Title: General Relativity 6 - The Star Called a Black Hole

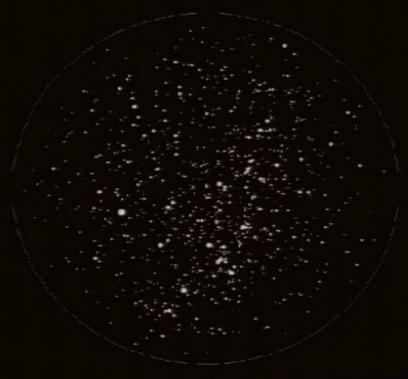
Date: Aug 02, 2006 09:15 AM

URL: http://pirsa.org/06080034

Abstract: The physical attributes of a black hole and what types of physical evidence astronomers use the locate them.<br/>
br> Learning Outcomes:<br/>  $\hat{a} \in \phi$  What are the physical requirements for a star to become a black hole, and what properties of that star remain after the black hole is formed?<br/>
formed?<br/>
The types of black holes, including: the Schwarzschild black hole, the Reissner-Nordström black hole, the Kerr black hole, and the Kerr-Newman black hole.<br/>
Frage What a traveller would experience if he orbited one of these more general black holes, or fell through to the singularity.

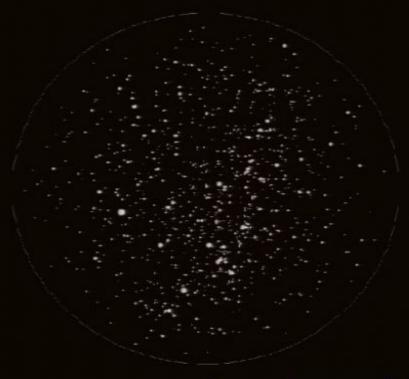
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# Looking Up



$$r = 12 \tan \left( \frac{5\pi}{6} \sqrt{1 - \frac{C_b}{C}} \right)$$

# Looking Up



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### Looking Up



$$r = 12 \tan \left( \frac{5\pi}{6} \sqrt{1 - \frac{C_b}{C}} \right)$$



Sideways

Backwards

Thanks to Ute Kraus, Mass-Planak-Institute

Forward

Sideways

Backwards







r = 100 rs

Thanks to Ute Kraus, Max-Planak-Institute

Forward



Backwards







 $r = 20 \, rs$ 

Thanks to Ute Kraus, Max-Planak-Institute

Forward



Sideways







 $r = 4.5 \, rs$ 

Thanks to Ute Kraus; Max-Planak-Institute

Sideways

Backwards





 $r = 2.5 \, rs$ 

Thanks to Ute Kraus, Max-Planak-Institute

Sideways

Backwards





r = 1.5 rs

Thanks to Ute Kraus, Max-Planak-Institute

Sideways

Backwards





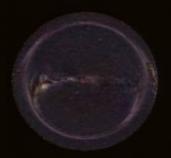
 $r = 1.2 \, rs$ 

×

Thanks to Ute Kraus, Masc-Planck-Institute

Sideways

Backwards



 $r = 1.05 \, rs$ 

\*

Thanks to Ute Kraus, Masc-Planak-Institute

Sideways

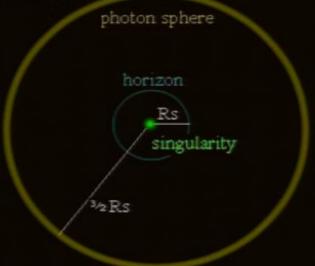
Backwards



 $r = 1.005 \, rs$ 

Thanks to Ute Kraus, Mass-Planck-Institute The Anatomy Summary

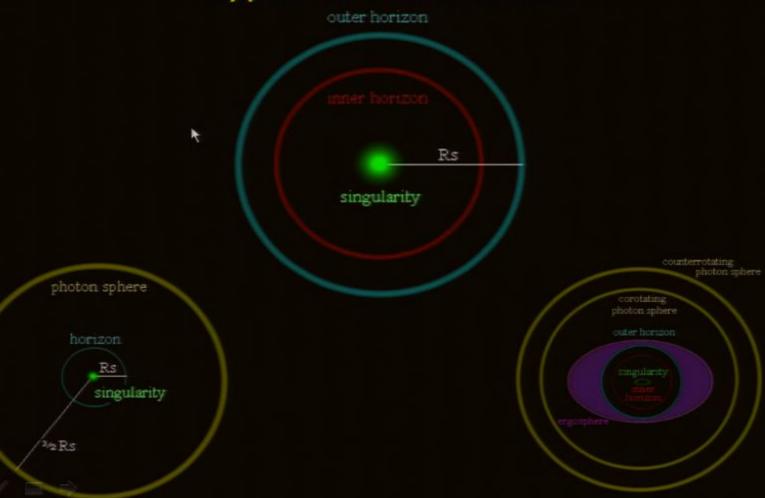
If you calculate the size of an object whose escape velocity is the speed of light, you get the "Schwarzschild radius", which defines the "event horizon". This is the formal size of a black hole (even though there is nothing at that location). It is given by R<sub>s</sub>=3km(M<sub>\*</sub>/M<sub>sun</sub>). It is the horizon over which you can see no more events. Outside that at 1.5 R<sub>s</sub> photons would orbit the hole (the photon sphere).



Far from the hole, the gravity is the same as it would be if the star were still there (so no "vacuum cleaner" effect). If the Sun collapsed to a BH, the Earth's orbit would be unaffected.

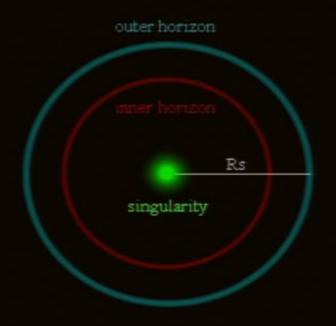


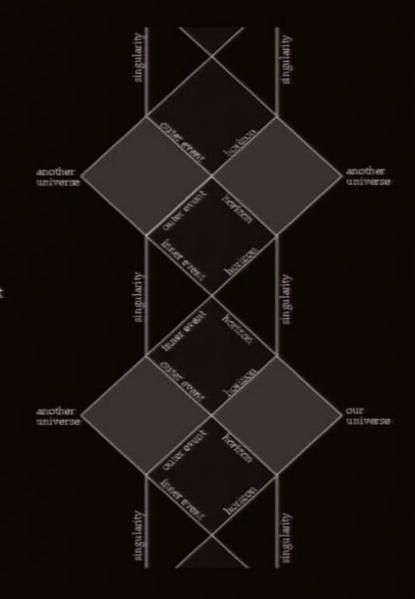
#### Types of Black Holes



#### Reissner-Nordström Black Hole

An electrically Charged Black Hole





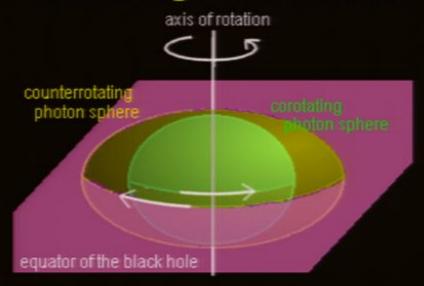


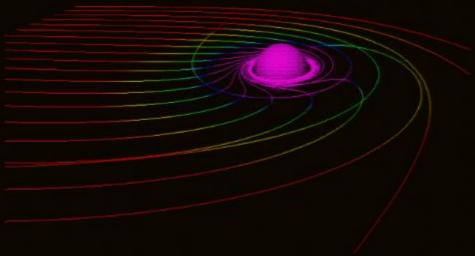
# corotating photon sphere outer horizon singularity miner horizon ergosphere

Rotation also leads

to "frame-dragging"

#### Rotating Black Holes

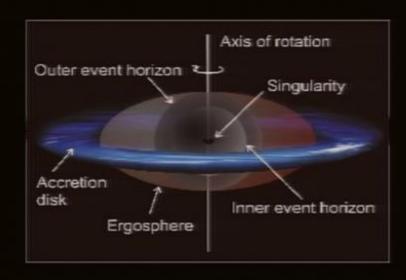


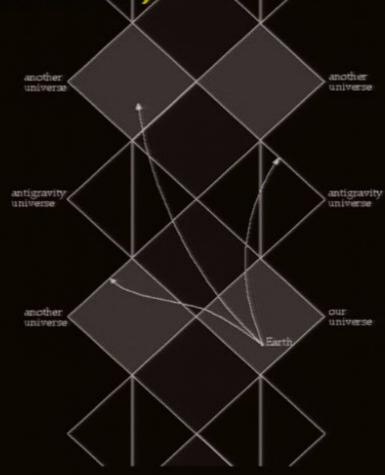




A Rotating Black Hole (The Kerr Black Holes)

Why is it called a rotating black hole? The event horizon doesn't rotate---it's just a boundary-line

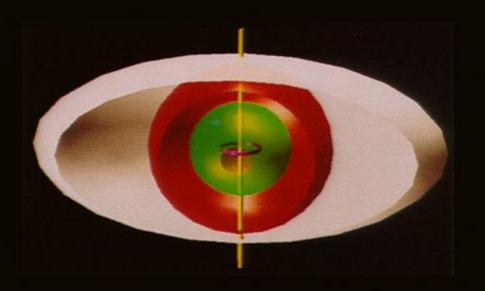






#### Kerr-Newman Black Hole

Same Structure as Kerr Black Hole. But now it has a charge as well as a rotation.







