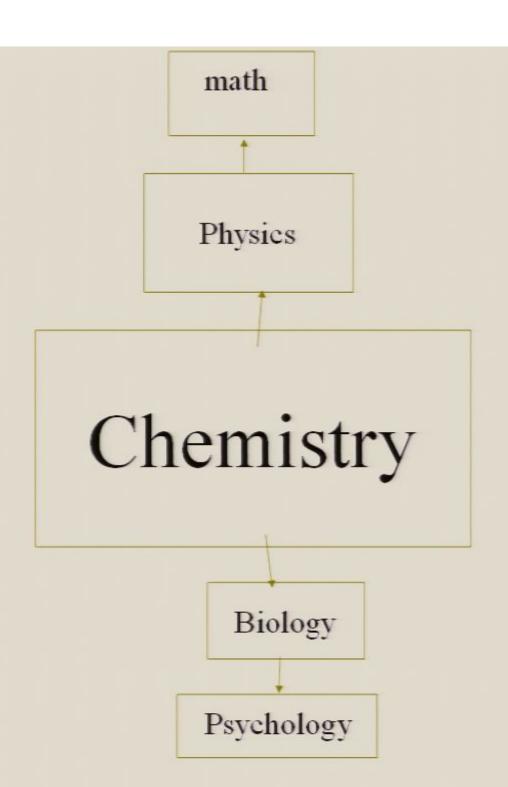
Title: Black hole evaporation and information

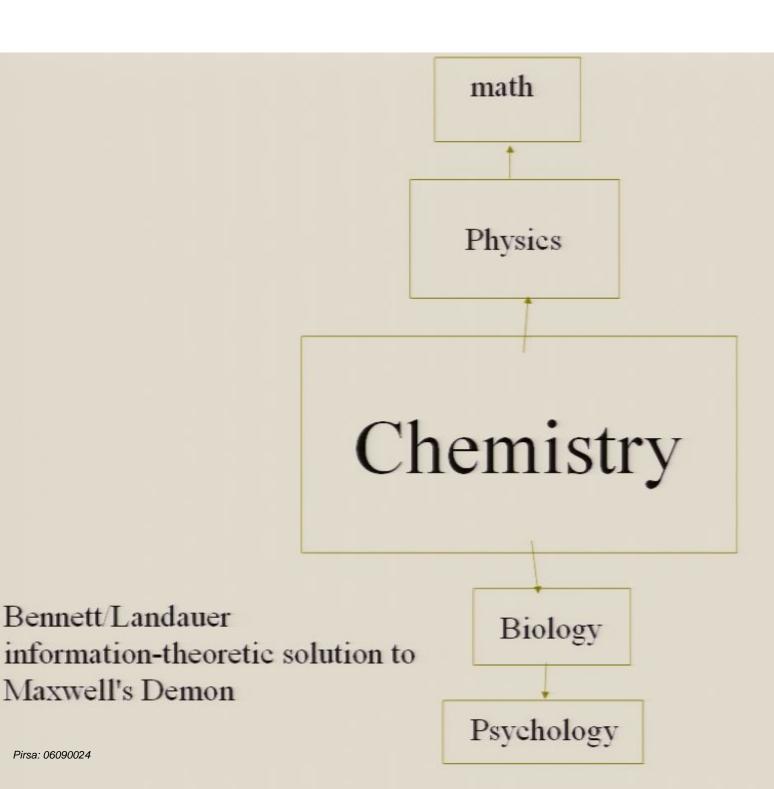
Date: Sep 05, 2006 06:00 PM

URL: http://pirsa.org/06090024

Abstract:



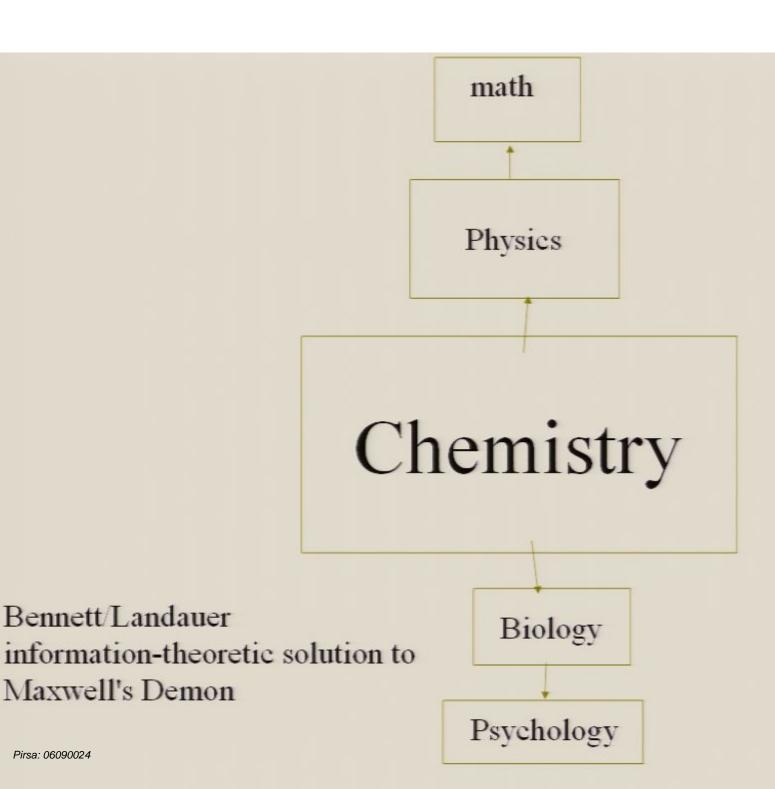




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Bennett/Landauer

Maxwell's Demon



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Bennett/Landauer

Maxwell's Demon

Black hole information loss paradox is **the** reason quantum mechanics and general relativity are not unified.

1: New result in quantum information theory: Locking

2: What is the information loss paradox?

3: How locking helps solve the paradox

4: Some problems with the approach

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Black hole information loss paradox is **the** reason quantum mechanics and general relativity are not unified.

1: New result in quantum information theory: Locking Eating our vegetables

2: What is the information loss paradox?

3: How locking helps solve the paradox

Hopefully it's almost obvious by now

4: Some problems with the approach

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Classical mutual information:

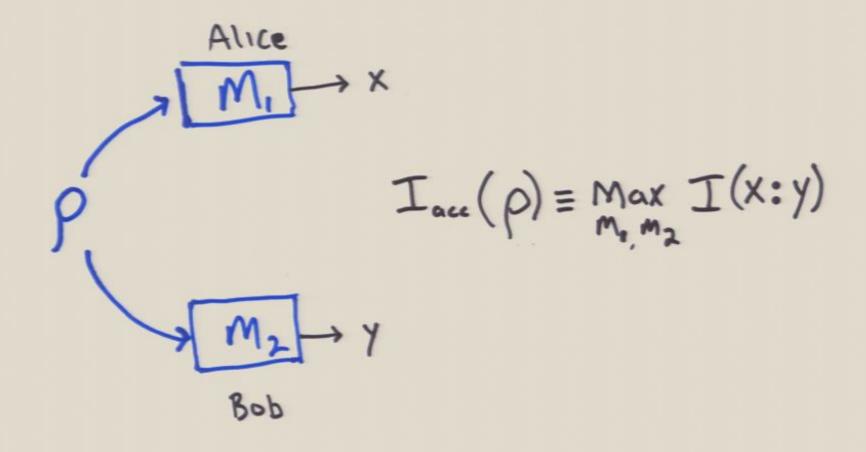
$$I(X:Y) = H(X) + H(Y) - H(XY)$$

$$I(XZ:Y) - I(X:Y) \le I(Z:Y) \le H(Z)$$

That is to say, the additional information about Y one gets by getting Z can be no more that the information in Z about Y.

So if someone sends you a bit, your information increases by at most 1 bit---can be less if you already knew the information.

