

Title: Professional development for physics teachers: A UK perspective

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URL: <http://pirsa.org/06050012>

Abstract: <kw> Institute of Science, A levels, boring, difficult, support for schools, girls in physics, advancing physics, physics education, celestia, videshell, quantum atomica, walter fend, physics lab, phet, modellus, clea, lancaster particle physics, many paths, seismology, stellarium, Alice law, warp, wintreb, datapoint, tinycad, falstad applets, spektrus, emanim </kw>

(Science) Education in the (UK)

Age	Subject	Number	Assessment
5- 7	Science	~ 500K	National Test
7-11	Science	~ 500K	National Test
11-14	Science	~ 500K	National Test
14-16	Sciences(?) Core, Applied, Academic	~ 250K (A-C) Physics ~33K	GCSE
16-18	Physics	~ 28K	AS A2

Physics in Schools (Numbers)

- **Physics A-level in 2005,**
 - 28,119 UK candidates (12th most popular)
 - 21,922 (6th most popular (M))
 - 6197 (19th most popular (F))

Physics in Schools (Attitudes)

- **Physics is difficult**
 - Research backs this up but probably not to the extent that it is perceived
- **Physics is irrelevant**
 - Poor careers advice and lack of contemporary physics is an issue
- **Physics is boring**
 - Non specialist teaching, plus a didactic approach to learning

Physics Teachers (Numbers)

- **Trainee Teachers 2005 (England)**

- 301 Physics (not all physicists),
- 434 Chemistry
- 862 Biology
- 977 Science (almost no physicists)

~ 15% of trainee science teachers are physicists

- **Physics Teachers in Schools**

- 19% of science teachers are physicists
- No physicist in 26 per cent of 11-16 schools

What the education department does

- Support for schools
 - Professional Development for Teachers
 - Affiliation Scheme
 - Teacher Network
 - Websites
 - Email lists
 - Publications
 - Current and forthcoming projects

Professional Development for Teachers

- **Physics Teacher Network**
 - 33 local physicists coordinating a variety of training opportunities
- **Physics Updates**
 - 3 day residential courses
 - April, July and December
- **Meetings**
- **Email Lists**
 - PTNC (Physics Teachers News and Comment)
 - PEPD (Physics Early Professional Development)

SPT Project

- professional development programmes
- supported by a series of CD-ROMs
 - **Electricity & magnetism**
 - **Forces & motion**
 - **Light & sound**
 - **The Earth & beyond**
 - **Energy resources and energy transfer**
- each CD ROM covers a chunk of the 11-14 physics curriculum

Practical Physics

- www.practicalphysics.org
A website listing practical activities, equipment and advice on how to carry out practicals and build them into a teaching sequence
- New section being developed to help teach 'How science works' through practical activities.
- Please Contribute!

Girls in Physics

- Girls are under represented at A-level – everyone knows this but no one has found a way of changing this.
- Two research studies have been commissioned:
 - Case studies from schools that are successful in recruiting girls to A-level physics
 - A review of the research that has been done in this area

Institute *of* **Physics**

Teaching Advanced Physics

www.tap.iop.org

- **Contains detailed ideas and resources for teaching physics to students aged 16-19.**
- **Aims to help those new to teaching this age group**
 - Electricity
 - Mechanics
 - Vibrations and Waves
 - Atomic and Nuclei
 - Fields
 - Energy
 - Astronomy

Advancing Physics AS

- **Physics in Action**

- Communication is about imaging and instrumentation.
- Designer materials looks at the way the world is made

- **Understanding Processes**

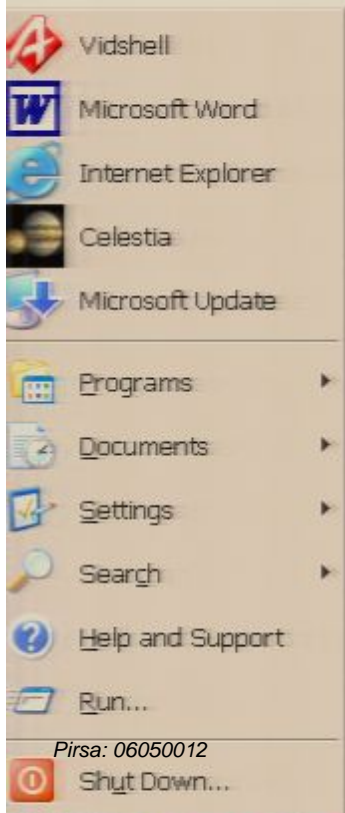
- Waves and quantum behaviour introduces some simple quantum ideas
- Space and time moves from vectors to modelling

Advancing Physics A2

- **The Rise and Fall of the Clockwork Universe**
 - **Models and Rules** introduces the use of mathematical modelling in physics
 - **Matter in Extremes** develops a bridge between random events and empirical laws
- **Field and particle pictures**
 - **Fields** develops ideas about magnetic and electric fields
 - **Fundamental particles** considers the structure and binding of atoms and nuclei and introduces ideas about risk
- **Advances in Physics**
 - Draws together and uses physics ideas from the whole course in looking at new contexts

Advancing Physics A2

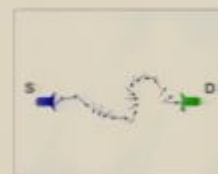
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7: Quantum Behaviour



Fundamental particles of matter – electrons, quarks, neutrinos, and others – act like nothing you have seen before. But they all share the same remarkable style of acting: 'quantum behaviour'. Photons, the fundamental particles of light, do it too. In this chapter we will:

- explain why light needs to be thought of as photons
- use photons as a simple example of quantum behaviour

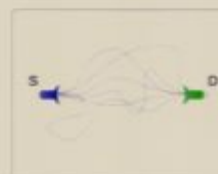
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7: Quantum Behaviour
Teacher's Guide

Teacher's Guide

Overview

Aims

Introducing quantum ideas early

We believe that a first introduction to quantum ideas is an essential part of a physics A-level course, to be started as early as possible. We want students to get a clear and simple picture of what is characteristic of quantum behaviour, so as to be able to understand what they read in newspapers and magazines, perhaps detecting misinformed views. Quantum ideas, in various guises, have entered our culture and its debates about the nature of things, and physics students above all ought to be given the means to join in. And they need such an introduction so that their physics is not cut off from what they hear in other sciences, especially in chemistry.

Going for essentials

We want students to grasp the essentials of the quantum world view, with at this stage the minimum of technicalities. One way of achieving this was worked out by Richard Feynman, in his popular book 'QED' (1985 Penguin). This little book is required – and easy – reading for any teacher embarking on this part of the course. And in this book Feynman concentrates on the deepest essential features of quantum physics, in a strikingly simple way. A deep unification is that quantum behaviour is shown both by photons and electrons; by in classical terms both 'radiation' and 'matter'.

Localised or non-localised – 'Try all paths'

One essential is that in quantum physics the classical distinction between non-localised things like fields and waves which are 'everywhere', and things like particles which are localised in space, simply doesn't apply. Quantum objects manage both at once. Feynman's picture brings this out by making the fundamental command to particles (such as photons or electrons) be 'Try all paths'. The imagined world is one in which quantum particles go everywhere and try everything all at once.



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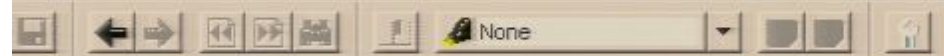
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Section 7.1: Quantum behaviour

Teaching time: 4 hours

What we are trying to achieve

The aim of this section is to introduce 'quantum behaviour' as its own unique self, as a new way of imagining the behaviour of quantum objects.

The first step is to go beyond the wave picture of chapter 6 by insisting on the lumpy nature of the energy exchanges between photons and matter, and the random nature of the arrival of that energy. The simple experience of listening to gamma ray photons arriving in a detector one by one, at random, is key. This is supported by a sample of evidence for the Planck relationship $E = hf$.

The second step is to begin to see how, given the lumpy nature of the exchange of quantum energy, quantum behaviour still accounts for superposition effects. So we introduce the 'try all paths' story, but with just two paths. The photon has to be thought of as trying **both** paths. The way to calculate phase differences for the paths, with rotating arrows, carries over, but not the commitment to waves which previously lay behind it.

The case of two paths is just a start on the many paths construction, to be deepened in section 7.2. Discussion about the implications can begin, but it is too soon to reach for conclusions.

What to cover

You need to cover the lumpy nature of photons, and one example of the evidence for the relation $E = hf$, for instance measuring the minimum striking potential difference of LEDs of different colour. Other evidence might be discussed, but more important is to show a clear case of lumpy behaviour such as the detection of gamma ray photons, which brings out well the random nature of their arrival. At the other extreme, with microwaves, you need to remind students that photons still superpose.

Secondly, you need to show how the 'try all paths' argument works, getting an arrow direction for each path, and adding



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Section 7.2: 'Try all paths' at work

Teaching time: 4 hours

What we are trying to achieve

We want students to grasp something at once very simple and quite difficult: the fact that the rule of quantum behaviour 'try all paths' accounts for well known familiar optical phenomena, such as reflection, in which quite to the contrary the photons appear to take just one path, not many. Examples are reflection, refraction, straight line propagation, the design of lenses and curved mirrors, and the action of a grating.

The explanation always has the same form: only along paths very close to a special one do the phasor arrows line up. All other paths, though tried, add up to a negligible contribution.

In the *Advancing Physics AS* student's book, these examples are presented briefly, one after the other, to give an impression of the scope of the argument, of the wide range of phenomena it accounts for.

In class, because the argument is subtle and seemingly self-contradictory (photons ending up at one place **because** they take all paths), the opposite strategy is likely to be better, looking at one or two examples in detail and arguing them through carefully. After that, readings from the *Advancing Physics AS* student's book may suffice.

For a first run through, we suggest:

1. The case of reflection at a plane mirror, including the least time principle.
2. Rectilinear propagation.
3. The reflection grating, showing that the photons do in fact try all paths.
4. Engineering a curved mirror, putting the least time principle to work.

In this guide these are therefore described first, followed by the other examples which you may wish to consider, or on which you may want to draw.

With more experience, you may want to make a different selection. Alternatively, you may wish to divide the various



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That gratings get certain definite paths to have phasors which line up by removing paths which when added in make the phasors 'curl up'.

Reflection at a mirror

Remind the class, perhaps by flashing a beam of light around the room, that the photons take the special path for which the angle of reflection is the same as the angle of incidence. Tell them that, strangely, this fact follows from the photons trying all paths, as the rule of quantum behaviour insists they do.

Make a start with paper-and-pencil work with a trundle wheel, showing that the phasor arrows from paths near the 'true' path more or less line up, whilst paths for which the angles are unequal tend to curl up. Move to the same thing shown in software, where the effects can be explored much more quickly.

Three regions dominate: two are the ends which produce arrows that curl up, and the third is the middle which produces arrows that line up.

The software written in Modellus uses a defined set of paths, a few at a time, concentrating on identifying adjacent paths that produce arrows which line up or curl up. The other software, due to Edwin Taylor, is more flexible and allows freer exploration. The combination of both is very profitable.

Activities

- Activity 80E Experiment 'Calculating for a mirror on the bench'
- Activity 90S Software Based 'Calculating for a mirror on the screen'

Question

- Question 40S Short Answer 'Three paths on a mirror'

Files

- File 50I Image 'Paths for a mirror'
- File 60I Image 'A hexagonal grid'
- File 70L Launchable File 'Many-paths software'



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- [Activity 90S](#) Software Based 'Calculating for a mirror on the screen'

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- [Question 40S](#) Short Answer Three paths on a mirror'

Files

- [File 50I](#) Image 'Paths for a mirror'
- [File 60I](#) Image 'A hexagonal grid'
- [File 70L](#) Launchable File 'Many-paths software'

Display Materials

- [Display Material 80P](#) Poster 'Many paths for a mirror'

The least time path

The minimum here is simply to draw students' attention to the fact that the path which obeys the law of reflection is also the one which takes the least time. They can see the minimum in the graphical output of a spreadsheet set up to calculate trip times, and have it pointed out that near the minimum the trip times change very little with change of path, so that their phasors should line up.

Display Materials

- [Display Material 90O](#) OHT 'Mirror: Contributions from different paths'
- [Display Material 100O](#) OHT 'Calculated trip times for a mirror'



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Resource Manager

Resource Manager

Listening to photons arriving

Activity 20D: Demonstration



When and where

You are seeking to find out more about the **photon**, so finding out more about where and when they travel is a good start. What we can know about these two aspects of their travel helps to determine the kind of theory we can develop about them.

You will need

- ✓ pure gamma source
- ✓ one or more Geiger-Müller tubes with counter with audio signal



What to do

1. Place the source and counter so you can hear a few counts a minute. Each count should be distinct.
2. Listen carefully. Can you predict the arrival of the next photon?
3. Alter the separation of the counter and source. Can you describe the changes in what you hear?



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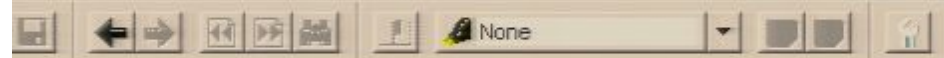
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 - A photon explores two h
 - Calculating for a mirror c
 - Trip times for a mirror
 - A few mirror paths
 - Many photons make a b
 - Photons propagating
 - Checking the ends of th
 - Engineering a focusing r
 - Least time and refractor
 - Least time and refractor
 - Engineering a lens
 - Photons explore a narrow

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Random arrival of photons

Activity 30S: Software Based



What does randomness look like?

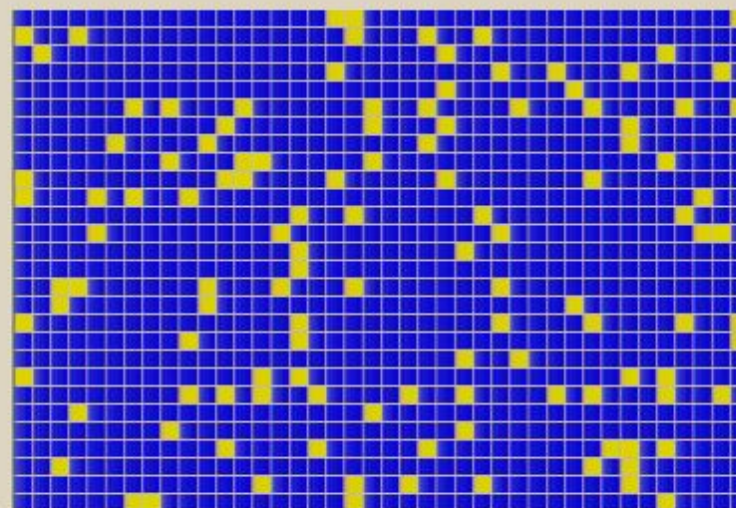
This computer model shows you what the random arrival of photons at a screen looks like. You will see that randomly placed photons are far from evenly placed!

This activity is designed for use with [File 10L](#) 'A model for the random arrival of photons'.

You will need

- ✓ computer with *Advancing Physics* CD-ROM

Photons come at random





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 How exploring paths leads to arrows

How exploring paths leads to arrows

Activity 50S: Software Based



Paths contribute

Photons explore all paths from source to detector and each path contributes to the **probability** of finding a **photon** at the detector. Each path contributes an arrow. This activity explores how different paths contribute and how the contributions of those paths are aggregated. Each step makes one simple point only. Do not spend a long time on any one step.

This activity is designed for use with [File 20L](#) 'Getting arrows from paths' and [File 40L](#) 'Getting arrows from more complex paths'.

You will need

- ✓ PC running Modellus
- ✓ four or six model files

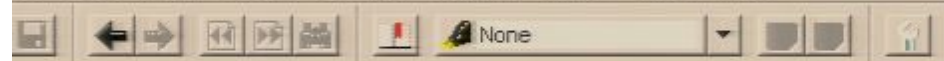
One path contributes

This first step shows how a single path generates a single arrow.

1. Launch the model. Make sure you can see the Animation Window.

One path contributes





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 - Exploring paths with changing kinetic energy

From small to large

Small things, whether **electrons** or other particles, have their own special **quantum nature**. At larger scales, collections made of these same small things behave in ways that are much more familiar. This activity shows how increasing the **kinetic energy** makes their quantum nature less evident.

You will need to remember that the **phasor** rotates at a **frequency** given by

$$f = \frac{E_K}{h}$$

This activity is designed to go with **File 180L** 'Changing kinetic energy'.

You will need

- ✓ computer running Modellus
- ✓ the Modellus model 'Changing kinetic energy'

Looking at propagation

Propagation of electrons





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You will need to remember that the **phasor** rotates at a **frequency** given by

$$f = \frac{E_K}{h}$$

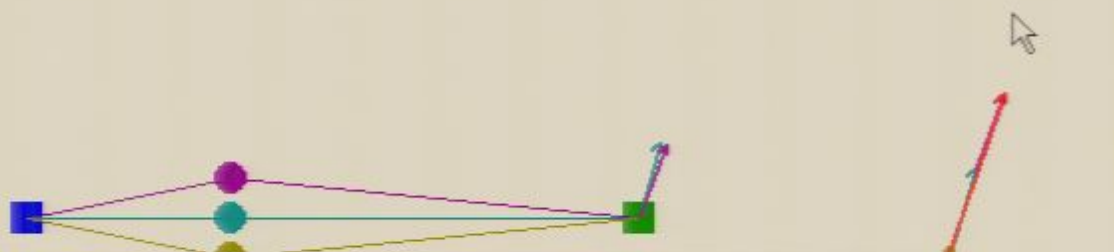
This activity is designed to go with **File 180L** 'Changing kinetic energy'.

You will need

- ✓ computer running Modellus
- ✓ the Modellus model 'Changing kinetic energy'

Looking at propagation

Propagation of electrons





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 - Least time and refractor
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How exploring paths leads to arrows

Activity 50S: Software Based



Paths contribute

Photons explore all paths from source to detector and each path contributes to the **probability** of finding a **photon** at the detector. Each path contributes an arrow. This activity explores how different paths contribute and how the contributions of those paths are aggregated. Each step makes one simple point only. Do not spend a long time on any one step.

This activity is designed for use with [File 20L](#) 'Getting arrows from paths' and [File 40L](#) 'Getting arrows from more complex paths'.

You will need

- ✓ PC running Modellus
- ✓ four or six model files

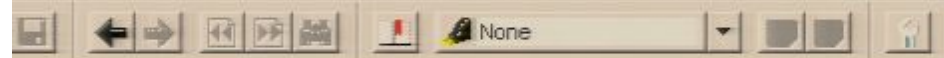
One path contributes

This first step shows how a single path generates a single arrow.

1. Launch the model. Make sure you can see the Animation Window.

One path contributes





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2. Run the model once using the 'go' button in the Control Window.

Note how the **phasor**, the spinning, rotating arrow, marks the exploration of the path by the photon. Only when the phasor reaches the detector does the arrow freeze.

3. You can place the waypoint in different positions by clicking on the 'case' buttons on the menu bar of the Animation Window. To see how the arrows are generated for these different cases, click on the 'go' button in the Control Window with the case you want to see selected.

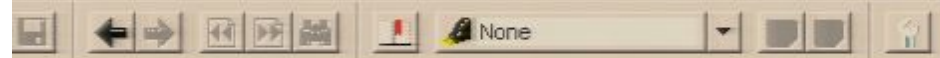
The greater the distance to be covered the longer the trip time to explore the path. The longer the trip time the more rotations for the phasor. The final angle of the phasor fixes the arrow. For this simple model the length of the phasor does not change as the path is explored.

Varying one path and seeing the contribution change

Now cut out the spinning and translating. The number of rotations can be worked out from the trip time, which in turn can be calculated from the path length. So now you can see how the arrow responds to changes in path length.

4. Launch the model. Make sure you can see the Animation Window.

A path contributes



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5. Run the model once using the 'go' button in the Control Window.

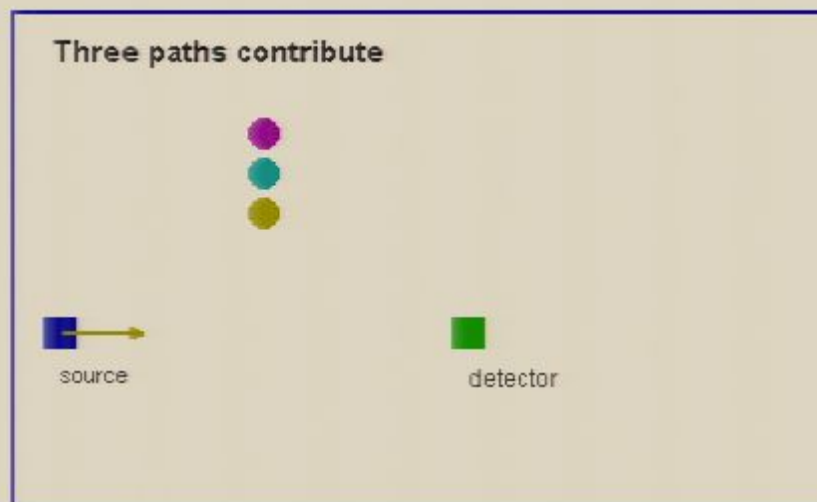
6. Drag the waypoint to different positions and see how the arrow responds.

At this stage you might be looking for regions from which the arrows tend to line up. These will be places where the trip times for adjacent waypoints are not so different.

Three paths contribute

As a further step we consider how the exploration of three paths generates three arrows. These three arrows are then summed to make an amplitude (imagine that there are only three paths – this reduces the problem of sampling all paths to manageable proportions). This will turn out to be a useful trick.