#### No-hair theorem (well maybe 3 hairs)

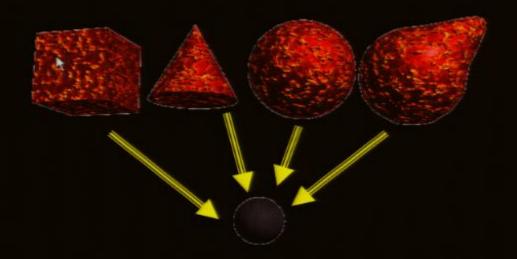
- A black hole has no hair; its only 'hair' are its
  - mass
  - angular momentum
  - electric charge





John A. Wheeler (51911)

### Stars of all Shapes and Sizes





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#### Types of black holes

- Schwarzschild (1916)
  - mass
- Reissner-Nordström (1916, 1918)
  - mass, electric charge
- Kerr (1963)
  - mass, angular momentum
- Kerr-Newman (1965)
  - mass, angular momentum, electric charge



#### Entropy of Black Hole

- •Black hole presents us with a problem: What happens to the information when a particle falls inside a Black Hole?
- Remember only 3 parameters are required to describe a Black Hole (charge, mass, and angular momentum).
- In order to describe a physical system, we need entropy (a measure of disorder).
- Hawking had no problem with this "entropy eater"
- Hawking (after changing mind) and Bekenstein produced laws of Black Hole mechanics that bore an amazing resemblance to laws of thermodynamics.
- •The 2nd law of thermodynamics "Entropy (randomness) increases
- You replace "Horizon Area" with "Entropy"

.

$$S(entropy) = \frac{kAc^3}{4hG}$$



#### Entropy of Black Hole

New problem: if the Black Hole has an entropy, it must have a temperature too.

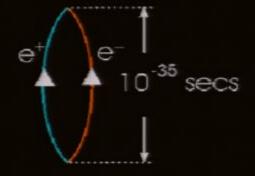
Worse Problem: If is has a temperature it must radiate, but in classical definition, nothing can escape a Black Hole.

#### Crisis:

- · Several ways to picture how a black Hole evaporates
  - •Some ways correspond to different ways of formulating laws of quantum fields.
  - Some ways correspond to String Theory
  - •Some ways correspond to Quantum Gravity
  - •Some ways make no sense at all.

#### Vacuum Fluctuations

- A vacuum is a place which is anything but empty.
- •It is a place of continuous creation and destruction. Pairs of virtual particles are born live a short but happy life then die.
- •Possible by Heisenberg's uncertainty principle: The energy of a vacuum, that we suppose to be zero, can be defined with an uncertainty of  $\Delta E$  during at  $\Delta T$ .
- ATX AE ≈h
- •Therefore particles/antiparticles, with  $\pm \Delta E$  are constantly being created.
- One particle has positive energy, one particle has negative energy.
- •The particles live momentarily on fluctuational energy "borrowed" from neighbouring regions of space.



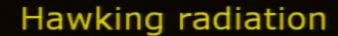


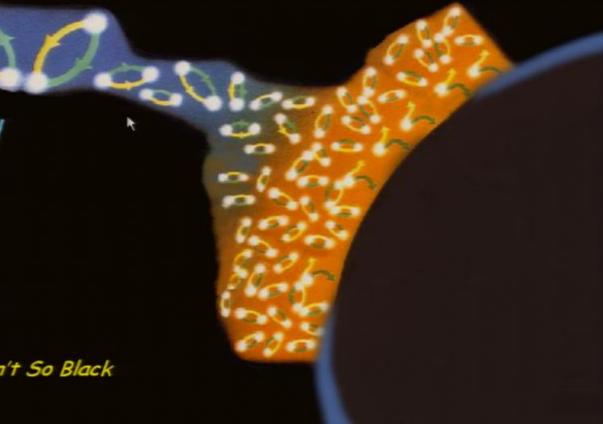
#### Stephen W. Hawking (b1942)

Virtual photon is its own antiparticle

virtual particles

Black Holes Ain't So Black





#### Hawking Radiation

- The Hawking Radiation theory states that virtual particle-antiparticle pairs are sometimes created outside the event horizon of a black hole. Three things can happen to a pair of particles just outside the event horizon:
  - Both particles are pulled into the black hole.
  - Both particles escape from the black hole.
  - One particle escapes while the other is pulled into the black hole.
- For the third possibility, the particle that has escaped becomes real and can therefore be observed from Earth. The energy to separate the two virtual particles (thus making them real particles) is taken from the horizon, thus reducing the energy of the Black Hole.
- The wavelength of the particle/wave that enters the a hole will be of 25% of the hole's circumference.
- For Example: A black hole of 2 Solar Masses with a circumference of about 35K will emit a wavelength of

$$\frac{35}{4} \approx 9 \text{ km}$$



#### Hawking Radiation

•The larger (more massive) the Black Hole the lower the temperature and the longer it takes to evaporate.

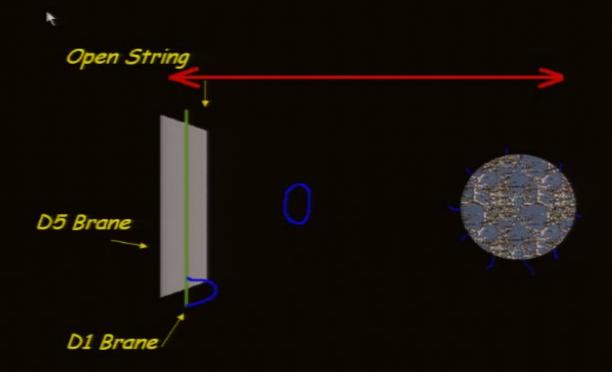
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$$Temp_K \approx \frac{6 \times 10^{-8}}{M_{\oplus}}$$

$$\tau ime \approx 10^{66} \cdot \left[ M_{\oplus} \right]^3$$

•Remember the age of the Universe is 10 10 years give or take 3 days.

#### Superstring Method



One of the most dramatic recent results in string theory is the derivation of the Bekenstein-Hawking entropy

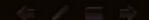


#### Are Black Holes Real

What are we going to look for if Black Holes are Real

#### Observational Evidence for Black Holes

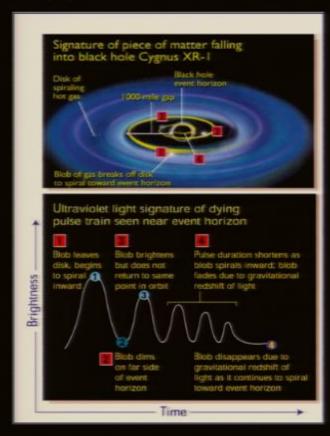




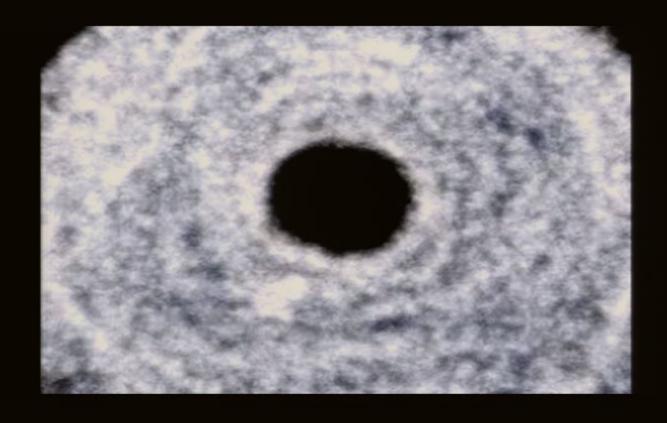
# Finding Black Holes Ultraviolet and X-rays Seeing Matter Disappear

Hubble observed pulses of UV light emitted by material as it fell into a black hole.

- Pulses arise from material orbiting around intense gravity of the black hole.
- Light pulses, lasting 0.2 s, are red-shifted from X-ray to UV, as they fall into gravity of the black hole.



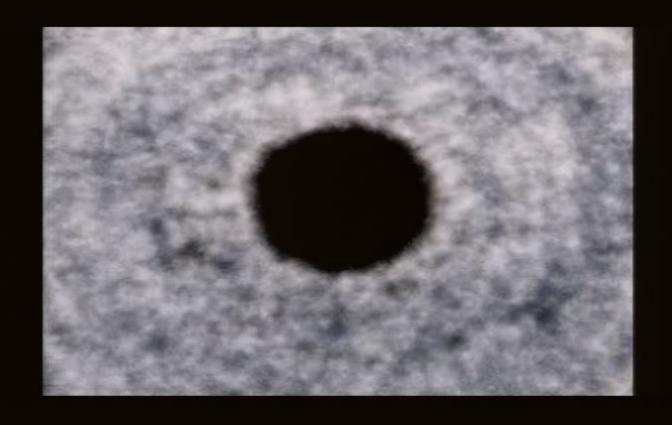




 $\cdot$  Zel'dovich speculated that if a black hole was surround by gas and this gas was orbiting the black hole in an accretion disk, it Would give off x-rays

