

Title: Is the emergence of macroscopic behavior due to a quantum mechanical indeterminacy of the structure of the Einsteinian Space-Time?

Date: Jul 19, 2006 05:15 PM

URL: <http://pirsa.org/06070048>

Abstract: According to a widely accepted view, the emergence of macroscopic behavior is related to the loss of quantum mechanical coherence. Opinions on the possible cause of this loss diverge. In the present talk it will be shown how a small, assessable amount of indeterminacy in the structure of space-time may lead to the emergence of macroscopic behavior, in agreement with empirical evidence.

Is the Emergence of Macroscopic Behavior due to
a Quantum Mechanical Indeterminacy of the Einsteinian &
Space-time Structure?

Károlyházy (1964-72)

Cell length a_c of a single microparticle

$$a_c \approx \left(\frac{L}{\lambda} \right)^2 L \quad \left(L = \frac{h}{cm} \right)$$

$$e \begin{pmatrix} 1 \\ 2 \\ 3 \end{pmatrix}$$

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Space-time Structure?

Is the Emergence of Macroscopic Behavior due to
a Quantum Mechanical Indeterminacy of the Einsteinian Λ
Space-time Structure?

Subtitles:

2. A Compromise between Quantum Mechanics and General Relativity
 3. Lower Bound of the Indeterminacy in the Length of a Time Interval, and in the Synchronization of Moments of Time
 4. A Tentative Expression of $(\Delta_L T)_{\text{syn}}$ through Vacuum Expectation Values
 5. Indeterminacy in the Relative Phases of a Quantum State
 6. The Coherence Cell and the Cell Length of a Single Microparticle
 7. The Coherence Cell and the Cell Length of Homogeneous, Spherical Solid Bodies
 8. The Transition from Quantum (micro) to Classical (Macro) Behavior
-
- Taking into Account the Surroundings

- Submacroscopic Decay of a Superposition in a Gas
 - A Remark on Superconductivity
 - Anomalous Brownian Motion
 - Smoothing the jumps (GRW)
 - Comment on the Non-Choice of a Preferred Basis
15. Concluding Remark

2. A Compromise Between

GR

$$\Delta x \cdot \Delta p \geq \frac{\hbar}{2}$$

Δ in the structure
of space-time

$$\Delta \phi(x, x', t)$$

$$\phi(x, x', t) = \psi(x', t) - \psi(x, t)$$

in the Ψ of any physical
system.

$$x = (x_1, \dots, x_N)$$

N : number of microparticles
constituting the system

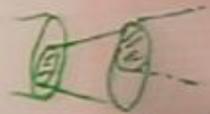
If $\Delta \phi(x, x', t) \geq \pi$, the coherence between
 x' and x is lost.

Massive bodies influenced strongly, microparticles
weakly

When the Ψ of a massive body expands in
the course of its Schrödinger time evolution,
non-overlapping, incoherent domains begin
to appear in the configuration space.

Károlyházy assumed that if this happens, then

ψ stochastically localizes itself to one of the domains inside which $\Delta\phi \leq \pi$, thereby counterbalancing the deterioration of its coherence. From that domain ψ expands again as dictated by the Schrödinger equation, until the deterioration of its coherence calls for the next self-localization, and so on.



A stochastic self-localization triggered by the appearance of non-overlapping incoherent domains results to a breakdown of the superposition principle independently of the localization. This is a non-unitary component of the time evolution as objective as the stochastic Schrödinger equation.

It turns out that for domains composed of a moderate number of particles $\Delta\phi$ remains much smaller than π and no self-localization is required.

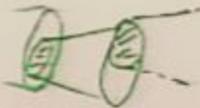
Ψ stochastic localizes itself to one of the domains inside which $\Delta\phi \leq \pi$, thereby counterbalancing the deterioration of its coherence. From that domain Ψ expands again as dictated by the Schrödinger equation until the deterioration of its coherence calls for the next self-localization, and so



A stochastic self-localization triggered by the appearance of non-overlapping incoherent domains leads to a breakdown of the principle of independence of observation. This is a non-deterministic component of Ψ , as objective deterministic

systems
 states composed
 of particles $\Delta\phi$
 incoherent
 self-local-

$\Delta\phi \leq \pi$,
 ...variance, the deteriora-
 tion of its coherence. From that domain
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A stochastic self-localization triggered by
 the appearance of non-overlapping inco-
 herent domains amounts to a **breakdown**
 of the superposition principle inde-
 pendently of any observation. This is
 a non-unitary, stochastic component
 of the time evolution of Ψ , as objec-
 tive as the unitary, deterministic
 Schrödinger evolution.

It turns out that for microsystems
 (i.e. the microsystems, and for systems composed
 of a large number of microparticles) $\Delta\phi$
 remains small, $\Delta\phi \ll \pi$, incoherent
 domains do not develop, no self-loc-
 alization occurs.

Schrödinger, until the acceleration of its coherence calls for the next self-localization, and so on.



A stochastic self-localization triggered by the appearance of non-overlapping incoherent domains amounts to a **breakdown** of the superposition principle independently of any observation. This is a non-unitary, stochastic component of the time evolution of Ψ , as objective as the unitary, deterministic Schrödinger evolution.

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It turns out that for microsystems (for the microparticles, and for systems composed of a moderate number of microparticles) $\Delta\phi$ remains much less than π , incoherent domains do not develop, no self-localization is needed.

5

This is how QM and GR seem to cooperate. to preserve the coherence of the Ψ 's of the microparticles, and to prevent the excessive expansion of the Ψ 's of macroscopic bodies. By the same token the development of large indeterminacies in the space-time structure is also prevented. This balance is reached by giving up the perfect sharpness of the structure of space-time in GR, and the absolute validity of the superposition principle in QM.

3. Lower Bound on the Indeterminacy in the Length of the Time Interval and in the Position of Moments of Time

Studying the indeterminacies in the length of the time interval put into an object (non-relativistic) space-time, and the indeterminacy of the position of moments of time, the Einsteinian theory of relativity. The order of magnitude of these indeterminacies should be estimated, and a relation should be found in the form of a relation between the indeterminacy of a time interval ΔT and the indeterminacy of the position of moments of time Δx .