This is obvious, indeed it is why mutual information is designed this way



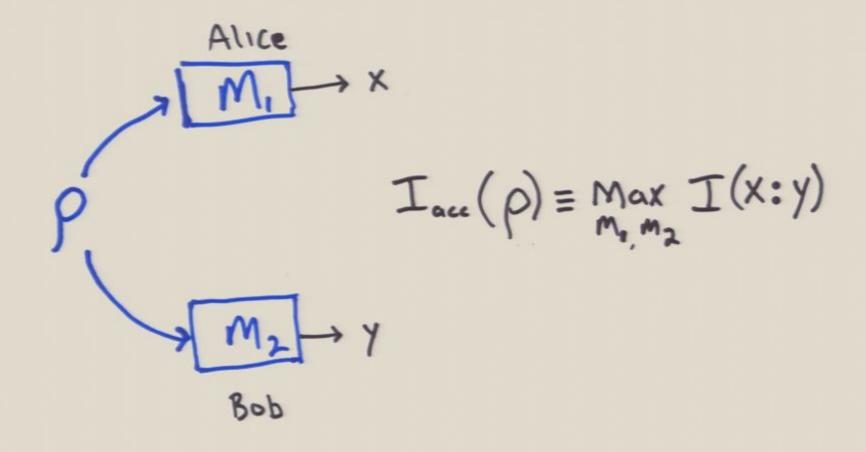
$$|\uparrow\rangle$$
 $|10\rangle$
 $|\downarrow\rangle$
 $|\downarrow\rangle$

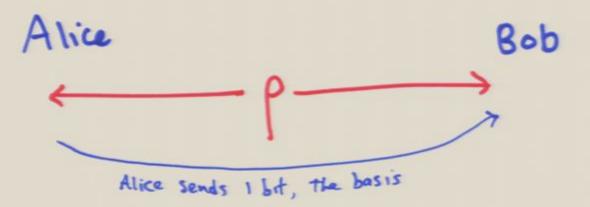
111

from BB84 Quantum Cryptography

, Not an orthonormal basis!

Alice has 2 bits, Bob learns nothing about the basis, gets the "bit" half the time





Bob finds out the bit by measuring in appropriate basis. Now he has 2 bits

$$I_{acc}(p') = 2$$

$$I_{acc}(p') - I_{acc}(p) = \frac{3}{2} > 1$$

$$P = \frac{1}{20} \sum_{i=1}^{2} \sum_{j=1}^{2} |i \times i|_{A} \otimes |j \times j|_{A} \otimes |ij \times ij|_{B}$$
Still not a basis

$$I_{acc}(p') - I_{acc}(p) = \frac{\log D}{2} + 1 \gg 1$$

Arbitrarily large violation!

$$\rho = \frac{1}{dn} \sum_{i=1}^{d} \sum_{j=1}^{n} |i\rangle\langle i|_A \otimes |j\rangle\langle j|_A \otimes U_j |i\rangle\langle i|_B U_j^{\dagger}$$

$$n = (\log d)^3 + k$$
$$I_{\text{acc}}(\rho) < \delta = C^{-k}$$

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$$I_{
m acc}(
ho') pprox \log d$$

$$\rho = \frac{1}{dn} \sum_{i=1}^{d} \sum_{j=1}^{n} |i\rangle\!\langle i|_A \otimes |j\rangle\!\langle j|_A \otimes U_j |i\rangle\!\langle i|_B U_j^\dagger$$

$$\underset{\text{message}}{\uparrow}$$
we have the property of the property

$$n = (\log d)^3 + k$$
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Pirsa: 060900
$$^{14}_{
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 message key encrypted message

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Pirsa: 060900
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$$\underset{\text{message}}{\text{message}} \qquad \underset{\text{key}}{\text{encrypted message}}$$

$$\rho' = \frac{1}{dn} \sum_{i=1}^{d} \sum_{j=1}^{n} |i\rangle\!\langle i|_{A} \otimes |j\rangle\!\langle j|_{A} \otimes U_{j}|i\rangle\!\langle i|_{B} U_{j}^{\dagger} \otimes |j\rangle\!\langle j|_{B}$$

$$n = (\log d)^3 + k$$
$$I_{acc}(\rho) < \delta = C^{-k}$$

Pirsa: 060900
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$$n = (\log d)^3 + k$$
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Pirsa: 060900
$$I_{
m acc}(
ho') pprox \log d$$

Quantum Mechanics is Unitary

An initial pure state ψ will remain forever pure.

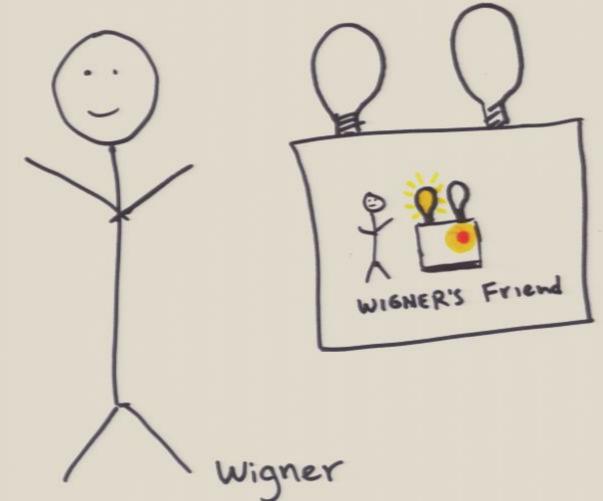
Except for measurement?

Wigner's friend

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Black holes are not unitary

Irreversible—Things that fall in don't come out

Not so bad. This is a restriction on the evolution of states, like superselection, rather than a total breakdown.

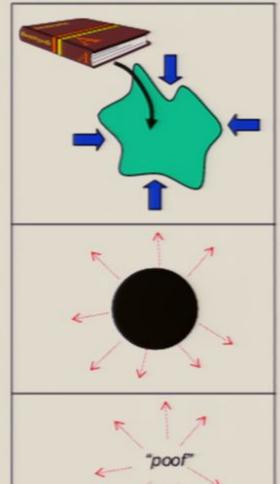
But, black holes evaporate (Hawking 1975) (and the way they do is nonunitary)

$$S(\psi) = \rho_{thermal}$$

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Black hole evaporation

Suppose we prepare a quantum state, encoding some information, as pressureless dust on the brink of gravitational collapse.



It collapses, and begins to emit Hawking radiation. This radiation is featureless, not dependent on the information encoded in original collapsing body.

Eventually, all the mass is radiated away, and the black hole disappears. What happened to the information?

If it is gone, unitarity fails and quantum mechanics is wrong. If it comes out, GR must be wrong.

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Whereas Stephen Hawking and Kip Thorne firmly believe that information swallowed by a black hole is forever hidden from the outside universe, and can never be revealed even as the black hole evaporates and completely disappears,

And whereas John Preskill firmly believes that a mechanism for the information to be released by the evaporating black hole must and will be found in the correct theory of quantum gravity,