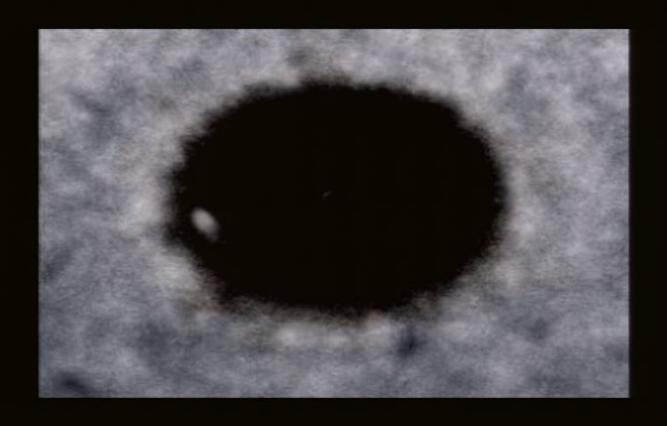
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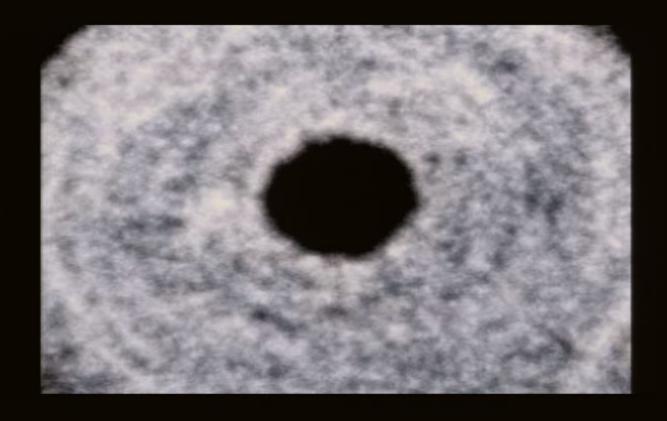
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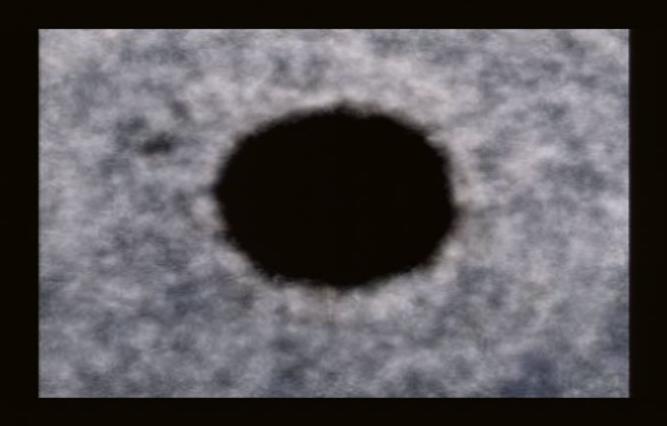
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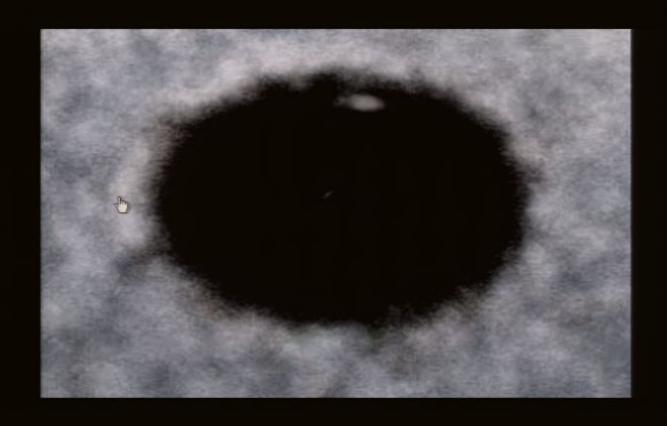
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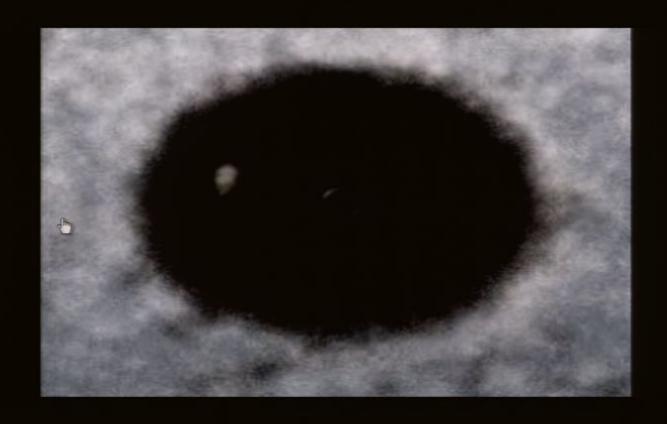
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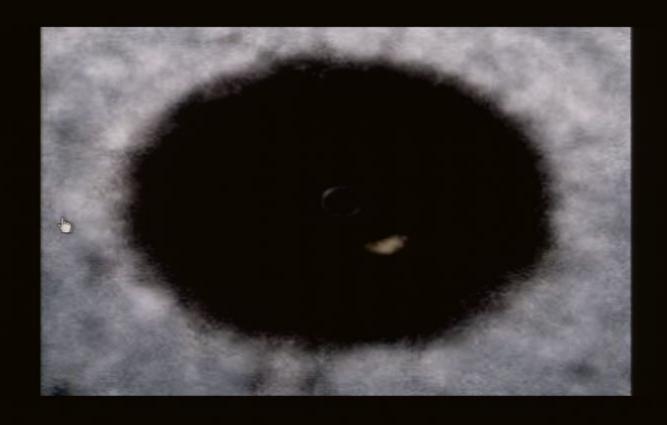
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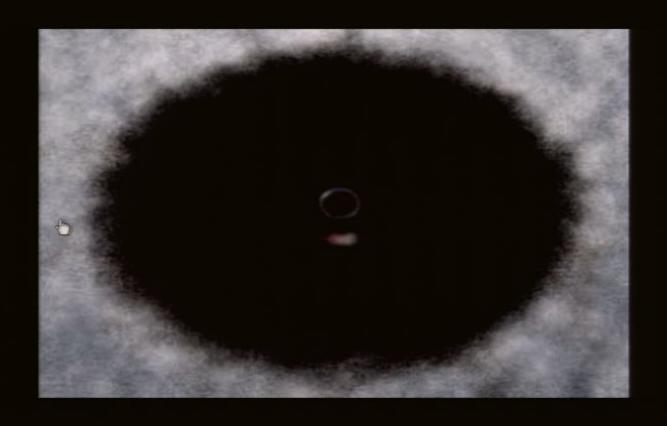
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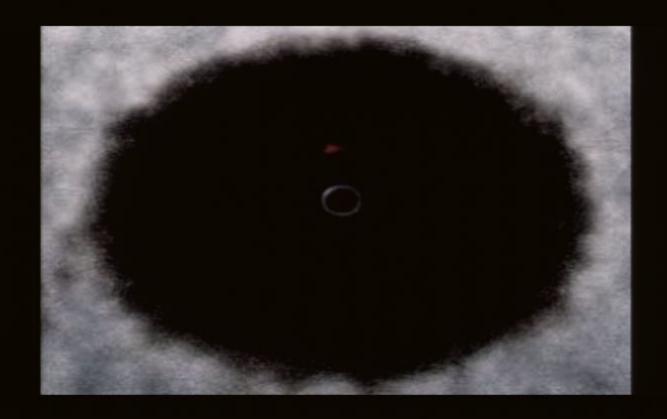
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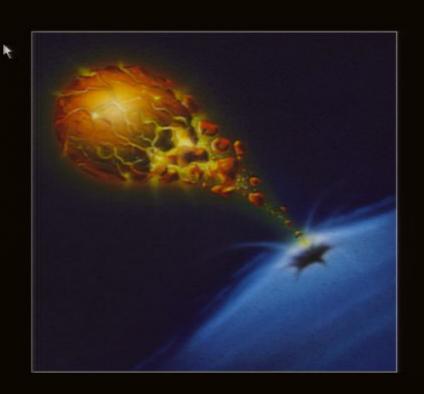


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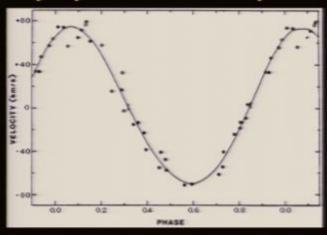


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### Seeing Holes

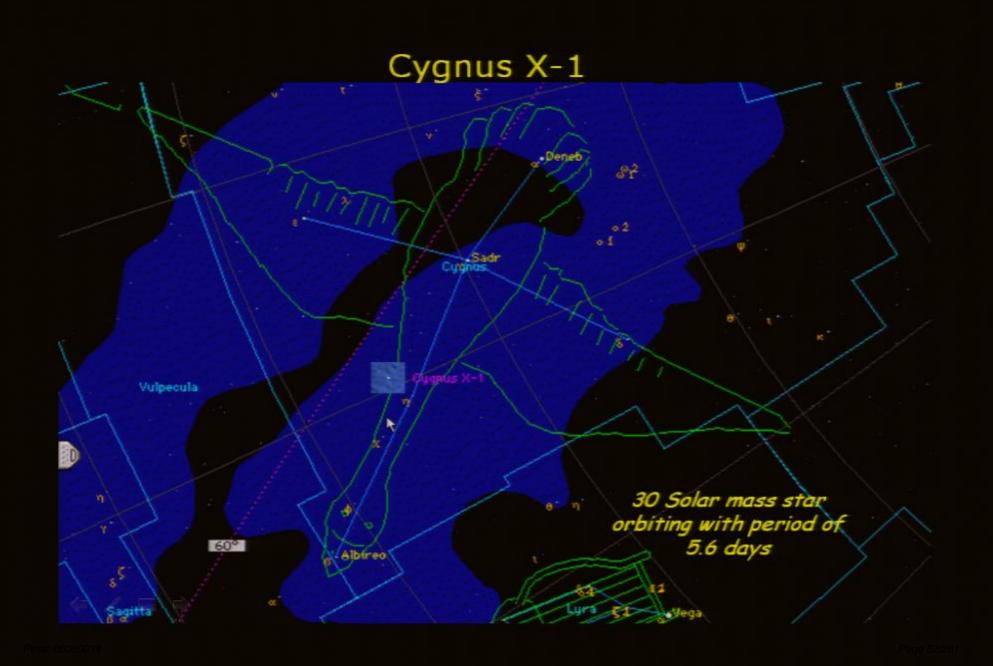


- Can't see black hole itself, but can see matter falling into a hole.
- Gravitational forces stretch and rip matter: heats up.
- Very hot objects emit in Xrays (interior of Sun)





