mode, hep-th/0510041 P.Benincasa,AB:

$$\omega = -iDq^2 = -i\Gamma_{\eta} \frac{q^2}{2\pi T_0}$$
, where  $\Gamma_{\eta} = \frac{1}{2} \left( 1 + 120\gamma \right)$ 

According to hydro consistency relation  $\frac{\Gamma_{\eta}}{2\pi T_0}=D$  ( the shear diffusion constant)

$$\frac{\eta}{s} = DT = \frac{\Gamma_{\eta}}{2\pi} \frac{T}{T_0} = \frac{1}{4\pi} \left( 1 + 120\gamma \right) \left( 1 + 15\gamma \right) = \frac{1}{4\pi} \left( 1 + 135\gamma \right) + \mathcal{O}(\gamma^2)$$

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Consider a sound wave mode, hep-th/0510041 P.Benincasa,AB:

$$\omega = v_s q - i \, \Gamma_{sound} \, \frac{q^2}{2\pi T_0}, \qquad \text{where} \qquad \Gamma_{sound} = \frac{1}{3} \left( 1 + 120 \gamma \right) + \cdots$$

$$v_s = \frac{1}{\sqrt{3}} \left( 1 + 0 \cdot \gamma + \cdots \right) \iff \text{does not receive } \alpha' \text{ corrections}$$

Consistency of hydro

$$\frac{\Gamma_{sound}}{2\pi T_0} = \frac{2}{3T} \frac{\eta}{s} \left( 1 + \frac{3\xi}{4\eta} \right) = \frac{2}{3T} \frac{1}{4\pi} \left( 1 + 135\gamma \right) \left( 1 + \frac{3\xi}{4\eta} \right)$$

OF

$$3 \Gamma_{sound} \frac{T}{T_0} = \left(1 + 135\gamma\right) \left(1 + \frac{3\xi}{4\eta}\right)$$

$$\updownarrow$$

$$3 \frac{1}{3} \left(1 + 120\gamma\right) \left(1 + 15\gamma\right) = \left(1 + 135\gamma\right) \left(1 + \frac{3\xi}{4\eta}\right) \Rightarrow \xi = 0 \cdot \gamma$$

Again,  $\eta$  computation is not sensitive to  $\alpha'$  corrections in matter sector (in this case 5-form flux) of IIB SUGRA, while sound dispersion relation (which depends on  $\eta$ ) is sensitive

### Applications B

- shear viscosity at RHIC
- Experimental data at RHIC suggest very fast thermalization of the quark-gluon plasma produced in heavy ion collisions
- $\blacksquare$  (for review: hep-ph/0510232, Kovchegov) the thermalization time  $\tau_T$

$$au_T \propto \left(\frac{\eta}{T^3}\right)^{4/3}$$
 $au_T$ 

small thermalization times  $\iff$  small  $\frac{\eta}{s}$ 

## Drag force vs. jet quenching parameter

ullet For a massive quark  $m\gg\sqrt{\lambda}T$ 

$$\frac{dp}{dt} = -\frac{\pi\sqrt{\lambda}T^2}{2} \, \frac{v}{\sqrt{1 - v^2}}$$

Problem: in SUGRA approximation  $\lambda = \infty$ 

ullet Consider a light-like Wilson loop C with large extension  $L^-$  in  $x^-$  direction and small extension L in transverse direction. Introduce a 'jet quenching parameter'  $\hat{q}$ 

$$\langle W^A(C) \rangle = \exp\left(-\frac{1}{4}\hat{q}L^-L^2 + \mathcal{O}(L^4)\right)$$

Problem: in SUGRA approximation  $\lambda = \infty$ 

## Shear viscosity in the presence of chemical potential

ullet Recently (A.Starinets, D.Son, $\cdots$ ) it was shown that introducing R-charge chemical potential for N=4 SYM leads to

$$\frac{\eta}{s} = \frac{1}{4\pi} \, \frac{\hbar}{k_B}$$

£77

Can it be generalized?

- ullet (with J.Liu) We constructed new gauged supergravities by gauging U(1) isometry of  $Y^{p,q}$  manifolds
- Leads to <u>new</u> examples of SUGRA dual to CFT plasmas with a chemical potential
- Dual (rotating/charged) black hole solution is found analytically

$$\frac{\eta}{s} = \frac{1}{4\pi} \, \frac{\hbar}{k_B}$$

em)

- Claim: Above result implies universality for shear viscosity even with a chemical potential
- Proof: (with P.Benincasa and R.Naryshkin) hep-th/0610145

### Conclusions

• fun to study non-equilibrium AdS/CFT correspondence

■ In the future:

· relation between different approaches to jet quenching in plasma

photon and dilepton production in non-conformal QGP

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