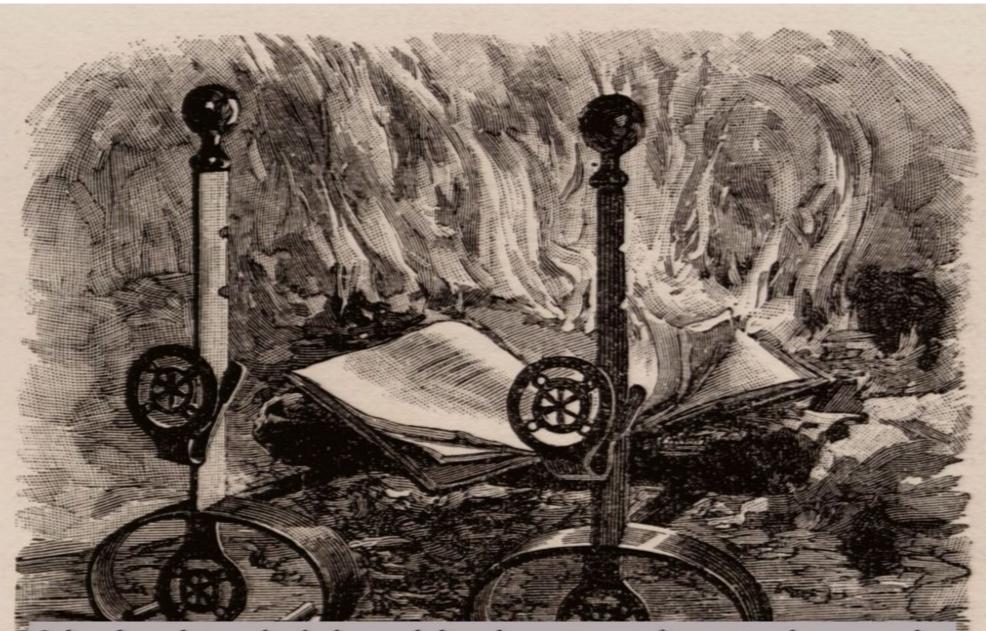
Therefore Preskill offers, and Hawking/Thorne accept, a wager that:

When an initial pure quantum state undergoes gravitational collapse to form a black hole, the final state at the end of black hole evaporation will always be a pure quantum state.

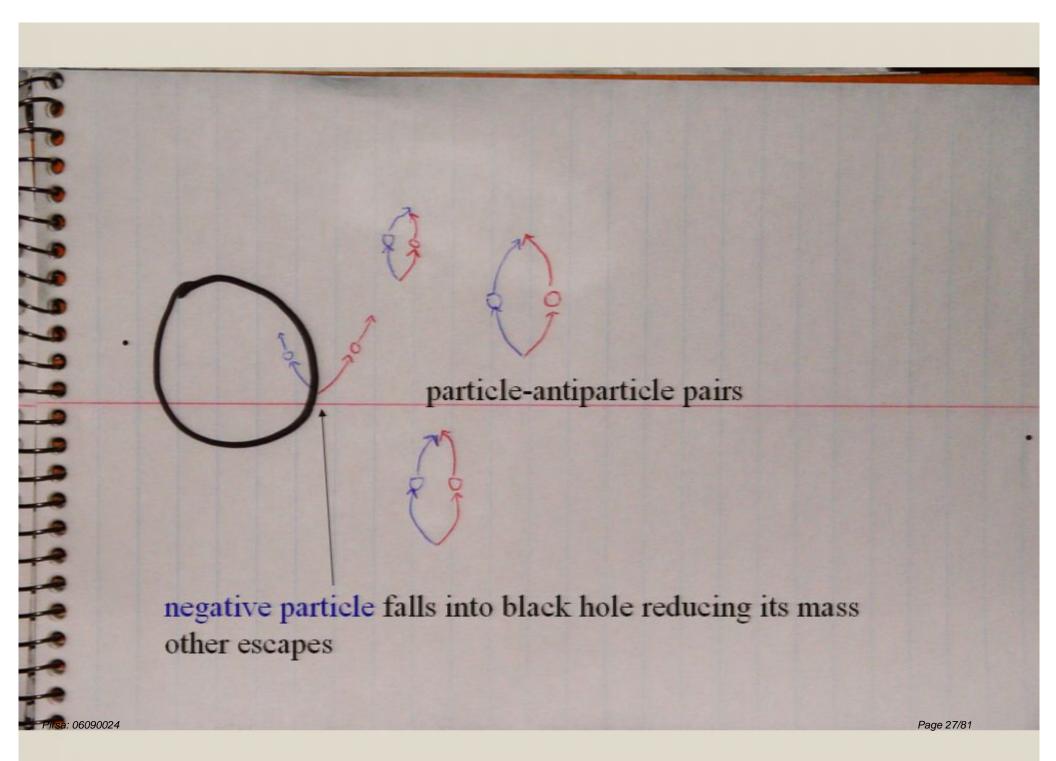
The loser(s) will reward the winner(s) with an encyclopedia of the winner's choice, from which information can be recovered at will.

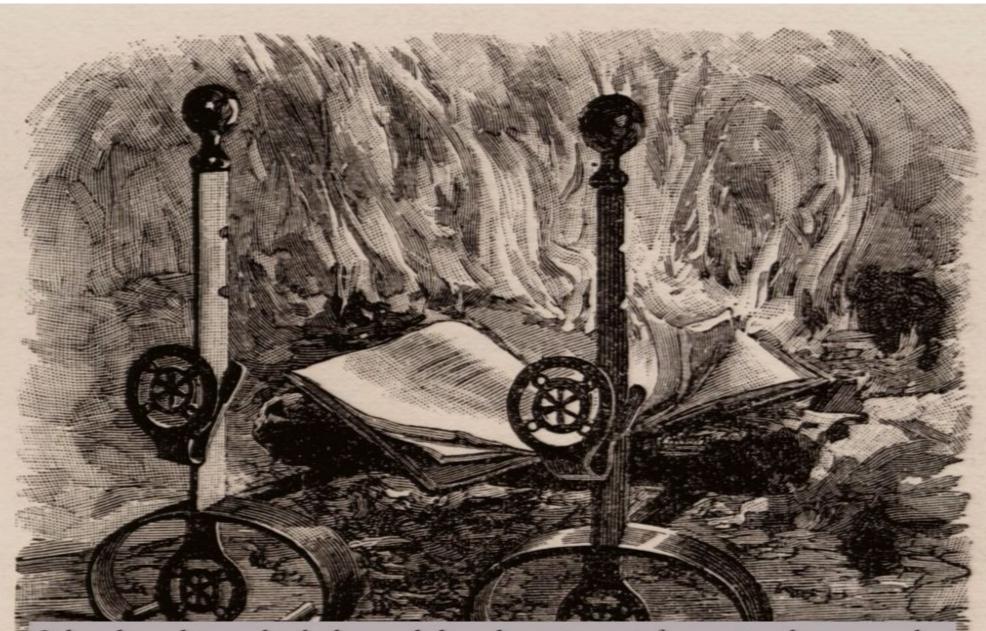
Stephen W. Hawking, Kip S. Thorne, John P. Preskill Pasadena, California, 6 February 1997

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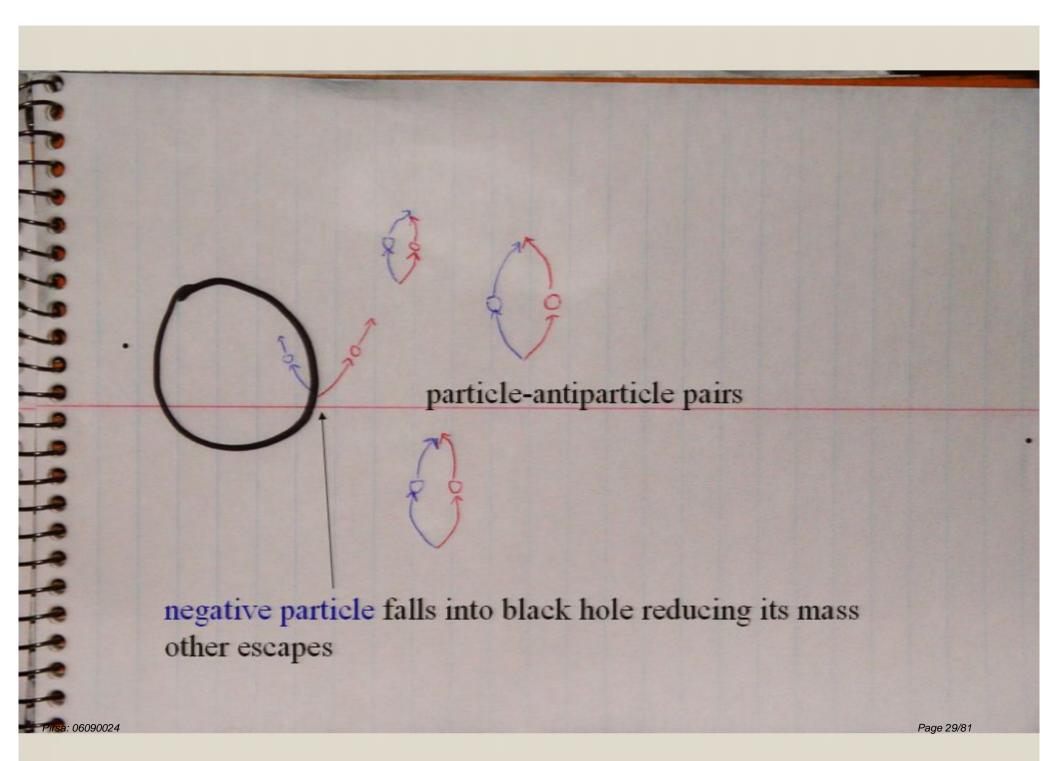


Other hot objects look thermal, but there is no information-loss paradox. Why are black holes counted as a special problem?

