

Title: The chaotic evolution of Newton's universe

Date: Mar 09, 2006 04:00 PM

URL: <http://pirsa.org/06030010>

Abstract: In this expository talk, I describe how "chaotic behavior" not only was discovered in the study of the Newtonian N-body problem, but also is responsible for several strange appearing motions. Then, a mathematical outline of the general evolution of the universe, under Newton's laws, is provided. No prior background in dynamics or the mathematics of the N-body problem is needed to follow this lecture

# **Chaotic** Evolution of the Universe

Don Saari  
Institute for Mathematical Behavioral Sciences  
University of California, Irvine  
[dsaari@uci.edu](mailto:dsaari@uci.edu)

# **Chaotic** Evolution of the Universe

Don Saari  
Institute for Mathematical Behavioral Sciences  
University of California, Irvine  
[dsaari@uci.edu](mailto:dsaari@uci.edu)

# ***Chaotic* Evolution of the Universe**

## Newtonian N-body problem

Don Saari

Institute for Mathematical Behavioral Sciences

University of California, Irvine

[dsaari@uci.edu](mailto:dsaari@uci.edu)

# ***Chaotic* Evolution of the Universe**

## Newtonian N-body problem

Don Saari

Institute for Mathematical Behavioral Sciences

University of California, Irvine

[dsaari@uci.edu](mailto:dsaari@uci.edu)

Shepherds, ancient astronomers and astrologers

# **Chaotic** Evolution of the Universe

## Newtonian N-body problem

Don Saari

Institute for Mathematical Behavioral Sciences

University of California, Irvine

[dsaari@uci.edu](mailto:dsaari@uci.edu)

Shepherds, ancient astronomers and astrologers

World's **Oldest** profession!

# **Chaotic Evolution of the Universe**

## Newtonian N-body problem

Don Saari

Institute for Mathematical Behavioral Sciences

University of California, Irvine

[dsaari@uci.edu](mailto:dsaari@uci.edu)

Shepherds, ancient astronomers and astrologers

World's **Second** **Oldest** profession!

# **Chaotic Evolution of the Universe**

## Newtonian N-body problem

Don Saari

Institute for Mathematical Behavioral Sciences

University of California, Irvine

[dsaari@uci.edu](mailto:dsaari@uci.edu)

Shepherds, ancient astronomers and astrologers

World's **Second Oldest** profession!

To introduce a research problem

# **Chaotic Evolution of the Universe**

## Newtonian N-body problem

Don Saari

Institute for Mathematical Behavioral Sciences

University of California, Irvine

[dsaari@uci.edu](mailto:dsaari@uci.edu)

Shepherds, ancient astronomers and astrologers

World's **Second Oldest** profession!

To introduce a research problem

**Does the Earth go about the Sun, or does  
the Sun go about the Earth?**



**on of the Universe**

**body problem**

n Saari

atrical Behavioral Sciences

California, Irvine

@uci.edu

onomers and astrologers

nd

est

est profession!

research problem

**Does the Earth go about the Sun, or does  
the Sun go about the Earth?**

# Geocentric vs. heliocentric approach

# Geocentric vs. heliocentric approach



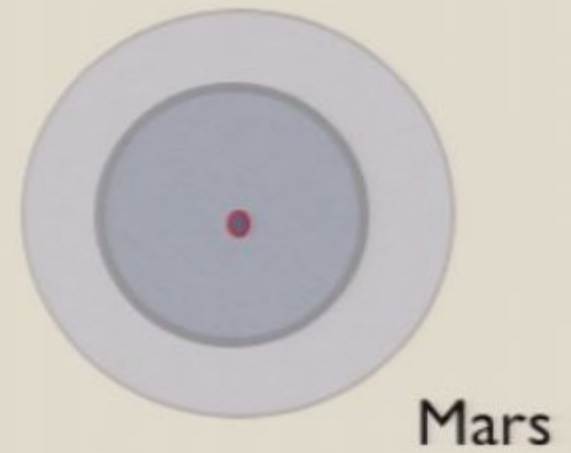
## Geocentric vs. heliocentric approach

One year, distance AU



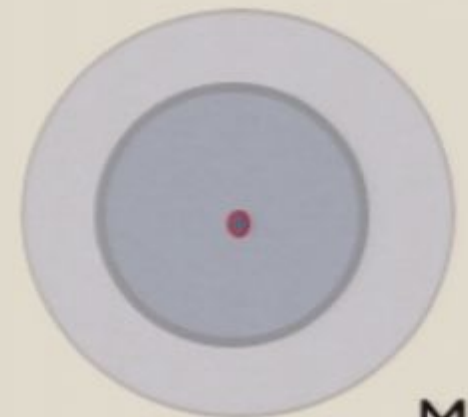
## Geocentric vs. heliocentric approach

One year, distance AU



# Geocentric vs. heliocentric approach

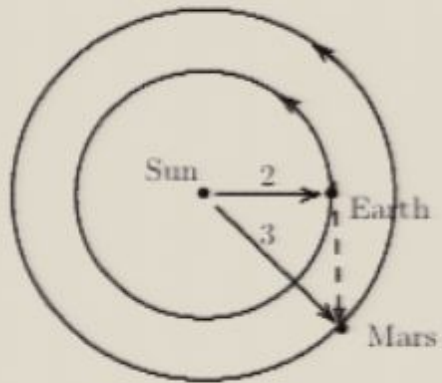
One year, distance AU



Mars

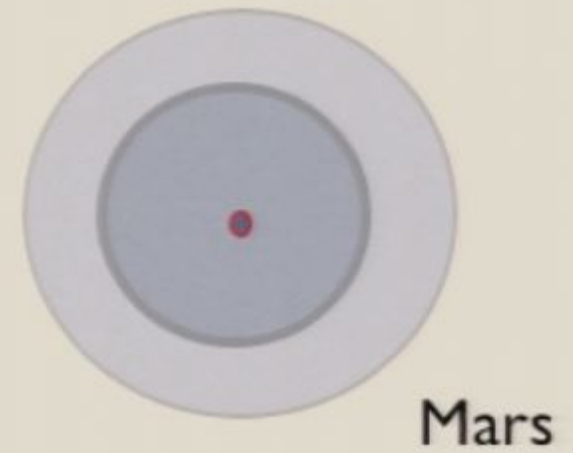
Two years, 1.5 AU

## Geocentric vs. heliocentric approach



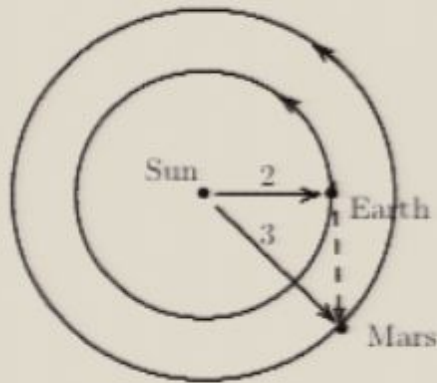
The Sun-Earth-Mars system in half-astronomical units