This is how QM and GR seem to cooperate. to preserve the coherence of the ψ 's of the microparticles, and to prevent the excessive expansion of the ψ 's of macroscopic bodies. By the same token the development of large indeterminacies in the space-time structure is also prevented. This balance is reached by giving up the perfect sharpness of the structure of space-time in GR, and the absolute validity of the superposition principle in QM. 5

Bound of the Indeterminacy
length of a Time Interval
Synchronization of Moments

relations between the in-
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3. Lower Bound of the Indeterminacy in the Length of a Time Interval and in the Synchronization of Moments of Time

Studying the relations between the indeterminacies induced by a QM-al object put into an otherwise ^{originally} empty (Minkowskian) space-time, K. noticed the existence of a lower bound of the indeterminacy of the Einsteinian space-time structure. The order of magnitude of this bound could be estimated, and conveniently expressed in the form of a relation between the length T of a time interval, and the smallest possible value $\Delta_L T$ of the indeterminacy of T . ("stands for "Lower bound".)

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This is to prevent the cooperation of the micro particles, the macroscopic bodies. By the same token the development of large structure is also prevented. This balance is reached by giving up the perfect sharpness of the structure of space-time in G.R., and the absolute validity of the superposition principle in Q.M.

3. Lower Bound of the Indeterminacy in the Length of a Time Interval and in the Synchronization of Moments of Time

Studying the relations between the indeterminacies induced by a Q.M.-al object put into an otherwise ^{completely} empty (Minkowski) space-time, K. noticed the existence of a lower bound of the indeterminacy of the Einsteinian space-time structure of order of magnitude of this bound could be estimated, and conveniently expressed in the form of a relation between the length T of a time interval, and the smallest possible value ΔL of the indeterminacy of T . (This stands for "Lower bound".)

None, it turns out that

6

$$\Delta_2 T \approx T_P^{2/3} T^{1/3}, \quad * (a)$$

where

$$T_P = 5.4 \times 10^{-44} \text{ sec}$$

is the Planck time.

$\approx \rightarrow$ Equal, up to a numerical constant of order 1, possibly 10. Cannot be calculated at present. Fortunately not important in what follows.

Salient point in the deduction of *:

A quantum object induces indeterminacy not only through Δx , but through Δp , too, and Δx and Δp work against each other; $\Delta x \cdot \Delta p \geq \hbar/2$.

Two restrictions on T:

i) **Non-relativistic** QM \rightarrow
 \rightarrow T along $|\underline{v}| \ll c$ worldlines
(in a frame in which 2.7°K isotropic)

ii) $T \gg T_P$,

because for $T \leq T_P$ the concept of the space-time structure is questionable.

* QM-al, not field theoretical!

Notice that
• for given Δx , Δp is an **absolute** lower bound.
(Compare with $\Delta x \cdot \Delta p \geq \hbar/2$)

• $\Delta L T$ is very small. E.g. for $T = 1 \text{ sec}$

$$\Delta L T \approx 10^{-29} \text{ sec.}$$

Still, $\Delta L T$ is large enough to harm the coherence of the Ψ 's of bodies perceived empirically macroscopic ("classical"), and small enough to leave the coherence of the Ψ 's of microparticles nearly perfect.

A remark on \hbar :

\hbar involves only the universal constant(s)

$$= \left(\frac{G \hbar}{c^5} \right)^{1/2},$$

the constant of gravitation. The
(mass, velocity) of the quantum
deduction of \hbar dropped out.

$$\Delta_{LT} \approx \frac{1}{P}$$

Notice that
• for given T , Δ_{LT} is an **absolute** lower bound.
(compare with $\Delta x \cdot \Delta p \geq \hbar/2$)

• Δ_{LT} is very small. E.g. for $T = 1$ sec

$$\Delta_{LT} \approx 10^{-29} \text{ sec.}$$

Still, Δ_{LT} is large enough to harm the coherence of the Ψ 's of bodies perceived empirically macroscopic ("classical"), and small enough to leave the coherence of the Ψ 's of microparticles nearly perfect.

A remark on $\#$:

$\#$ involves only the universal constant(s)

$$\bar{T}_P = \left(\frac{G\hbar}{c^3} \right)^{1/2},$$

where G is the constant of gravitation. The physical parameter (mass, velocity) of the quantum object used in the deduction of $\#$ dropped out.

A space-time relation independent of any particular property of matter can be, perhaps even must be attributed to space-time itself. We know that the expectation concerning the simplicity of ground states, or "vacuums", had to be given up in the quantum world. Examples are the zero point energy of the oscillator, the Dirac sea, the vacuum of the electro-weak interaction.

It says that the compatibility of GR with QM demands, as a minimum, a structure of the empty space-time with a slight indeterminacy, therefore cannot be exactly Minkowskian. Our understanding of this property cannot be reached without a synthesis of Quantum Physics and Gravity.

The lower bound Δ_{LT} of the length of time intervals of worldlines involves the lower bound $(\Delta_{LT})_{syn}$ of the synchronization of the two such worldlines.

any particular property of matter can be, perhaps even must be attributed to space-time itself. We know that the expectation concerning the simplicity of ground states, or "vacuums", had to be given up in the quantum world. Examples are the zero point energy of the oscillator, the Dirac sea, the vacuum of the electroweak interaction.

says that the compatibility of GR with QM demands, as a minimum, that the structure of the empty space-time had a slight indeterminacy, therefore it cannot be exactly Minkowskian. A deeper understanding of this property probably cannot be reached without a unified theory of Quantum Physics and General Relativity.

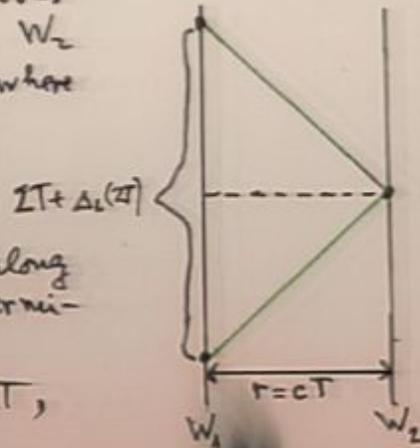
The existence of the lower bound Δ_{LT} of the indeterminacy in the length of time intervals along the $|v| \ll c$ worldlines involves the existence of a lower bound $(\Delta_{LT})_{syn}$ of the indeterminacy in the synchronization of the moments of time on two such worldlines.