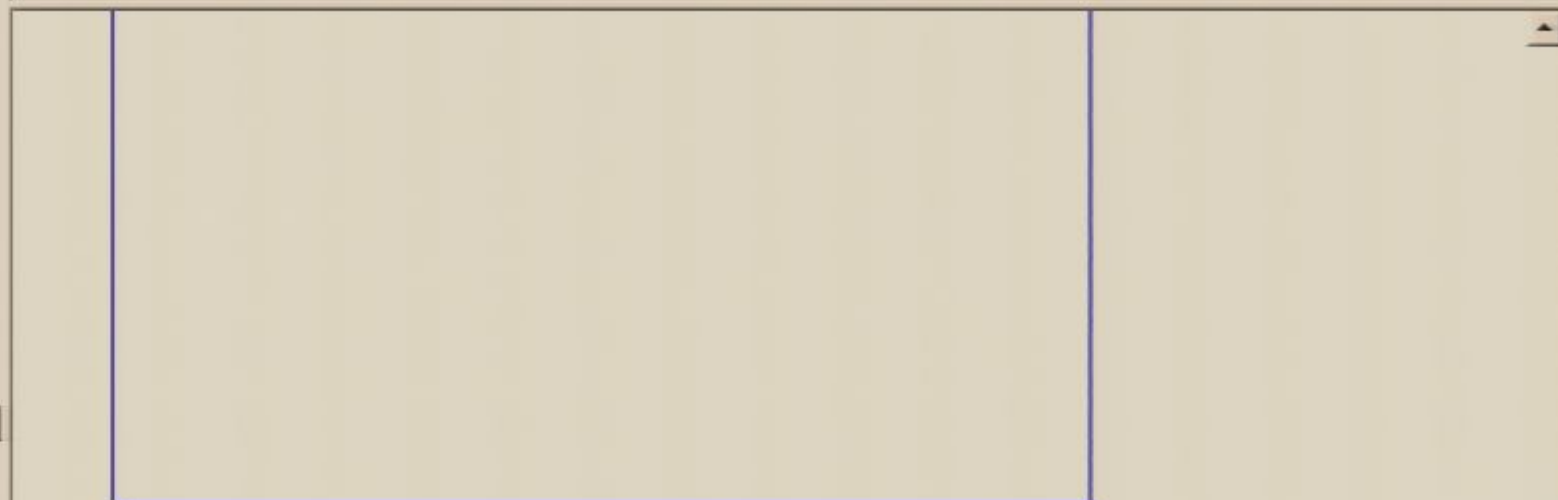
7. Launch the model. Make sure you can see the Animation Window.





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11. Run the model once using the 'go' button in the Control Window.

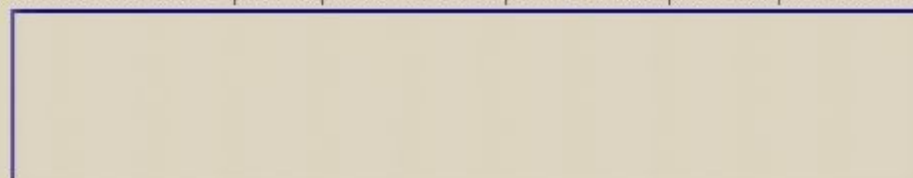
12. Drag the waypoints to different positions and see how the arrows respond and how the amplitude changes.

In this more unstructured exploration it is harder to find significant areas – but it does show the need for a planned attack on exploring all paths.

Extension:

Paths with two waypoints

The contribution of the path depends on the trip time. The trip time depends on the path length.

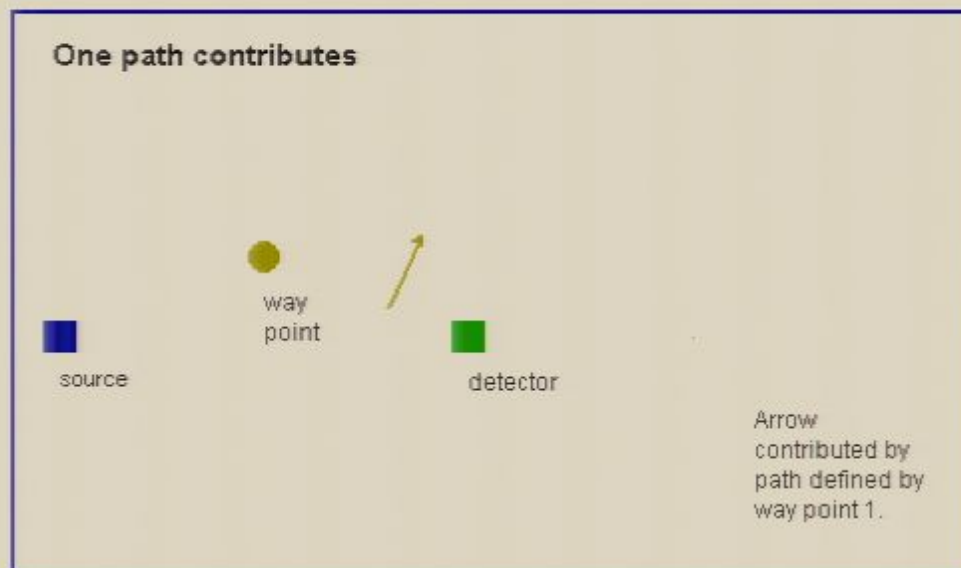




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14. Drag the waypoints and see the path **vector** change smoothly.
15. Look at the model. Only the first line – which is heavy Pythagoras – should be at all intimidating. Perhaps compare it with this one:



Extension:



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- Looking Inside Materials
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11. Run the model once using the 'go' button in the Control Window.
 12. Drag the waypoints to different positions and see how the arrows respond and how the amplitude changes.
- In this more unstructured exploration it is harder to find significant areas – but it does show the need for a planned attack on exploring all paths.

Extension:

Paths with two waypoints

The contribution of the path depends on the trip time. The trip time depends on the path length.



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4. Launch the model. Make sure you can see the Animation Window.



5. Run the model once using the 'go' button in the Control Window.

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At this stage you might be looking for regions from which the arrows tend to line up. These will be places where the trip times for adjacent waypoints are not so different.

Three paths contribute

As a further step we consider how the exploration of three paths generates three arrows. These three arrows are then summed to make an amplitude (imagine that there are only three paths – this reduces the problem of sampling all paths to manageable proportions). This will turn out to be a useful trick.

7. Launch the model. Make sure you can see the Animation Window.



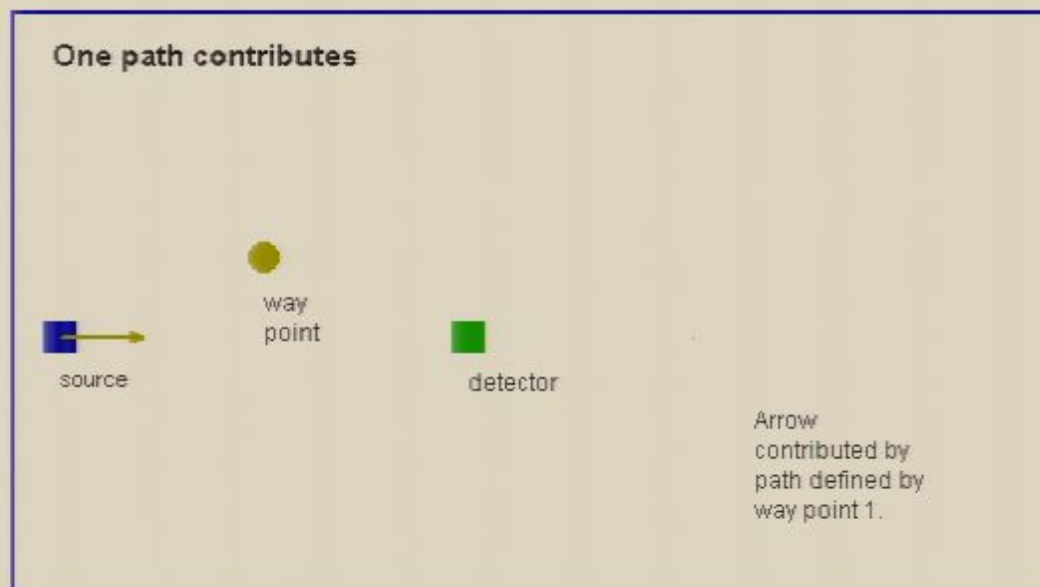
- Imaging
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One path contributes

This first step shows how a single path generates a single arrow.

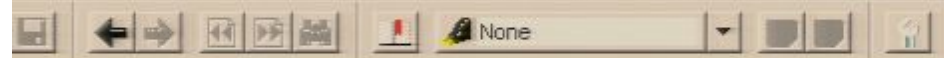
1. Launch the model. Make sure you can see the Animation Window.



2. Run the model once using the 'go' button in the Control Window.

Note how the **phasor**, the spinning, rotating arrow, marks the exploration of the path by the photon. Only when the phasor reaches the detector does the arrow freeze.

3. You can place the waypoint in different positions by clicking on the 'case' buttons on the menu bar of the Animation



- Imaging
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detector. Each path contributes an arrow. This activity explores how different paths contribute and how the contributions of those paths are aggregated. Each step makes one simple point only. Do not spend a long time on any one step.

This activity is designed for use with [File 20L](#) 'Getting arrows from paths' and [File 40L](#) 'Getting arrows from more complex paths'.

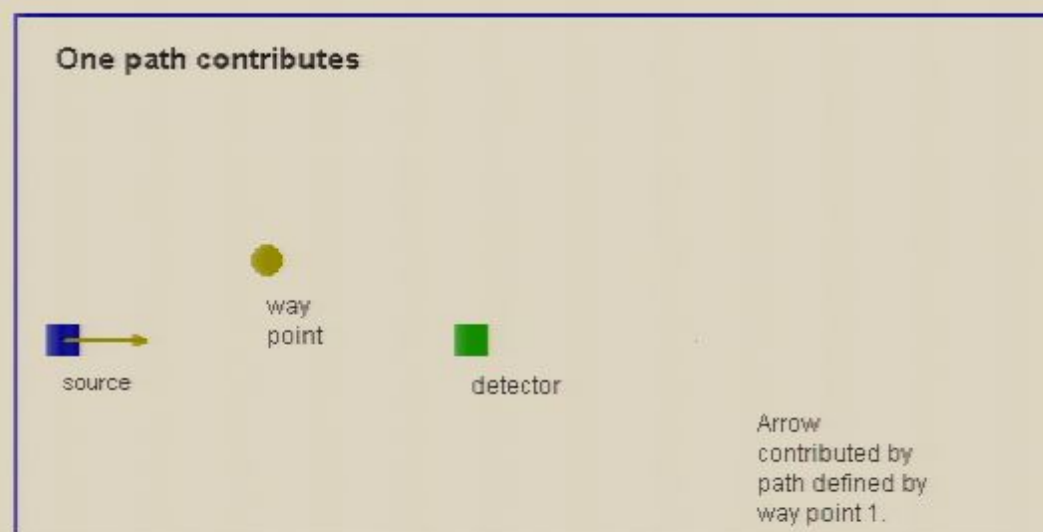
You will need

- ✓ PC running Modellus
- ✓ four or six model files

One path contributes

This first step shows how a single path generates a single arrow.

1. Launch the model. Make sure you can see the Animation Window.





- Files
- Questions
- Readings
- Student's Checklist
- Mapping Space and Time
- Computing the Next Move
- How to
- Teacher and Technician Information
- Specification
- A-Z of Key Terms

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 A model for the random arrival of photons

[Teaching Notes](#) | [Key Terms](#) | [Resources](#)

Getting arrows from paths

File 20L: Launchable File



A single arrow

The first pair looks at the process of getting a single arrow from a single explored path. Each is a simple moving model to make one point.

This file is designed to go with [Activity 50S](#) 'How exploring paths leads to arrows'.

An arrow from a path

One path contributes





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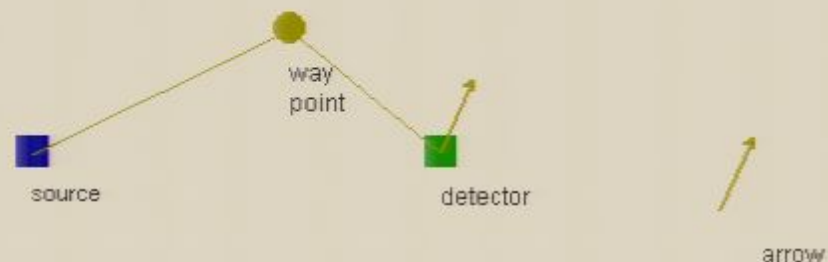
This model runs through the process of a photon exploring a path using a phasor. You see the path being explored.



Open the Modellus file

One arrow from dragged waypoints

A path contributes



Here you see only the result of the path being explored. As you alter the path, so you see the resulting arrow. The longer the path, the more spins for the phasor.



Open the Modellus file

Three spinning

MODELLUS - UNTITLED

Edit Case Window Help

del

\sqrt{x}

π

e

Δx

$x = y$

$\ln x$

Interpret

Control

t = 0.00

020

Options...

Initial Conditions

Parameters

case 1

Modellus

Interactive Modelling with Mathematics

Authors:

Vitor Duarte Teodoro

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Filipe Costa Clérigo

Faculty of Sciences and Technology

New University of Lisbon, Portugal

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<http://phoenix.sce.fct.unl.pt/modellus>
modellus@mail.fct.unl.pt

Advancing Physics




Advancing achievement

<http://post16.iop.org/advphys>

Model

x^n \sqrt{x} π e Δx $x \sim$ $\frac{d}{dt}x$

Interpret

;path calculations

$$hyp = \sqrt{(ywp1)^2 + (xwp1)^2}$$

$$\text{if}(x1 \leq xwp1) \text{ and } (x1 < detectx) \text{ then } \left(y1 = v \times \frac{ywp1}{hyp} \times t \right)$$

$$\text{if}(x1 > xwp1) \text{ and } (x1 < detectx) \text{ then } \left(y1 = 2 \times ywp1 - \left(v \times \frac{ywp1}{hyp} \times t \right) \right)$$

$$\text{if}(x1 < detectx) \text{ then } \left(x1 = v \times \frac{xwp1}{hyp} \times t \right)$$

;spinning



$$\text{if}(x1 < detectx) \text{ then } (ply = A1 \times \sin(2 \times \pi \times fl \times t)) \text{ and } (plx = A1 \times \cos(2 \times \pi \times fl \times t))$$





case 3	case 4	case 5
1.00	100.00	100.00
1.00	200.00	200.00
00	100.00	120.00
00	20.00	20.00
0	4.00	4.00
1.00	100.00	100.00

Control

t = 0.00

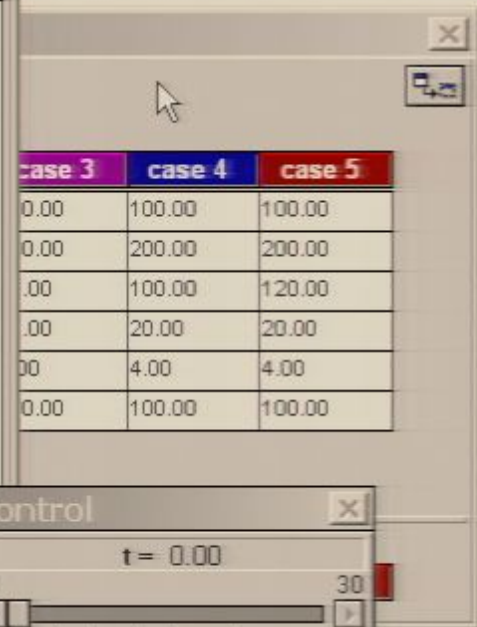
0 30

Options...

Use the case buttons to find the arrows controlled by different paths, defined by different way points.



Notes

Model

Press the stair

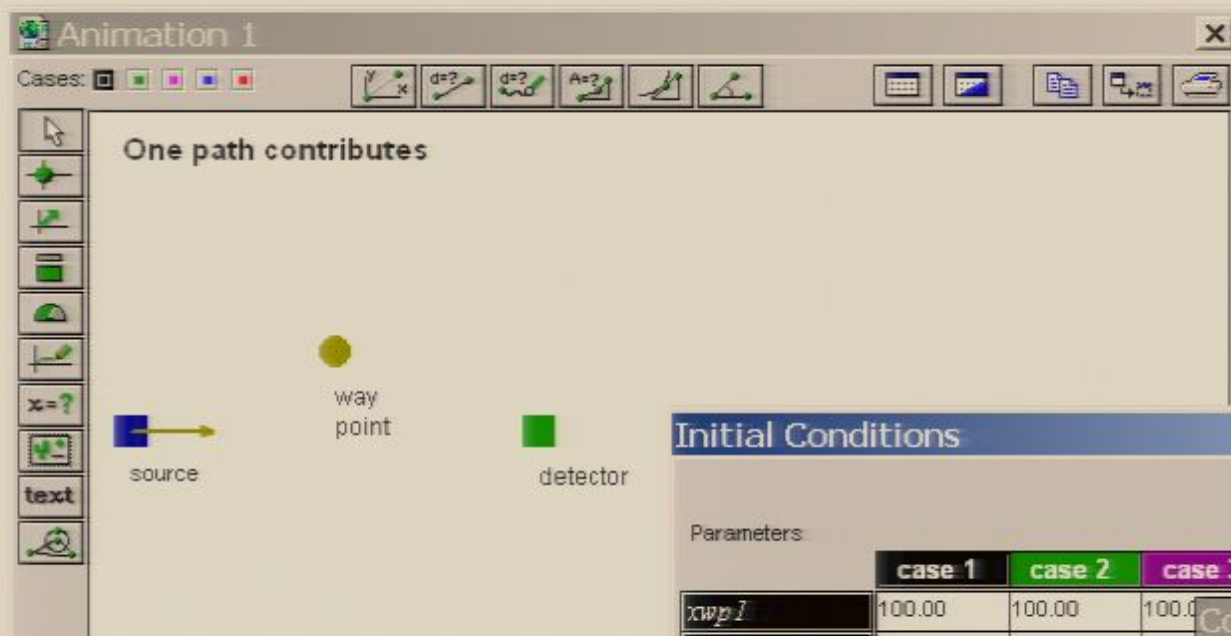
Use the case
different paths

path calculations

$$hyp = \sqrt{(ywp\ l)^2 + (xwp\ l)^2}$$

if $(x1 \leq xwpl)$ and $(x1 < detectx)$ then $y1 = v \times \frac{ywpl}{hyp} \times t$

if $(x_l > x_{wpl})$ and $(x_l < detect_x)$ then $y_l = 2 \times y_{wpl} - \left(v \times \frac{y_{wpl}}{hyp} \times t \right)$

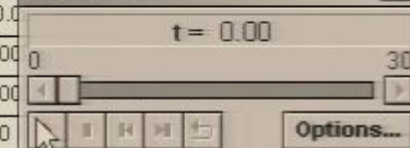


Initial Conditions

Parameters:

	case 1	case 2	case 3	case 4	case 5
x_{wp1}	100.00	100.00	100.00		
$detectxc$	200.00	200.00	200.00		
y_{wp1}	40.00	60.00	80.00		
v	20.00	20.00	20.00		
AI	4.00	4.00	4.00		
$f1$	100.00	100.00	100.00	100.00	100.00

Control



Options...

Initial values:

case 1	case 2	case 3	case 4	case 5
--------	--------	--------	--------	--------

Notes

Press the start

Use the case
different paths:

Model

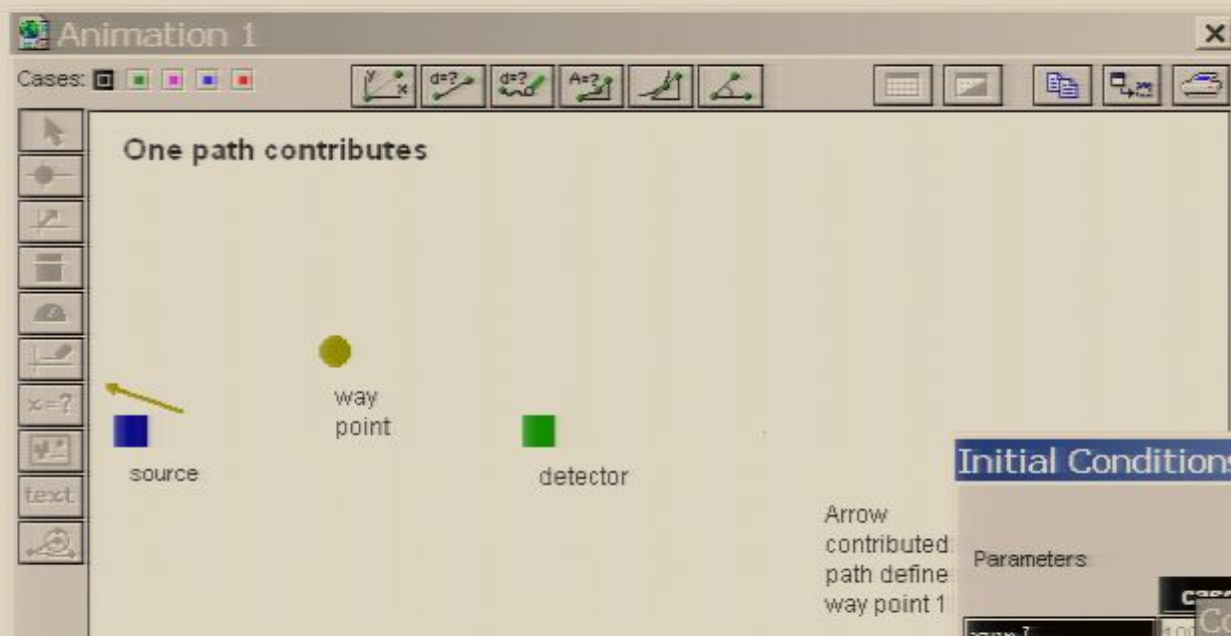
x^n \sqrt{x} π e Δx $x \sim$ $\ln x$

path calculations

$$hyp = \sqrt{(y_{wp1})^2 + (x_{wp1})^2}$$

$$\text{if } (x1 \leq x_{wp1}) \text{ and } (x1 < detectxc) \text{ then } \left\{ y1 = v \times \frac{y_{wp1}}{hyp} \times t \right\}$$

$$\text{if } (x1 > x_{wp1}) \text{ and } (x1 < detectxc) \text{ then } \left\{ y1 = 2 \times y_{wp1} - \left(v \times \frac{y_{wp1}}{hyp} \times t \right) \right\}$$