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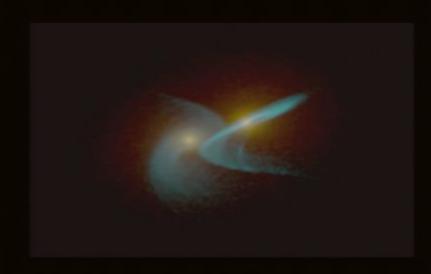












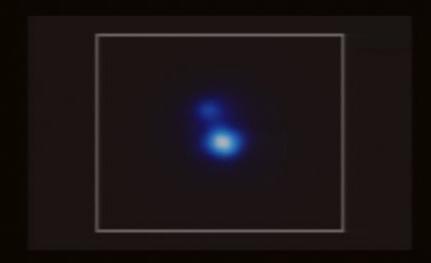
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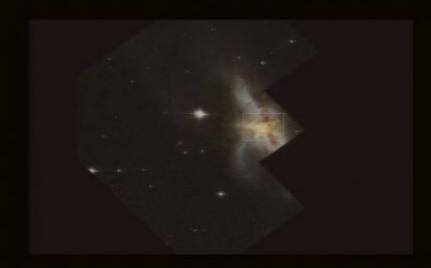






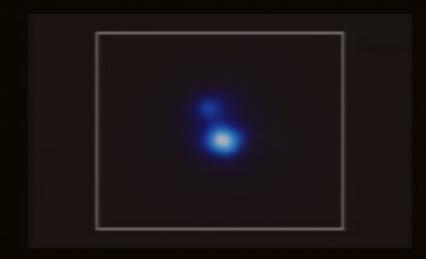




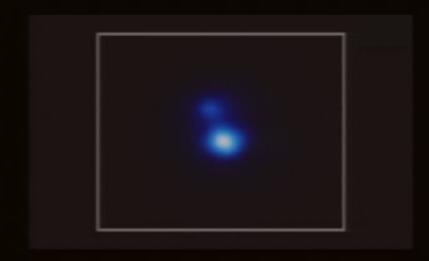


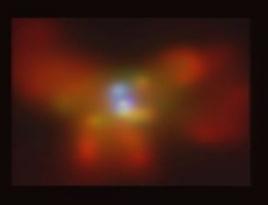


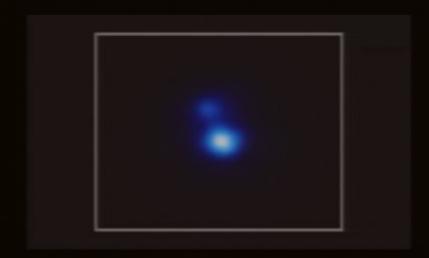


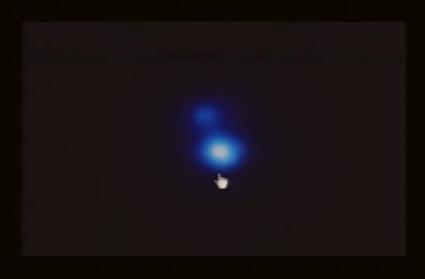


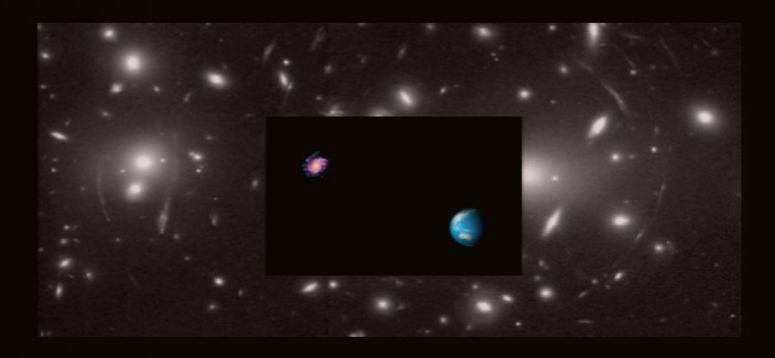


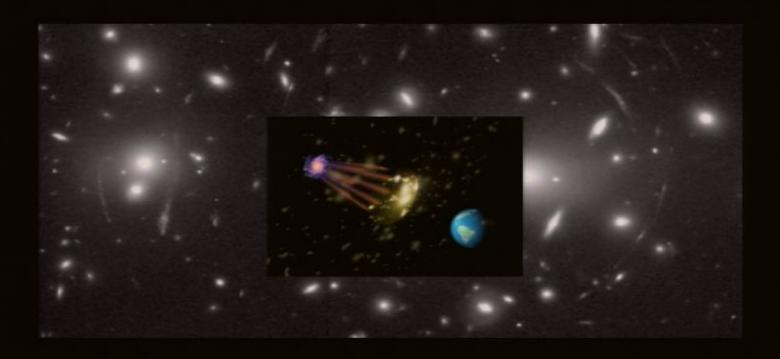


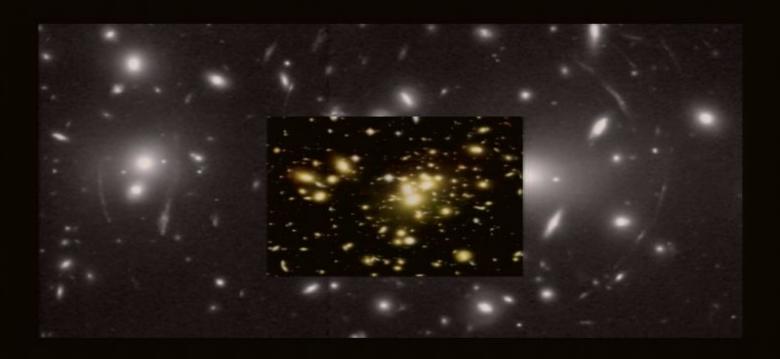




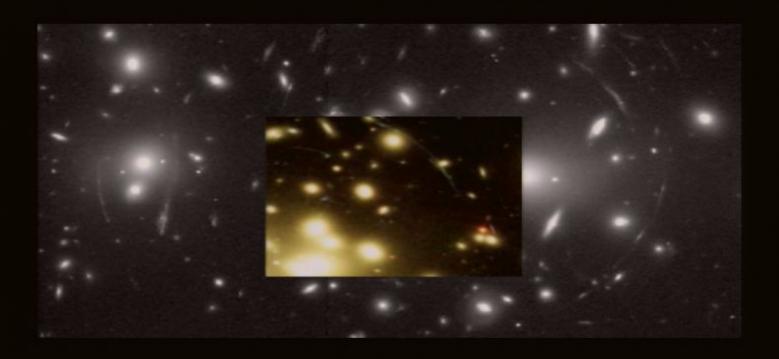


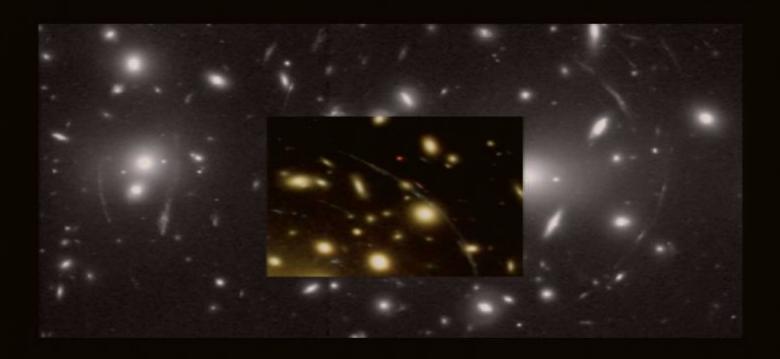


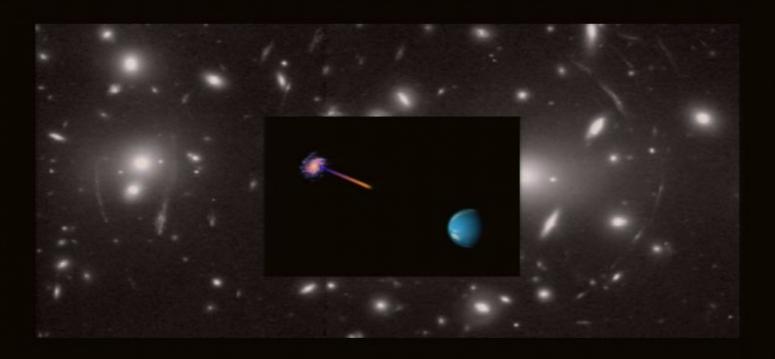


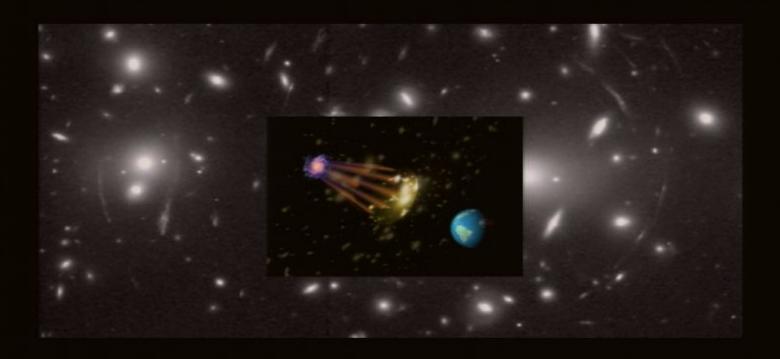


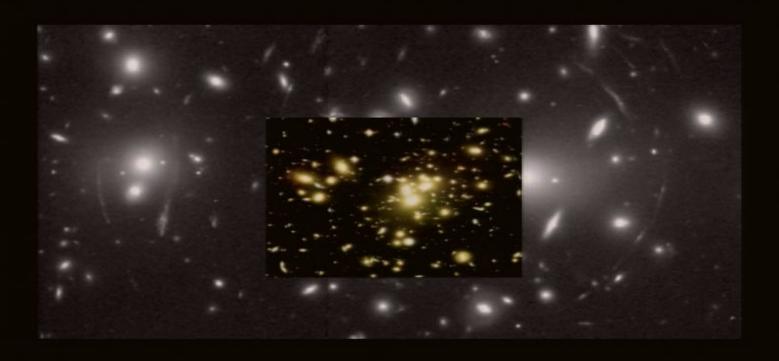


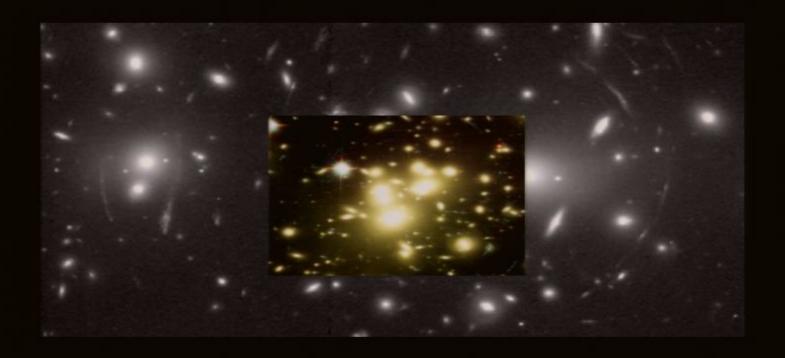






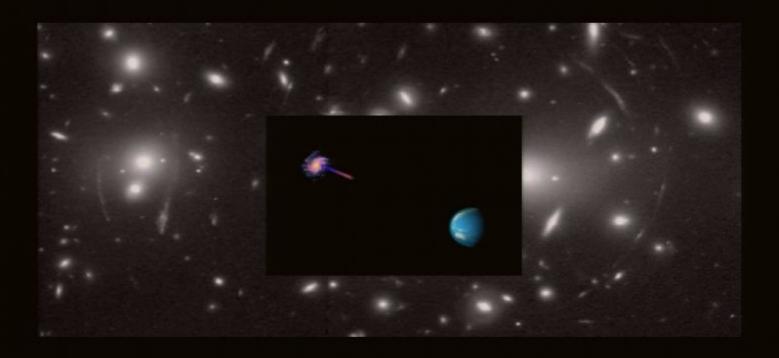


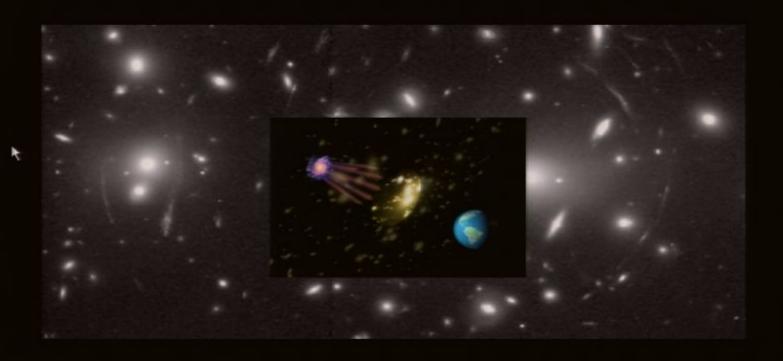