Consider, first in Minkowskian space-time, two $|v|=0$ worldlines W_1 and W_2 at a distance r - each other. A light signal emitted on W_1 , arrives back to W_1 , from W_2 after a time $2T$, where

$$T = \frac{r}{c}$$

The time interval along W_1 has the indeterminacy

$$\Delta L(2T) \approx \Delta L T,$$



Therefore the moment of the arrival of the signal to W_2 cannot be synchronized with a corresponding moment on W_1 better than with the indeterminacy

$$(\Delta L T)_{\text{sync}} \approx T_P^{2/3} \left(\frac{r}{c}\right)^{1/3} = \frac{\Lambda^{2/3}}{c^{1/3}}$$

**

where

$$\Lambda = c T_P = 1.6 \times 10^{-33} \text{ cm}$$

is the Planck length.

Consider, first in Minkowskian space-time, two $|v|=0$ worldlines W_1 and W_2 at a distance r from each other. A light signal emitted on W_1 arrives back to W_1 from W_2 after a time $2T$, where

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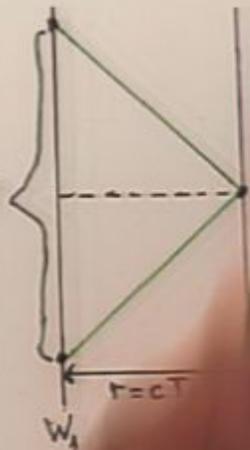
therefore the moment of the signal to W_2 cannot be seen as a corresponding moment on W_1 with the indeterminacy

$$(\Delta L T)_{\text{sig}} \approx T_P^{\frac{2}{3}} \left(\frac{r}{c}\right)^{\frac{1}{3}}$$

where

$$\Lambda = c T_P =$$

is the Planck length

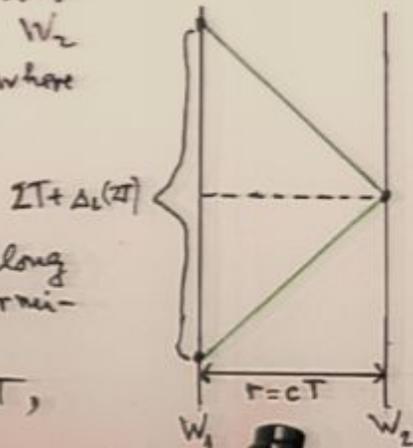


Consider two $|v| = 0$ worldlines W_1 and W_2 at a distance r from each other. A light signal emitted on W_2 arrives back to W_1 from W_2 after a time $2T$, where

$$T = \frac{r}{c}$$

The time interval along W_1 has the indeterminacy

$$\Delta L(2T) \approx \Delta LT,$$



therefore the moment of the arrival of the signal to W_2 cannot be synchronized with a corresponding moment on W_1 or than with the indeterminacy

$$(\Delta LT)_{\text{sync}} \approx T_P^{2/3} \left(\frac{r}{c}\right)^{1/3} = \frac{\Lambda^{2/3}}{c} \quad **$$

where

$$\Lambda = c T_P = 1.6 \times 10^{-33} \text{ cm}$$

is the Planck length

So, as to ΔLT , $(\Delta LT)_{syn}$ is an attribute of the space-time structure, therefore it cannot be beaten by clocks, however exact. Notice that $(\Delta LT)_{syn}$ is small. The relative lower bound

$$\frac{(\Delta LT)_{syn}}{r/c} \approx \left(\frac{\Lambda}{r}\right)^{2/3} \rightarrow 0, \quad r \rightarrow \infty$$

and for $r = 10$ meter ($T_{syn} \approx 10^{-8}$ sec) it is already as small as 10^{-14} .

4. A Tentative Expression for $(\Delta LT)_{syn}$ (and for ΔLT) through Vacuum Expectation Values

In the future of QM and GR, the indeterminacy will presumably be resolved in the ground state of the vacuum for $(\Delta LT)_{syn}$. A general formula for $(\Delta LT)_{syn}$ can be deduced as follows:

Since Δ_{LT} , $(\Delta_{LT})_{\text{syn}}$ is an attribute of the space-time structure, therefore it cannot be beaten by clocks, however exact. Notice that $(\Delta_{LT})_{\text{syn}}$ is small. The relative lower bound

$$\frac{(\Delta_{LT})_{\text{syn}}}{r/c} \approx \left(\frac{\Lambda}{r}\right)^{2/3} \xrightarrow{r \rightarrow \infty} 0,$$

and for $r = 10$ meter ($T = \frac{r}{c} \approx 10^{-8}$ sec) it is already as small as 10^{-24} .

4. A Tentative Expression for $(\Delta_{LT})_{\text{syn}}$ (and for Δ_{LT}) through Vacuum Expectation Values

In the future unified theory of QM and GR, the indeterminacies Δ_{LT} and $(\Delta_{LT})_{\text{syn}}$ will presumably appear as spreads in the ground state of time. An expression for $(\Delta_{LT})_{\text{syn}}$ can be directly to the general relativity for $\Delta\phi$ than the original one can be obtained

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In the future unified theory of QM and GR, the indeterminacies Δ_{LT} and $(\Delta_{LT})_{\text{syn}}$ will presumably appear as spreads in the ground state of space-time. An expression for $(\Delta_{LT})_{\text{syn}}$, leading more directly to the general formula of K for $\Delta\phi$ than the original deduction (1972), can be obtained as follows:

Similarly to Δ_{LT} , $(\Delta_{LT})_{syn}$ is an attribute of the space-time structure, therefore it cannot be beaten by clocks, however exact. Notice that $(\Delta_{LT})_{syn}$ is small. The relative lower bound

$$\frac{(\Delta_{LT})_{syn}}{r/c} \approx \left(\frac{\Lambda}{r}\right)^{2/3} \xrightarrow{r \rightarrow \infty} 0,$$

and for $r = 10$ meter ($T = \frac{r}{c} \approx 10^{-8}$ sec) it is already as small as 10^{-24} .

4. A Tentative Expression for $(\Delta_{LT})_{syn}$ (and for Δ_{LT}) through Vacuum Expectation Values

In the future unified theory of General Relativity, the indeterminacies Δ_{LT} and $(\Delta_{LT})_{syn}$ will presumably appear as spreads in the ground state of space-time. An expression for $(\Delta_{LT})_{syn}$, leading more directly to the general formula of K for Δ_{LT} , was originally deduced (1972), and is as follows:

Therefore Δ_{LT} cannot be beaten by clocks, however exact. Notice that $(\Delta_{LT})_{\text{sym}}$ is small. The relative lower bound

$$\frac{(\Delta_{LT})}{r/c} = \left(\frac{\Lambda}{r} \right)^{2/3} \xrightarrow{r \rightarrow \infty} 0,$$

and for $r = 10$ meter ($T = \frac{r}{c} \approx 10^{-8}$ sec) it is already as small as 10^{-24} .