Notes

Press the star

Use the case different paths

Model

;path calculations

$$hyp = \sqrt{(ywp1)^2 + (xwp1)^2}$$

$$\text{if}(x1 \leq xwp1) \text{ and } (x1 < detectx) \text{ then } \left\{ y1 = v \times \frac{ywp1}{hyp} \times t \right\}$$

$$\text{if}(x1 > xwp1) \text{ and } (x1 < detectx) \text{ then } \left\{ y1 = 2 \times ywp1 - \left(v \times \frac{ywp1}{hyp} \times t \right) \right\}$$

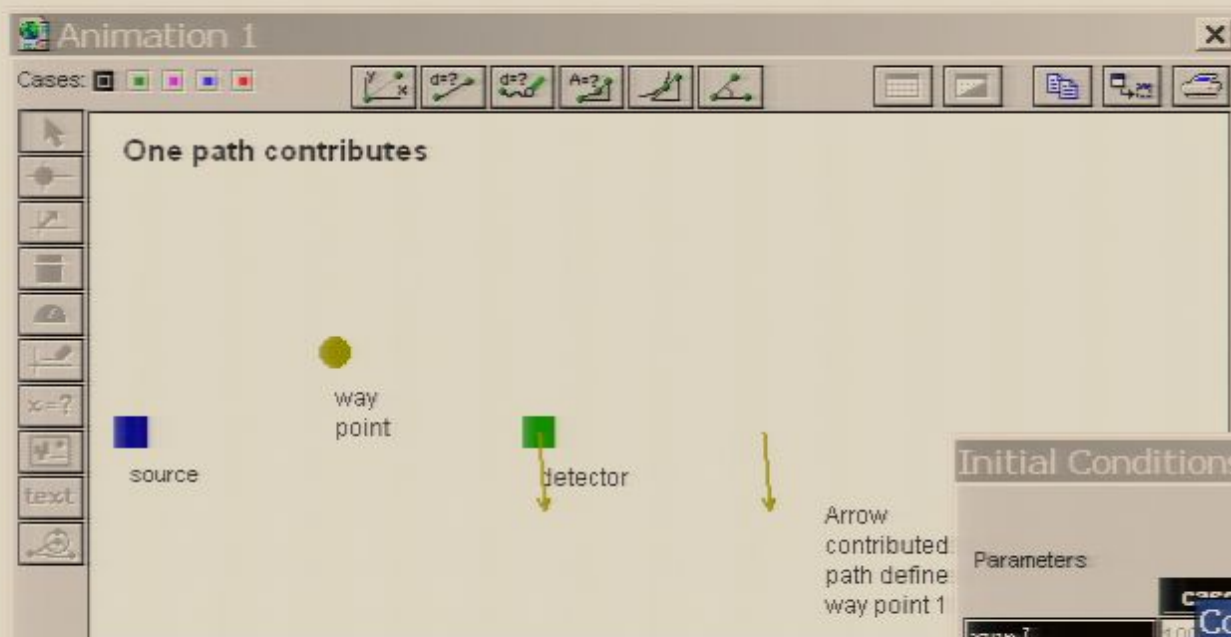
Initial Conditions

Parameters:

	case 1	case 2	case 3	case 4	case 5
xwp1	100.00	100.00	100.00	100.00	100.00
detectx	200.00	200.00	200.00	200.00	200.00
ywp1	40.00	40.00	40.00	40.00	40.00
v	20.00	20.00	20.00	20.00	20.00
A1	4.00	4.00	4.00	4.00	4.00
f1	100.00	100.00	100.00	100.00	100.00

Initial values:

case 1	case 2	case 3	case 4	case 5
--------	--------	--------	--------	--------



Notes

Press the star

Use the case different paths

Model

;path calculations

$$hyp = \sqrt{(ywp1^2 + (xwp1)^2)}$$

$$\text{if}(x1 \leq xwp1) \text{ and } (x1 < detectx) \text{ then } \left\{ y1 = v \times \frac{ywp1}{hyp} \times t \right\}$$

$$\text{if}(x1 > xwp1) \text{ and } (x1 < detectx) \text{ then } \left\{ y1 = 2 \times ywp1 - \left(v \times \frac{ywp1}{hyp} \times t \right) \right\}$$

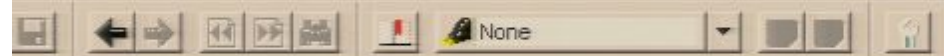
Initial Conditions

Parameters:

	case 1	case 2	case 3	case 4	case 5
xwpl	100	100	100	100	100
detectx	200	200	200	200	200
ywpl	40	40	40	40	40
v	20	20	20	20	20
A1	4.00	4.00	4.00	4.00	4.00
f1	100.00	100.00	100.00	100.00	100.00

Initial values

case 1	case 2	case 3	case 4	case 5
--------	--------	--------	--------	--------



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- Teacher and Technician Information
- Specification
- Index of Key Terms

- 7: Quantum Behaviour
- Resource Manager
- Files
- Launchable Files
- Getting arrows from paths

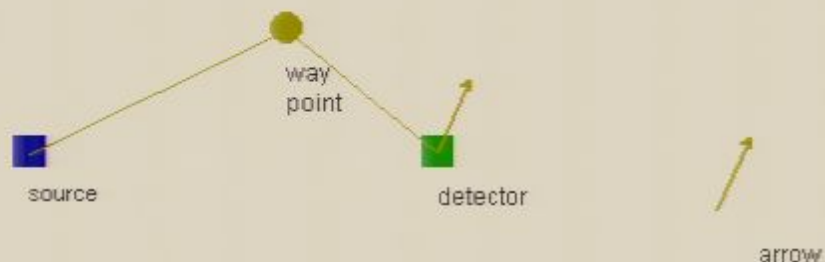
This model runs through the process of a photon exploring a path using a phasor. You see the path being explored.



Open the Modellus file

One arrow from dragged waypoints

A path contributes



Here you see only the result of the path being explored. As you alter the path, so you see the resulting arrow. The longer the path, the more spins for the phasor.



Open the Modellus file

Three spinning

Advancing Physics A2

- **The Rise and Fall of the Clockwork Universe**
 - [Models and Rules](#) introduces the use of mathematical modelling in physics
 - [Matter in Extremes](#) develops a bridge between random events and empirical laws
- **Field and particle pictures**
 - [Fields](#) develops ideas about magnetic and electric fields
 - [Fundamental particles](#) considers the structure and binding of atoms and nuclei and introduces ideas about risk
- **Advances in Physics**
 - Draws together and uses physics ideas from the whole course in looking at new contexts

IoP Teachers Network



Institute of Physics
Teaching Physics

Pirsa: 06050012

the international physics teaching journal

Physics*education*

Volume 41 Number 2 March 2006

ISSN 0031-9120

Physics*education*
www.iop.org/journals/physed

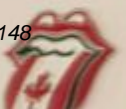


Dramatic Demonstrations

Amaze your students with simple equipment

News: ASE '06 report, 'Visualise' Frontline: Retraction, Spectra, Resonance

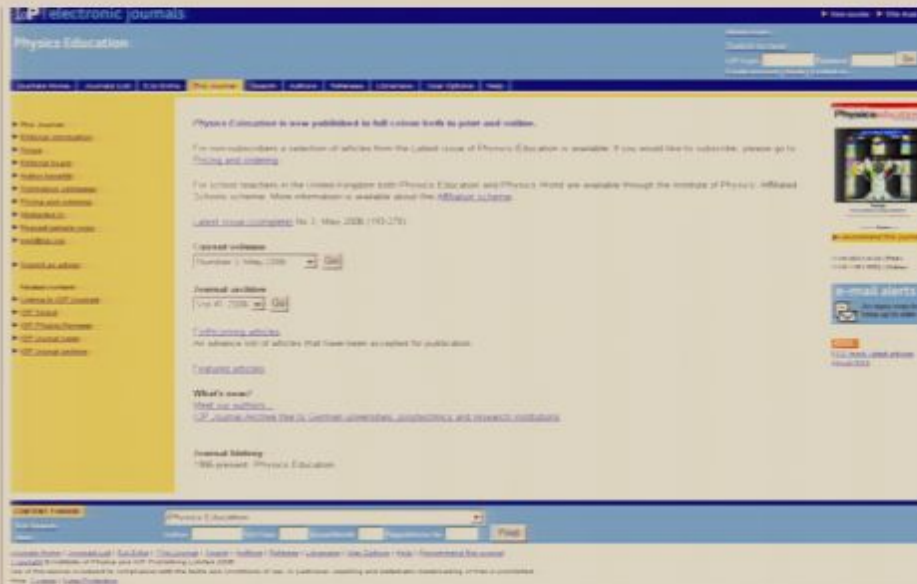
Institute of Physics PHYSICAL SOCIETY



the international physics teaching journal

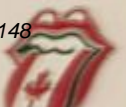
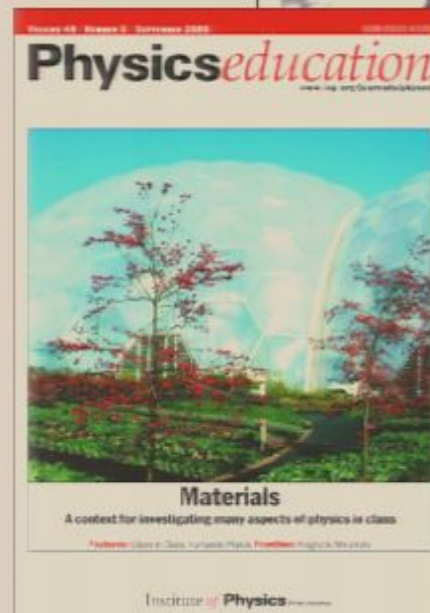
Physics*education*

Six issues a year

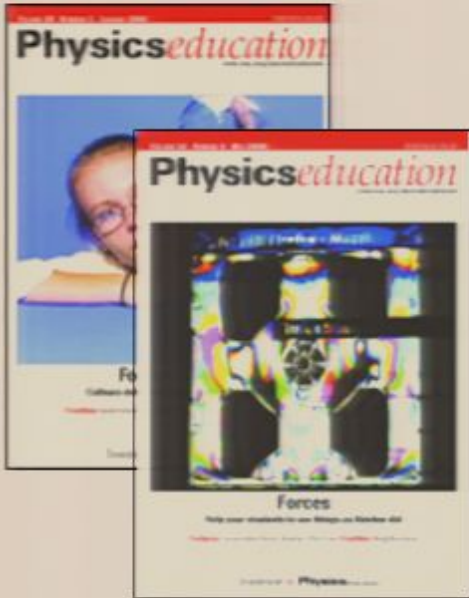


Access to 10 years of
electronic archive

www.iop.org/journals/physed



the international physics teaching journal

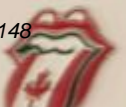


Can you write?

2006 individual rate - \$137
(US\$122)

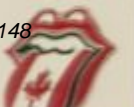
**YOU WILL NOT SEE THIS
ADVERTISED**

E-mail: ped@iop.org



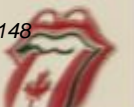
IoP Teachers Network

Institute of Physics
Teaching Physics



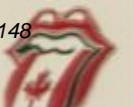
Aim..

Institute of Physics
Teaching Physics



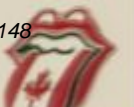
Aim..

- The aim of the Physics Teacher Network is to provide support for those involved with the teaching of physics, at an individual level.

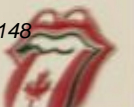


Aim..

- The aim of the Physics Teacher Network is to provide support for those involved with the teaching of physics, at an individual level.
- We want to engage with as many science and physics teachers and technicians as we can, helping them to improve the experience that students associated with them undergo.

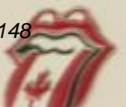


And....



- “...at an individual level” doesn’t mean that we can cater for every individuals needs.

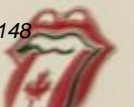
And....



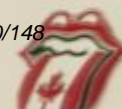
- “...at an individual level” doesn’t mean that we can cater for every individuals needs.

And....

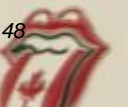
- It does mean that we can try and cater for the local needs of those we come in contact with
- Different regions will have different needs and require different support



Now...

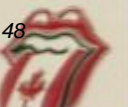


So what do we do...?



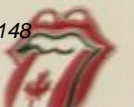
So what do we do...?

- Coordinators own sessions



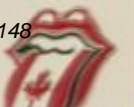
So what do we do...?

- Coordinators own sessions
- Centrally resourced sessions



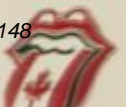
So what do we do...?

- Coordinators own sessions
- Centrally resourced sessions
- Newsletters



So what do we do...?

- Coordinators own sessions
- Centrally resourced sessions
- Newsletters
- Anything and everything!



The National Physical Laboratory and A-Level Physics

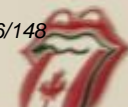
Monday 3rd July 2006

09:30 – 16:00

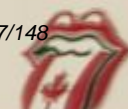
The National Physical Laboratory, Teddington



Institute of Physics
Teaching Physics



The National Physical Laboratory and
Monday 3rd July 2000
09:30 – 16:00
The National Physical Laboratory

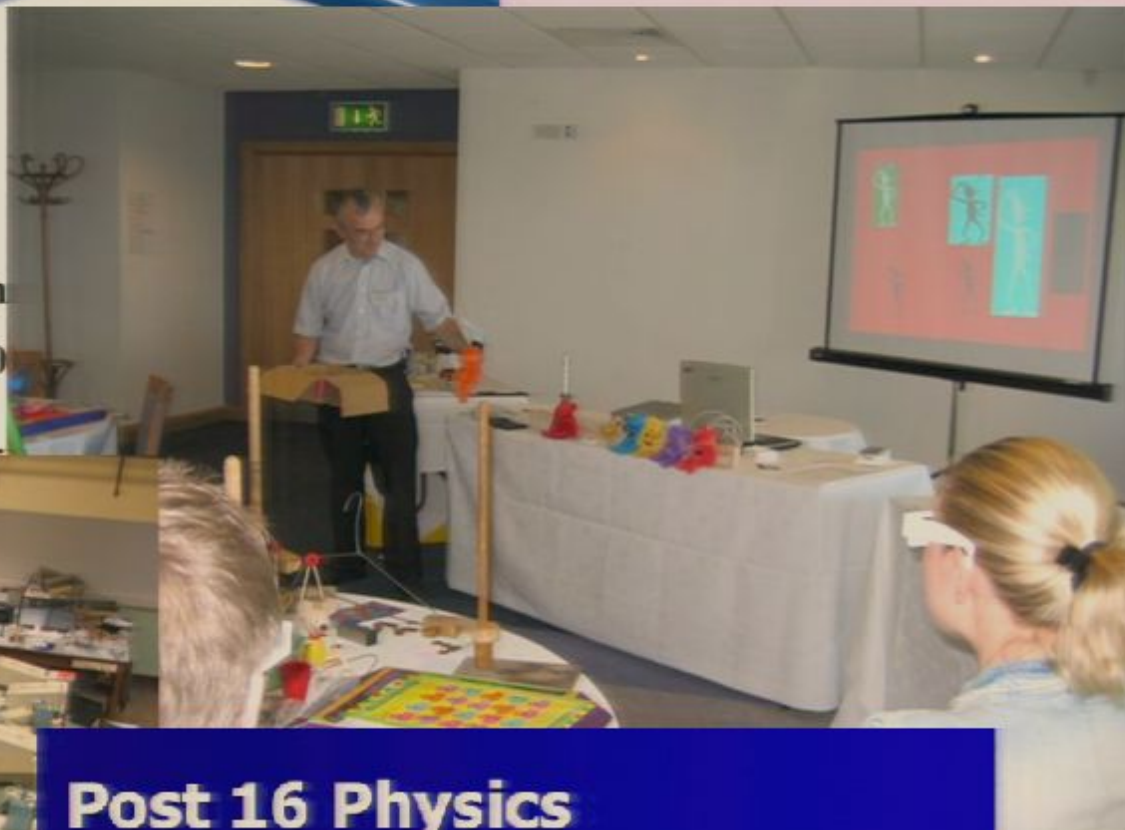


The National Physical Laboratory and

Monday 3rd July 200

09:30 – 16:00

The National Physical Laboratory

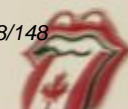


Post 16 Physics

Helen Pollard

6 December 2005

University of Nottingham



The winners from the Y10 category were three girls from Edgbaston High School



The winners from the Y9 category were from King Edward V Handsworth School,

with runners up from Edgbaston High School



Physics

d
2005
f Nottingham



The winners from the Y10 category were three girls from Edgbaston High School



Rocket Day at Swanshurst School June 30th 2005



Customising the designs.....



The winners from the Y9 category were Edward VI Handsworth School,

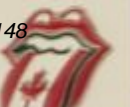


with runners up from Edgbaston High School



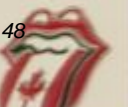
Displaying the finished rockets.....

Centrally resourced sessions



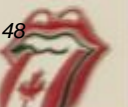
Centrally resourced sessions

- Software for Skint Schools



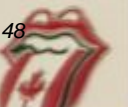
Centrally resourced sessions

- Software for Skint Schools
- New Ideas?



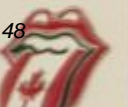
Centrally resourced sessions

- Software for Skint Schools
- New Ideas?
- Son of New Ideas?



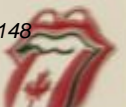
Centrally resourced sessions

- Software for Skint Schools
- New Ideas?
- Son of New Ideas?
- SEP - Energy Boards



Centrally resourced sessions

- Software for Skint Schools
- New Ideas?
- Son of New Ideas?
- SEP - Energy Boards
- SEP - Thermochromic sheet

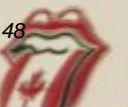


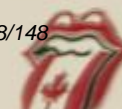
Centrally resourced sessions

- Software for Skint Schools
- New Ideas?
- Son of New Ideas?
- SEP - Energy Boards
- SEP - Thermochromic sheet

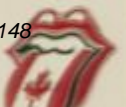
In the pipeline:

- Shocked and Stunned
- New Ideas Rides Again?
- More from SEP

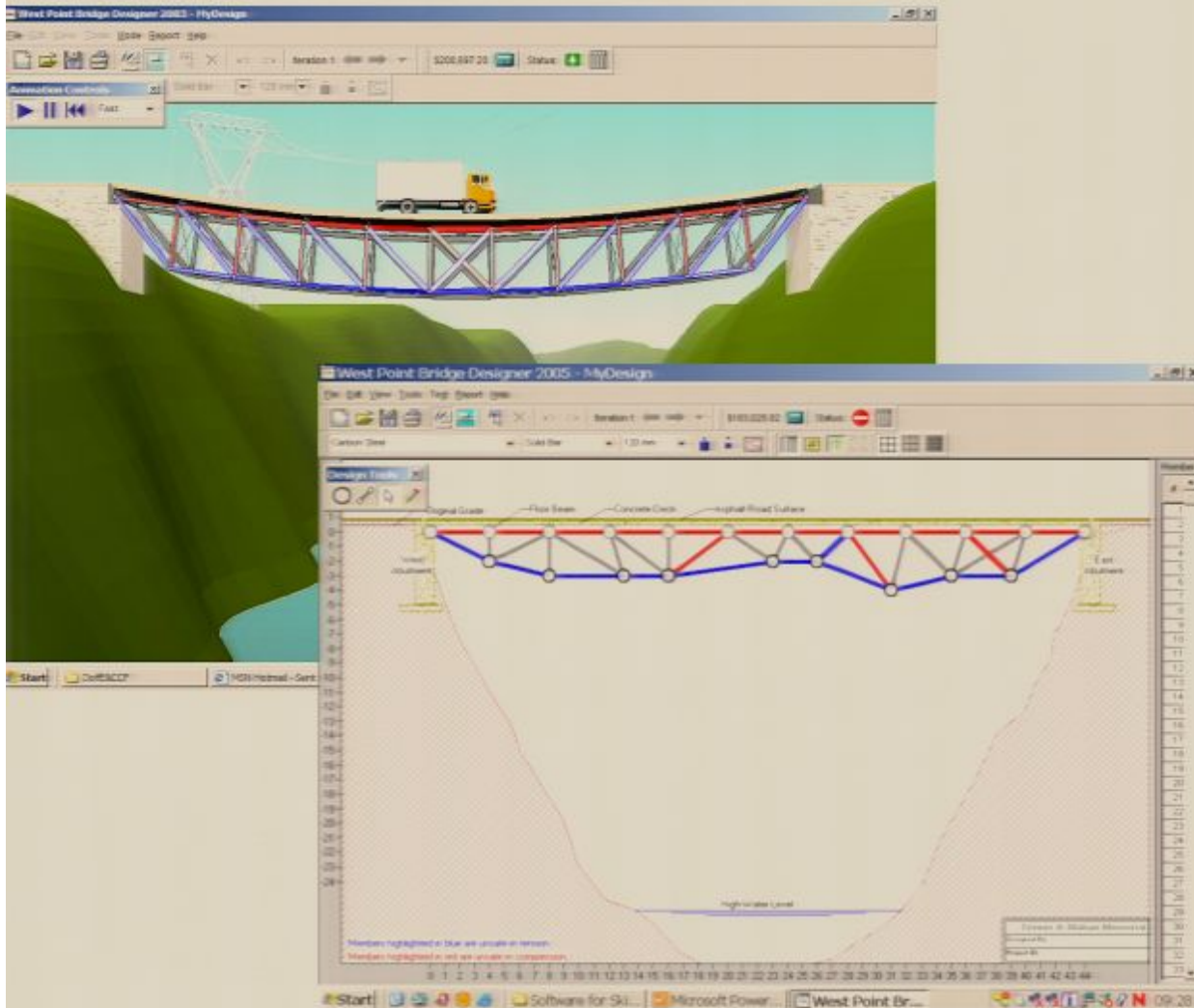




Software for
Skint Schools



Westpoint Bridge Designer

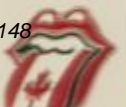


From Westpoint
Military

Academy. Allows
you to design,
build and test
bridges.

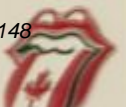
It also has
information
about the
materials used and
their cost.

Easy to use and
very stable, very
popular with all
ages from 4 up.