#### What else should we look for?



Jets



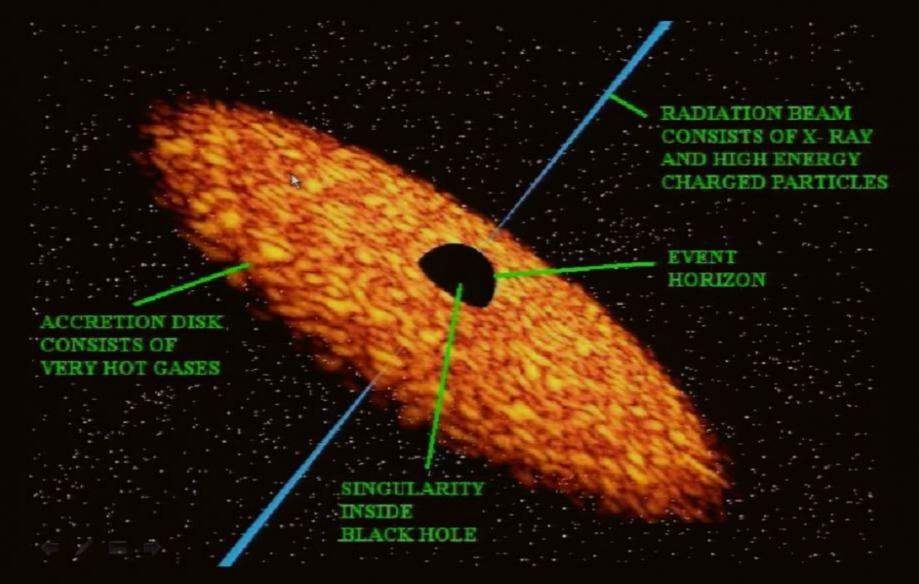
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#### What else should we look for?



Jets





#### Radio Jets from Black Holes

- Many black holes emit jets.
  - Material in jet moving at 0.9c.
  - Jet likely composed of electrops and positrons.
- Magnetic fields surrounding black hole expel material and form the jet.
  - Interaction of jet material with magnetic field gives rise to Radio emission.



## Jets

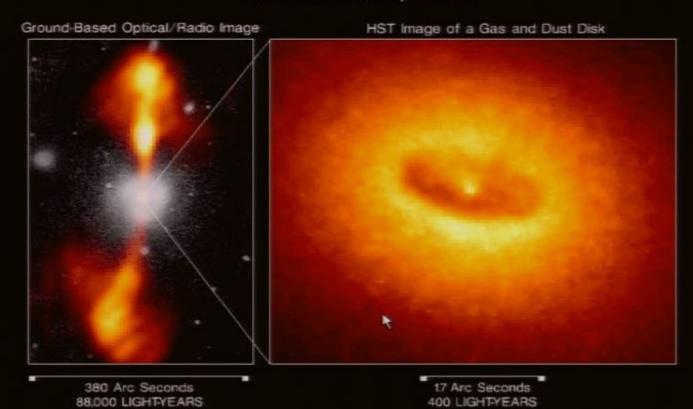




#### Core of Galaxy NGC 4261

#### Hubble Space Telescope

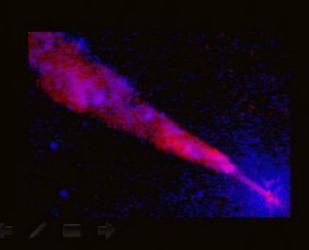
Wide Field / Planetary Camera

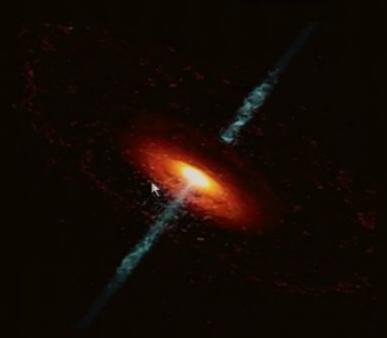


## Jets





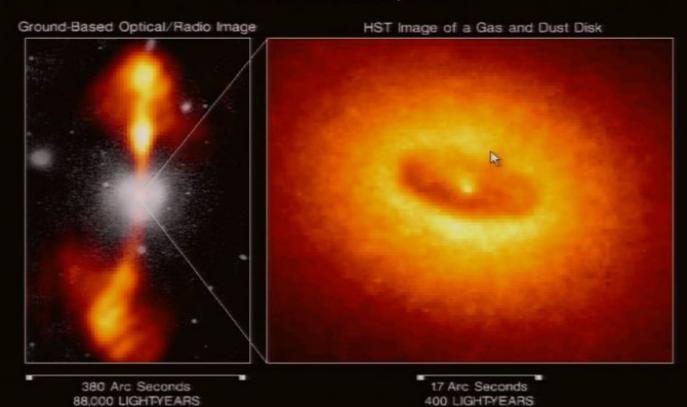




#### Core of Galaxy NGC 4261

#### Hubble Space Telescope

Wide Field / Planetary Camera





Thanks to Chandra Team at Nasa









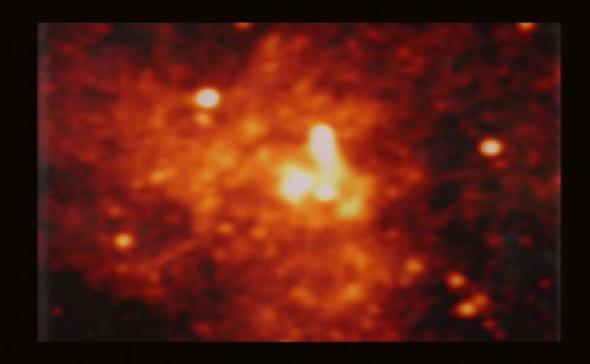
What about inside our own Galaxy?