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In the future unified theory of QM and GR, the indeterminacies Δ_{LT} and $(\Delta_{LT})_{\text{syn}}$ will presumably appear as spreads in the ground state of space-time. An expression for $(\Delta_{LT})_{\text{syn}}$, leading more directly to the general formula of K for $\Delta\phi$ than the original deduction (1972), can be obtained as follows:

Consider a $t = \text{constant}$ hyperplane in Minkowskian space-time. For any pairs of points $\underline{x}, \underline{x}'$ on such a plane

$$t_{\underline{x}'} = t_{\underline{x}} = t$$

that is

$$t_{\underline{x}'} - t_{\underline{x}} = 0 \quad (a)$$

In a space-time

(a)

' $t_{\underline{x}'}$ ' = t ... inacy. It seems that its value, $v_{\underline{x}'}$

and is $(\Delta t)_{\text{syn}}^2 = \frac{\Delta x^2}{c^2} - |\underline{x}' - \underline{x}|^2$

(An example obtained in a similar way in this talk.)

In a space-time where

$$(\Delta_{LT})_{\text{syn}} \neq 0,$$

$t_{x'}$ is an indeterminacy. It seems reasonable to assume that its VEV equals the Minkowskian zero value,

$$\langle t_{x'} - t_x \rangle = 0,$$

and its squared spread equals $(\Delta_{LT})_{\text{syn}}^2$:

$$\langle (t_{x'} - t_x - \langle t_{x'} - t_x \rangle)^2 \rangle = (\Delta_{LT})_{\text{syn}}^2 \approx \frac{\Delta t^3}{c^2} |x' - x|^{\frac{2}{3}}$$

(An expression for $(\Delta_{LT})^2$ can be obtained in a similar way, but it is not needed in this talk.)

(a)

In a space-time where

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$$\langle t_{x'} - t_x \rangle = 0,$$

and its squared spread equals $(\Delta_{LT})_{\text{sym}}^2 =$

$$\langle (t_{x'} - t_x - \langle t_{x'} - t_x \rangle)^2 \rangle = (\Delta_{LT})_{\text{sym}}^2 \approx \frac{\hbar^2}{c^2} |x' - x|^2$$

(An expression for $(\Delta_{LT})^2$ can be obtained in a similar way, but it is not needed in this talk.)

then $\phi(t) \rightarrow \phi(x,t) = -\frac{c^2}{\hbar} \sum_{e=1}^N M_e t_{x_e}$,
 because the coordinate of the particle with mass M_e
 is x_e . The relative rest energy phase between
 x and x'

$$\begin{aligned} \phi(x,x',t) &= \phi(x',t) - \phi(x,t) \\ &= -\sum_{e=1}^N M_e (t_{x'_e} - t_{x_e}) \end{aligned}$$

Remark: is not clear what kind of a
 substitution is made for t in the

phase non-relativistic

contribution comes
 phase factor.

$$\begin{aligned} &0, \\ &0, \\ &\langle \phi(x,t) - 0 \rangle^2 = \\ &(-t_{x_e})(t_{x'_e} - t_{x_e}) \rangle. \end{aligned}$$

at energy phase between

$$\psi(x,t)$$

$$e^{i(\dots)}$$

show what kind of a
for t in the

→ ?

the non-relativistic

the

evolution comes

factor.

evolution of an
particles is des-

cribed by

$$\psi(x,t) = e^{-\frac{i}{\hbar} H t} \psi(x,0)$$

In Minkowskian space-time the evolution
is deterministic, the relative phases between
any pairs of points

$$x = (x_1, \dots, x_N)$$

and

$$x' = (x'_1, \dots, x'_N)$$

of the configuration space are sharply deter-
mined; ψ is and **remains** perfectly coherent

How does $**$ affect the relative phases?

The main influence shows up in the usually
omitted rest energy phase factor

$$e^{-\frac{i}{\hbar} \sum_{e=1}^N H_e c^2 t} = e^{i\phi(t)}$$

(Omitted because ϕ is independent of
 x and of $p = -i\hbar \nabla$, drops out)

But if $t_{x'} \neq t_x$

then:

at energy phase between

$$\phi(x,t) = \dots - Et$$

clear what kind of a
for t in the
relativistic
factor.

distribution comes
factor.

state phases

isolated system of N microparticles is described by

$$\psi(x,t) = e^{-\frac{i}{\hbar} H t} \psi(x,0)$$

In the Minkowskian space-time the evolution is deterministic, the relative phases between any pairs of points

$$x = (x_1, \dots, x_N)$$

and

$$x' = (x'_1, \dots, x'_N)$$

of the configuration space are sharply determined; ψ is and **remains** perfectly coherent.

How does $x \leftrightarrow x'$ affect the relative phases?

The main influence shows up in the usually omitted rest energy phase factor

$$e^{-\frac{i}{\hbar} \sum_{l=1}^N H_0 c^2 t} := e^{i\phi(t)}$$

(Omitted because it is independent of x and of position (x_1, \dots, x_N) , drops out)
But if $t_{x'} = t_x$,
then

x and x' ... rest energy ... with mass M_e ... energy phase between

$$\begin{aligned}\phi(x, x', t) &= \phi(x', t) - \phi(x, t) \\ &= \frac{1}{\hbar} \sum_{i=1}^{2N} M_e (t_{x'_i} - t_{x_i})\end{aligned}$$

Remark: it is not clear what kind of a substitution should be made for t in the

$$e^{-\frac{i}{\hbar} H t}$$

→ ?