One year, distance AU



Two years, 1.5 AU

## Geocentric vs. heliocentric approach

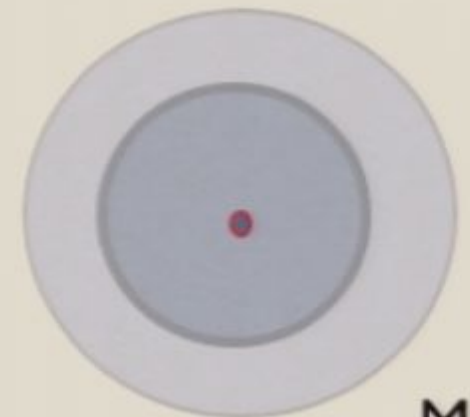


$$\mathbf{r}_E = 2e^{2\pi it}$$

$$\mathbf{r}_M = 3e^{\pi it}$$

The Sun-Earth-Mars system in half-astronomical units

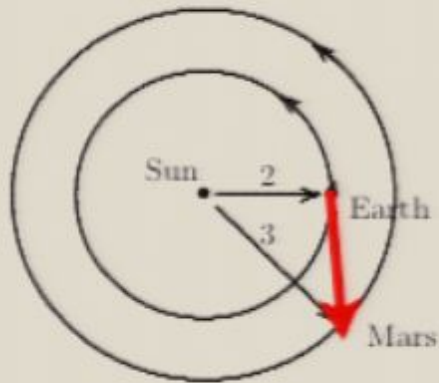
One year, distance AU



Mars

Two years, 1.5 AU

# Geocentric vs. heliocentric approach

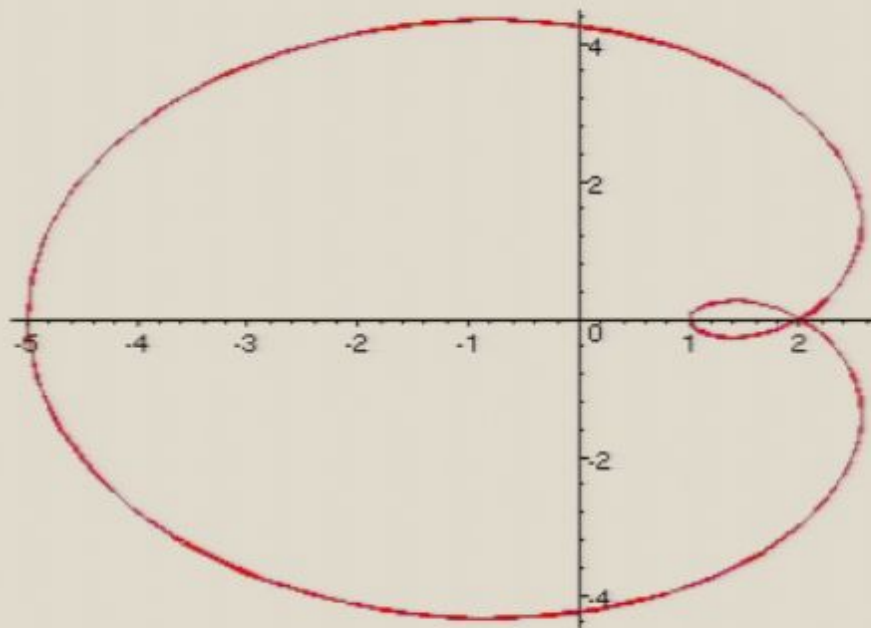


$$\mathbf{r}_E = 2e^{2\pi it}$$

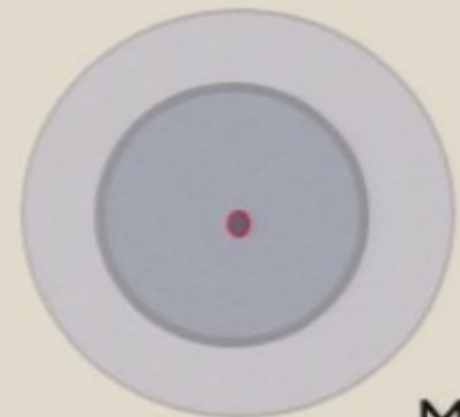
$$\mathbf{r}_M = 3e^{\pi it}$$

$$\mathbf{r}_M - \mathbf{r}_E = 2 + [3 - 4\cos(\pi t)]e^{\pi it}$$

The Sun-Earth-Mars system in half-astronomical units



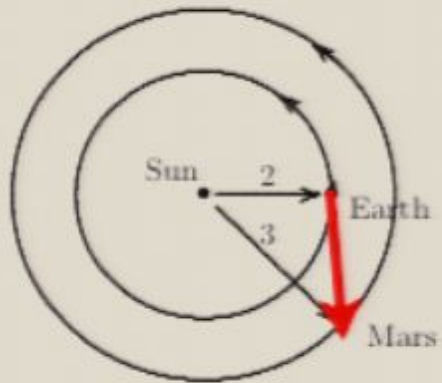
One year, distance AU



Mars

Two years, 1.5 AU

# Geocentric vs. heliocentric approach

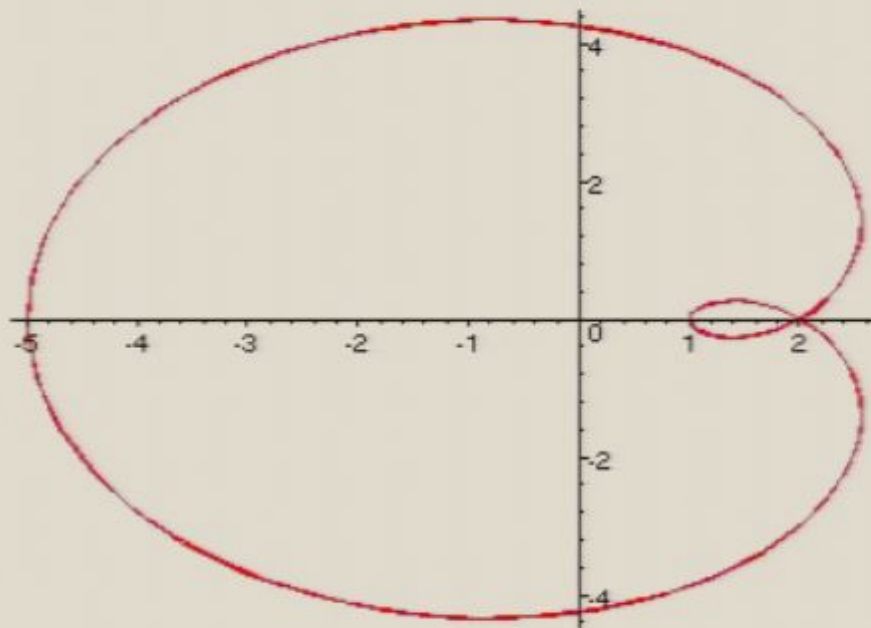


$$\mathbf{r}_E = 2e^{2\pi it}$$

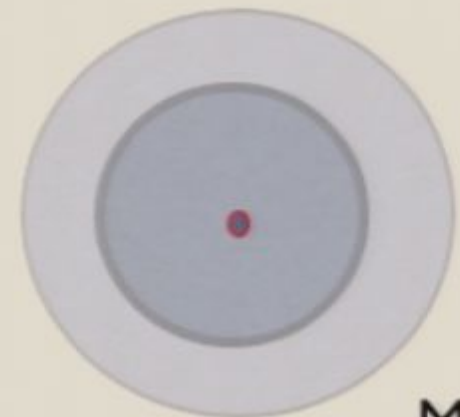
$$\mathbf{r}_M = 3e^{\pi it}$$

$$\mathbf{r}_M - \mathbf{r}_E = 2 + [3 - 4\cos(\pi t)]e^{\pi it}$$

The Sun-Earth-Mars system in half-astronomical units



One year, distance AU



Mars

Two years, 1.5 AU

# Ptolomy and his epicycles

# Ptolomy and his epicycles

Aristotle:

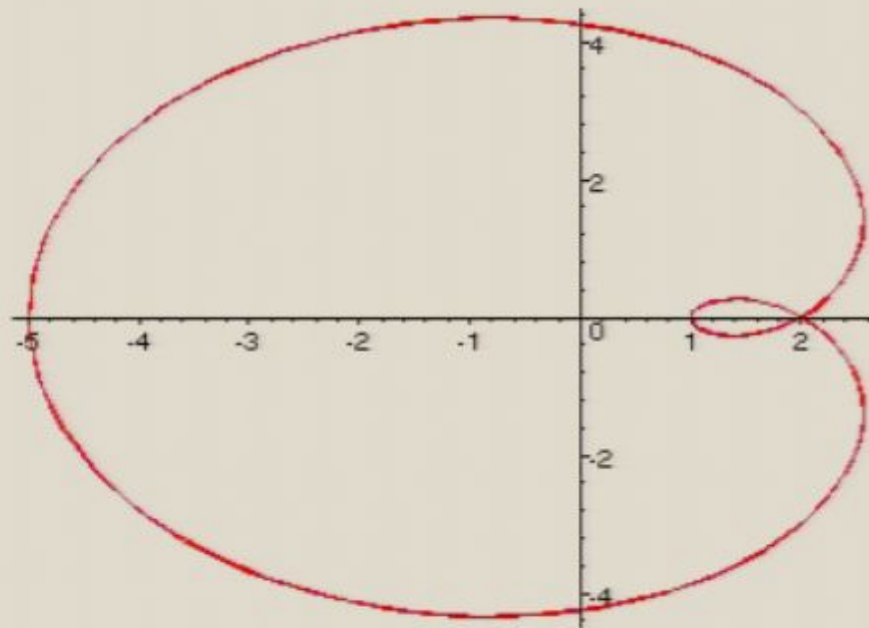
## Ptolomy and his epicycles

Aristotle: Uniform motion

# Ptolomy and his epicycles

Aristotle: Uniform motion

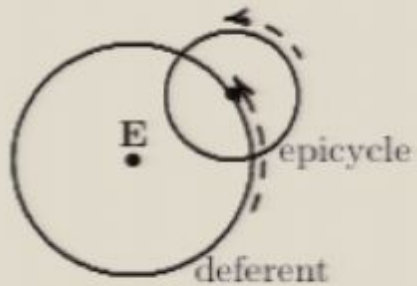
Circular



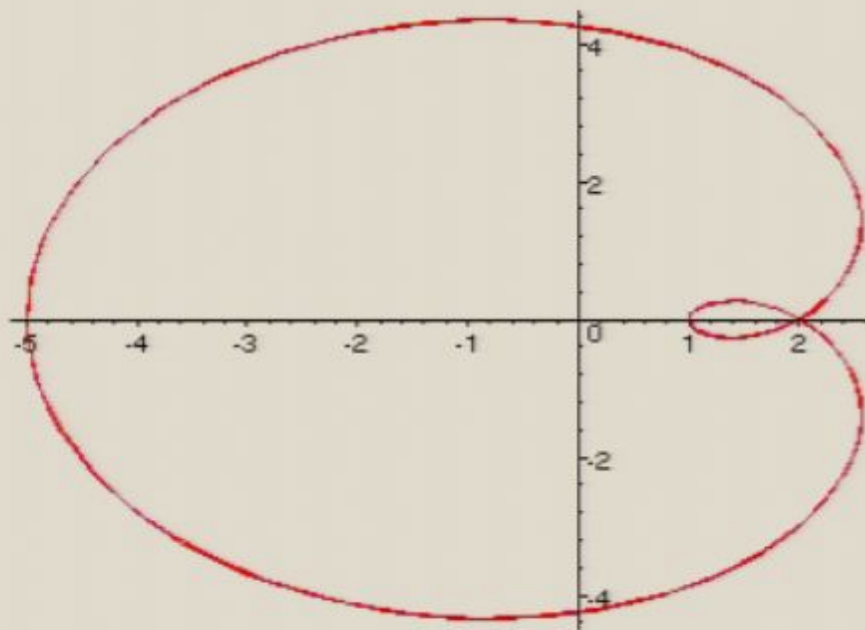
# Ptolomy and his epicycles

Aristotle: Uniform motion

Circular



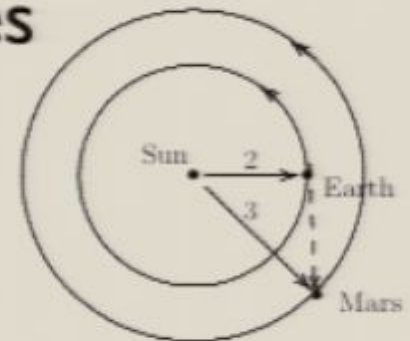
Epicycles



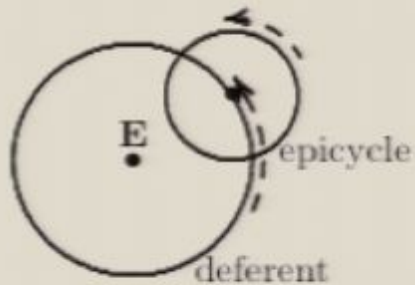
# Ptolomy and his epicycles

Aristotle: Uniform motion

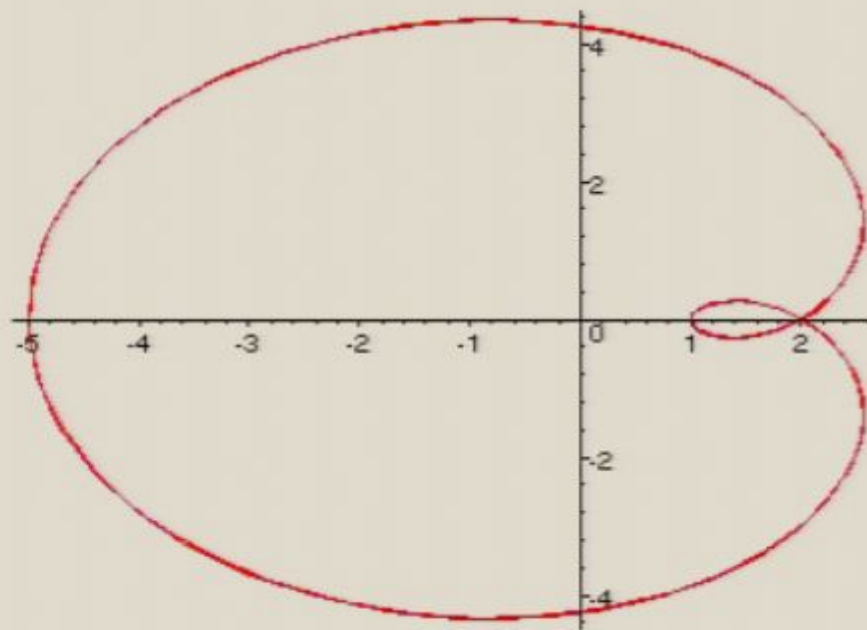
Circular



The Sun-Earth-Mars system in half-astronomical u



Epicycles



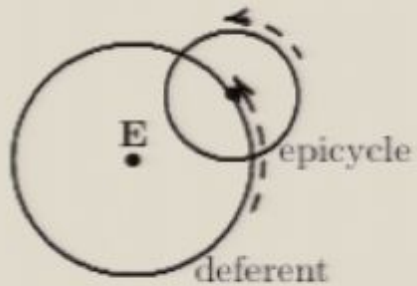
# Ptolomy and his epicycles

Aristotle: Uniform motion

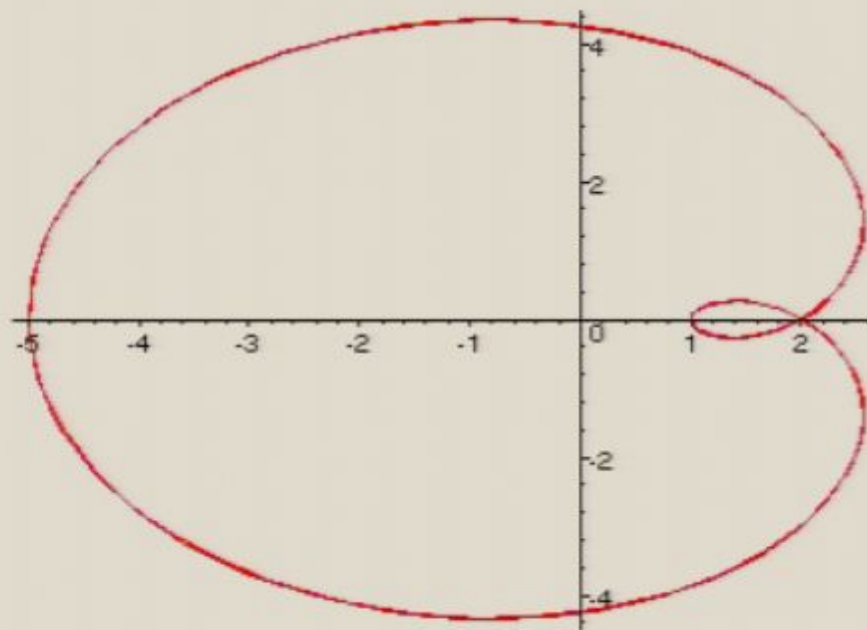
Circular



The Sun-Earth-Mars system in half-astronomical u



Epicycles



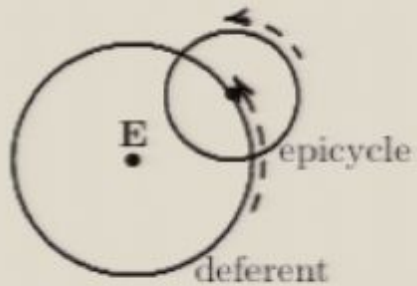