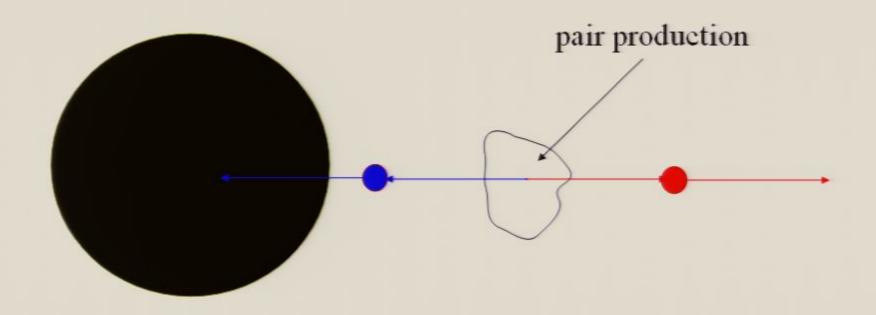




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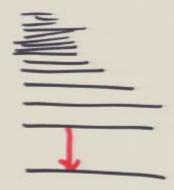


Since the creation happens outside the event horizon, the outgoing particle has no information about what's inside the black hole.



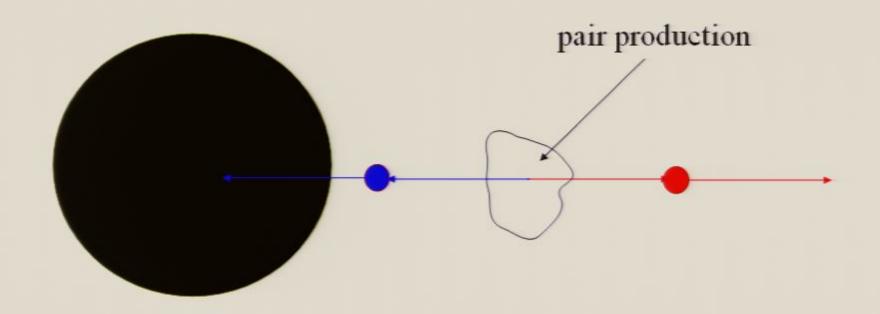
Thermal spectrum comes from a cross-section calculation involving wavelength vs. black hole size in full GR.

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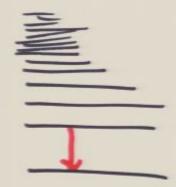
Har production cantled photon

QED View of all radiation System only "eats" certain wavelengths Since the creation happens outside the event horizon, the outgoing particle has no information about what's inside the black hole.



Thermal spectrum comes from a cross-section calculation involving wavelength vs. black hole size in full GR.

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Har production emitted photon

QED View of all radiation System only "eats" certain wavelengths

Evolution to a thermal state is bad for **two** reasons:

- 1) Initial pure state evolves into highly mixed state (anomalous entropy production)
- 2) Final state is uncorrelated with the initial state (information loss paradox)

Any unitary theory will disagree with Hawking on these so we will assume his calculation is only true approximately, and for large black holes.

Hydrogen atom analogy explains away both points, but is satisfactory only for point 1 as we shall see

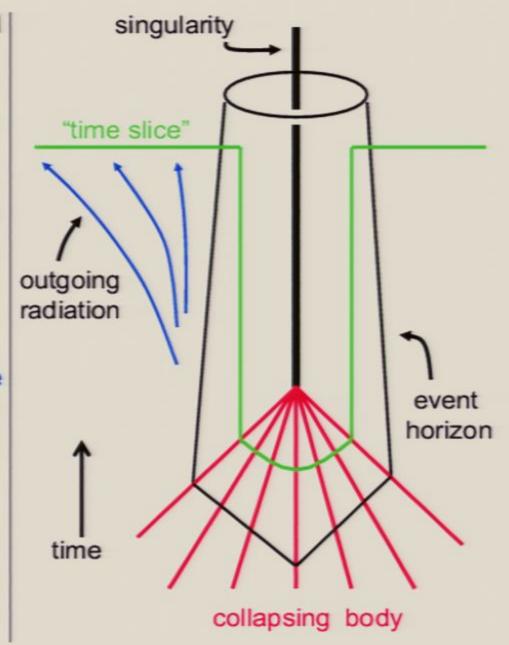
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Black hole: a quantum cloning machine?

Suppose that the information about the collapsing body is subtly encoded in correlations among the quanta in the Hawking radiation; the information is *thermalized*, not destroyed.

The green time slice crosses both the collapsing body behind the horizon and the radiation outside the horizon. Thus the same information is in two places at the same time.

A quantum cloning machine has operated, which is not allowed by the linearity of quantum mechanics.



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Quantum Information cannot be cloned.



Bah!



Unitarity implies no cloning

Quantum Information cannot be cloned.



Bah!



Unitarity implies no cloning

We insist on unitarity

Therefore information escapes

Therefore we can clone?

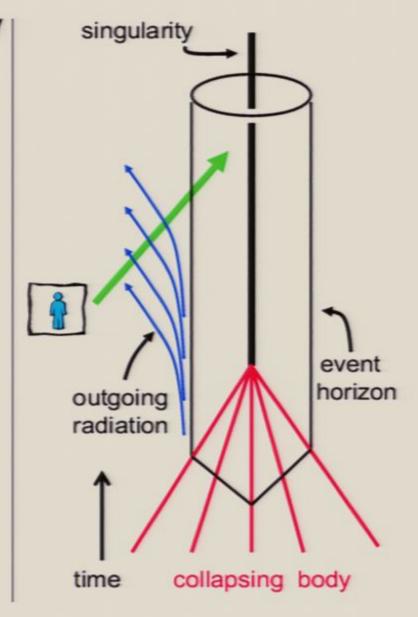
Black hole complementarity

An observer who reads the outgoing radiation concludes that the information must be erased as the collapsing body crosses the event horizon.

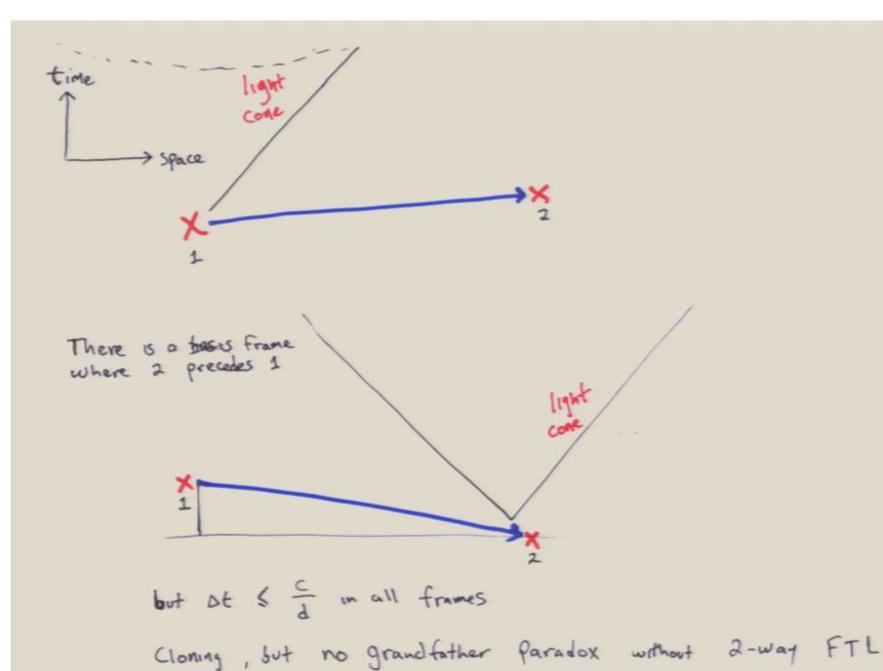
A freely falling observer who follows the collapsing body across the horizon knows otherwise.