phase factor. But in the non-relativistic case

$$|H| \ll \sum_{i=1}^N M_e c^2,$$

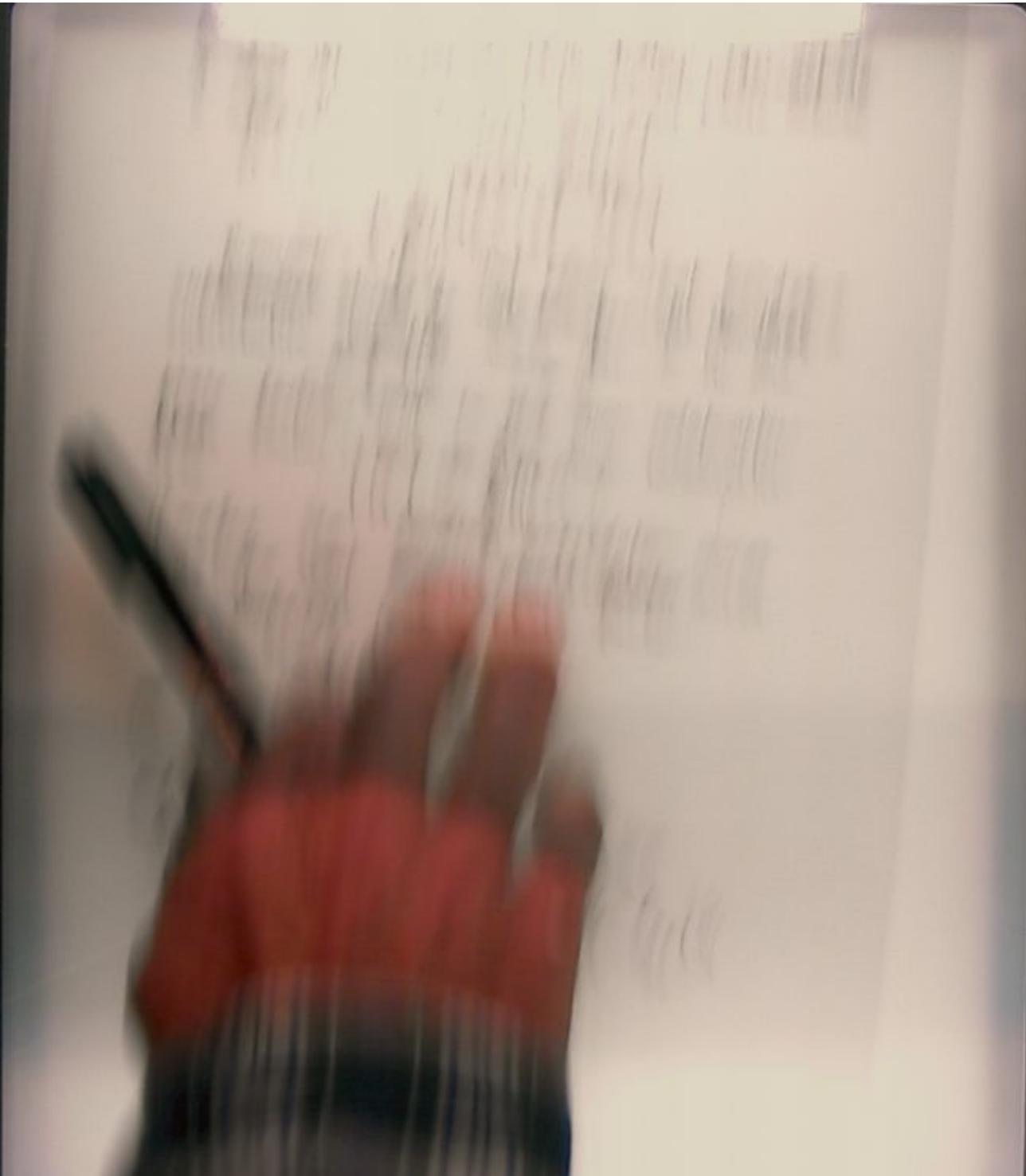
therefore the leading contribution comes from the rest energy phase factor.

$$-t_{x_e} \geq 0,$$

$$t_{x'_i} \geq 0,$$

$$\langle \phi(x, x', t) - 0 \rangle =$$

$$\langle (t_{x'_i} - t_{x_i})(t_{x'_e} - t_{x_e}) \rangle.$$



this: $\phi(t) \rightarrow \phi(x, t) = -\frac{c^2}{\hbar} \sum_{e=1}^N M_e t_{x_e}$,
 because in coordinate of the particle with mass M_e
 is x_e . The relative rest energy phase between
 x and x' is

$$\begin{aligned} \phi(x, x', t) &= \phi(x', t) - \phi(x, t) \\ &= -\frac{c^2}{\hbar} \sum_{e=1}^N M_e (t_{x'_e} - t_{x_e}) \end{aligned}$$

Remark: it is not clear what kind of a
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e^{-iHt} \rightarrow ?
 phase factor. But in the non-relativistic
 case

$$|H| \ll \sum_{e=1}^N M_e c^2,$$

therefore
 from

Re

& contribution comes
 phase factor.

\geq

$$\langle e^{-i(t'_{x'_e} - t_{x_e})} \rangle.$$

x or

$$\begin{aligned}\phi(x, x', t) &= \phi(x', t) - \phi(x, t) \\ &= - \sum_{i=1}^N M_i e^{i H_i (t_{x'_i} - t_{x_i})}\end{aligned}$$

Remark: it is not clear what kind of a substitution should be made for t in the

$$e^{-\frac{i}{\hbar} H t} \rightarrow ?$$

phase factor. But in the non-relativistic case

$$|H_i| \ll \sum_{i=1}^N M_i c^2,$$

therefore the leading contribution comes from the rest energy phase factor.

Remembering that

$$\langle t_{x'_i} - t_{x_i} \rangle = 0,$$

we see that

$$\langle \phi(x, x', t) \rangle = 0,$$

and

$$\begin{aligned}\Delta \phi^2(x, x', t) &= \langle (\phi(x, x', t) - 0)^2 \rangle = \\ &= \frac{c^4}{\hbar^2} \sum_{i, k=1}^N M_i M_k \langle (t_{x'_i} - t_{x_i})(t_{x'_k} - t_{x_k}) \rangle\end{aligned}$$

The expression under $\langle \rangle$ is an algebraic sum of full squares:

$$\frac{1}{2} \left\langle \left(\frac{1}{c} \dot{x}'_i - t_{x_e} \right)^2 + \left(t_{x_i} - t_{x'_e} \right)^2 - \left(t_{x_i} - t_{x_e} \right)^2 - \left(t_{x'_i} - t_{x'_e} \right)^2 \right\rangle.$$

Applying ** to these full squares, one finds:

$$\Delta \phi^2(x, x') \approx \Lambda^{4/2} \frac{c^2}{h^2} \sum_{i=1}^N M_i M_e \left(|x'_i - x_e|^{2/3} - \frac{1}{2} |x_i - x_e|^{2/3} - |x'_i - x'_e|^{2/3} \right)$$

$\Delta \phi$ increases with the masses and with the number of the particles. These fixed, it increases with the distances $|x'_i - x_e|$, that is with the separation between x and x' . These are encouraging properties from the point of view of the expected loss of coherence between "macroscopically distinct" components of the wave function of a massive body.