# Ptolomy and his epicycles

Aristotle: Uniform motion

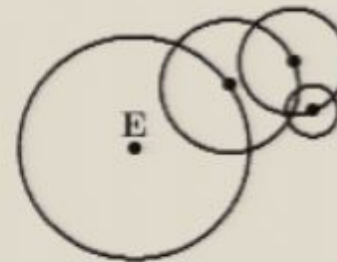
Circular



The Sun-Earth-Mars system in half-astronomical u



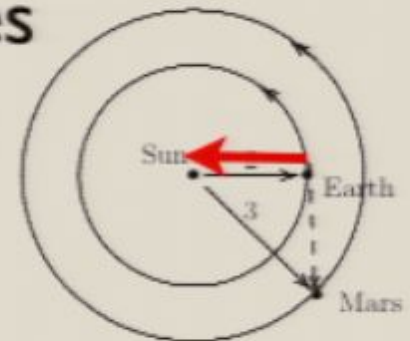
Epicycles



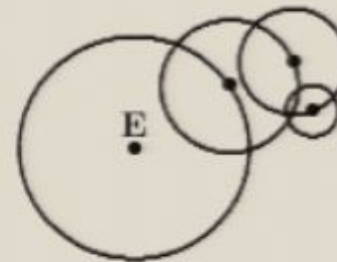
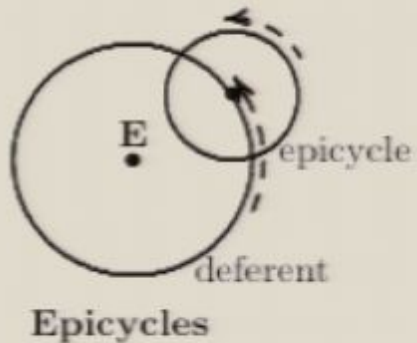
# Ptolomy and his epicycles

Aristotle: Uniform motion

Circular



The Sun-Earth-Mars system in half-astronomical u



Quasi-periodic motion

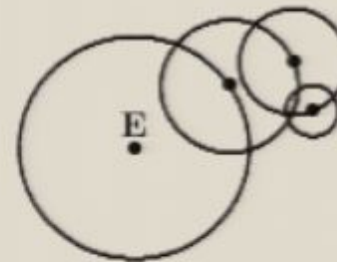
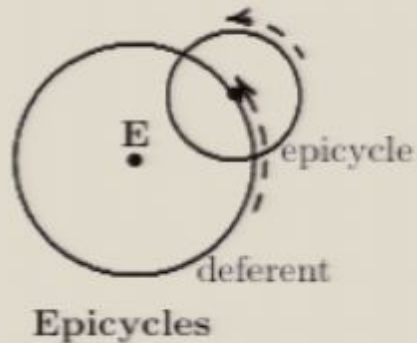
# Ptolomy and his epicycles

Aristotle: Uniform motion

Circular



The Sun-Earth-Mars system in half-astronomical u



Quasi-periodic motion

$$z(t) = \sum_{j=0}^{\infty} a_j e^{b_j i t}$$

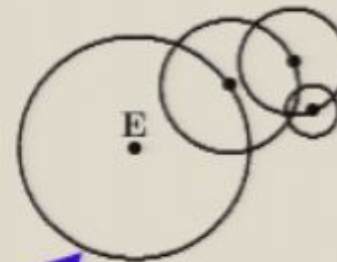
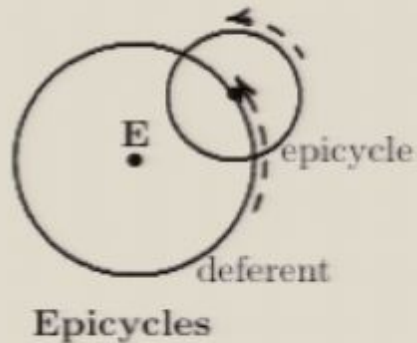
# Ptolomy and his epicycles