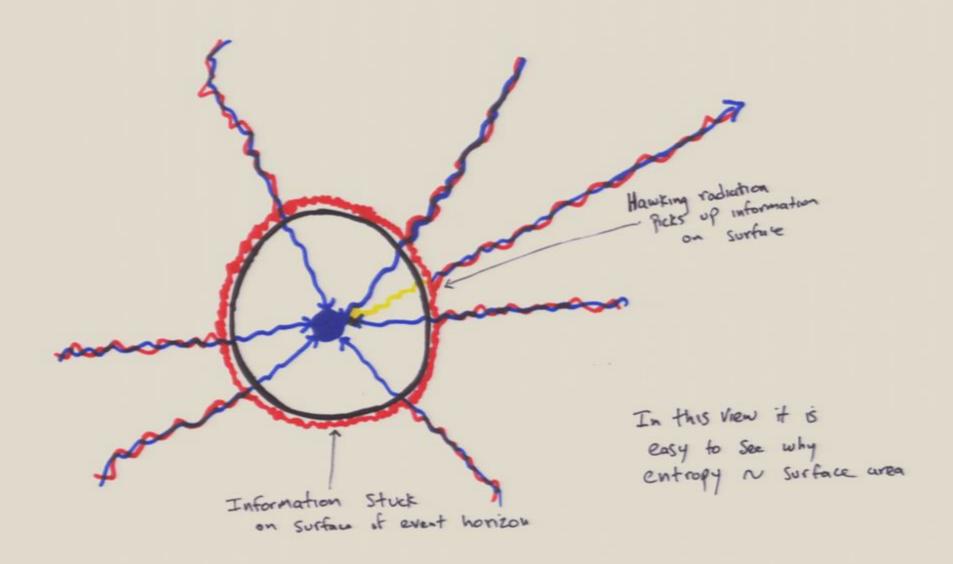
But they can never compare notes...

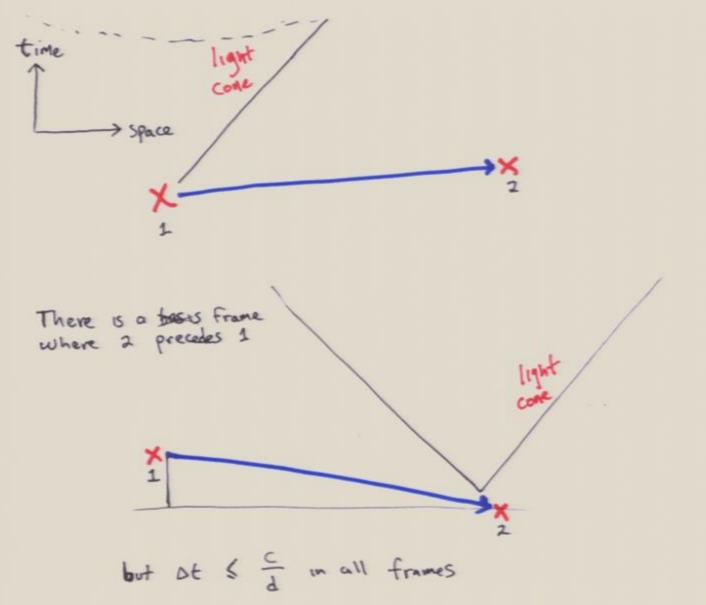
Perhaps it is okay for quantum information to be copied, if no one can ever find out!



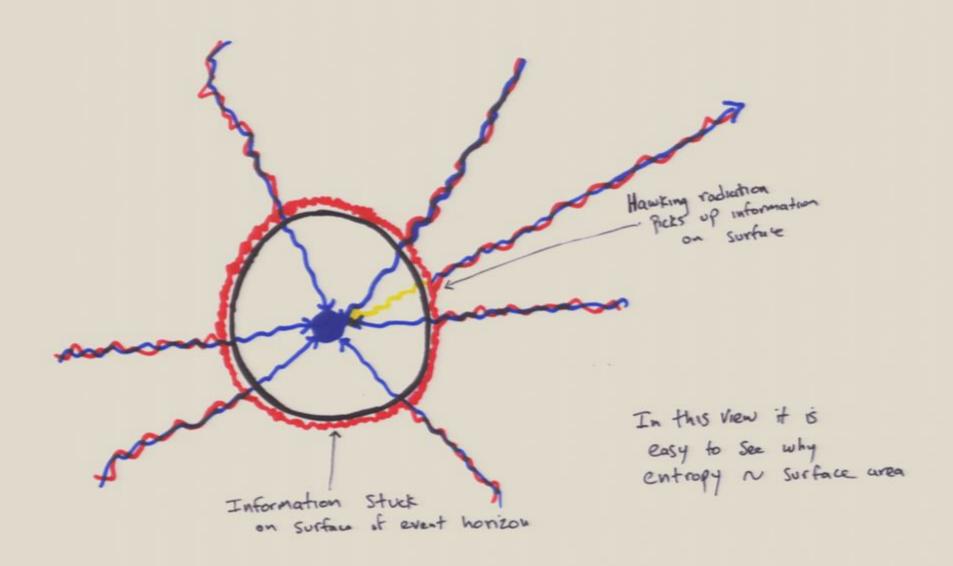
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Cloning, but no grandfather Paradox without 2-way FTL



Complementarity unsatisfying because we want a theory that works for everyone.

To preserve causality for information that really falls into the hole, we must insist that no information leaks out.

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Complementarity unsatisfying because we want a theory that works for everyone.

To preserve causality for information that really falls into the hole, we must insist that no information leaks out.

Loophole—Information all gets out at the end of the evaporation, when the black hole is Planck scale and semi-classical ideas surely fail

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Complementarity unsatisfying because we want a theory that works for everyone.

To preserve causality for information that really falls into the hole, we must insist that no information leaks out.

Loophole—Information all gets out at the end of the evaporation, when the black hole is Planck scale and semi-classical ideas surely fail

But this means all the information has to live in a very small hole

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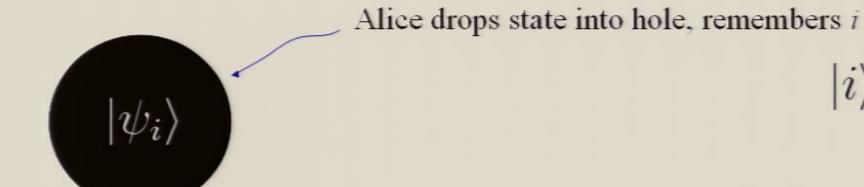
Tiny black holes shouldn't be able to hold large amount of information:

1: We have a black hole entropy forumula: $S=M^2$

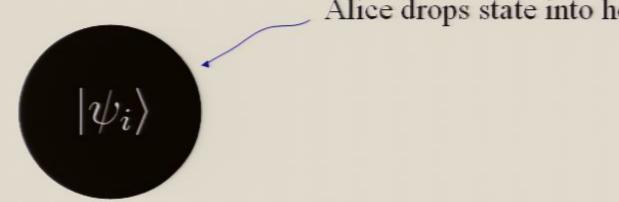
we're discarding semi-classical ideas anyway, so not too convincing

2: It takes a long time for such an information-dense hole to evaporate, of order M^4 Aharonov, Casher and Nussinov/Carliz and Willey (1987)

This is so long this is effectively a stable remnant. There would be $(M/M_p)^2$ different species. This degeneracy would change low-energy physics.