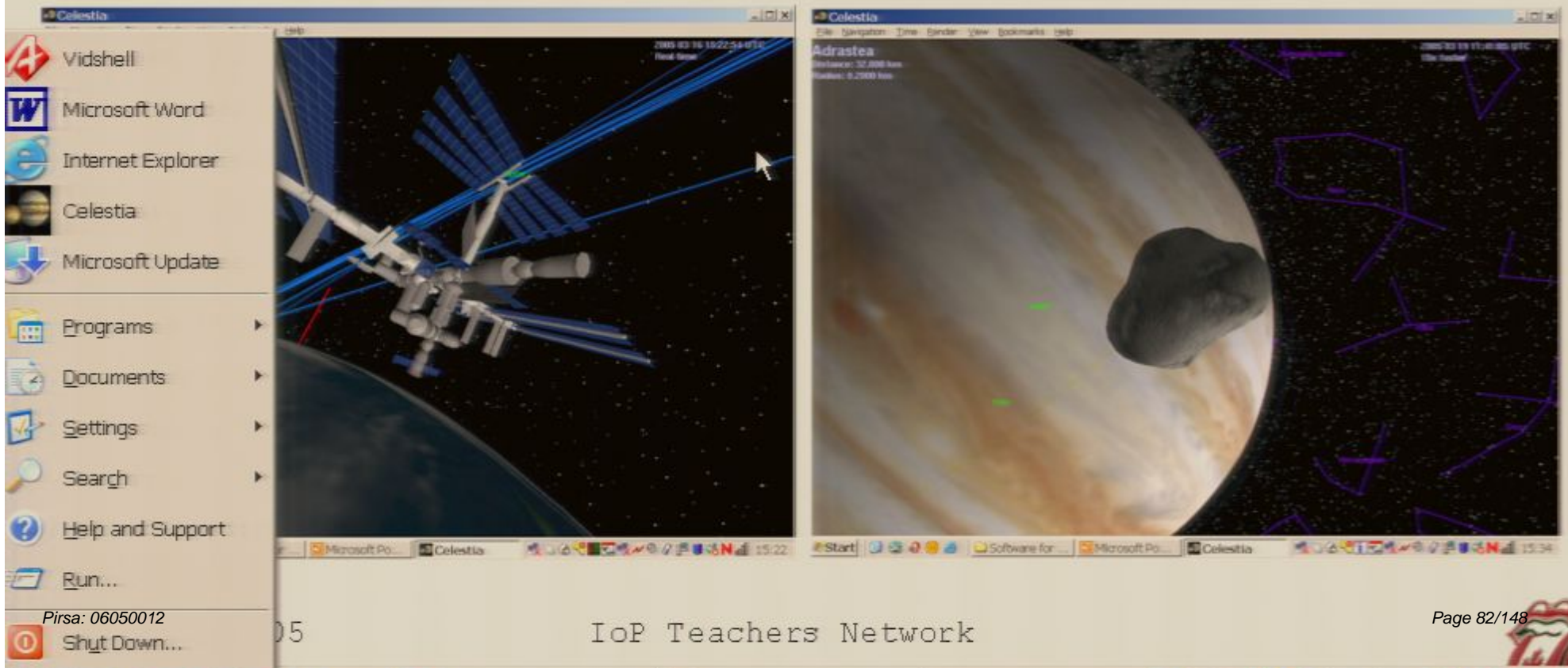
Celestia

A virtual chunk of the Galaxy. Really needs a good computer to do the software justice. You have to see this to appreciate how good it is.



Celestia

A virtual chunk of the Galaxy. Really needs a good computer to do the software justice. You have to see this to appreciate how good it is.

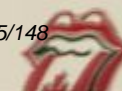
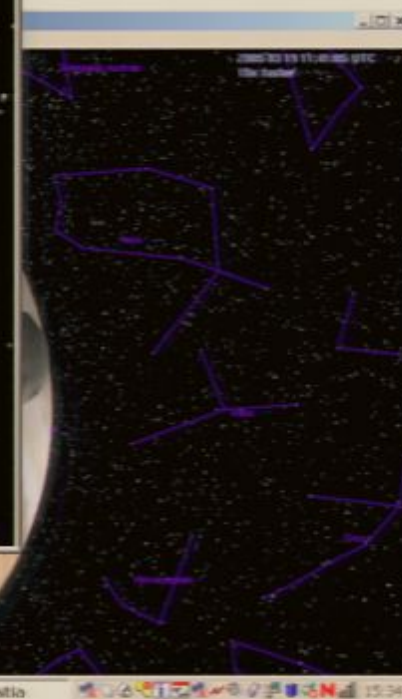


Celestia

A virtual chunk of the Galaxy. Really needs a good computer to do the software justice. You have to see this to appreciate how good it is.







2006 05 25 23:32:55 UTC
Real timeDistance: 7,286.4 km
Altitude: 1,821.6 km

Dazhbog Patera

Loki Patera

Pele

Pisa: 00050012

Speed: 0.000 m/s

Page 80/148
Follow to
FOV: 25 55' 33.0" (1.00 x)



Distance: 7,286.4 km

Altitude: 1,821.6 km

2006 05 25 23:32:59 UTC

Time stopped

Dazhbog Patera

Loki Patera

Beginning demo
Press ESC to end.

Pele

Pisa: 00050012

Page 89/148

Speed: 0.000 m/s

Follow to

FOV: 25 55' 33.0" (1.00 x)



Celestia

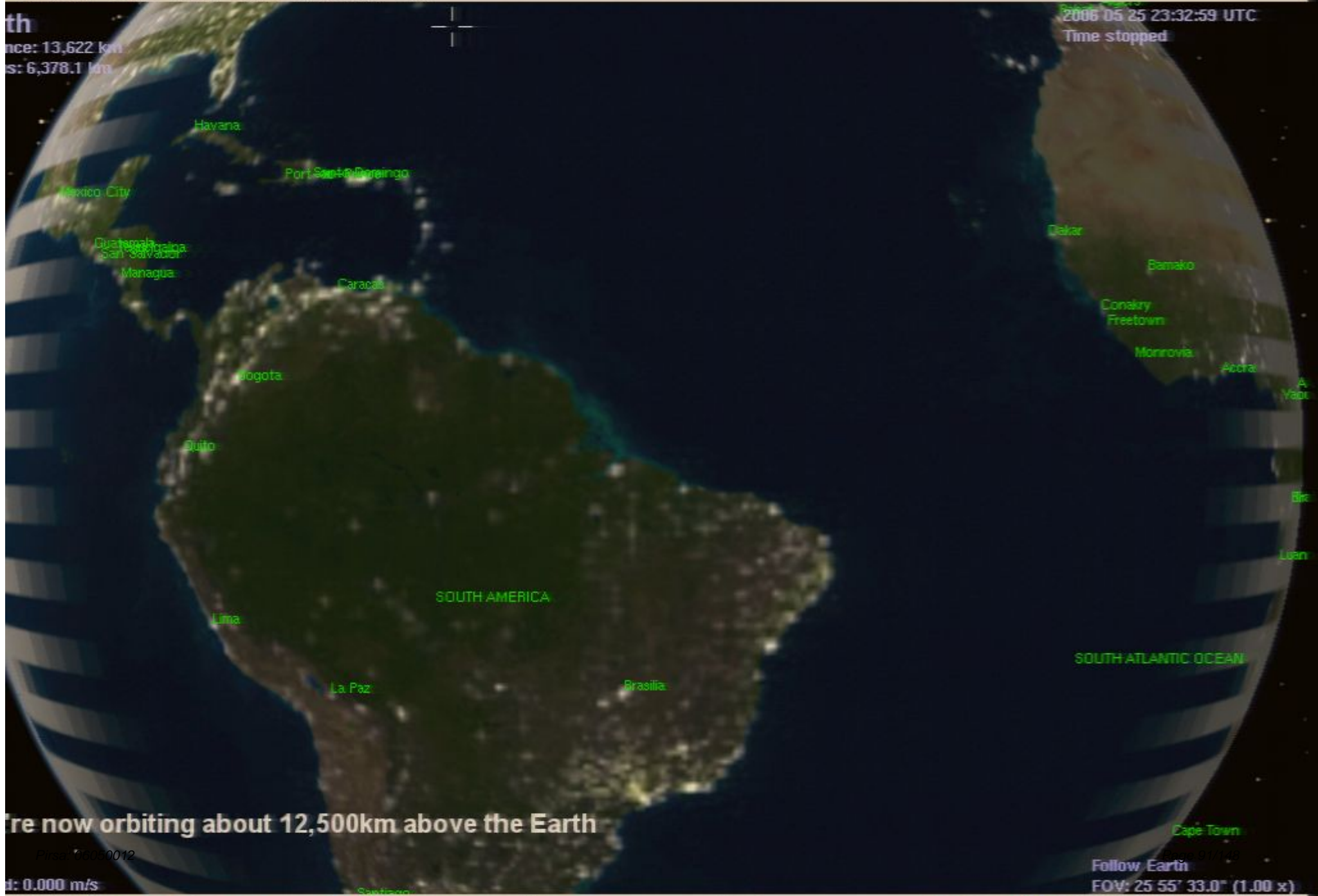
Navigation Time Render View Bookmarks Help

th

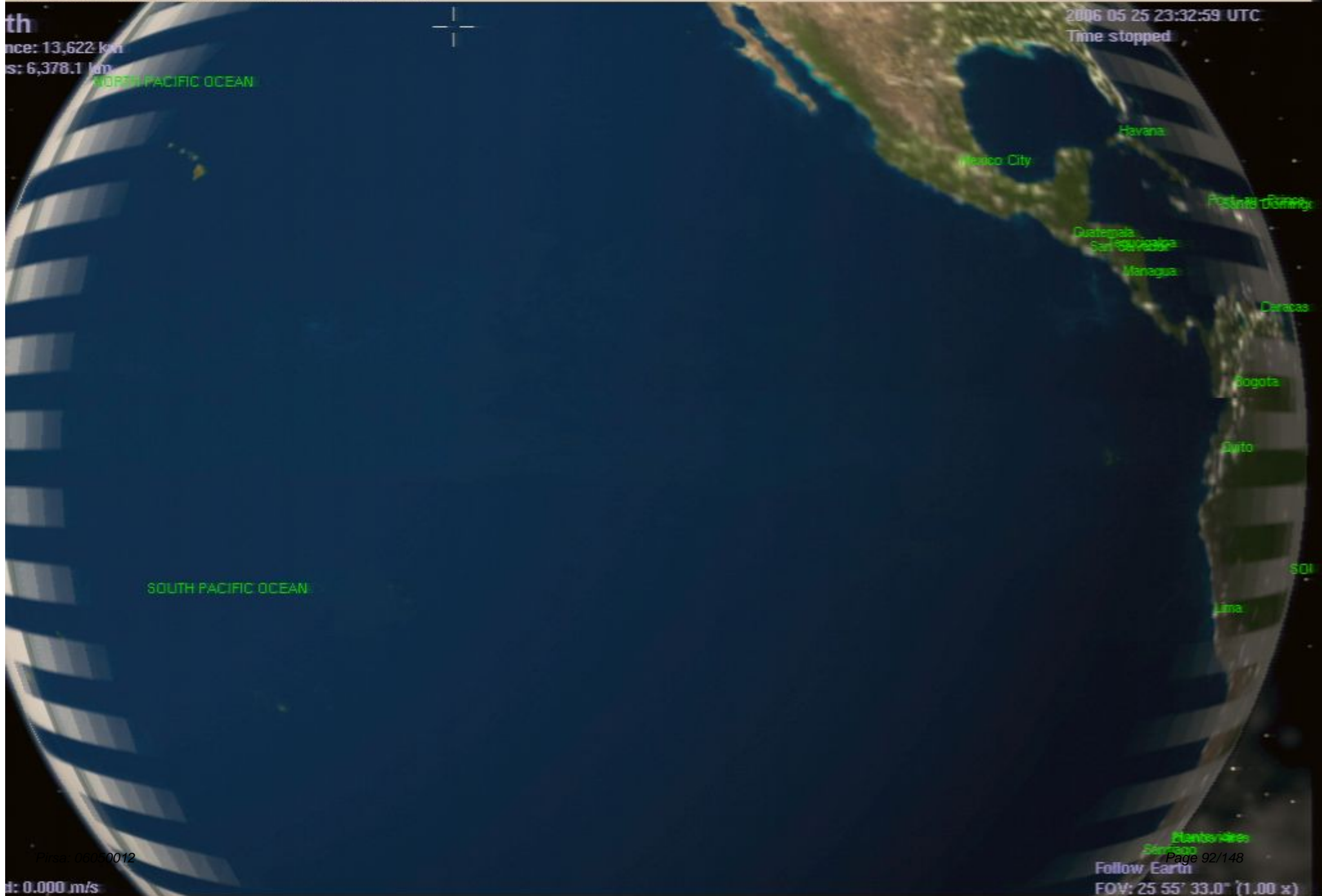
nce: 13,622 km

s: 6,378.1 km

2006 05 25 23:32:59 UTC
Time stopped



We're now orbiting about 12,500km above the Earth





th
nce: 13,622 km
s: 6,378.1 km

2006-06-25 23:32:59 UTC
Time stopped

Taipei
Victoria
Hanoi
Bangkok
Phnom Penh
Manila
Kuala Lumpur
Singapore
Jakarta

NORTH PACIFIC OCEAN

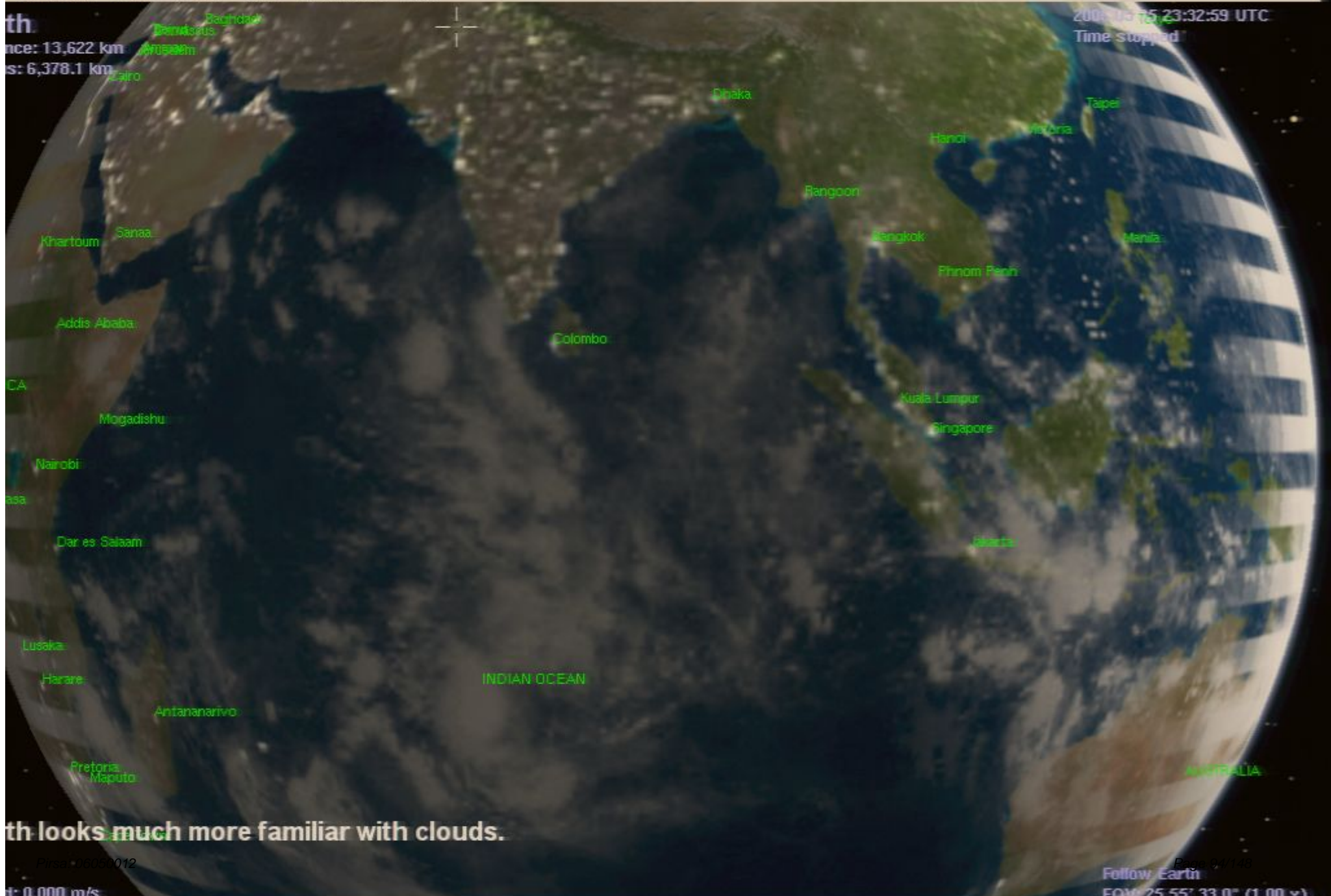
SOUTH PACIFIC OCEAN

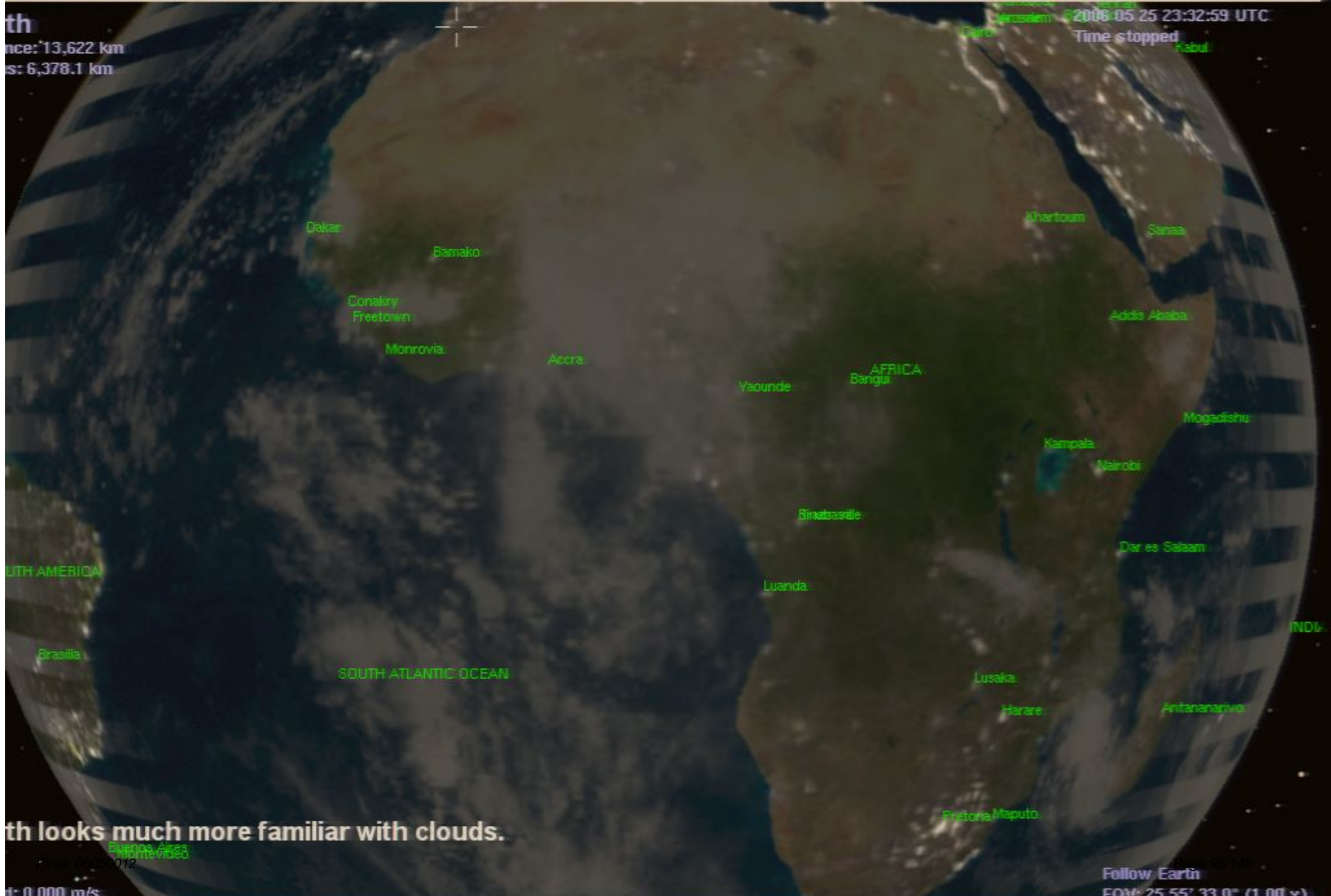
AUSTRALIA

Pisa: 06050012

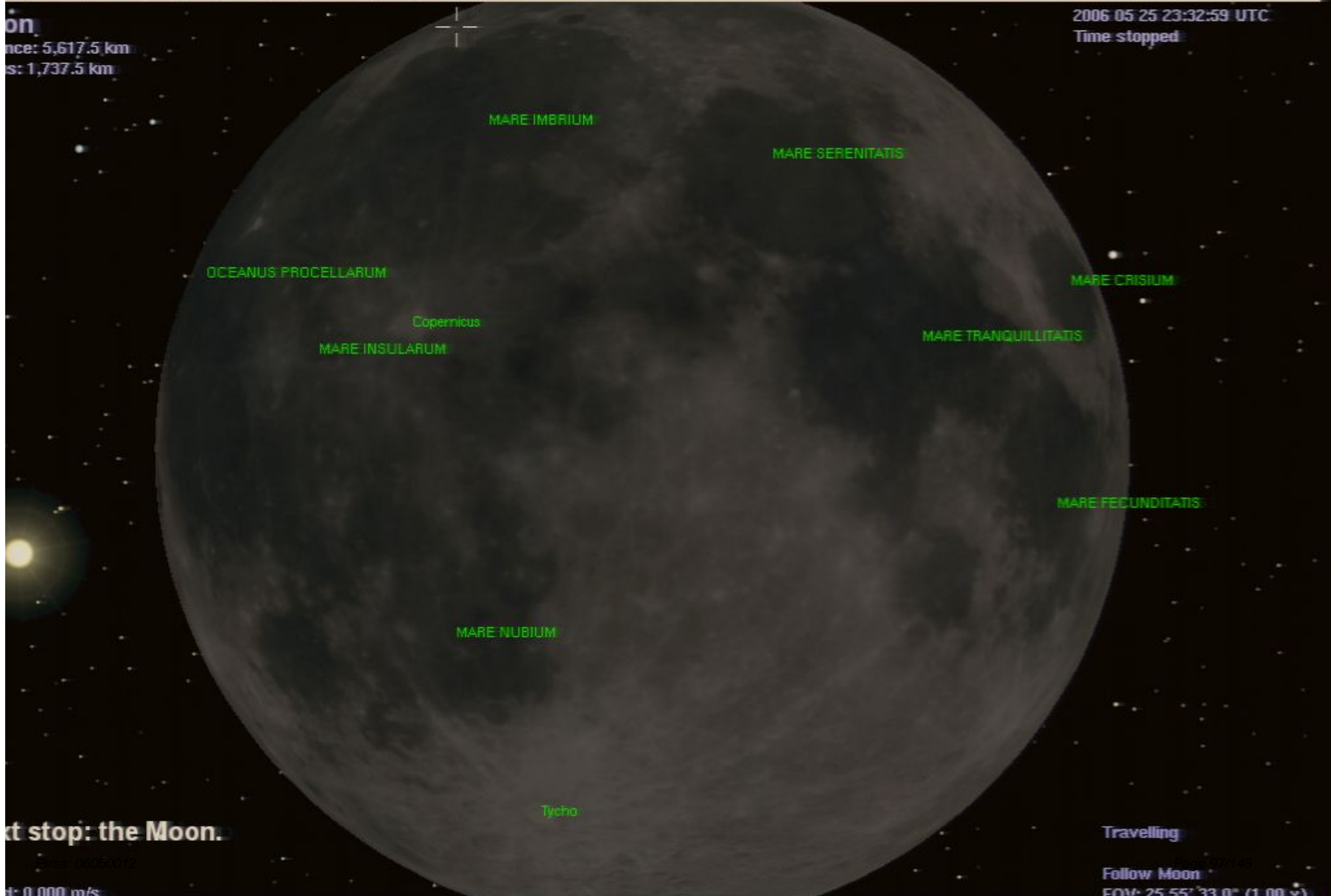
d: 0.000 m/s

Follow Earth
FOV: 25 55' 33.0° (1.00 x)









on
nce: 5,617.5 km
s: 1,737.5 km

2006 05 25 23:32:59 UTC
Time stopped

at stop: the Moon.