Aristotle: Uniform motion

Circular

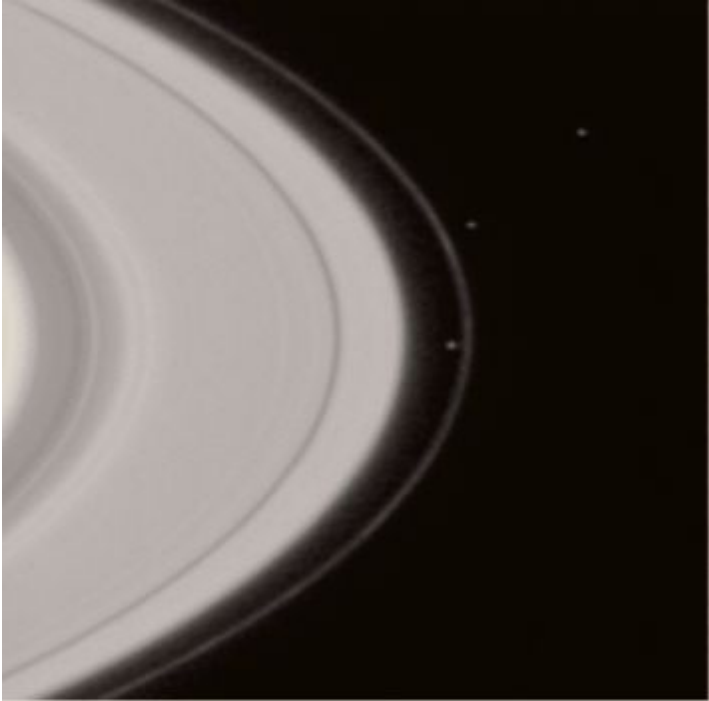


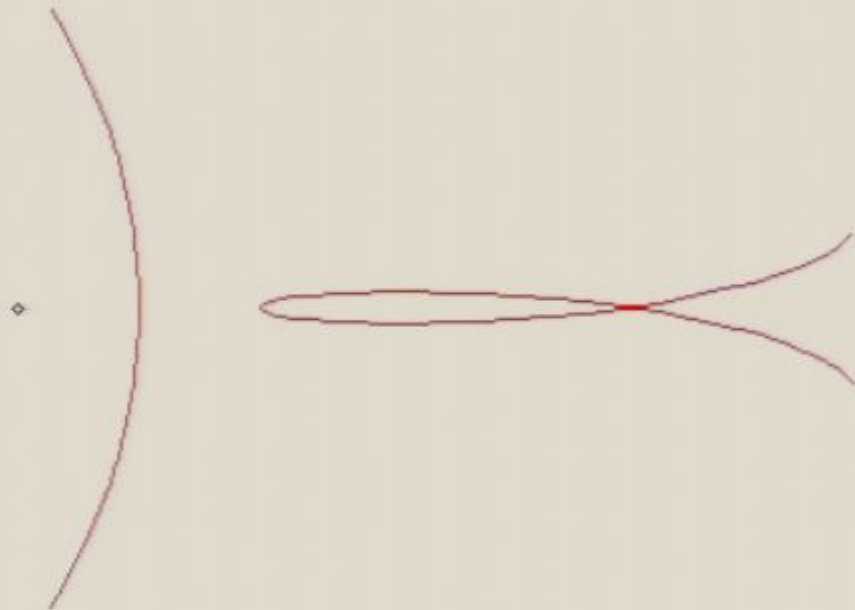
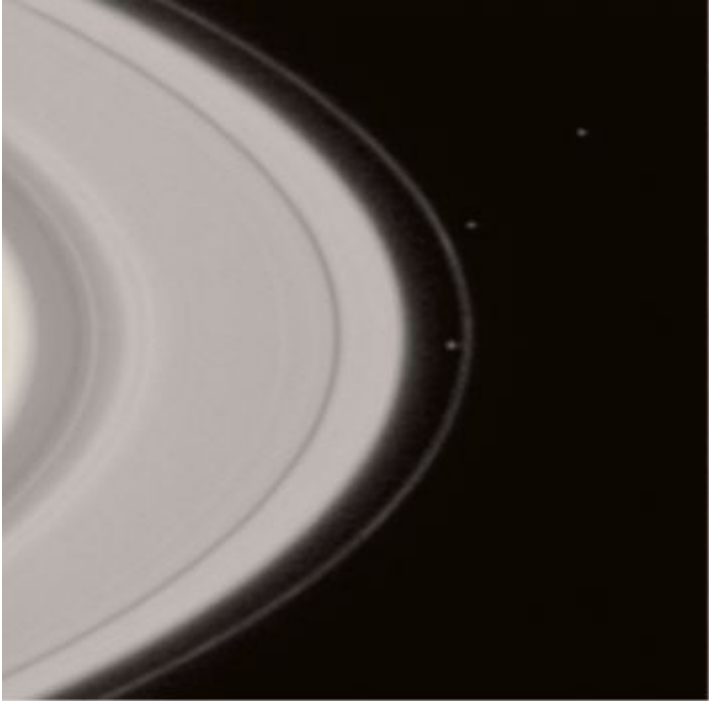
The Sun-Earth-Mars system in half-astronomical u

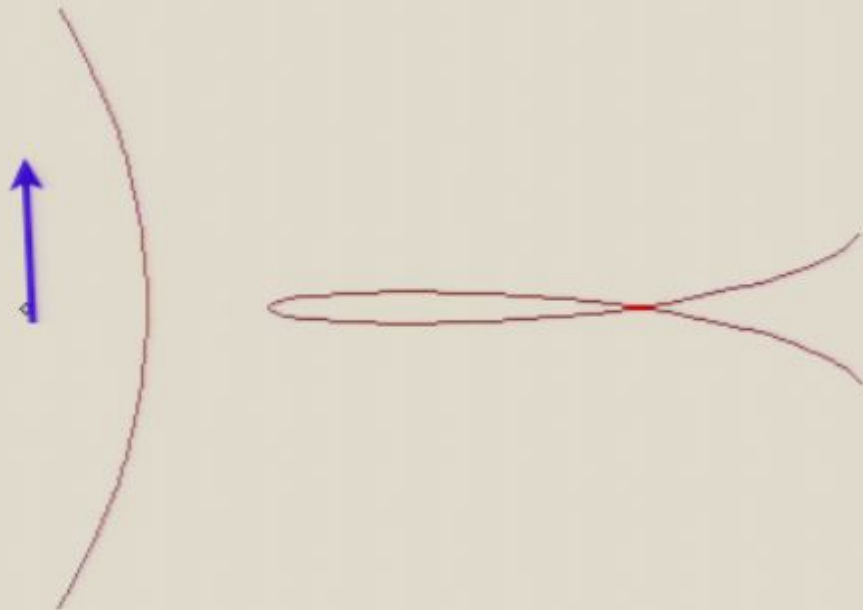
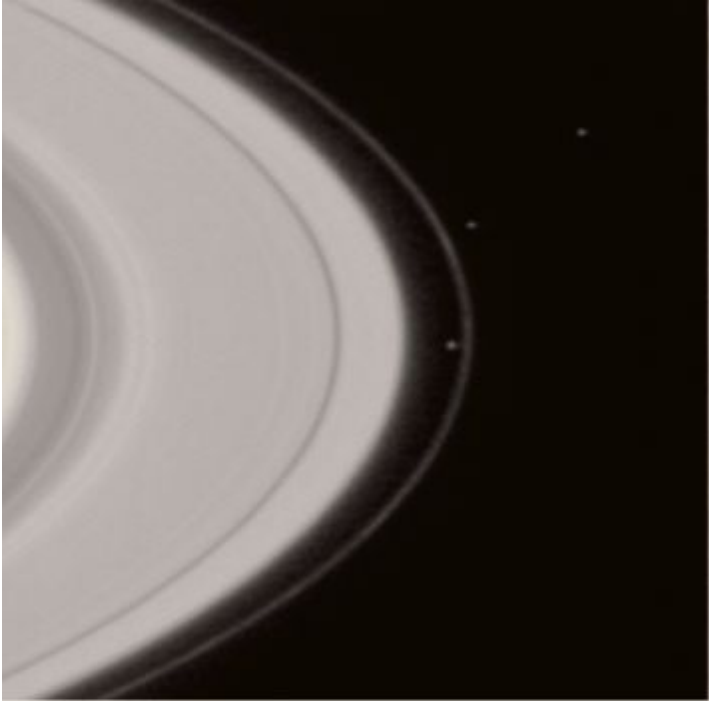