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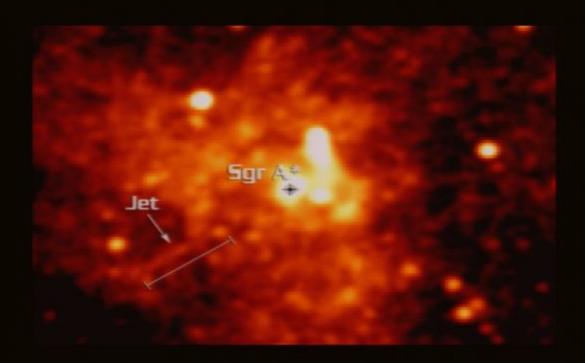


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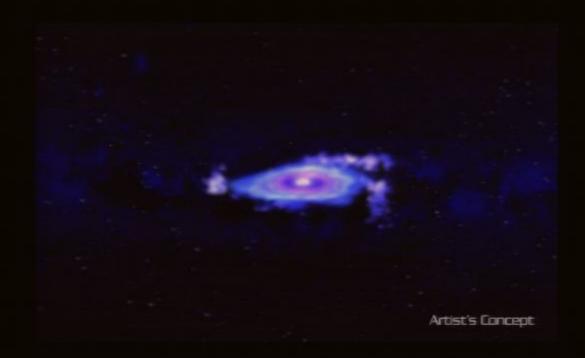


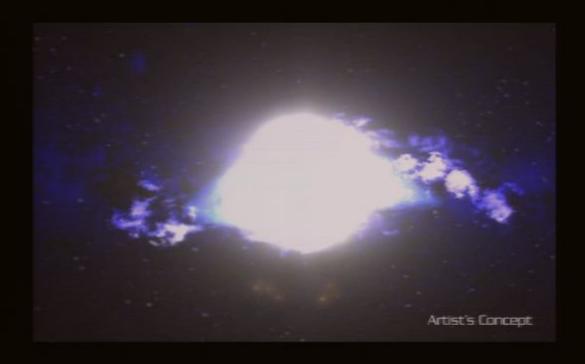
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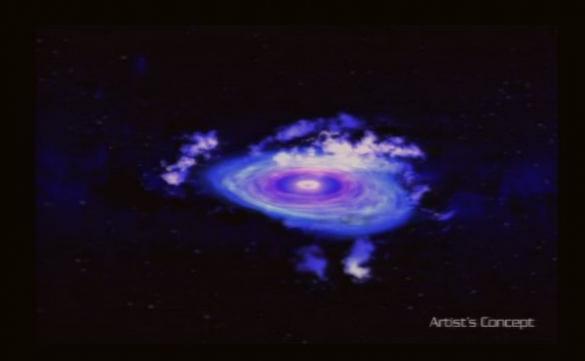


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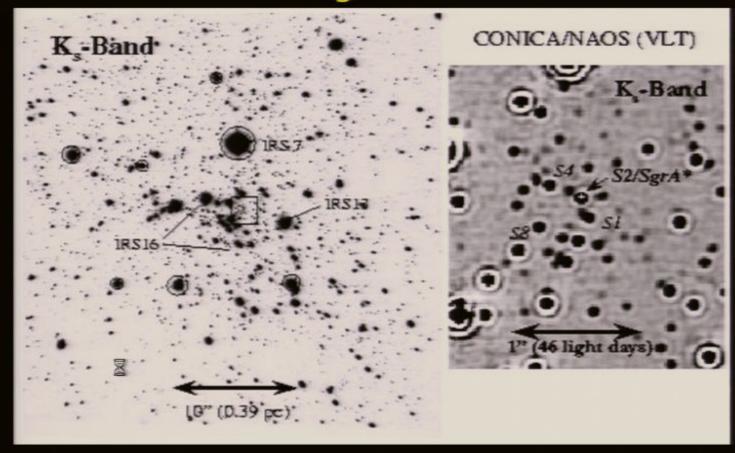






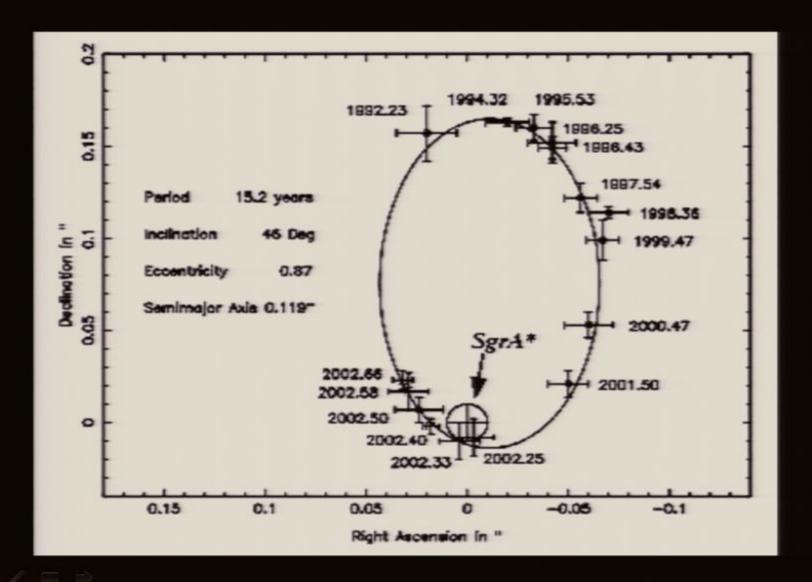






3,700,000 solar masses

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1992.6



1993.4



1994.3



1995.2



1996.4



1997.2



1998.2



1999.2



2000.2



2001 10 light days



2002

2002.9

1992

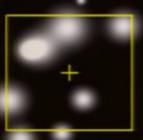


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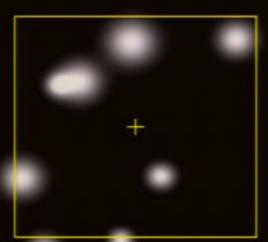


1992





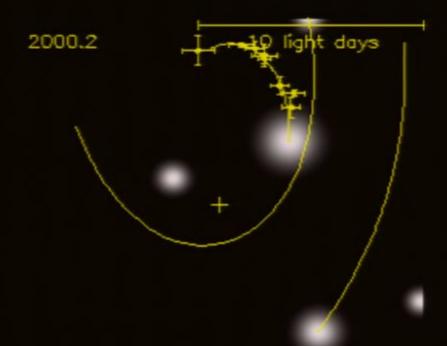
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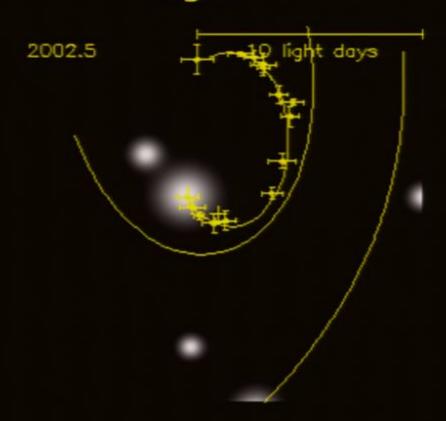


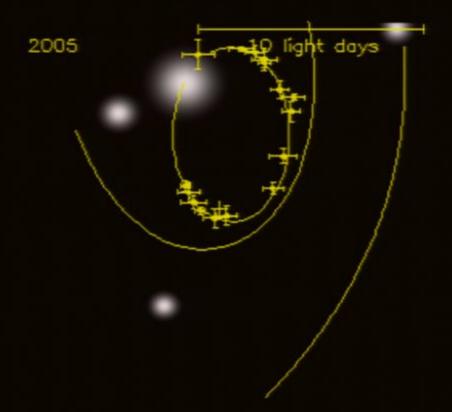


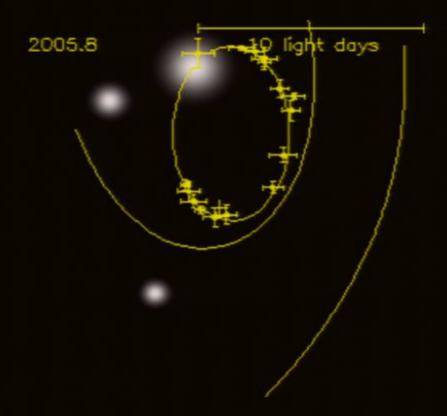


# SgrA\* 1997.8 Handle Sight days Handle Sight days









1992.8

1993.8

1994.6

1995.7

1996.7

1997.6

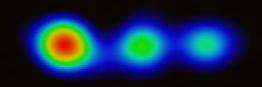
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2000

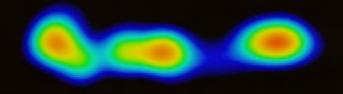
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2002

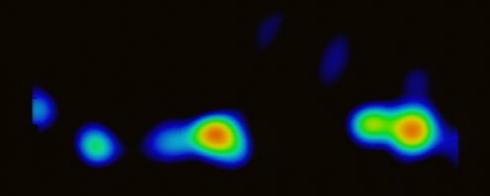


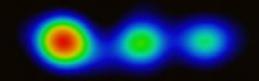


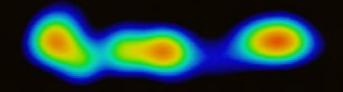


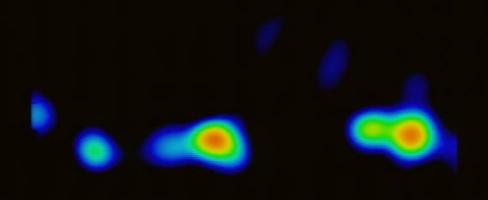






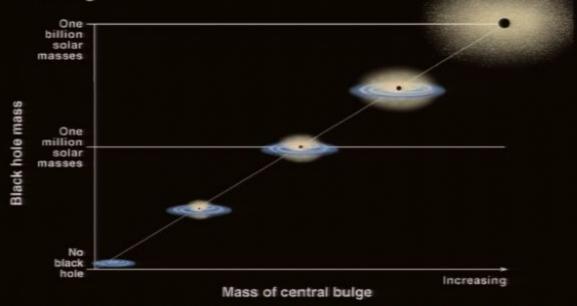






#### Speed of Gas and Black Holes

Correlation Between Black Hole Mass and Bulge Mass



- It discovered a correlation between a Black Hole's mass and the average speed of the stars in the galaxy's central bulge.
- The faster the stars are moving, the larger the black hole.
- The central Black Hole comprises 0.5% of mass of stars in the spheroid of the galaxy.
   (Magorrian Relation)
- Previously, black holes were seen as the endpoints of evolution, the final resting state of
  most or all of the matter in the universe. Now we believe black holes also play a critical role
  in the birth of galaxies."