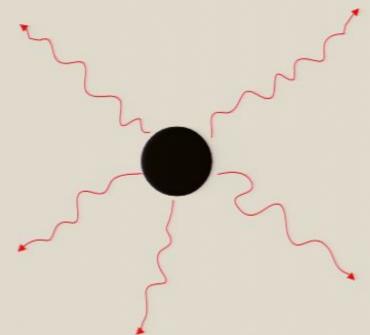


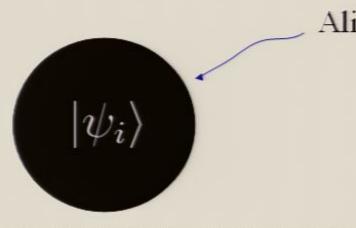
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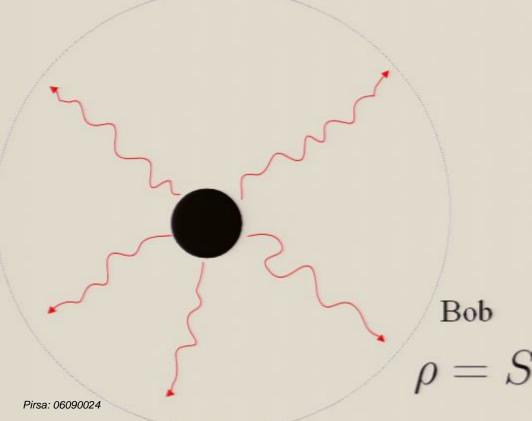
 $|i\rangle$

Alice $|i\rangle$



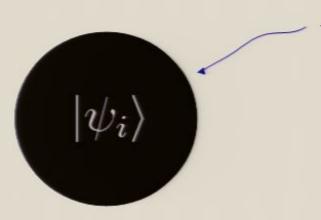


 $|i\rangle$

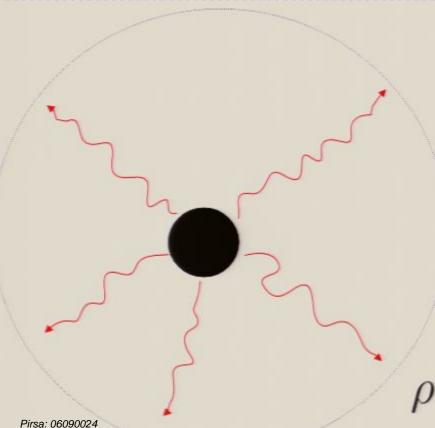


Alice $|i\rangle$

 $\rho = S_t(|\psi_i\rangle\langle\psi_i|)$



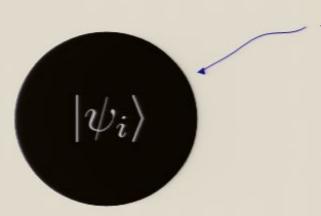
 $|i\rangle$



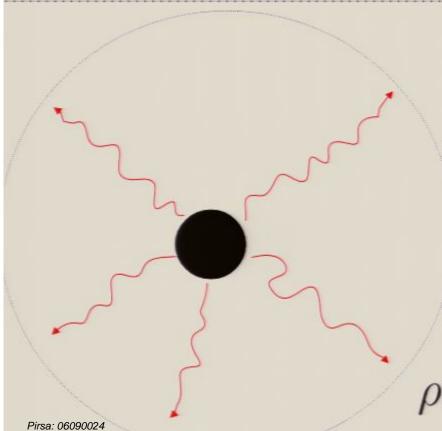
Alice $|i\rangle$

$$\rho_{AB} = \sum_{i} p_{i} |i\rangle\langle i|_{A} \otimes S_{t}(|\psi_{i}\rangle\langle\psi_{i}|)_{B}$$

$$\rho = S_t(|\psi_i\rangle\langle\psi_i|)$$



 $|i\rangle$



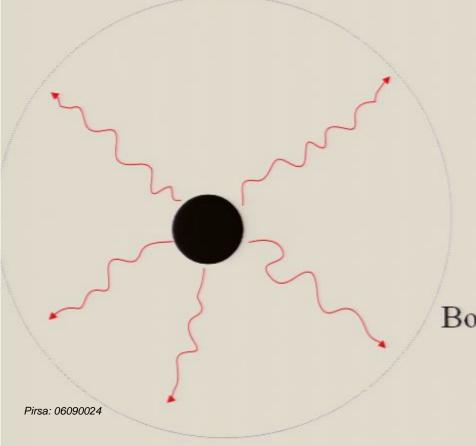
Alice $|i\rangle$

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To preserve causality we want:

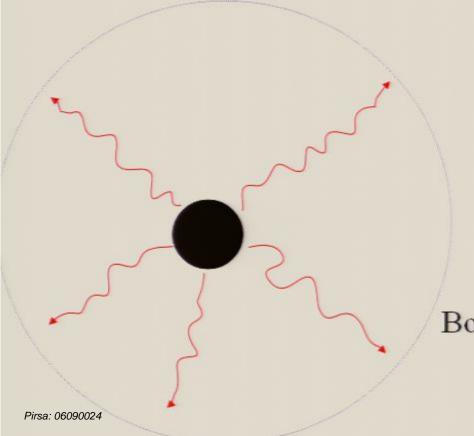
$$I_{\rm acc}(\rho_{AB}) \le \epsilon$$

$$\rho = S_t(|\psi_i\rangle\langle\psi_i|)$$



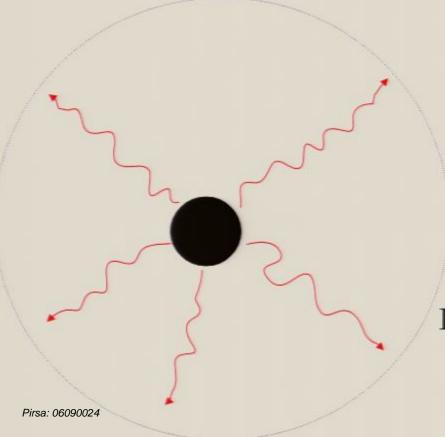
$$I_{\rm acc}(\rho_{AB}) \le \epsilon$$

$$S_t(|\psi\rangle\langle\psi|) = \frac{1}{n} \sum_{j=1}^n U_j |\psi\rangle\langle\psi| U_j^{\dagger}$$



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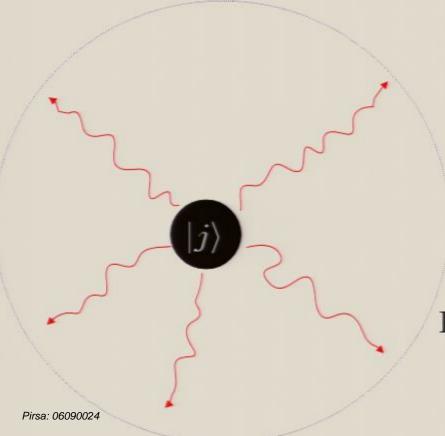


$$\rho_{AB} = \frac{1}{dn} \sum_{i=1}^{d} \sum_{j=1}^{n} |i\rangle\langle i|_{A} \otimes U_{j} |i\rangle\langle i|_{B} U_{j}^{\dagger}$$

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$$n = (\log d)^3 + k$$

$$S(n) \approx 3 \log \log d$$

Size of small black hole in information or Planck units

$$I_{\rm acc}(\rho') \approx \log d$$

All information that formed hole

$$I_{\rm acc}(\rho) < \delta = C^{-k}$$

Amount of information that leaks ahead of schedule

Objections (and solutions?)