id: 06050012

id: 0.000 m/s

Travelling

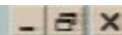
Follow Moon

FOV: 25 55' 33.0" (1.00 x)



elestia

Navigation Time Render View Bookmarks Help



on

nce: 5,212.6 km

s: 1,737.5 km

2006 05 25 23:32:59 UTC

Time stopped

MARE MOSCOVIENSE

MARE ORIENTALE

atch for the Earth and Sun as we orbit the Moon

Pisa: 06050012

d: 0.000 m/s

Page 99/140

Follow Moon

FOV: 25 55' 33.0" (1.00 x)



on
nce: 5,212.6 km
s: 1,737.5 km

2006 05 25 23:32:59 UTC
Time stopped

MARE CRISIUM

MARE MOSCOWENSE

Tsiolkovsky

MARE AUSTRALE

Watch for the Earth and Sun as we orbit the Moon

Speed: 0.000 m/s

Page 100/140
Follow Moon
FOV: 25 55' 33.0" (1.00 x)

2006 05 25 23:32:59 UTC
Time stopped

Distance: 0.45214 au
Apparent mag = 4.83 (-28.47)
Proximity: 0.999x Sun
Type: G2 V

toward the Sun.

Proximity: 0.999x Sun

Speed: 0.000 m/s

Travelling (3)

Page 101/148

Follow Sol

FOV: 25 55' 33.0" (1.00 x)

2006 05 25 23:32:59 UTC
Time stopped

Distance: 8,348,500 km
(app) mag = 4.83 (-33.01)
Proximity: 0.999x Sun
Type: G2 V

At this distance, dark sunspots are visible on the Sun's surface.

Pirsa: 06050012

Page 102/148

Speed: 0.000 m/s

Follow Sun
FOV: 25 55' 33.0" (1.00 x)

2007 01 10 21:48:48 UTC
2592000x faster

Distance: 2.7976 au
(app) mag = 4.83 (-24.51)
Density: 0.999x Sun
G2 V

Mercury

LMC

Venus

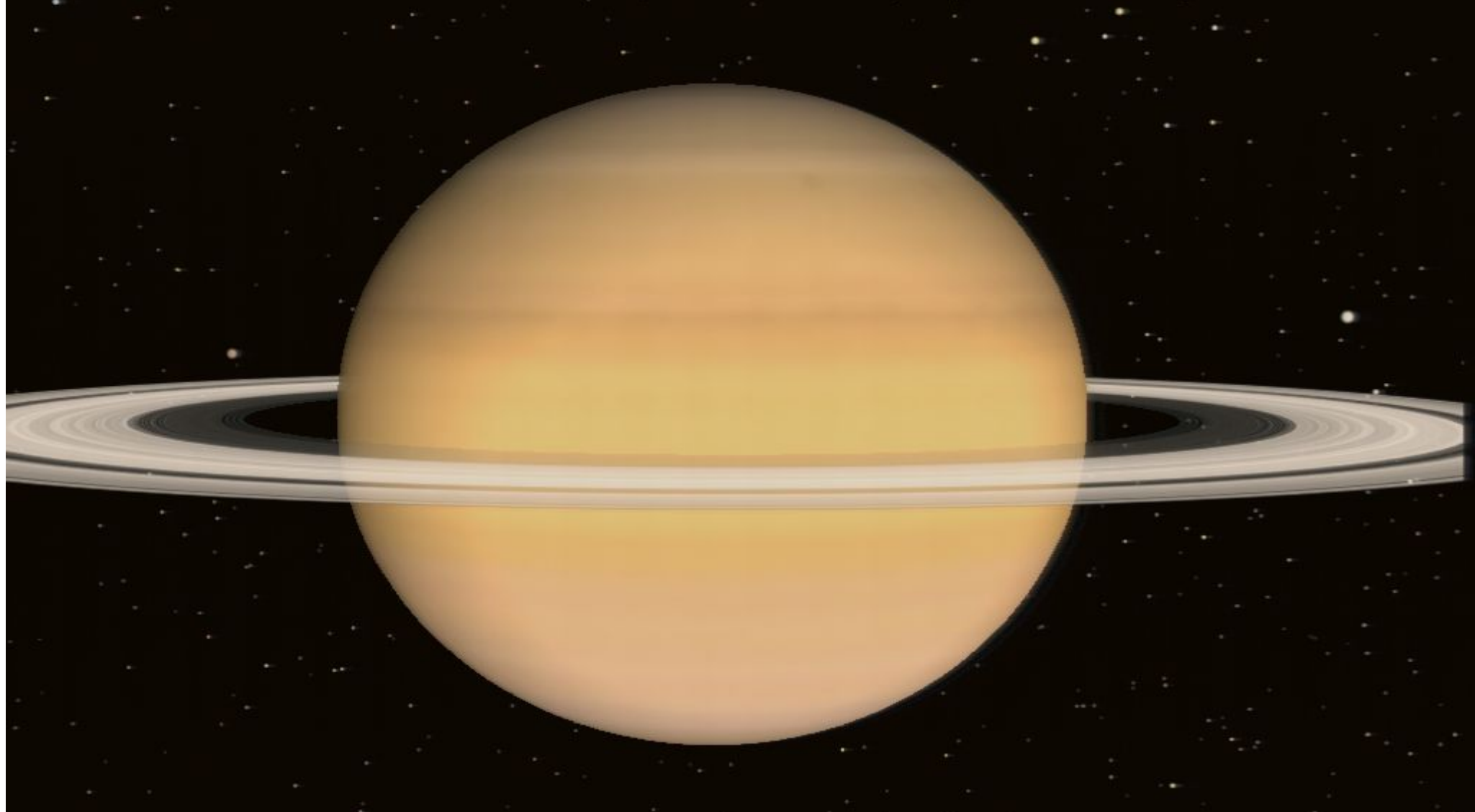
Earth

each second, a month of time elapses in the simulation.

Pos: 06050012

Vel: 0.000 m/s

Follow Sol
FOV: 25 55' 33.0" (1.00 x)



Polaris / Alruqaba / ALF UMi / 1 UMi / HD 8890 / HIP 11767 2007 08 18 22:49:09 UTC
Real time

Distance: 431.44 ly
Apparent magnitude: -3.64 (1.97)
Proper motion: 2,440 mas/yr
Spectral type: F7 I-b



help us get oriented in the sky, Celestia can draw constellation diagrams for us . . .

Pisa: 06050012 105/148

Alt: 0.000 m/s Follow Minus
FOV: 25 55' 33.0" (1.00 x)

celestia

Navigation Time Render View Bookmarks Help

Mimosa / BET Cru / HD 111123 / HIP 62434

Distance: 352.61 ly
Spectral type: (app) mag = -3.92 (1.25)
Proximity: 3,160x Sun
Type: B0 III

2007 08 18 22:49:23 UTC

Real time

Orion

Musca

Bigel Kentaurus A

Egena

Crux

Crux

Mimosa

Gacrux

The Southern Cross is a familiar sight in Southern Hemisphere skies.

Pisa: 06050012

Alt: 0.000 m/s

Follow Mimosa

FOV: 25 55' 33.0" (1.00 x)

Antares / Calbalacrab / ALF Sco / 21 Sco / HD 148478 / HIP 80763

2007 08 18 22:50:25 UTC

Real time

Distance: 693.75 ly

Milky Way

Apparent mag = -5.28 (1.36)

Density: 11,100x Sun

Type: M1 I-b

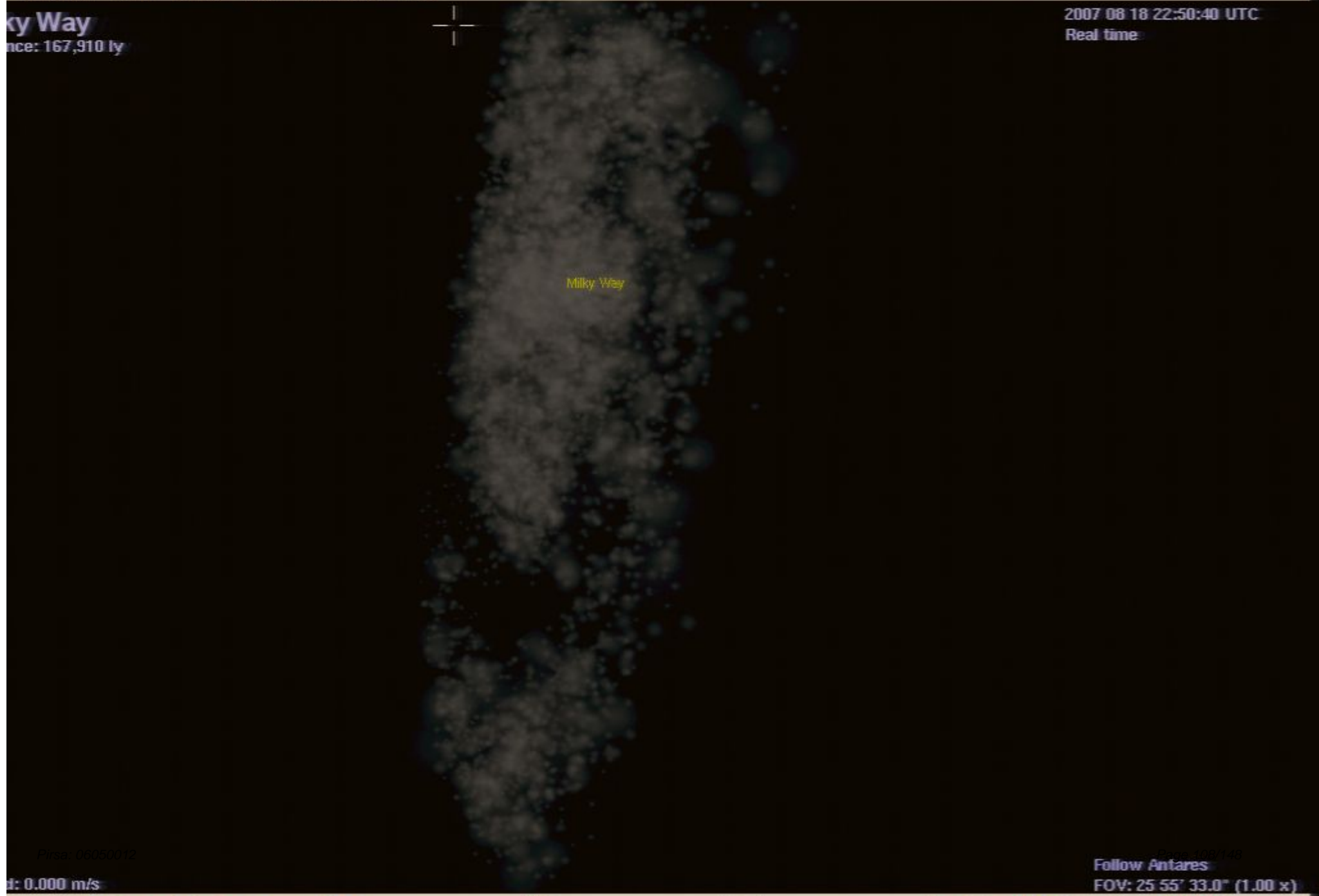
Zoom out and get the big picture . . .

Pirsa: 06050012

Speed: 0.000 m/s

Follow Antares

FOV: 25 55' 33.0" (1.00 x)



ky Way
nce: 167,910 ly

2007 08 18 22:50:40 UTC
Real time

Milky Way

Pisa: 06050012

0.000 m/s

Follow Antares
FOV: 25 55' 33.0" (1.00 x)

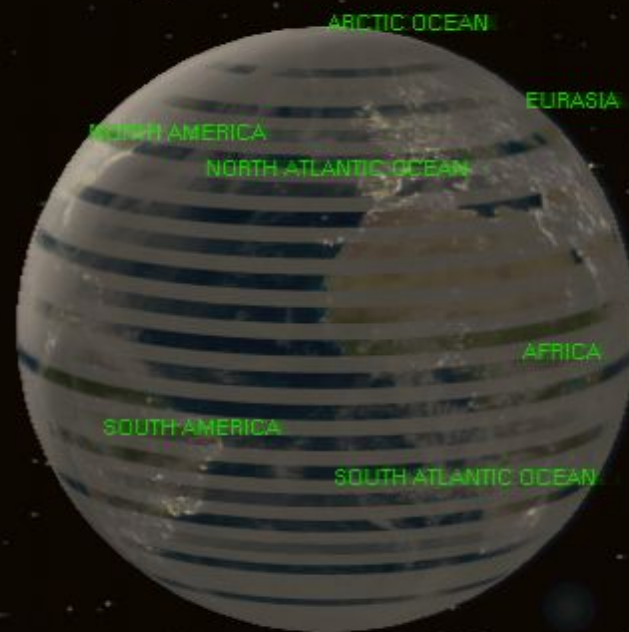
th

nce: 57,402 km

s: 6,378.1 km

2007 08 18 22:51:11 UTC

Real time



elestia

Navigation Time Render View Bookmarks Help

th

nce: 6,121,600 km

s: 6,378.1 km

2007 08 18 22:51:42 UTC

Real time

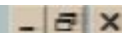
Pisa: 06050012

d: 2.855c

Follow Earth

FOV: 25 55' 33.0" (1.00 x)

elestia



Navigation Time Render View Bookmarks Help

th

nce: 214.47 ly

s: 6,378.1 km

2007.08.18 22:53:03 UTC

Real time

Pisa: 0000012

d: 41.809 ly/s

Page 11/148

Follow Earth

FOV: 25° 55' 33.0" (1.00 x)

Celestia



Navigation Time Render View Bookmarks Help

th

nce: 316.70 ly

s: 6,378.1 km

2007 08 18 22:53:05 UTC

Real time

Pisa: 06050012

d: 41.809 ly/s

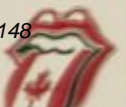
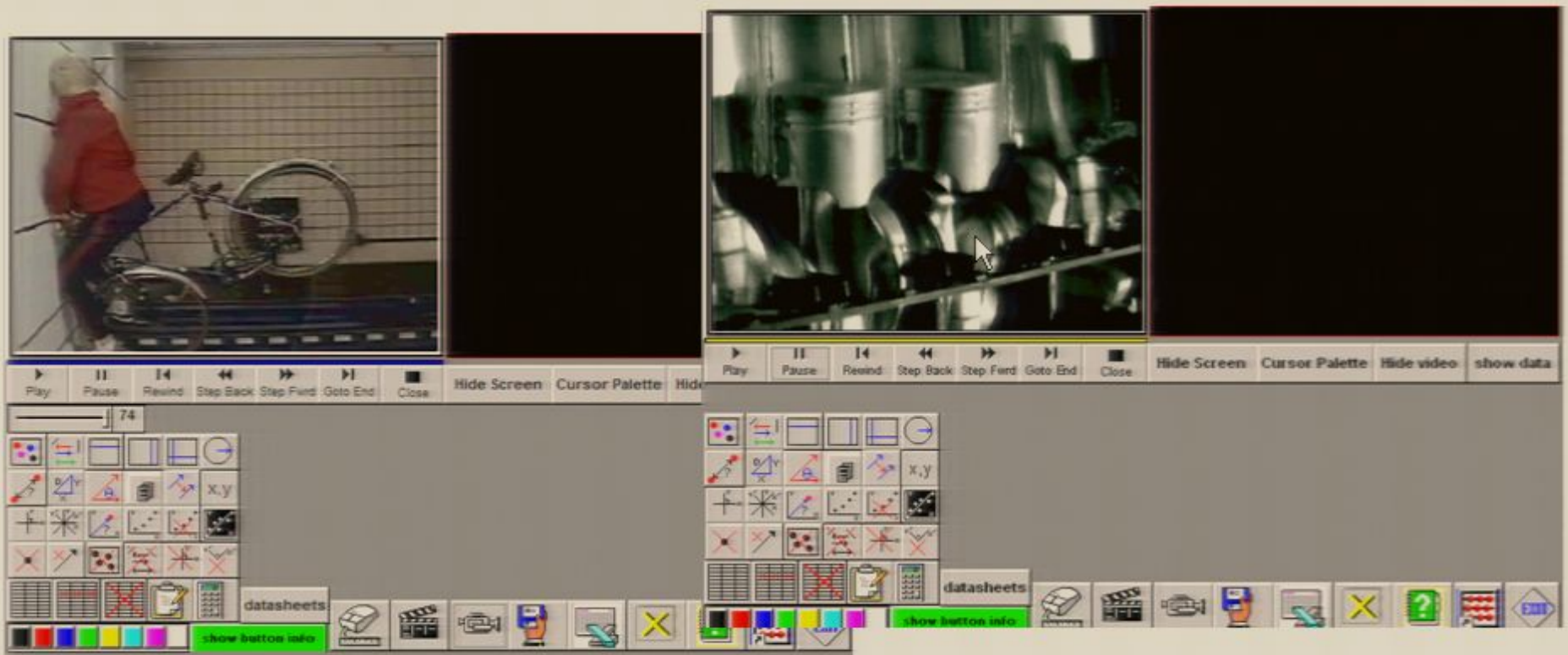
Page 112/119

Follow Earth

FOV: 25 55' 33.0" (1.00 x)

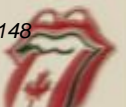
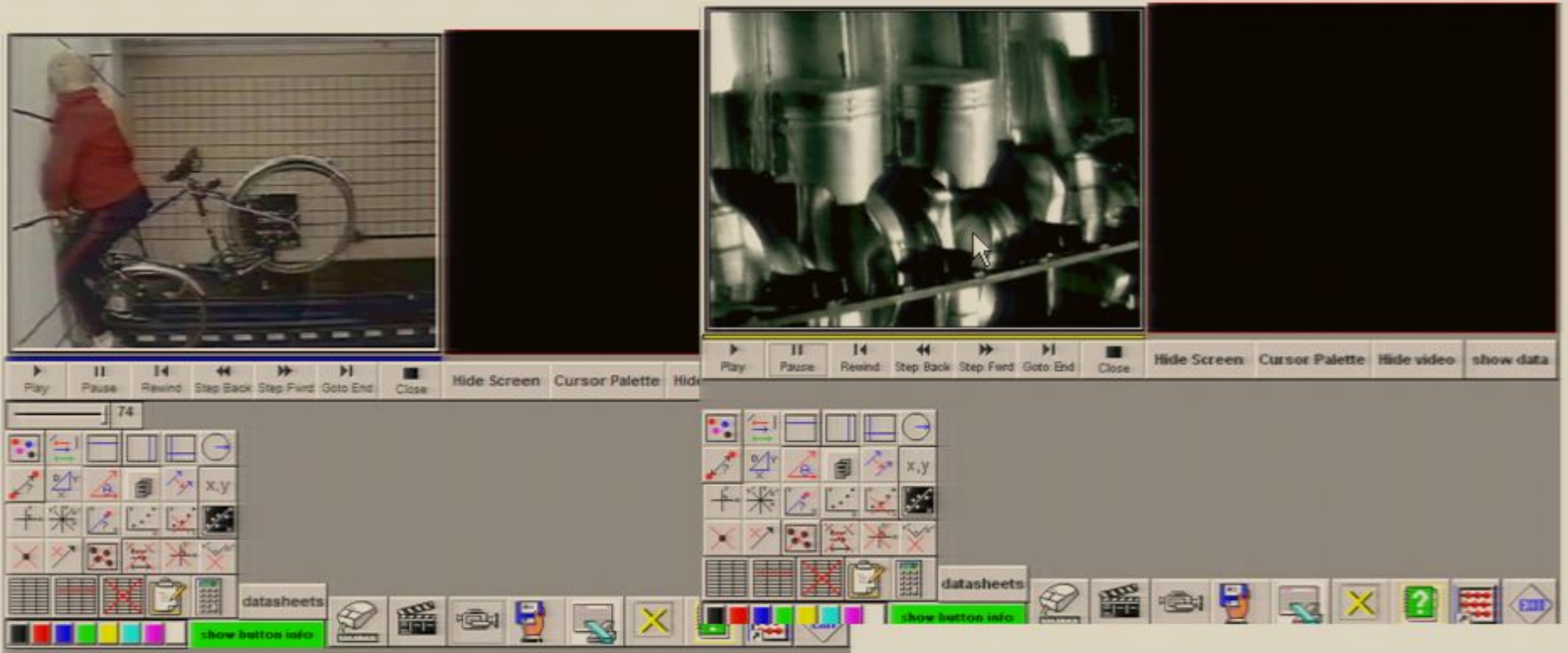
Vidshell

Motion analysis software that comes with video clips. It lets you measure and draw all sorts of things on the screen. You can also import your own video clips and export data to a spreadsheet. Written by Doyle V. Davis.