×







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Pires: 06/8/0034



Pene 15/266



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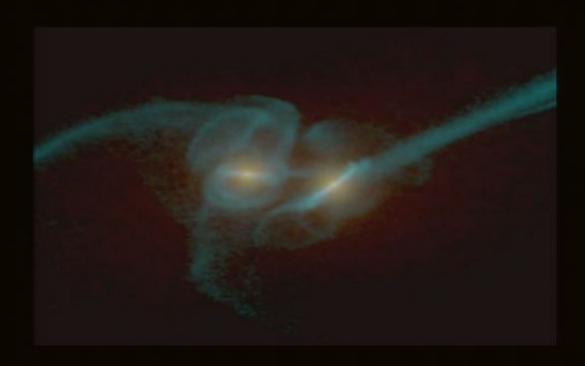
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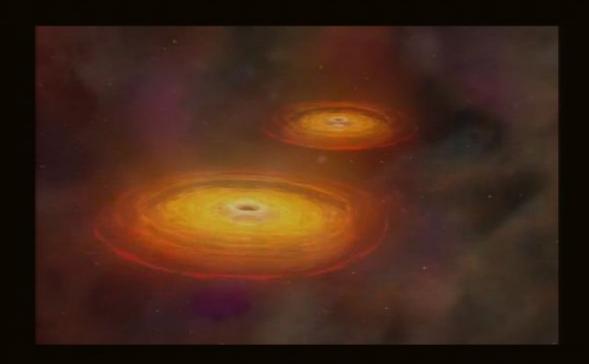
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Pirea: 06080034



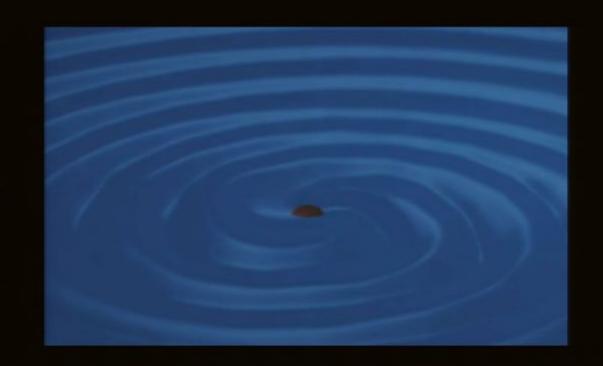
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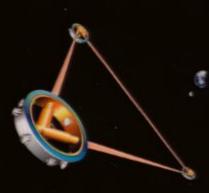




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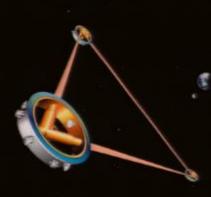


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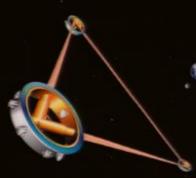
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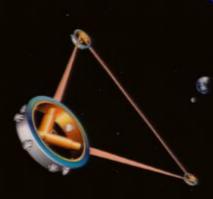


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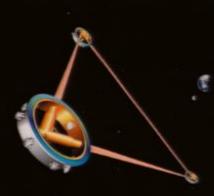


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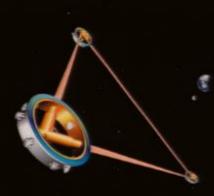


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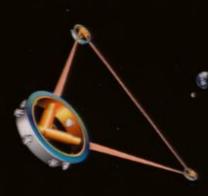
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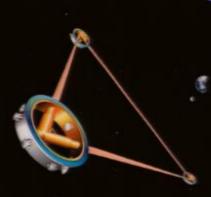
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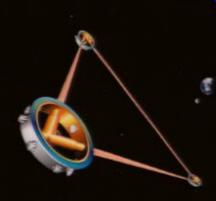
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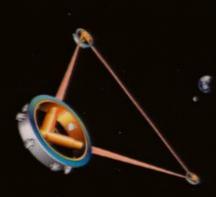
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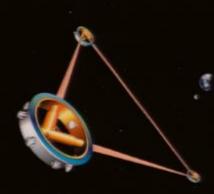
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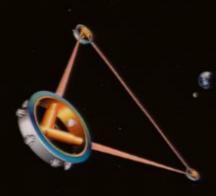
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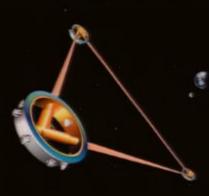
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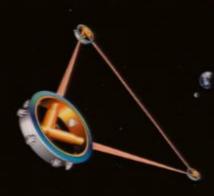


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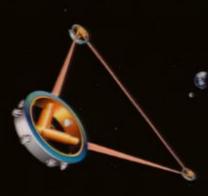
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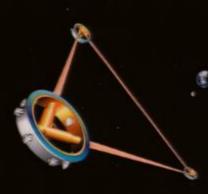








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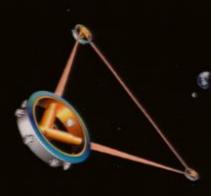
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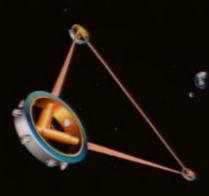


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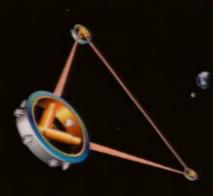




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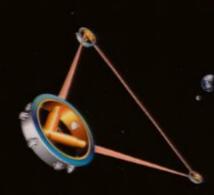




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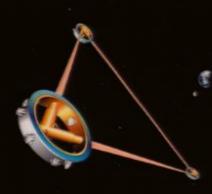
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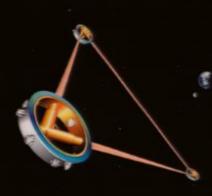




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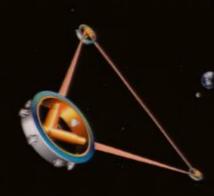
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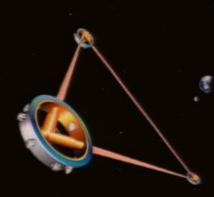


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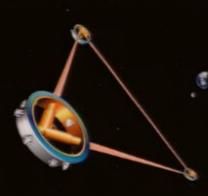




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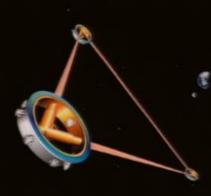




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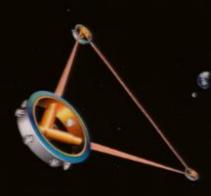
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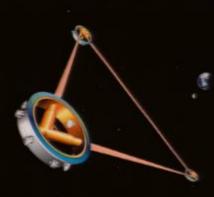
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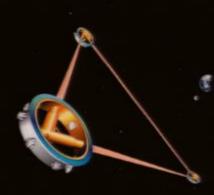




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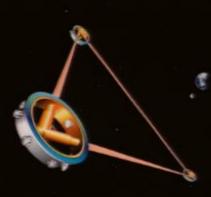


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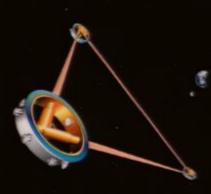




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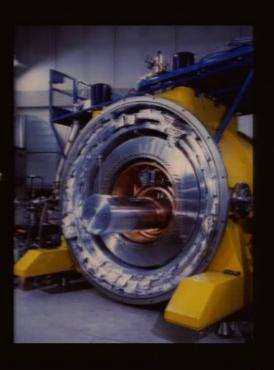




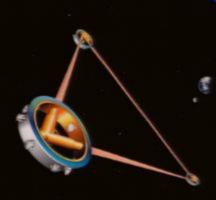


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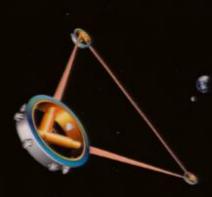


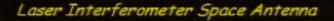
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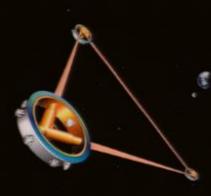




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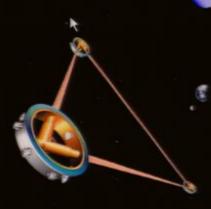
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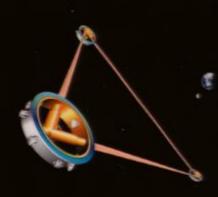




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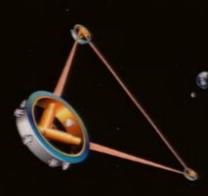
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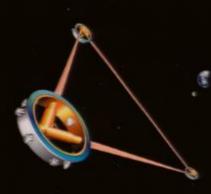
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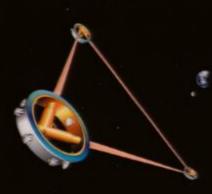
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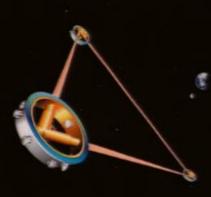


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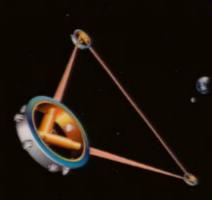




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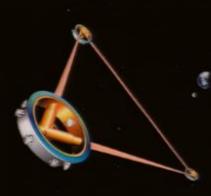
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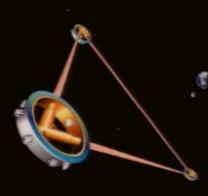
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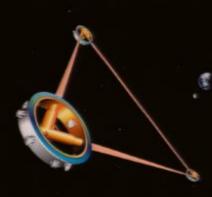
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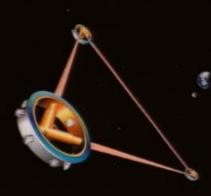
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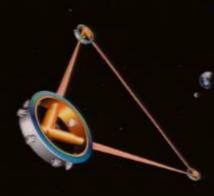


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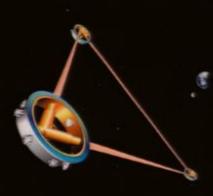