No mechanism:

Some information still leaks:

$$I_{\rm acc}(\rho) < \delta = C^{-k}$$

Objections (and solutions?)

No mechanism:

True, but at least the locking effect is very generic. Almost any random partitioning will work (and therefore many not-so-random ones as well).

Some information still leaks: $I_{\rm acc}(\rho) < \delta = C^{-k}$

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Objections (and solutions?)

No mechanism:

True, but at least the locking effect is very generic. Almost any random partitioning will work (and therefore many not-so-random ones as well).

Some information still leaks: $I_{\rm acc}(\rho) < \delta = C^{-k}$

Semi-classical Hawking calculation always allowed some leakage (actually much more than this) and this leakage is arbitrarily small

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$$n = (\log d)^3 + k$$

For some d this will be larger than Planck scale.

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$$n = (\log d)^3 + k$$

For some d this will be larger than Planck scale.

How large? For a solar mass black hole containing 10^{38} bits, $n=2^{20}$, or 20 Planck masses required.

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$$n = (\log d)^3 + k$$

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Cosmological answer: We don't have to worry about locking any amount of information, only all the information in entire universe:

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Lloyd estimates 10¹²⁰ bits of information in the universe, only 40 Planck masses needed.

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$$n = (\log d)^3 + k$$

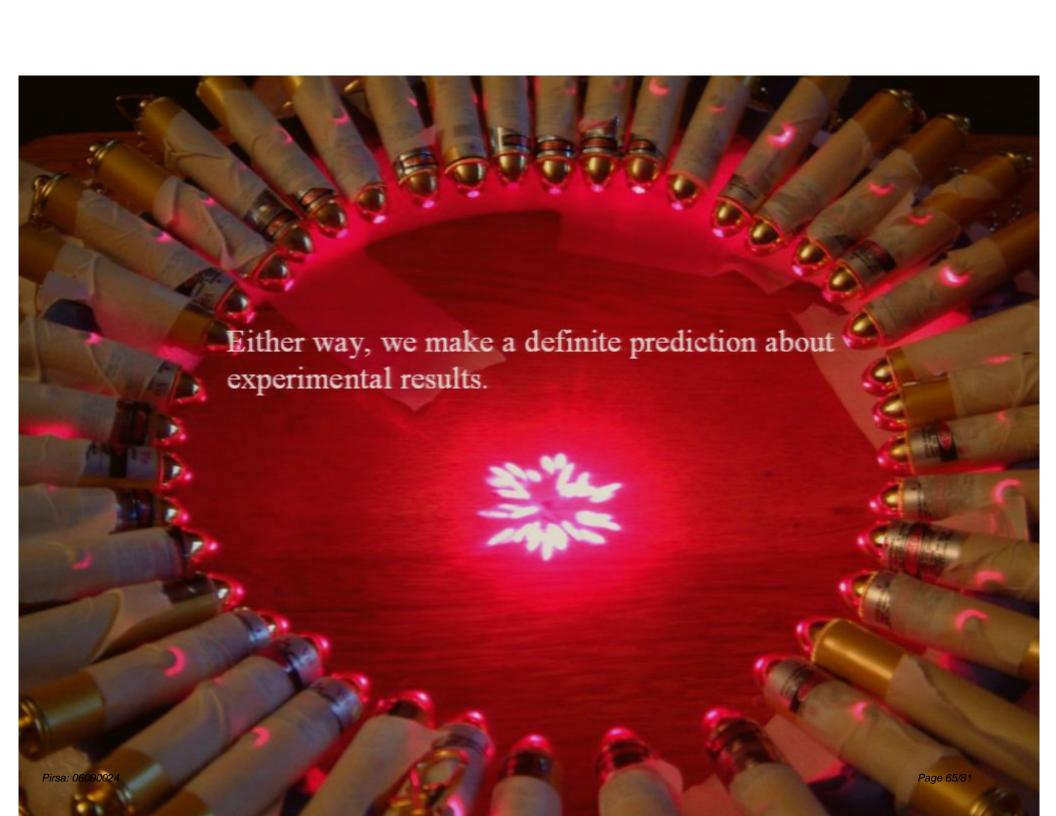
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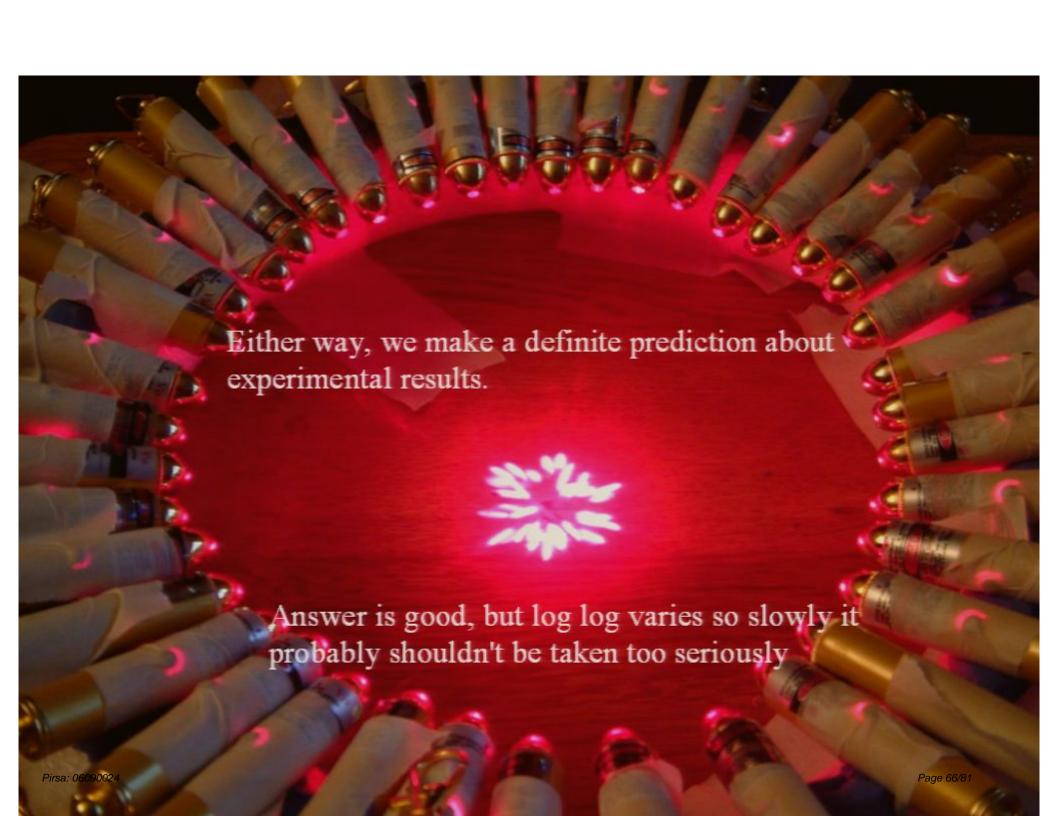
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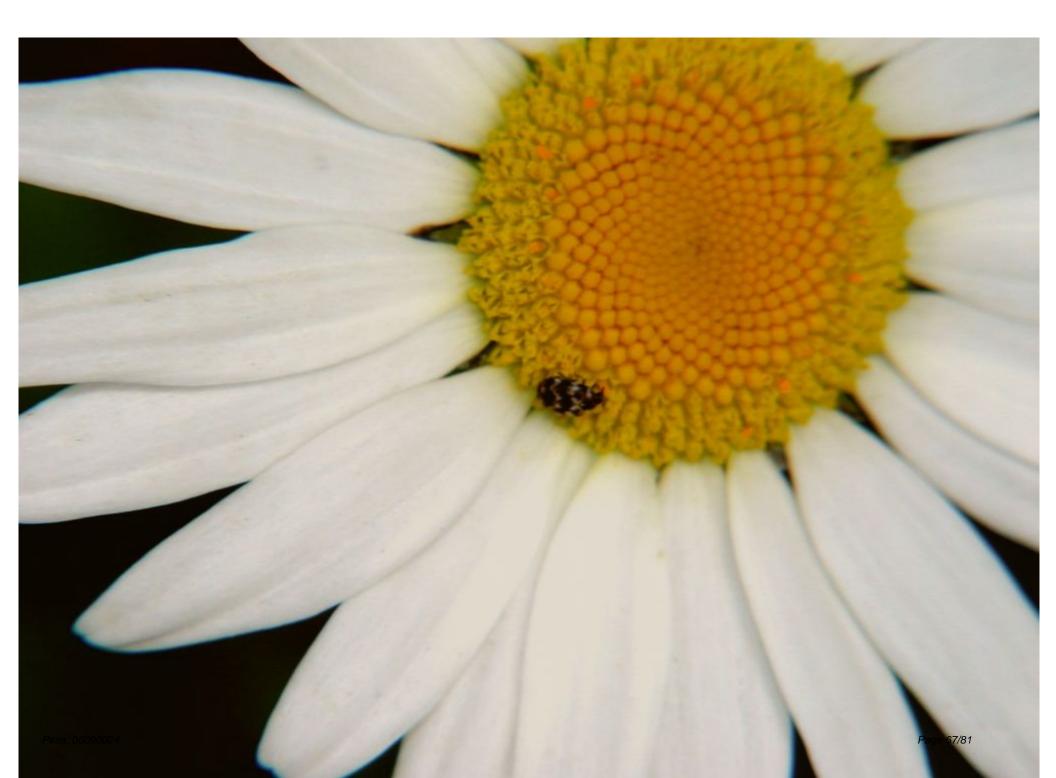
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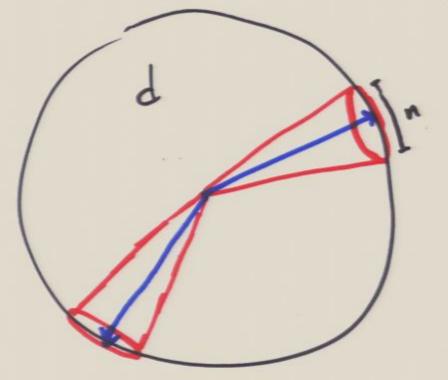
Even more fun (from (Karol Horodecki): Turn this on its head--size of universe determines Planck's constant.