Remark: $\Delta \phi(x, x')$ turned out to be independent of t , but it is weighted by $|\Psi(x, t)|^2 \cdot |\Psi(x', t)|^2$.

6. The i 's are Cell and the Cell length 15
of a single Microparticle

For $N=1$, $i=l=1$. In the $\Delta\phi$ only the first term survives. Dropping the index i of x_i, x_i' and one finds

$$\Delta\phi(x, x') \approx \Lambda^{2/3} \frac{cm}{h} |x' - x| = \frac{\Lambda^{2/3} a}{L},$$

where $L = \frac{h}{cm}$

is the Compton wavelength, and

$\Delta\phi = 0$ if with the separation is never perfect even

Let us consider "a". Because maximal for the surface

$a = a$ and calculate neutron

$$\Delta\phi(a)$$

Coherence Cell and the Cell Length
 = a Single Microparticle

For $N=1$, $i=l=1$. In the () of $\Delta\phi^2$
 only the first term survives. Dropping
 the index \pm of \underline{x} , \underline{x}' and M_1 , one finds

$$\Delta\phi(\underline{x}, \underline{x}') \approx \Lambda^{2/3} \frac{cM}{\hbar} |\underline{x}' - \underline{x}|^{1/3} = \frac{\Lambda^{2/3} a}{L}$$

where

$$L = \frac{\hbar}{cM}$$

is the Compton wavelength of the particle,
 and

$$a = |\underline{x}' - \underline{x}|.$$

$\Delta\phi = 0$ if $a=0$, and increases monotonically
 with the separation a . If $\underline{x}' \neq \underline{x}$ the coherence
 is never perfect, but it may remain nearly
 perfect even for large separation.

Let us consider a spherical volume of diameter
 "a". Because of the monotony of $\Delta\phi(a)$, $\Delta\phi$ is
 maximal for any diametrically opposed points on
 the surface of the sphere. Let

$$a = a_{\text{Earth}} \approx 10^9 \text{ cm},$$

and calculate $\Delta\phi(a_{\text{Earth}})$ for the neutron

$$L_{ne} = 1.8 \times 10^{-14} \text{ cm},$$

$$\Delta\phi_{ne}(a_{\text{Earth}}) \approx \frac{(10^{-33})^{2/3}}{10^{-14}} (10^9)^{1/3} = 10^{-5} \ll \pi.$$

$$* \quad \Delta_L T \approx T_P^{2/3} T^{1/3}; \quad T_P = 5.4 \times 10^{-44} \text{ sec}$$

$$** \quad (\Delta_L T)_{\text{syn}} \approx \frac{\Lambda^{2/3} r^{1/3}}{c}; \quad \Lambda = 1.6 \times 10^{23} \text{ cm}$$

$$\phi(x, x', t) := \phi(x, t) - \phi(x', t)$$

$$\Delta^2 \phi(x, x') \approx \Lambda^{4/3} \frac{c^2}{h^2} \left\{ \frac{1}{2} \left| \frac{x'_i - x_c}{\Lambda} \right|^{2/3} - \frac{1}{2} \left| \frac{x_i - x_c}{\Lambda} \right|^{2/3} - \dots \right\}$$

$$* \quad \Delta_L T \approx T_P^{2/3} T^{1/3}; \quad T_P = 5.4 \times 10^{-44} \text{ sec}$$

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$$\phi(x, x', t) := \phi(x, t) - \phi(x', t)$$

$$\Delta_L^2 \phi(x, x') \approx \Lambda^{4/3} \frac{c^2}{h^2} \sum_{i, l=1}^N M_i M_l \left(|x'_i - x'_l|^{2/3} - \frac{1}{2} |x_i - x_l|^{2/3} - |x_i - x_l|^{2/3} \right)$$

6. Coherence Cell and the Cell Length
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For $N=1$, $i=l=1$. In the () of $\Delta\phi^2$ only the first term survives. Dropping the index 1 of \underline{x} , \underline{x}' and M_1 , one finds

$$\Delta\phi(\underline{x}, \underline{x}') \approx \Lambda^{2/3} \frac{cM}{h} |\underline{x}' - \underline{x}|^{1/3} = \frac{\Lambda^{2/3} a}{L},$$

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$$L_{ne} = 1.8 \times 10^{-14} \text{ cm},$$

$$\Delta\phi_{ne}(a_{\text{Earth}}) \approx \frac{(10^{-33})^{2/3}}{10^{-14}} (10^9)^{1/3} = 10^{-5} \ll \pi.$$

where $L = \frac{h}{mv}$

is the Compton wavelength of the particle,
and $a = |\underline{x}' - \underline{x}|$.

$\Delta\phi = 0$ if $a=0$, and increases monotonically with the separation a . If $\underline{x}' \neq \underline{x}$ the coherence is never perfect, but it may remain nearly perfect even for large separation.

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$$a = a_{\text{Earth}} \approx 10^9 \text{ cm},$$

and calculate $\Delta\phi(a_{\text{Earth}})$ for the neutron

$$L^{ne} = 1.8 \times 10^{-14} \text{ cm},$$

$$\Delta\phi^{ne}(a_{\text{Earth}}) \approx \frac{(10^{-33})^{2/3}}{10^{-14}} (10^9)^{1/3} = 10^{-5}$$

where

$$L = \frac{h}{mv}$$

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and

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configuration -

$\Delta\phi(x, x') \leq \pi$ for any $x, x' \in \Omega_c$
 has been called "coherence cells" (K, 1972)
 For single microparticle a coherence
 cell is a sphere of diameter a_c
 such that

$$\Delta\phi(a_c) \approx \frac{\lambda^{1/3} a_c^{1/3}}{L} = \pi$$

Dropping the factor π^3 in a_c one finds

$$a_c \approx \left(\frac{\lambda}{\pi}\right)^2 L$$

For neutron

$$a_c \approx \left(\frac{10^{-14}}{10^{-33}}\right)^2 10^{-14} = 10^{29} \text{ cm.}$$