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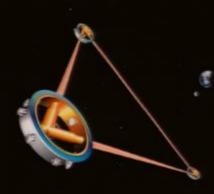




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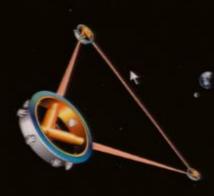


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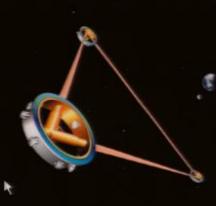


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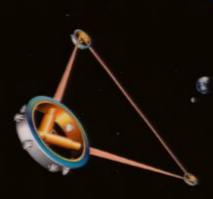
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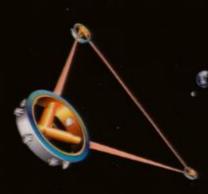
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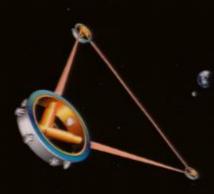




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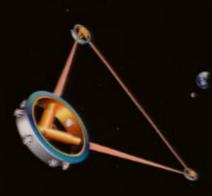




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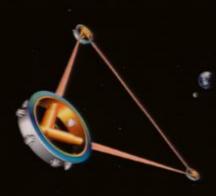


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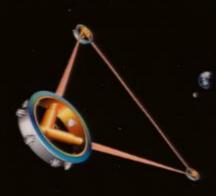
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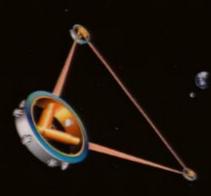




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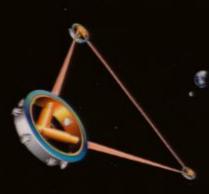




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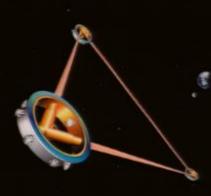








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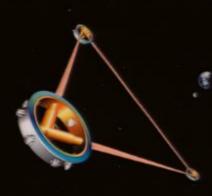
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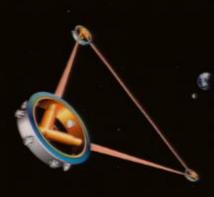
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Laser Interferometer Gravitational Wave Observatory (LIGO), Richland



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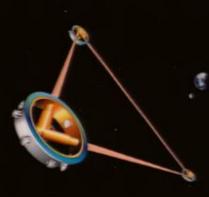




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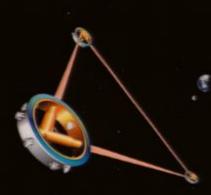








Nautilus



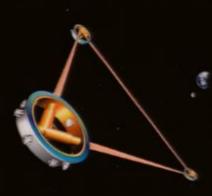
Laser Interferometer Space Antenna



Laser Interferometer Gravitational Wave Observatory (LIGO), Richland



Nautilus



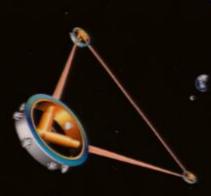


Laser Interferometer Space Antenna

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Nautilus





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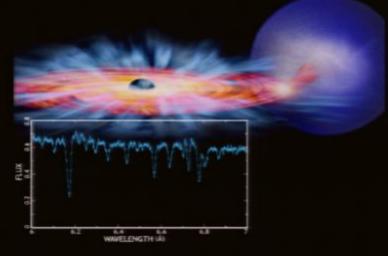


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### Strange Energy Out-Bursts

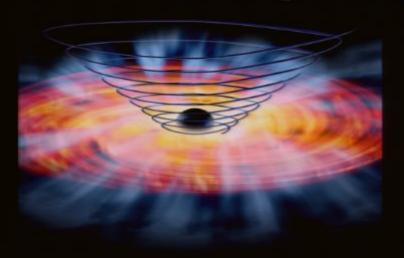


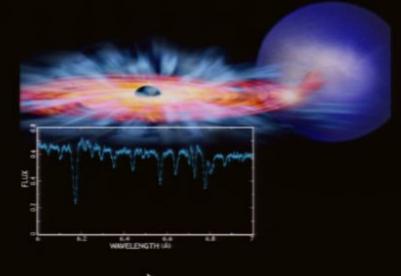


1655 is a binary system that harbors a black hole with a mass seven times that of the sun, which is pulling matter from a normal star about twice as massive as the sun. The Chandra observation revealed a bright X-ray source whose spectrum showed dips produced by absorption from a wide variety of atoms ranging from oxygen to nickel. A detailed study of these absorption features shows that the atoms are highly ionized and are moving away from the black hole in a high-speed wind.



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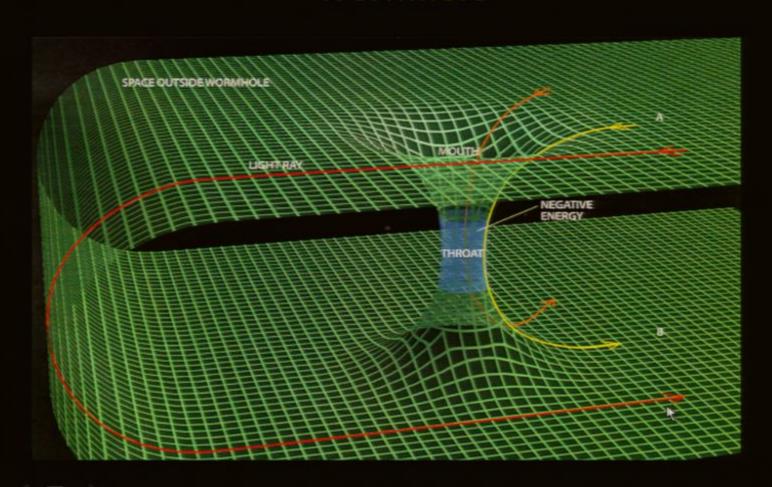




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# Wormhole



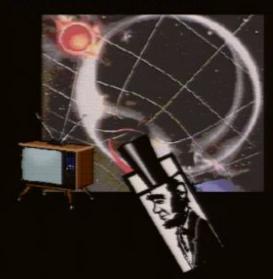
Wormholes are tunnels that connect two areas of space

Black hole White hole.

Can wormholes lead to time travel?

### Naked Singularity

Such a naked singularity would be a breakdown in the laws of physics. After that, you could no longer guess what would come out of the black hole--it could be anything (to quote William H. Press) "from television sets to busts of Abraham Lincoln."



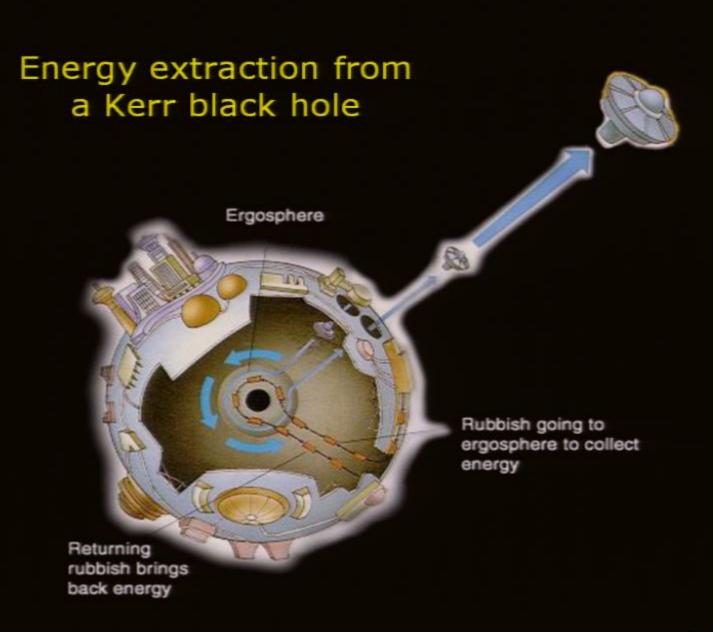
A singularity that is not inside a black hole (not surrounded by an event horizon), and therefore can be seen by someone outside it.



# Time Travel



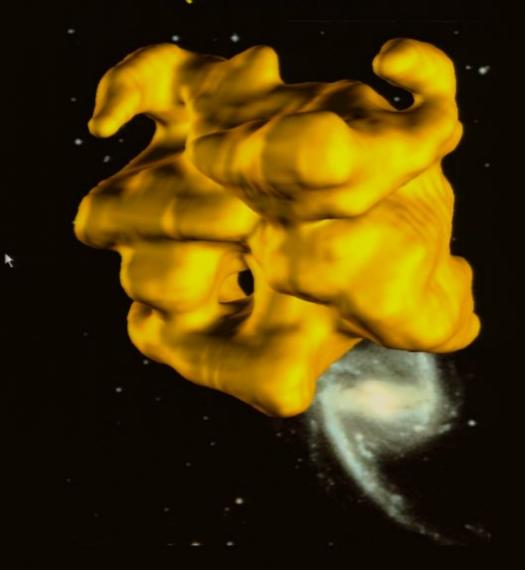






Roger Penrose (b1931)

# Quantum Foam





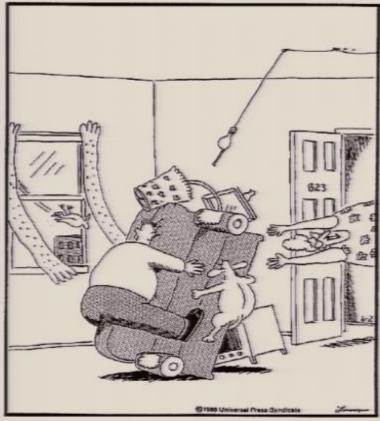






### THE FAR SIDE

### By GARY LARSON



Suddenly, through forces not yet fully understood, Darren Belsky's apartment became the center of a new black hole. End of slide show, click to exit.

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