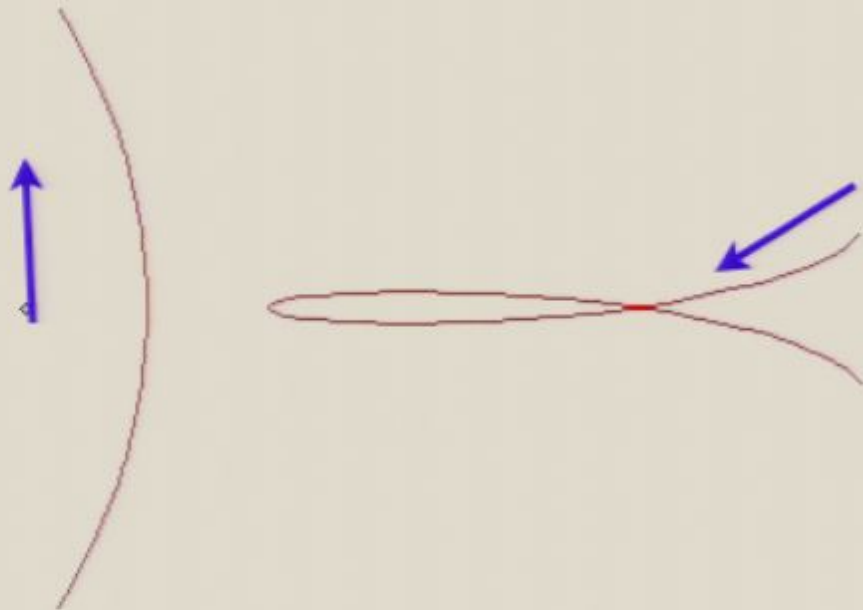
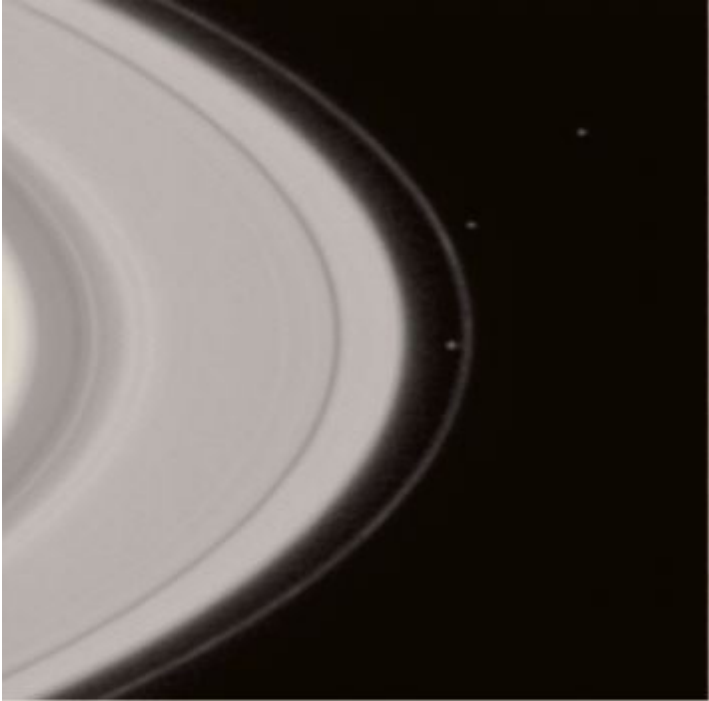
Quasi-periodic motion

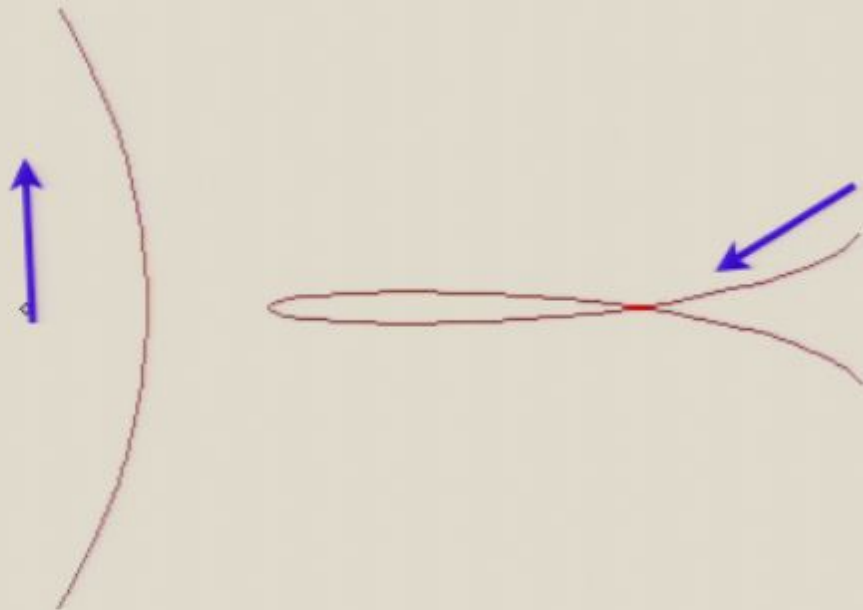
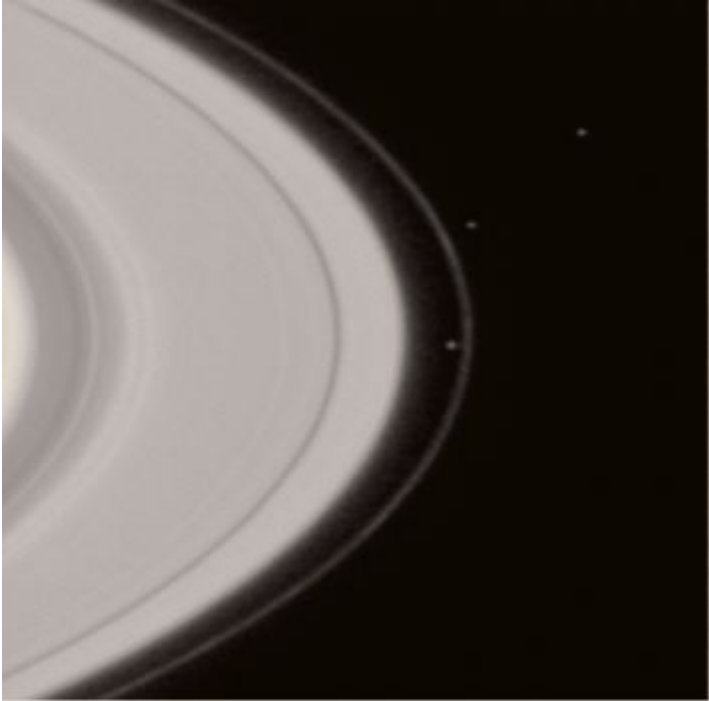
$$z(t) = \sum_{j=0}^{\infty} a_j e^{b_j i t}$$