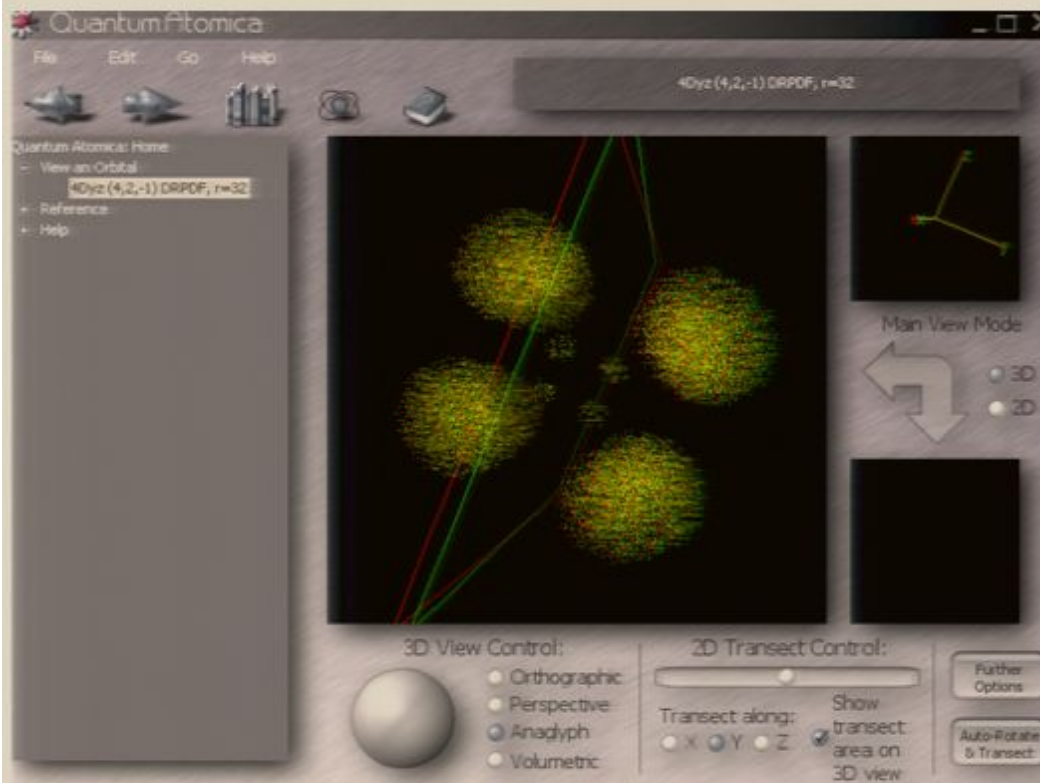


Vidshell

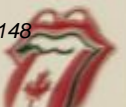
Motion analysis software that comes with video clips. It lets you measure and draw all sorts of things on the screen. You can also import your own video clips and export data to a spreadsheet. Written by Doyle V. Davis.



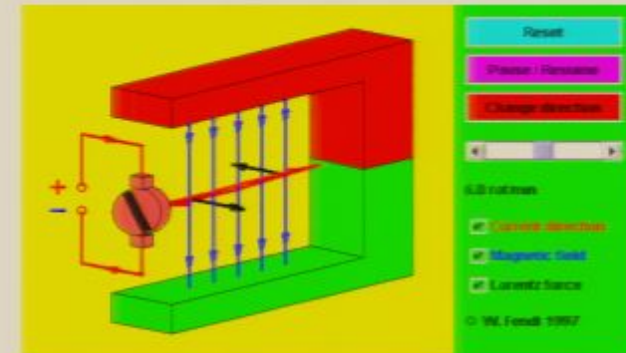
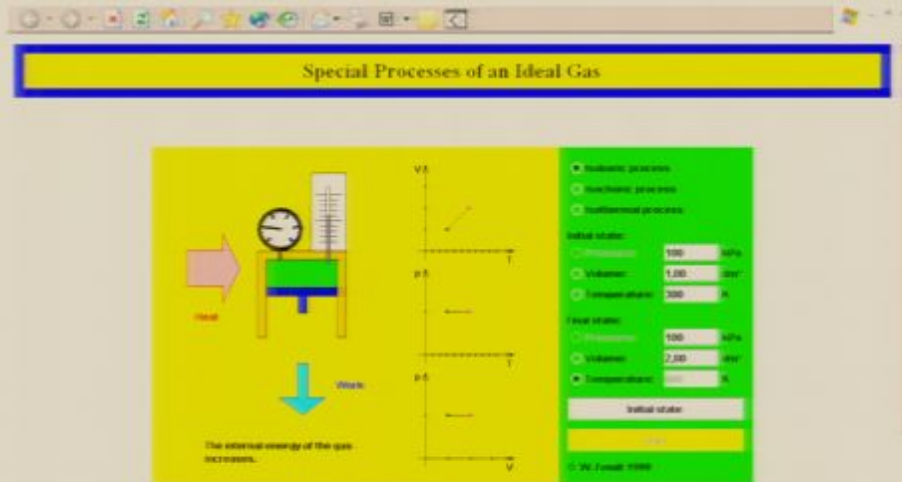
Quantum Atomica



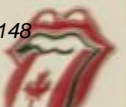
Shows the shapes of electron orbitals in 3D. Allows you to change the viewpoint and look at the orbitals with 3D glasses!



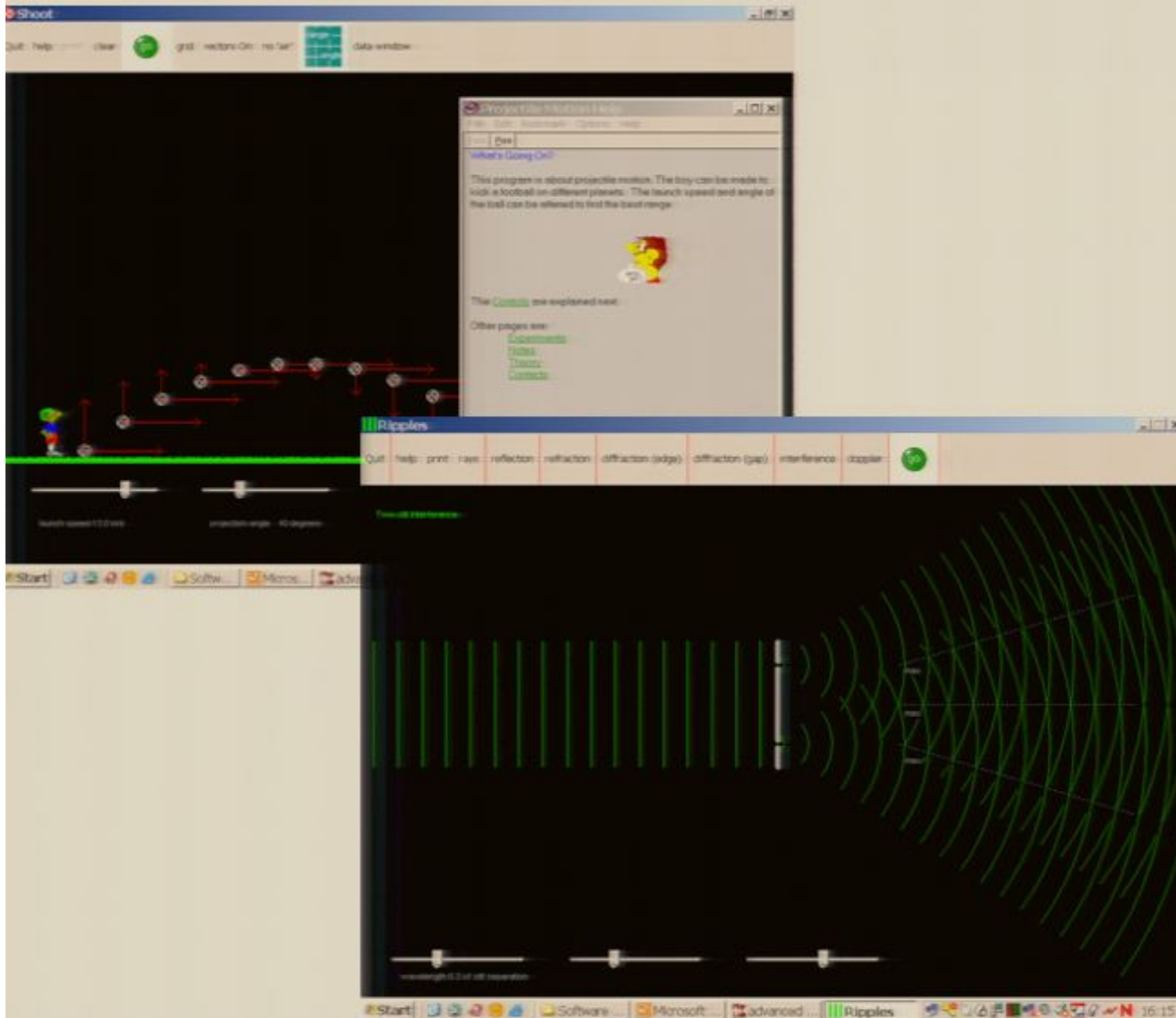
Walter Fendt



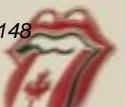
A whole load of applets that are ideal for displaying with a data projector. Highly rated.



Physics Lab

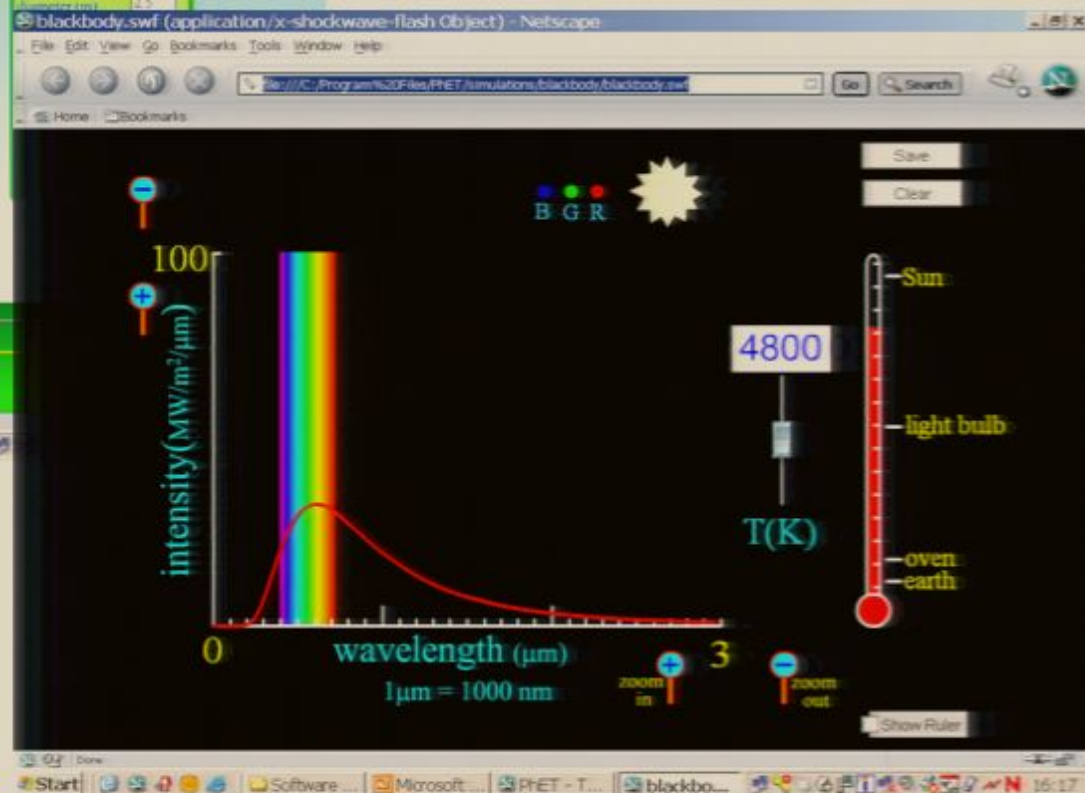
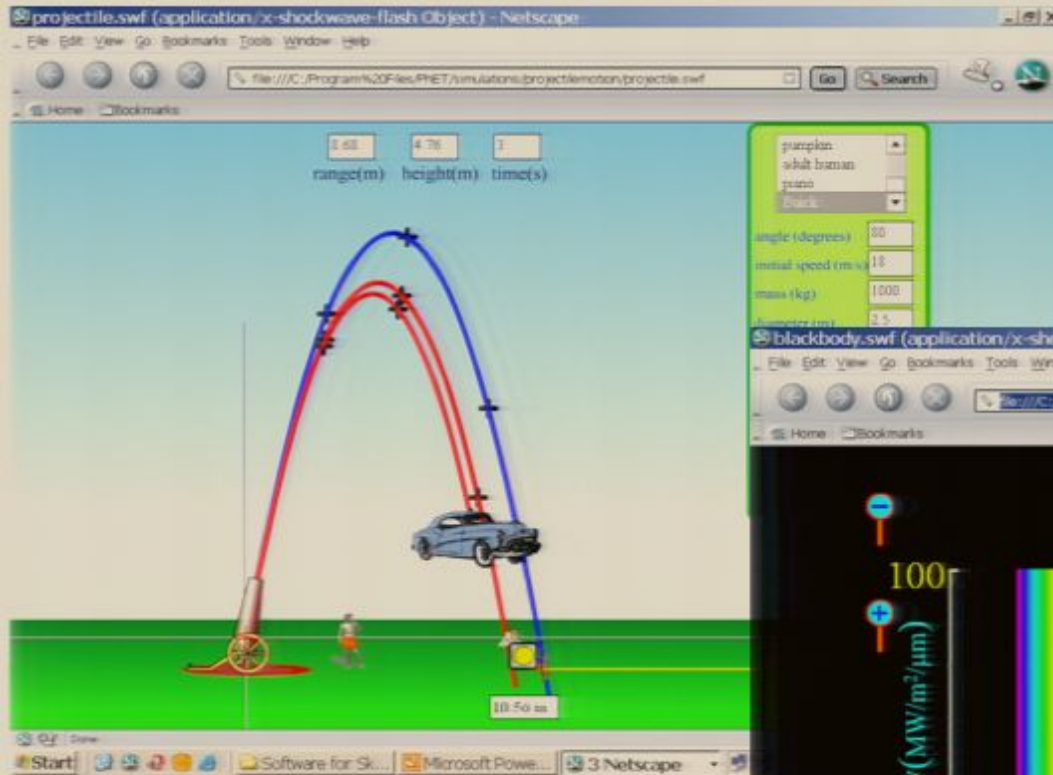


A set of
experiment
simulations
that pupils
can work
through -
clear and
entertaining.

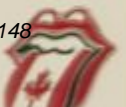


PhET

A set of applets and flash animations that cover a wide range of topics.

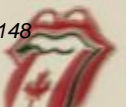
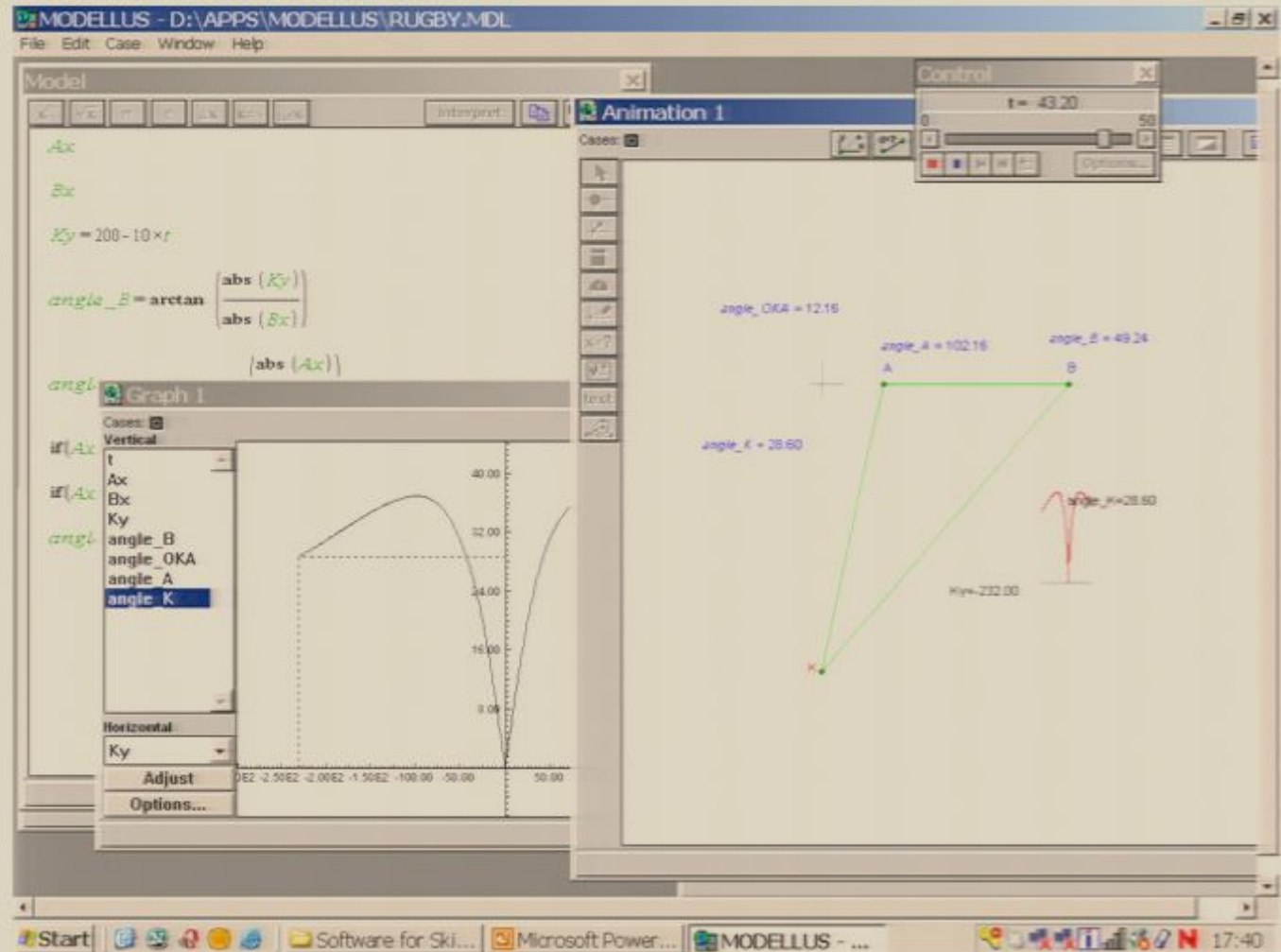


Can be
installed and
run offline.