Compare with

$$a_{\text{box}} \approx 10^{21} \text{ cm.}$$

The microparticle occupies only a tiny coherence cell, ψ is almost pure, no self-localization is needed. any microsystem (composed of number of microparticles, free in the "cell" .. length" expression "coherence length"

(972)
can " a sphere of diameter a_c " coherence
such that

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The ψ of a microparticle occupies only a tiny part of its coherence cell, it is almost perfectly coherent, no self-localization is needed. This is so for any microsystem (composed of a moderate number of microparticles, free or bound). a_c is the "cell" "length" of a particle, the expression "coherence length" be extended.

such that

parameter a_c

$$\Delta\phi(a_c) \approx \frac{\sqrt{2} a_c^{1/2}}{L} = \pi$$

Dropping factor π^3 in a_c one finds

$$a_c \approx \left(\frac{L}{\lambda}\right)^2 L$$

For the neutron

$$a_c \approx \left(\frac{10^{-14}}{10^{-13}}\right)^2 10^{-14} = 10^{29} \text{ cm.}$$

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The ψ of a microparticle occupies only a tiny part of its coherence cell, it is almost perfectly coherent, no self-localization is needed. This is so for any microsystem (composed of a moderate number of microparticles, free or bound). a_c is the "cell" "length" of the particle, the expression "coherence length" being patented.

7. The Coherence Cell and the Cell Length of
Homogeneous, Macroscopic Solid Bodies

17

One would think that if $N \approx 10^{23}$,
then nothing can be said about ψ and
about $\Delta\phi$. This is not so. I have no time
to go through the surprisingly simple
mathematics, I just quote the result.

Let the full mass of the body
be M , its diameter $2R$. The coherence
cell turns out to be a sphere of
diameter a_c in the center of mass coordi-
nate sub-space of the $3N$ dimensional
configuration space. The contribution to
 $\Delta\phi$, and hence to a_c of the $3(N-1)$
relative coordinates are negligible.
The formula for a_c breaks up into two
parts:

One would think that if $N \approx 10^{23}$,
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 configuration space. The contribution to
 $\Delta\phi$, and hence to a_c of the $3(N-1)$
 relative coordinates are negligible.
 The formula for a_c breaks up into two
 parts:

if $a_c \geq 2R$,

then $a_c \approx \left(\frac{R}{\Lambda}\right)^2 L$

This formula is the same as that for a microparticle. a_c depends on M , but not on R .

if $a_c \leq 2R$,

then $a_c \approx \left(\frac{R}{\Lambda}\right)^{2/3} L$

The formula is different from that for a microparticle, and besides M a_c depends on R too.

18

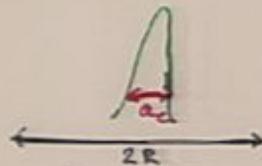
consider the cases

and

$a_c \ll 2R$

you expand
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main
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y

means that Ψ loses its coherence already when it expands in the c.m. subspace over a domain still much smaller than R , and then a self-localization counterbalances the loss of coherence.



uncertainty
of the
classical
behavior

Small indeterminacy in the position of the c.m. is characteristic of microscopic (quantum) behavior

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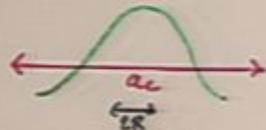
Let us consider the cases

$$a_c \gg 2R$$

and

$$a_c \ll 2R$$

means that Ψ can expand in the c.m. coordinate subspace over a domain much larger than the geometrical size R of the body without losing its coherence.



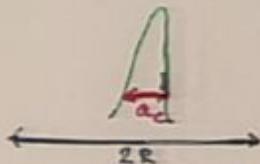
Large indeterminacy in the position of the c.m. is characteristic of **macroscopic** (quantum) behavior

if $a_c \leq 2R$,

then
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The formula is different from that for a microparticle, and besides M a_c depends on R too.

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Small indeterminacy in the position of the c.m. is characteristic of **microscopic** (quantum) behavior

Károlyházy (1964-72)

Cell length a_c of a single microparticle

$$a_c \approx \left(\frac{L}{\lambda}\right)^2 L \quad \left(L = \frac{\hbar}{\epsilon M}\right)$$

Cell length of spherical solid

if $a_c \geq 2R$

$$a_c \approx \left(\frac{L}{\lambda}\right)^2 L$$

$$\sim \chi_e^{(2)}(3)$$



a_c
 $\geq 2R$
 $a_c \gg 2R$

if $a_c \gg 2R$,

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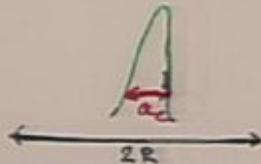
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Small indeterminacy in the position of the c.m. is characteristic of microscopic (quantum) behavior

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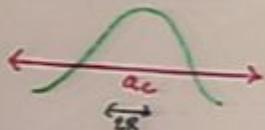
18

Let us consider the cases

$a_c \gg 2R$ and

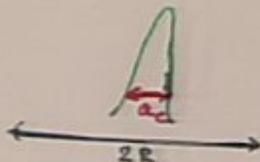
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Small indeterminacy in the position of the c.m. is characteristic of **microscopic** (quantum) behavior

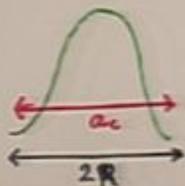
8. The Transition from Microscopic to Macroscopic Behavior

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It follows from the above that the transition takes place in the region where

$$a_c \approx 2R,$$

i.e. where the indeterminacy in the position of the c.m. equals the geometrical size of the object:



Let us estimate $a_c \approx 2R$ and M in the transition region. Inserting into the formula

$$a_c \approx \left(\frac{L}{\lambda}\right)^2 L$$

the expression

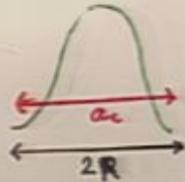
$$L = \frac{h}{cM} \approx \frac{h}{c\rho R^3},$$

for normal terrestrial density $\rho \approx 19/\text{cm}^3$
the condition $a_c \approx 2R$ gives

$$a_c^{\text{tr}} \approx 2R \approx 10^{-5} \text{ cm},$$

$$M^{\text{tr}} \sim \rho R^3 \approx 10^{-14} \text{ gram}.$$

i.e. where the indeterminacy in the position of the c.m. equals the geometrical size of the object:



Let us estimate $a_c \approx 2R$ and M in the transition region. Inserting into the formula

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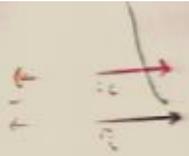
$$L = \frac{h}{cM} \approx \frac{h}{\rho R^3}$$

for normal terrestrial density $\rho \approx 19/\text{cm}^3$
the condition $a_c \approx 2R$ gives

$$a_c \approx 2R$$

$M \approx$

This is the size of colloidal grains
and



Let us say $a_c \approx 2R$ and M in the transition region. Inserting into the formula

$$a_c \approx \left(\frac{L}{\lambda}\right)^2 L$$

the expression

$$L = \frac{h}{cM} \approx \frac{h}{c\rho R^3}$$

for normal terrestrial density $\rho \approx 19/\text{cm}^3$ the condition $a_c \approx 2R$ gives

$$a_c^{tr} \approx 2R \approx 10^{-5} \text{ cm,}$$

$$M^{tr} \approx 10^{-14} \text{ gram.}$$

This is the mass region of colloidal grains and small particles.