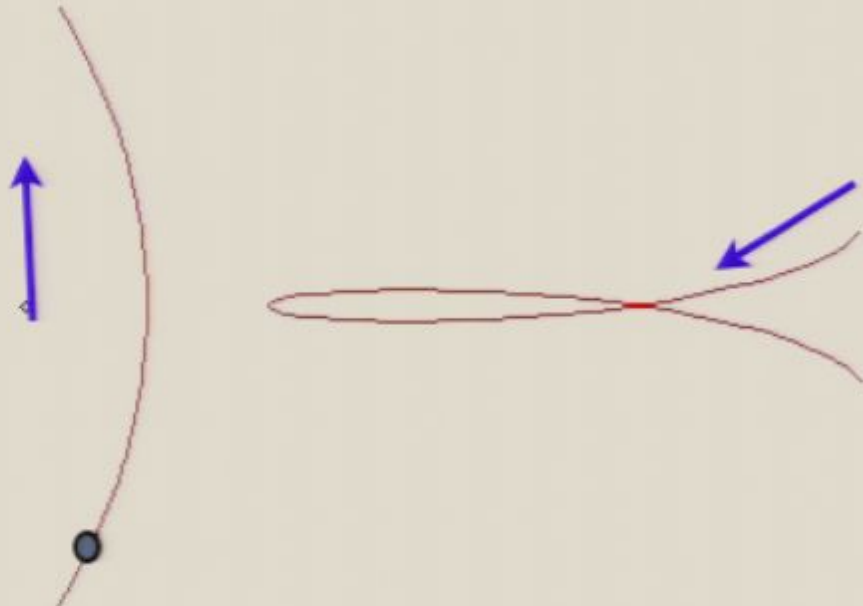
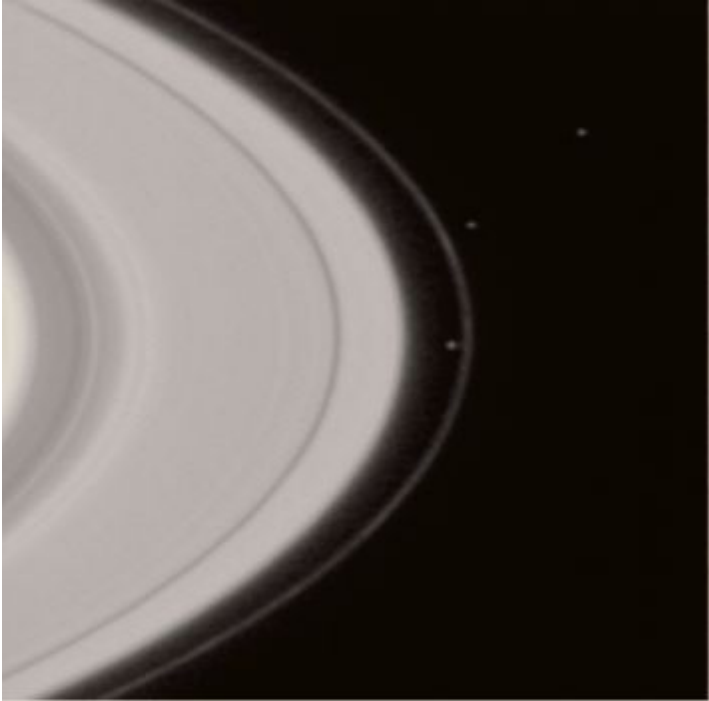


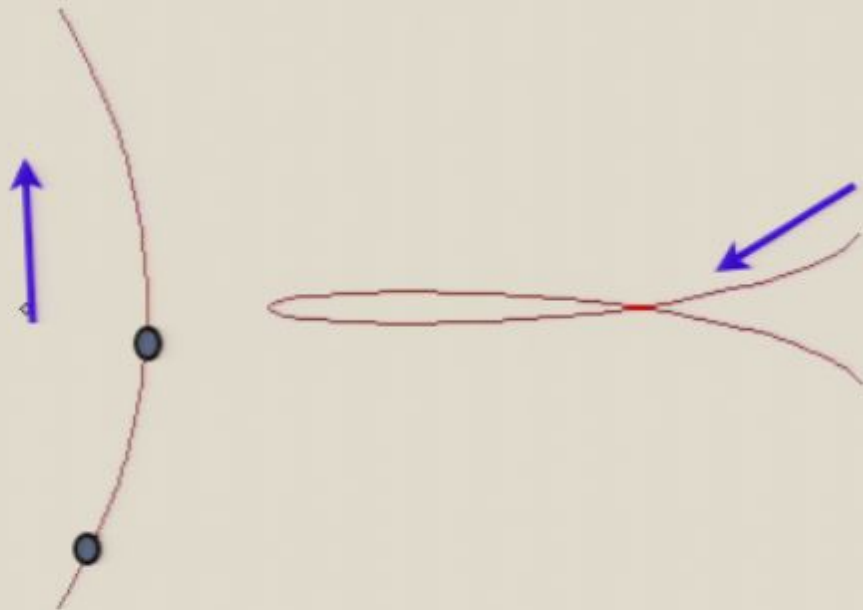
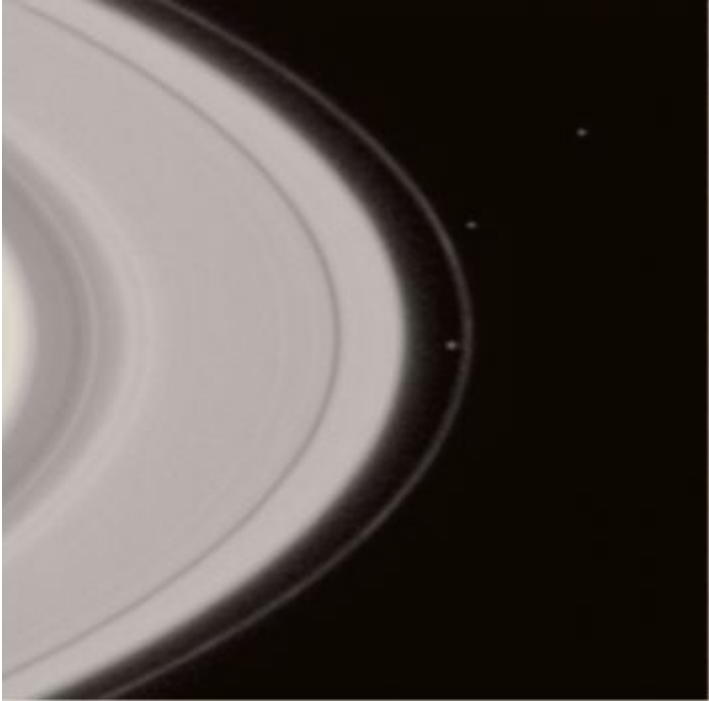