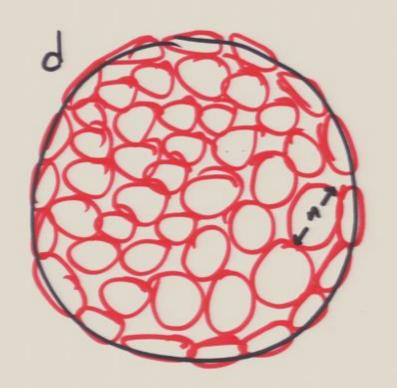
A State is mapped to a cone of size n



chance that two rundom States overlap is small

n = (log d)3

sphere is covered by roughly $\frac{d}{h} = k$ circles



log d - log n = log K

total Possible key size f

total Possible encoded

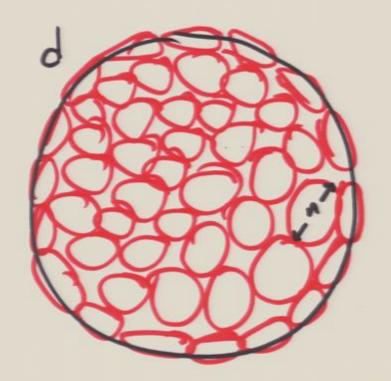
information encoded

bits

Packing Lemma (Devetak): if we send only $k = \frac{d}{n}$ Vectors, then $I_{acc}(P) = log k$

All of it / No locking!

Sphere is covered by roughly & = K Circles



log d - log n = log K total Possible Key Size Information

Packing Lemma (Devetak): if we send only $k = \frac{d}{n}$ Vectors, then Ince (P) = log K

All of it / No locking!

"Mule Rescue": Information Complementarity

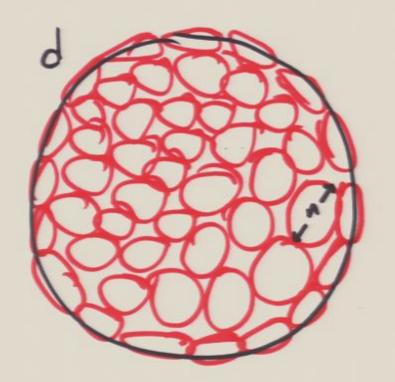
Alice's codebook

Howking Cadiation of Shir Codebook

Alice needs the codebook to "interpret" the state

But the Codebook & bigger than the black

sphere is covered by roughly $\frac{d}{h} = k$ circles



log d - log n = log K

total Possible key size f

total Possible encoded

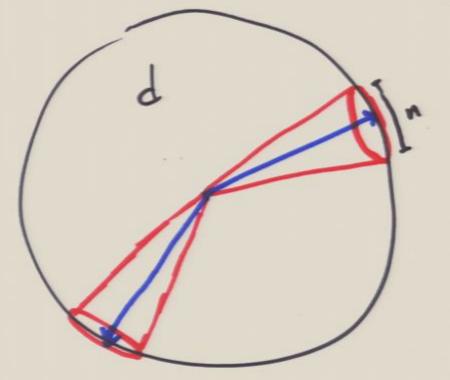
bits

Packing Lemma (Devetak): if we send only $k = \frac{d}{n}$ Vectors, then $I_{acc}(P) = log k$

All of it / No locking!

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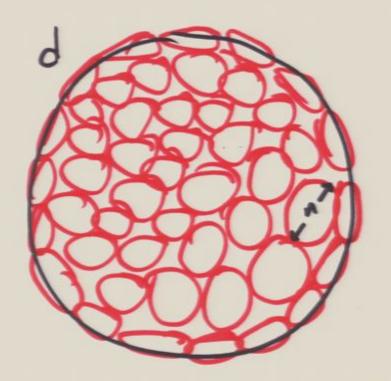
A State is mapped to a cone of size n



chance that two rundom States overlap is small

n = (log d)3

sphere is covered by roughly $\frac{d}{h} = k$ circles



log d - log n = log K

total Possible key size f

total Possible encoded

bits

Packing Lemma (Devetak): if we send only $k = \frac{d}{n}$ Vectors, then $I_{acc}(P) = log k$

All of it / No locking!

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"Mule Rescue": Information Complementarity

original infalling State

Alice's codebook

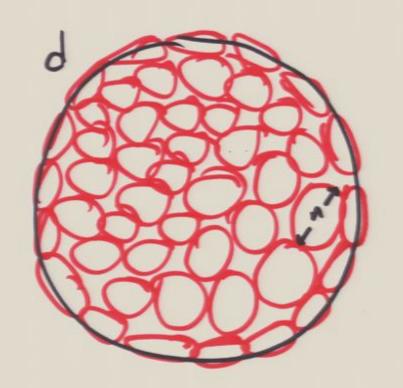
Alice needs the codebook to "interpret" the State

Howking radiation

Bob's Code book

But the Codebook & bigger than the black

sphere is covered by roughly $\frac{d}{h} = k$ circles



log d - log n = log K

total Possible key size f

total Possible encoded

bits

Packing Lemma (Devetak): if we send only $k = \frac{d}{n}$ Vectors, then $I_{acc}(P) = log k$

All of it / No locking!

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"Mule Rescue": Information Complementarity

Alice's codebook

Howking radiation of Shir codebook

Alice needs the codebook to "interpret" the state

But the Codebook & bigger than the black page 77/81 hole!

Conclusions and future work

The quantum "locking" effect helps solve the black hole information loss paradox.

Information can escape only at the end of the evaporation, but the small black hole need not hold very much.

Generalize locking:

Symmetric locking Locking other forms of information

Deal with coding problem—towards a full quantum gravity