Then ...
 responsible for the loss of the coherence,
 because if it were, then the transition
 from quantum classical behavior should
 take place the region of the Planck
 mass

$$M_P = \frac{\hbar}{c \lambda} = 2.2 \times 10^{-5} \text{ gram.}$$

However, a body of such a large mass is known
 to behave classically, it does not belong to
 the transition region.

As we have seen, only the amount of the
 indeterminacy of the space-time structure is
 fixed by M_P (or, which is the same, by
 $T_P = \hbar / M_P c^2$). The degree of the loss
 of coherence is given by $\Delta \phi$, which depends
 not only on M_P , but also on the physical
 composition of the system considered, in particu-
 lar on the masses of its constituents. This leads to

$$M_{tr} \approx 10^{-15} M_P,$$

a plausible value

For a tiny grain of mass $m \approx 10^{-10}$ gram $R \approx 10^{-6}$ cm,
 and $a_c \approx 10^{-8}$ cm

For a tiny grain

Take \hbar
mass

$$M_P = \frac{\hbar}{c \lambda} = 2.2 \times 10^{-5} \text{ gram.}$$

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$$M_{tr} \approx 10^{-15} \text{ gram} \ll M_P,$$

a plausible value.

For a tiny grain of $M \approx 10^{-18}$ gram $R \approx 10^{-6}$ cm,
and $a_c \approx 10^5$ km $\gg R$.

For a ball of $M \approx 10^{-16}$ gram $R \approx 1$ cm, and $a_c \approx 10^{-16}$ cm $\ll 2R$.

Take
mass

$$M_P = \frac{\hbar}{c \lambda} = 2.2 \times 10^{-5} \text{ gram.}$$

However, a . . . such a large mass is known to behave classically, it does not belong to the transition region.

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For a ball of $M \approx 1 \text{ g}$ $R \approx 1 \text{ cm}$, and
 $a_c \approx 10^{-16} \text{ cm} \ll .2R$.

Planck
mass

$$M_P = \frac{\hbar}{c \lambda} = 2.2 \times 10^{-5} \text{ gram.}$$

However, a $\frac{1}{2}$ such a large mass is known to behave classically, it does not belong to the transition region.

As we have seen, only the amount of the indeterminacy of the space-time structure is fixed by M_P (or, what is the same, by $T_P = \hbar / M_P c^2$). The degree of the loss of coherence is given by $\Delta \phi$, which depends not only on M_P , but also on the physical composition of the system considered, in particular on the masses of its constituents. This leads to

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 $a_c \approx 10^{-16}$ cm $\ll .2R$.

15. Concluding Remark

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According to the orthodox view, in order to extract observable results from QM one needs at least one agent which does not obey the laws of QM, but behaves "classically".

However, if one abandons the concept of the sharp space-time structure, and with it the unlimited validity of the superposition principle, the division of the world into objects with purely quantum mechanical or purely classical behavior is not needed. The time evolution of any object obeys a unified dynamics, combining the deterministic Schrödinger ~~time~~ evolution with the stochastic self-localizations. Because of "practical" reason microparticles do not undergo self-localization, but in principle they could. For solid macroobjects the self-localizations occur frequently, and they would keep the localization of the c.m. very tight even if the body would be completely isolated, and if nobody would observe it.

Károlyházy (1964-72)

cell length a_c of a single microparticle

$$a_c \approx \left(\frac{L}{\hbar} \right)^2 L \quad \left(L = \frac{\hbar}{cM} \right)$$

ψ



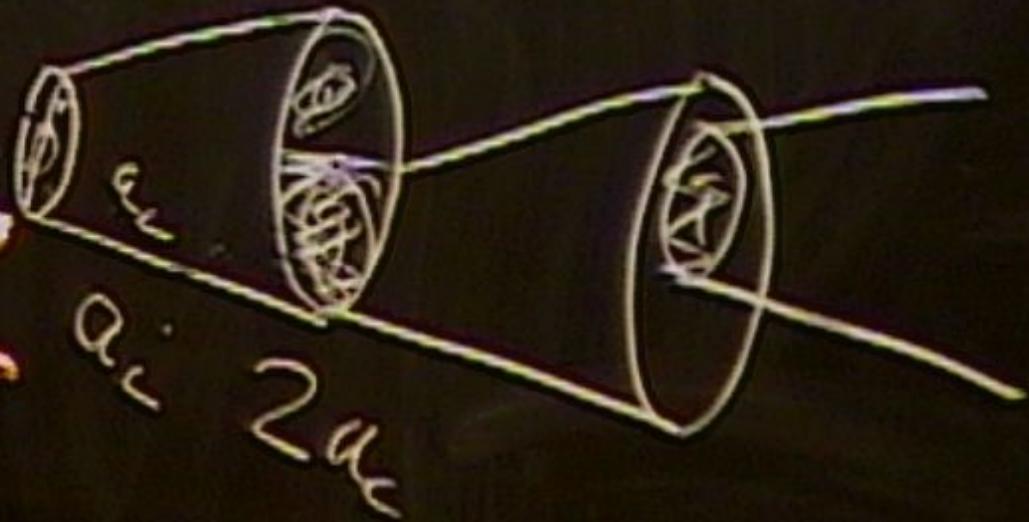
$\binom{2}{1} \binom{3}{3}$

Károlyházy (1964-72)

cell length a_c of a single microparticle

$$a_c \approx \left(\frac{L}{\lambda} \right)^2 L \quad \left(L = \frac{h}{cm} \right)$$

ψ



(2/3)