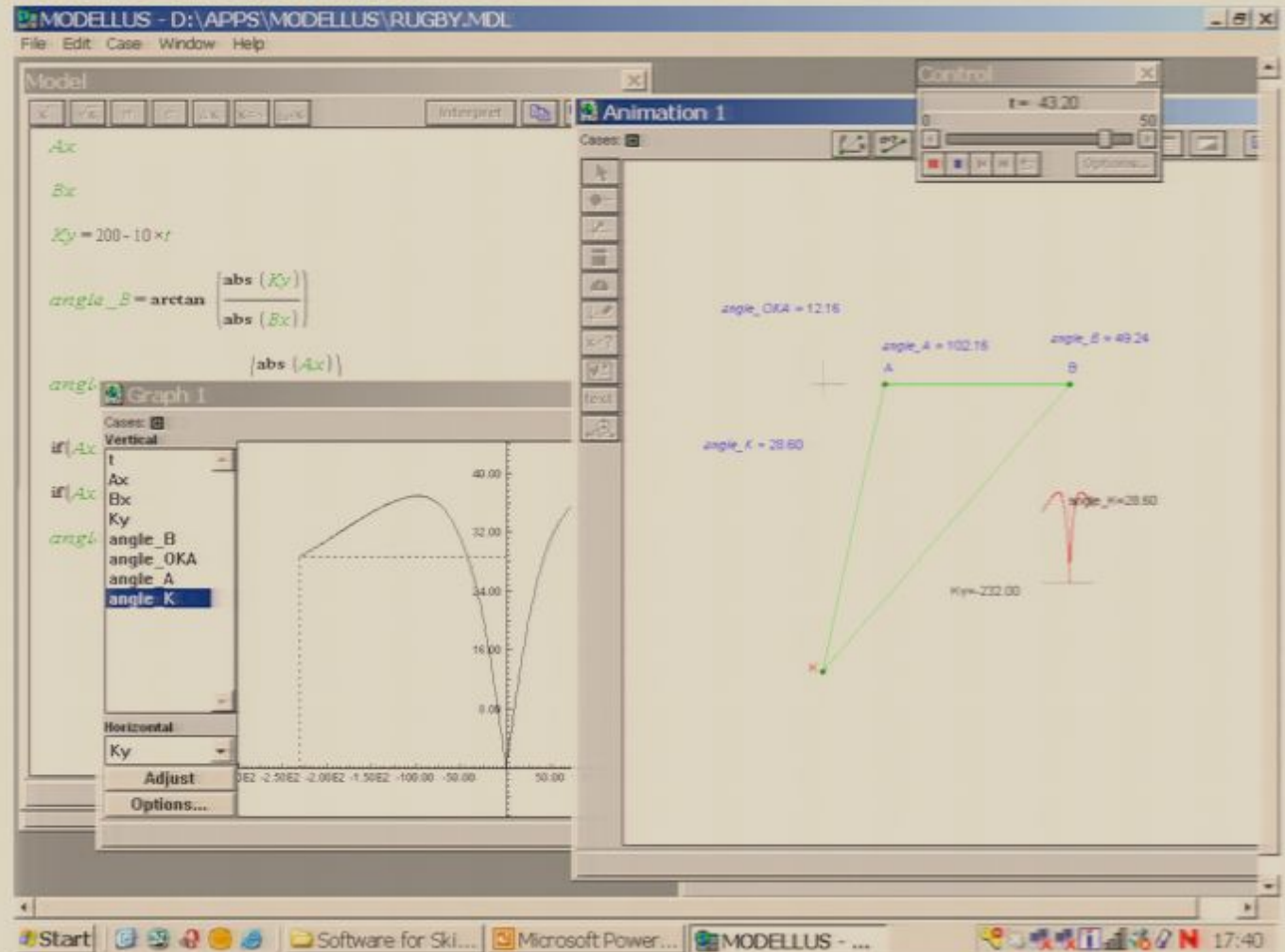
Modellus

Mathematical
modelling
software,
create
models and
draw graphs
and
animations.
Very
professional

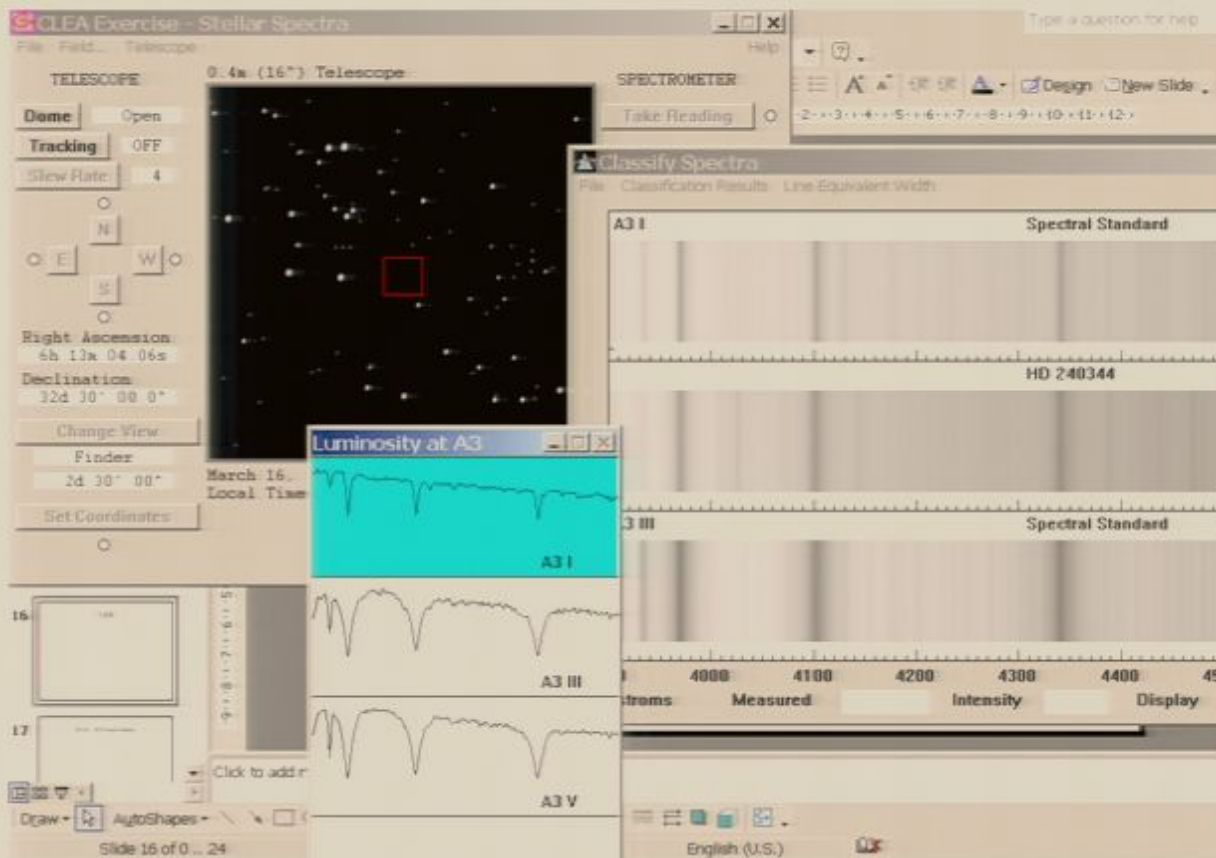


Modellus

Mathematical
modelling
software,
create
models and
draw graphs
and
animations.
Very
professional



CLEA



A number of very realistic astronomy simulations. More suited to older pupils.



Lancaster Particle Physics

A tutorial based
piece of
software that
explains the
theory and gives
some insight
into the
experimental
side of particle
physics

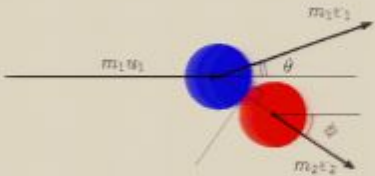
lancasteruniversity
particle physics package

home	pool	annihilation	magnetic field	lifetime	links
back	Introduction > Collisions > Masses > Calculation				next

A blue pool ball enters from the left hand side of the animation and collides with the stationary ball in the centre. Momentum and energy are transferred from the moving ball to the stationary ball. If the balls are considered to be hard spheres then the direction of momentum transfer is along the line joining the centres of the balls as they collide.

Collisions between hard spheres

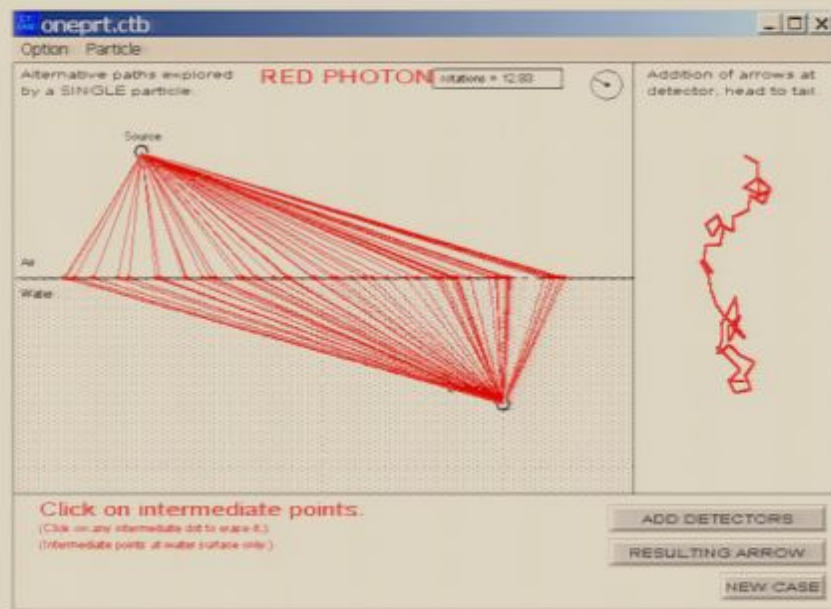
When two hard spheres collide they only touch at one point. This point lies on a line joining the centres of the spheres. Try this with two coins if you don't get it! If one ball is moving (blue) and hits a stationary ball (red) then the balls will exert equal size but opposite direction forces upon each other (Newton's third law). These forces act along the line joining the centres of the spheres. The blue ball will experience a force that makes it decelerate (Newton's first law) and the red ball will be accelerated by the equal but opposite force that it experiences.



Angle ϕ is set by the angle that the balls collide at.
Angle θ depends on ϕ and balls masses.



Many Paths

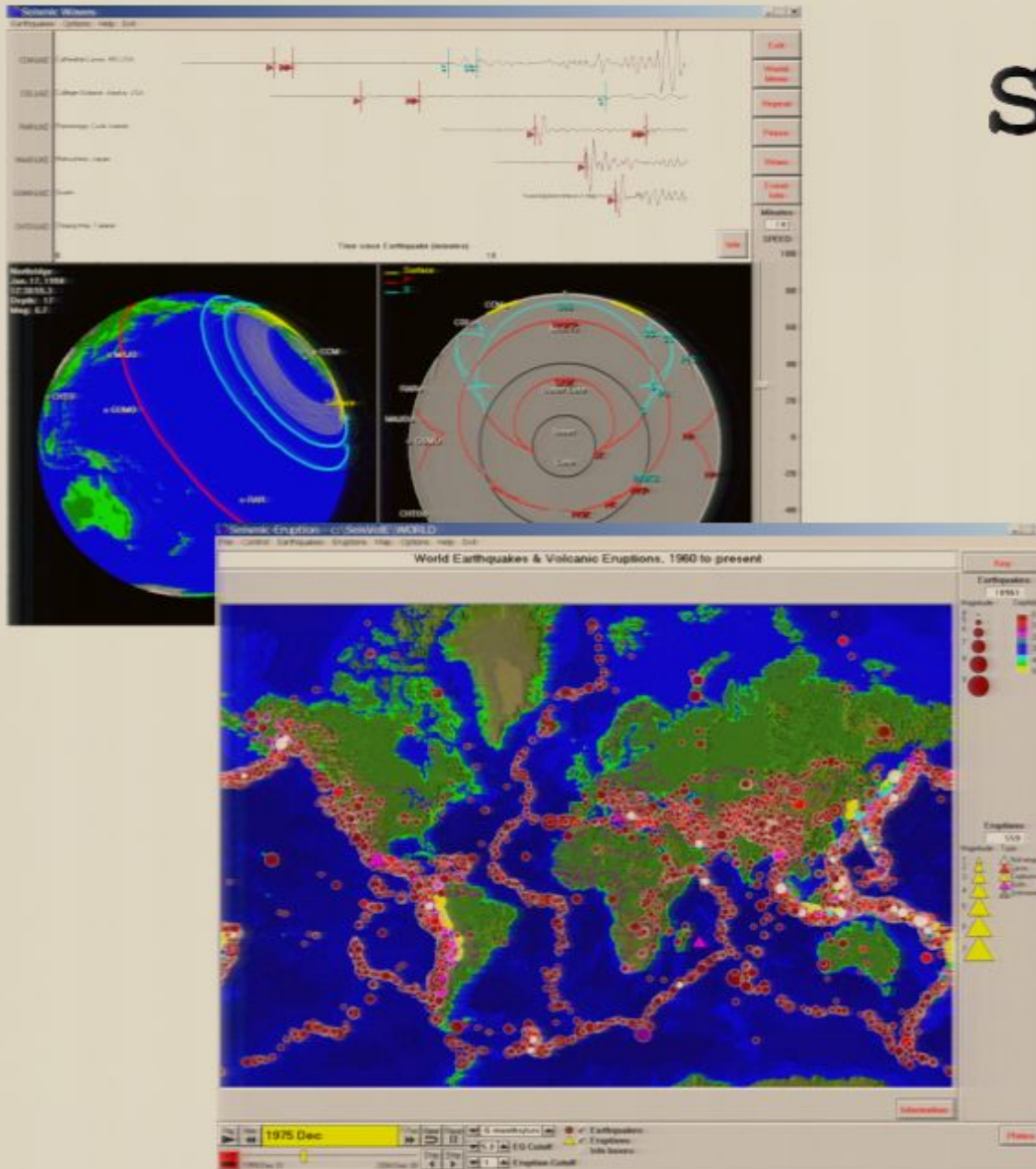


Draws all
the possible
paths of
photons when
teaching QED



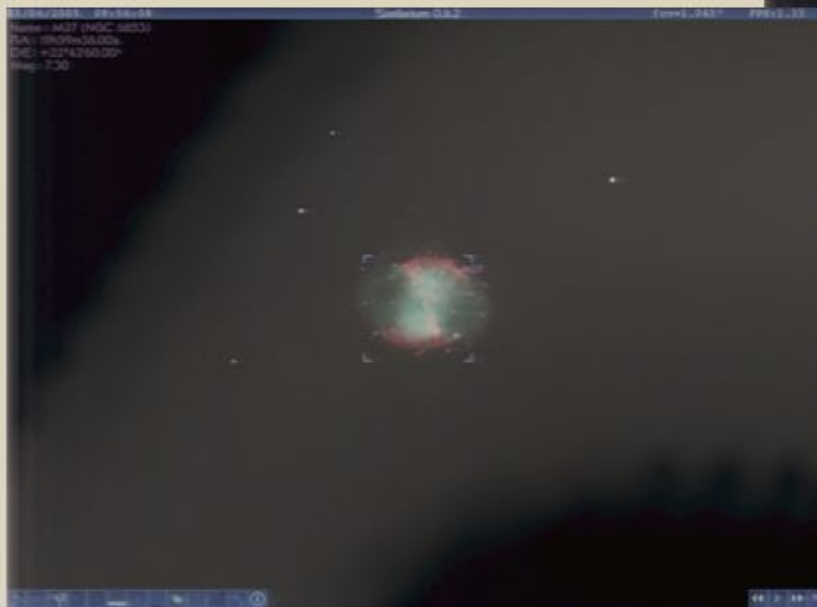
Seismology

Shows how P and S waves travel in the Earth and gives the locations of 1000's of earthquakes. Written by Alan L. Jones

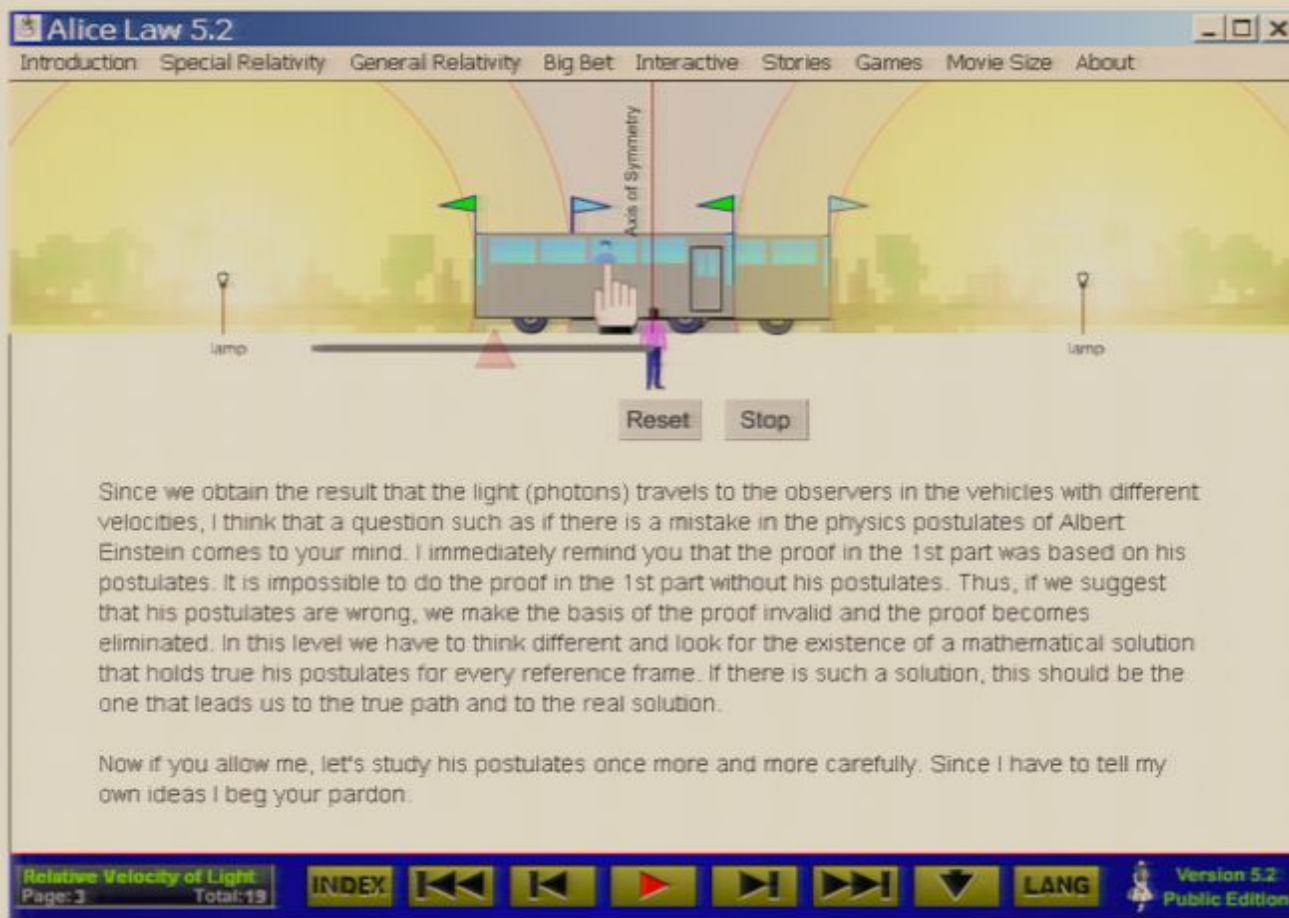


Stellarium

A planetarium
that allows you
to zoom in on
galaxies and
planets.



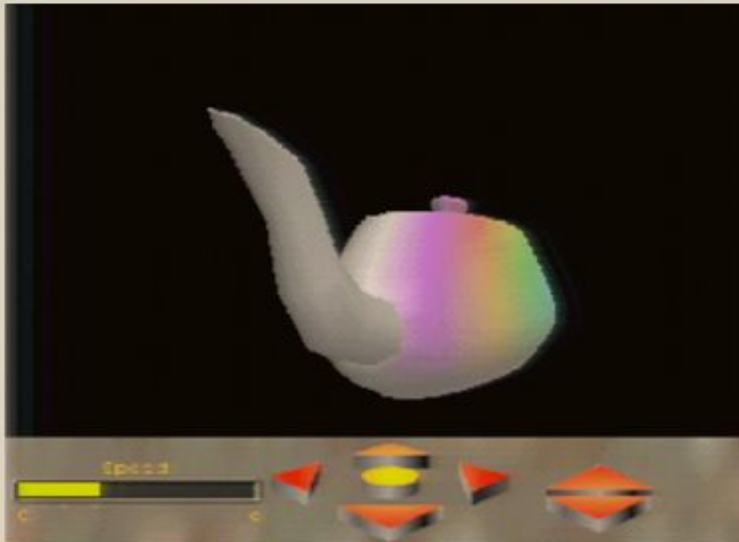
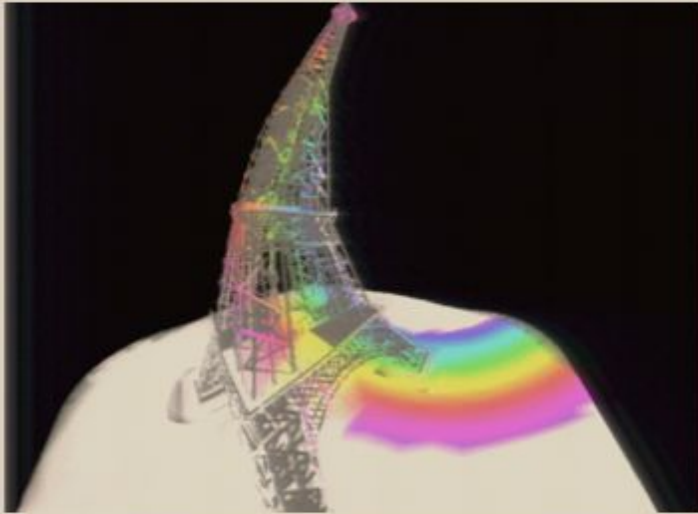
Alice Law



A tutorial that explains relativity using a variety of interactive animations.



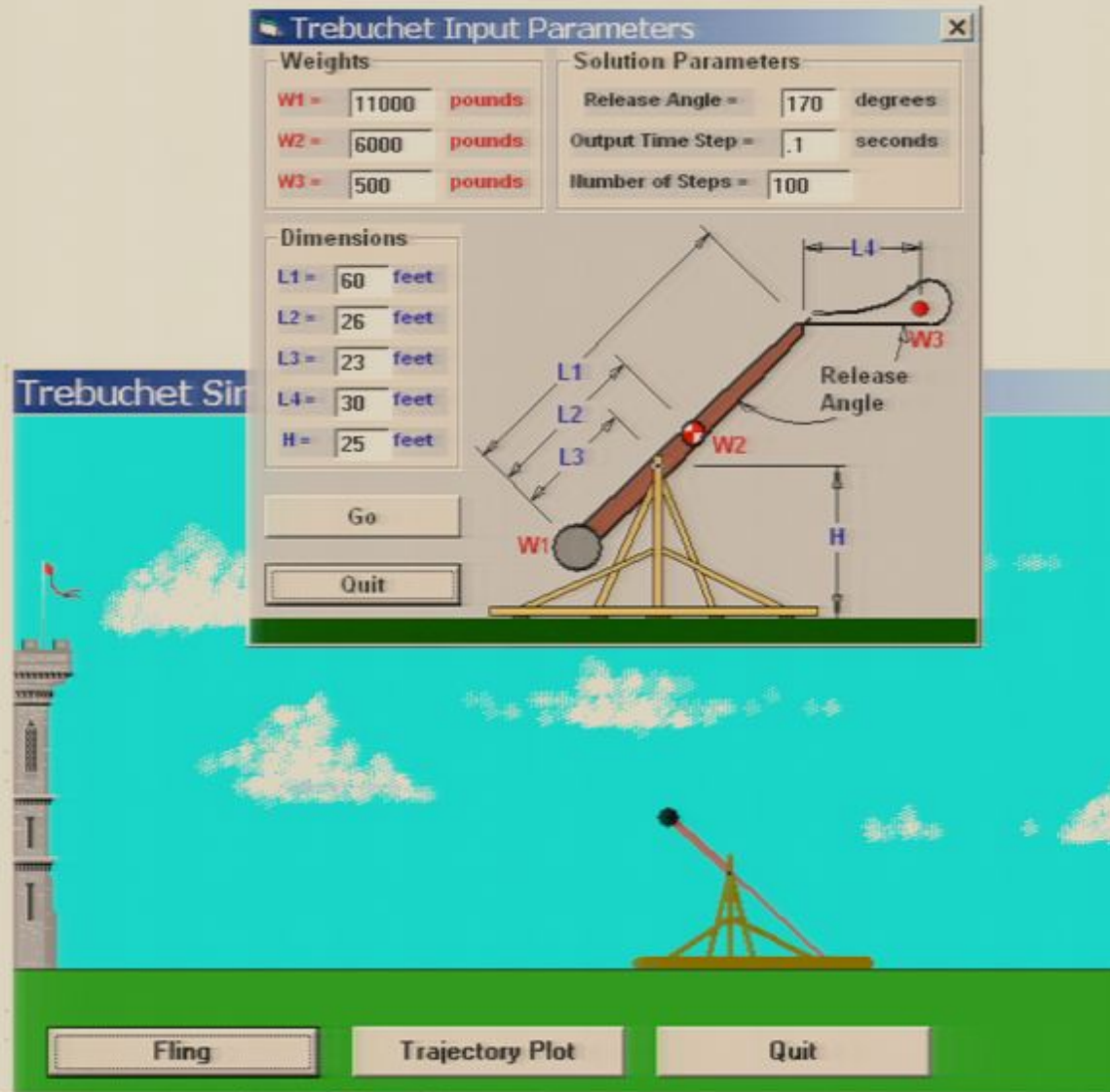
Warp



Simulates what
you might see
if you moved
past objects
while
travelling at
or near the
speed of light



WinTreb



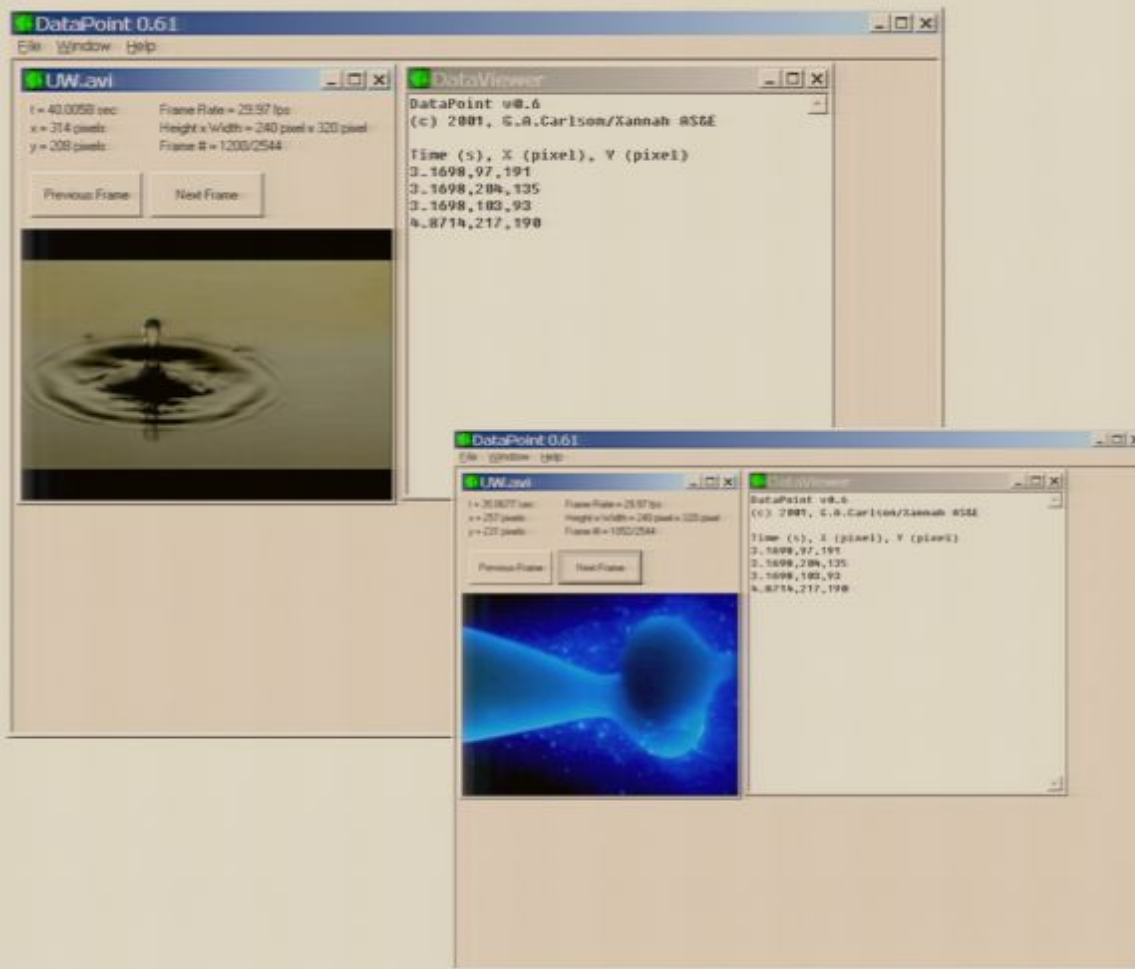
Change the dimensions of a trebuchet, fire it and measure the range and the maximum height of the projectile.

• **THERE IS NO SUPPORT AVAILABLE FOR THIS SIMULATION**

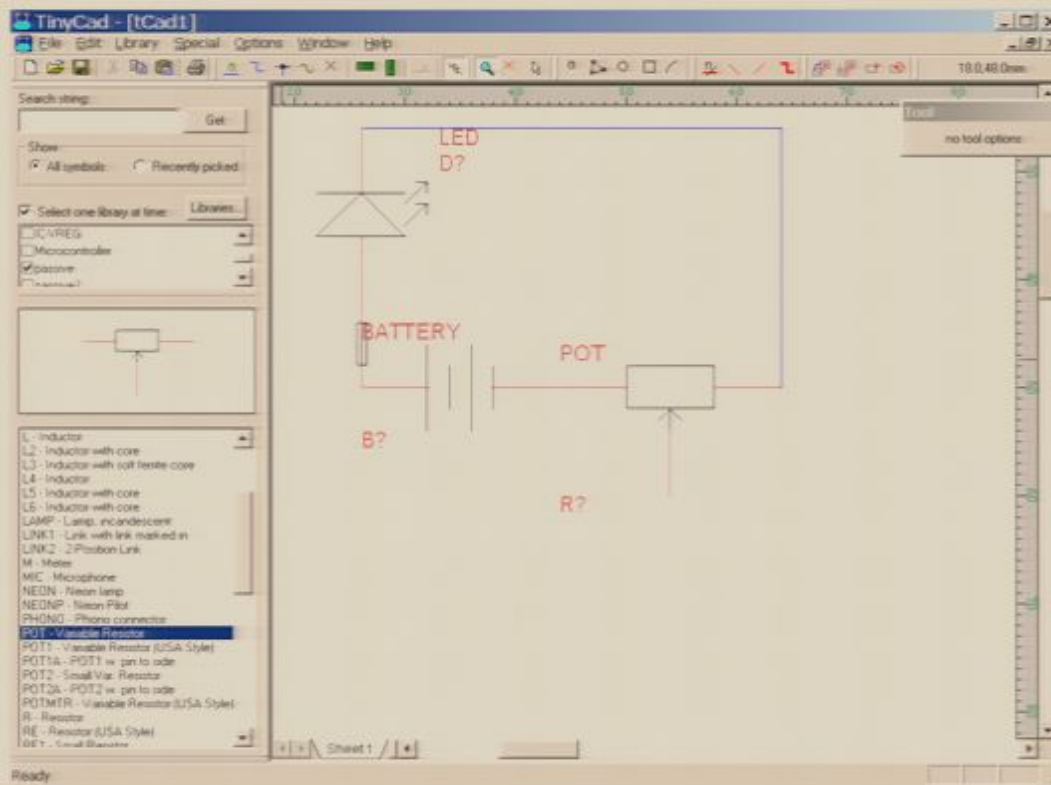


DataPoint

Allows you to move frame by frame through an avi file and collect data from the screen position.
Written by
Glenn A.
Carlson



TinyCAD

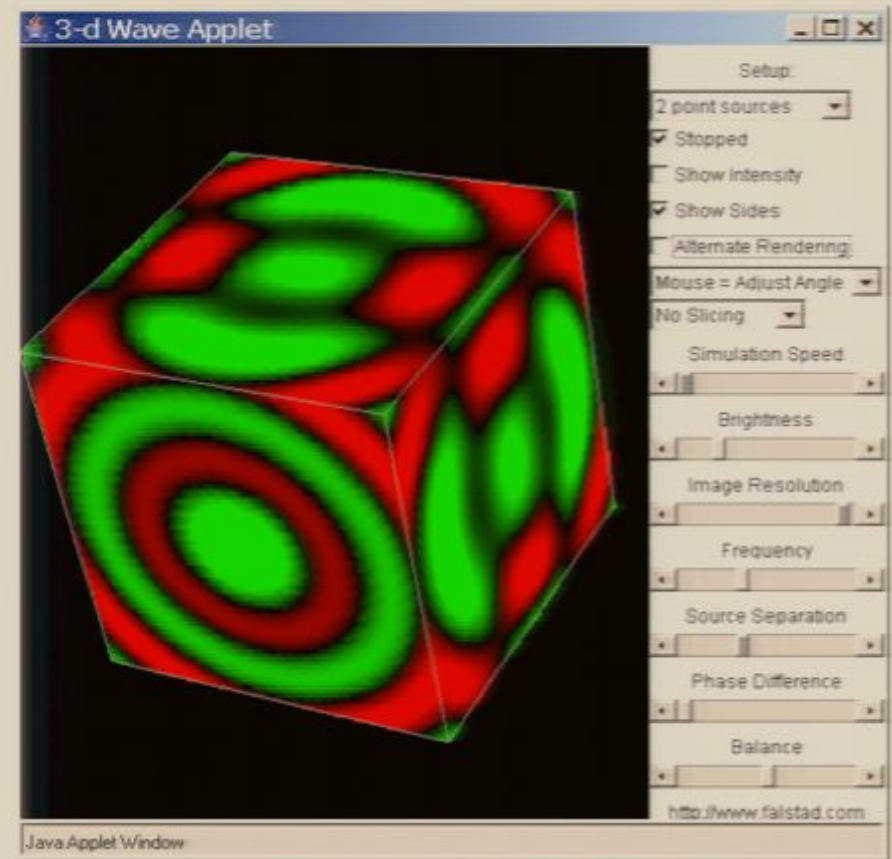


- Allows you to draw circuit diagrams, using a vast library, and to export them.



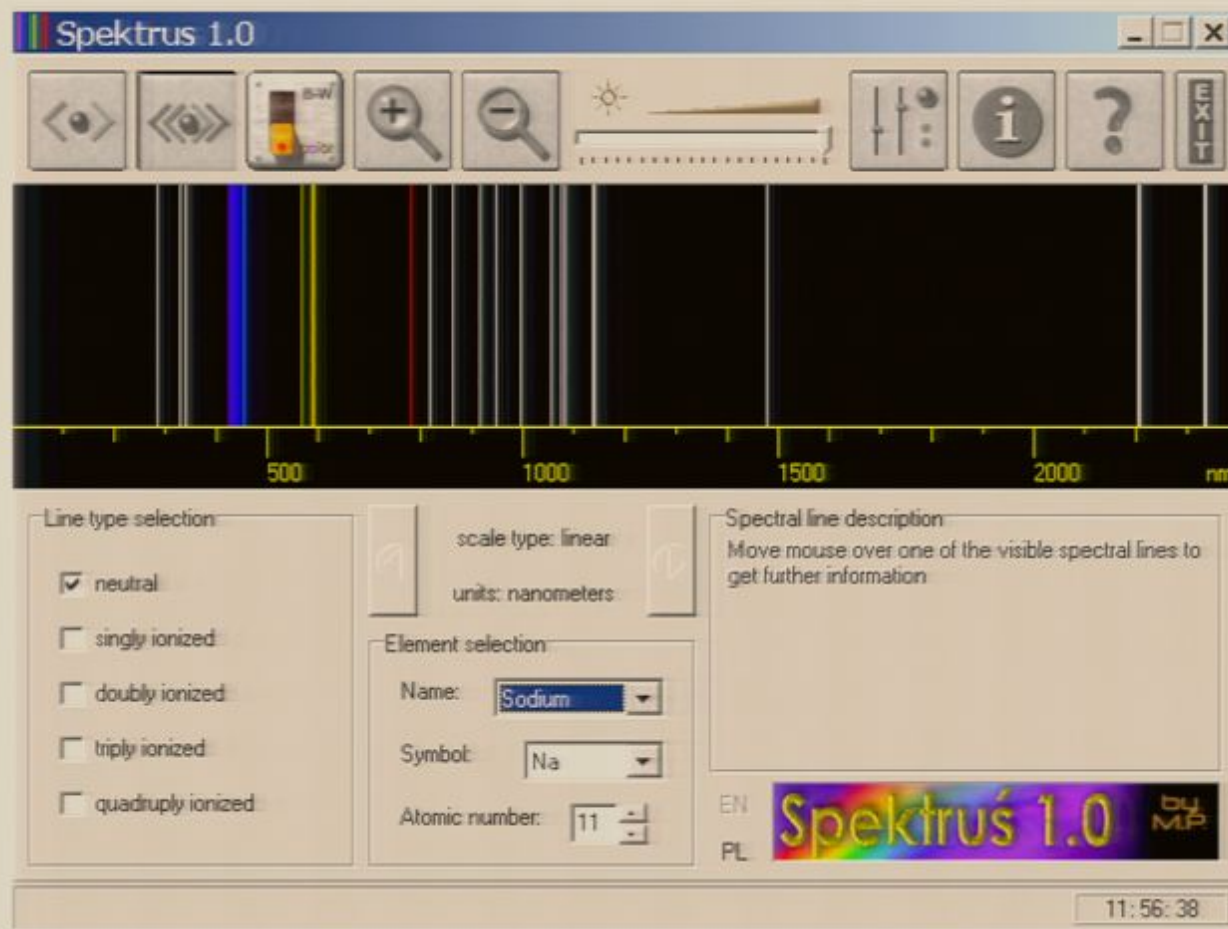
Falstad Applets

A set of applets that show a variety of things from acoustics to electrodynamics.



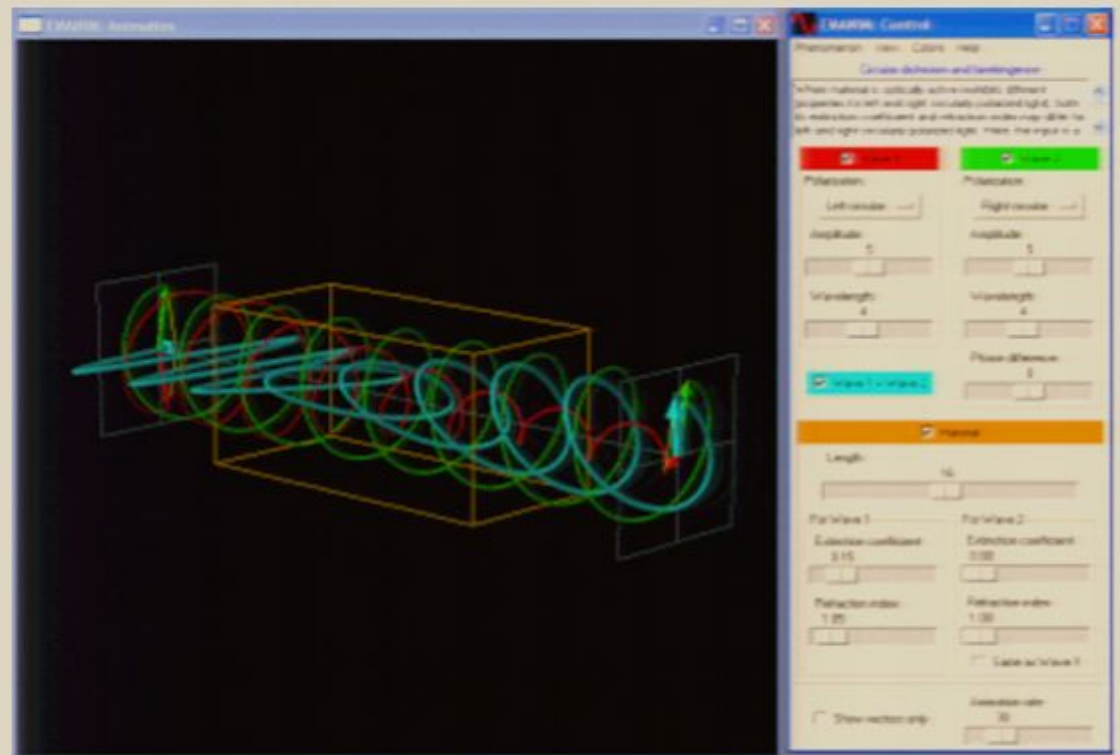
Spektrus

Shows the spectra of all of the elements and allows you to extract information about the emission lines.



Emanim

Visualises em
waves and
show
different
states of
polarisation,
interference,
refraction
etc



New Ideas?

A workshop from the
Institute of Physics
Teachers Network



Title:

Prism and OHP

Target:

Any

Equipment and suppliers:

All you need is an ohp, large prism and two sheets of card or books that can be used to make a slit on the top of the ohp. You'll also need a clamp, stand and boss to hold the prism in place, a G-clamp can be used with the right ohp.

If your ohp is sitting doing nothing then now is the time to press it back into service for making rainbows.

Prism

Slit



Title:

Bendy Pencils

Target:

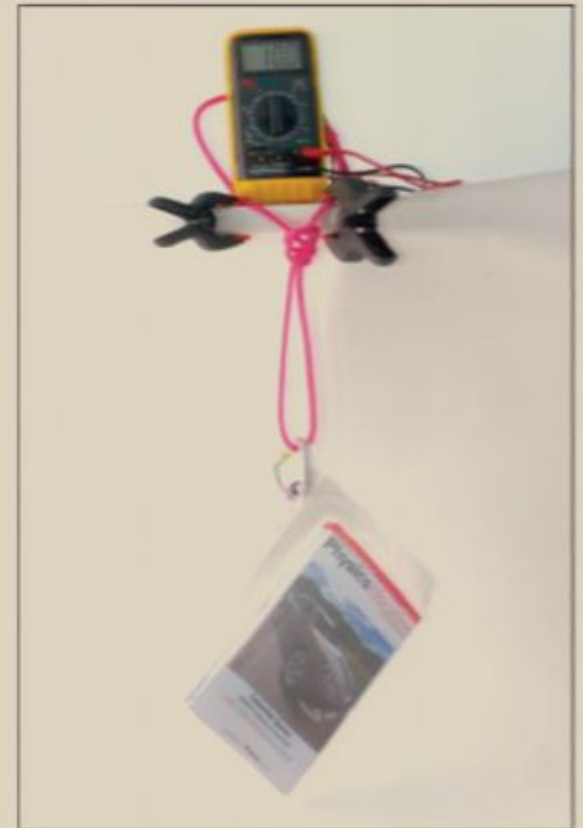
16 -18

Equipment and suppliers:

Try pound shops - the pencils come in packs of 8 - only 12.5p each. Longer pencils have high resistance and need a good ohmmeter.



Bendy pencils can be used to make a strain gauge. The pencils come in 2 lengths, 30cm and 200cm.



Centrally resourced sessions



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Prism

Slit



Son of New Ideas?

A workshop from the
Institute of Physics
Teachers Network



Title:

Wave changer

Target:

Any

Equipment and suppliers:

Marker pen,
stick,
wallpaper
or A3 sheets
taped
together,
old wire reel,
clamp stand



Sound is a longitudinal wave - so why do we show images of transverse waves and say they're sound?

Show pupils how longitudinal waves are represented by transverse waves with this simple gadget.



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Wave changer

Target:

Any

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Title:

Eggs

Target:

14-18

Equipment and suppliers:



Eggs supplied by chickens!
Tray from local supermarket - 79p
Using crème eggs is less messy

A variation on the bucket of water trick. Use a tray for boiling eggs and a length of rope. Make sure the folding handle is securely in place. Whirl round in circular fashion without breakages - but stopping is the difficult bit!



Title:

Butthead

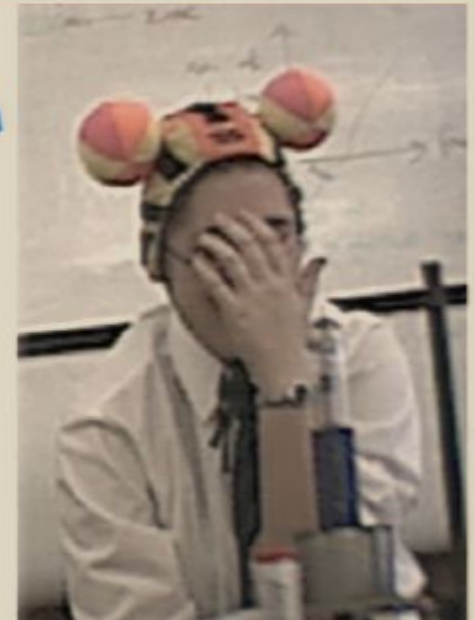
Target:

16-18

Equipment and suppliers:

Butt head game (£9.99 from Hawkin.com) or homemade Velcro equivalent, foam disc gun (£9.99 & £4.99) and an airzooka.

This game can be used to explain the photoelectric effect. The work function is the amount of energy needed to remove the ball from the head; the rest of the energy is transformed into KE.



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Thank You!