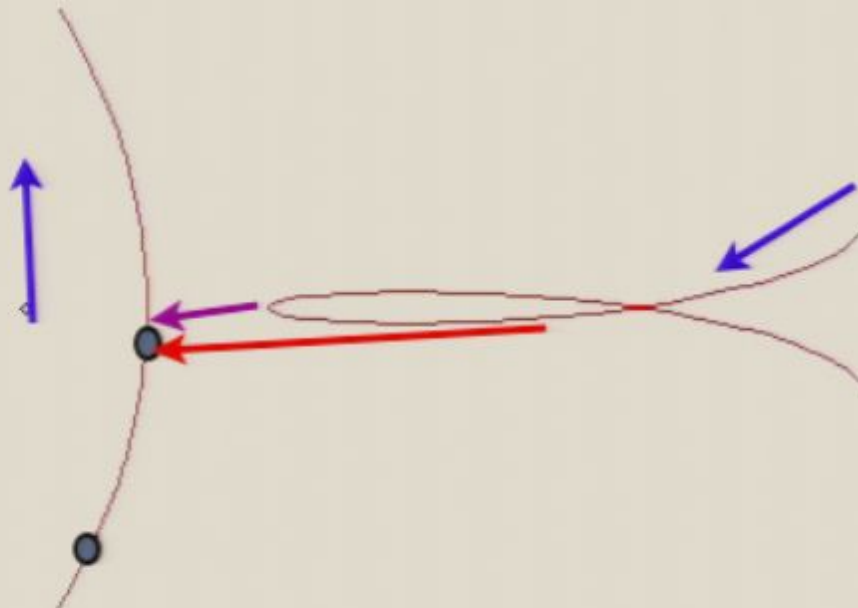
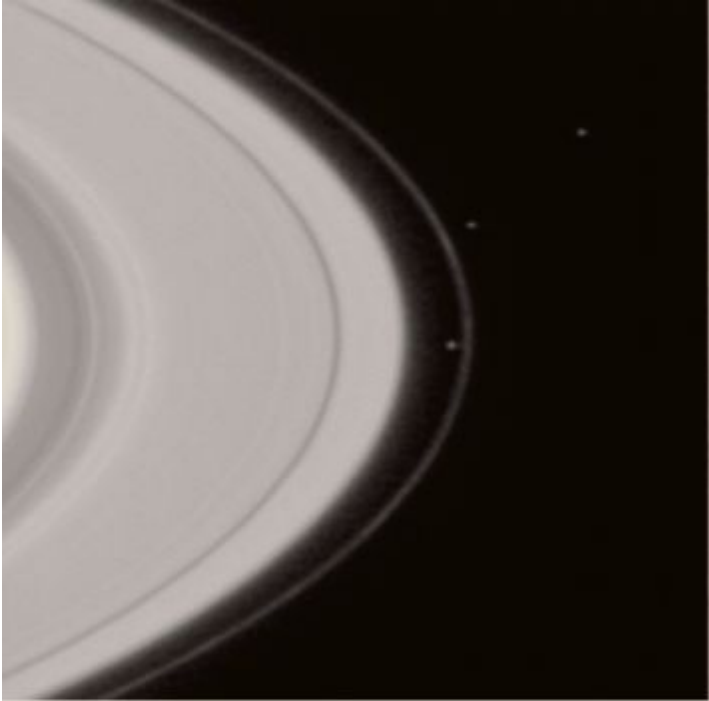


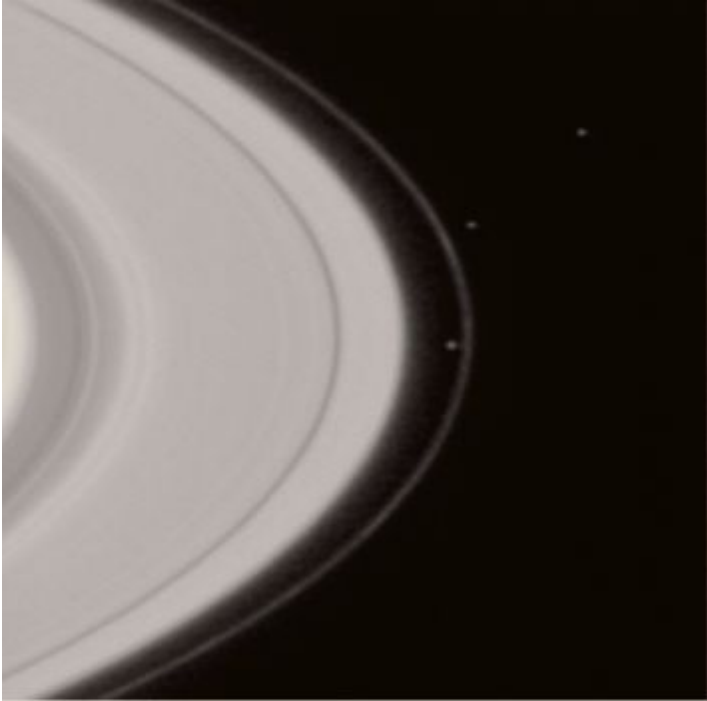




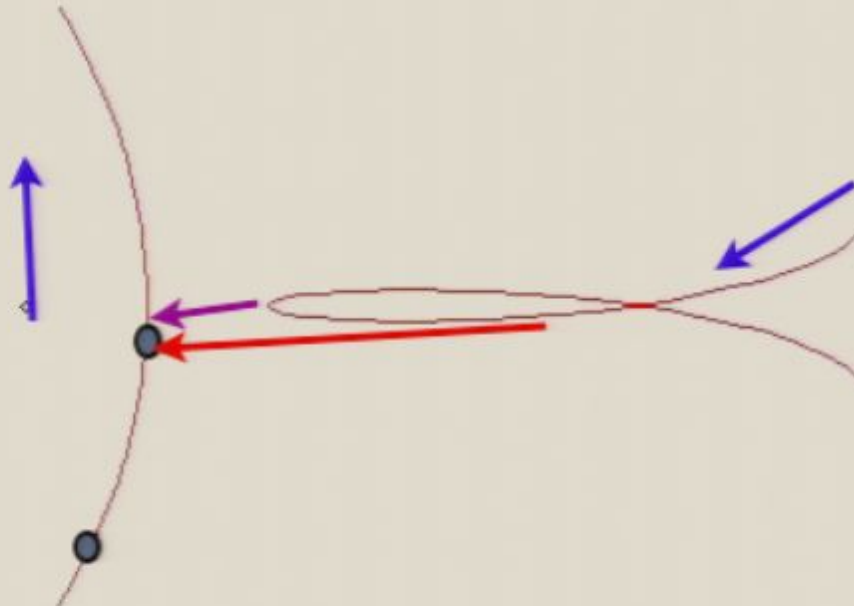


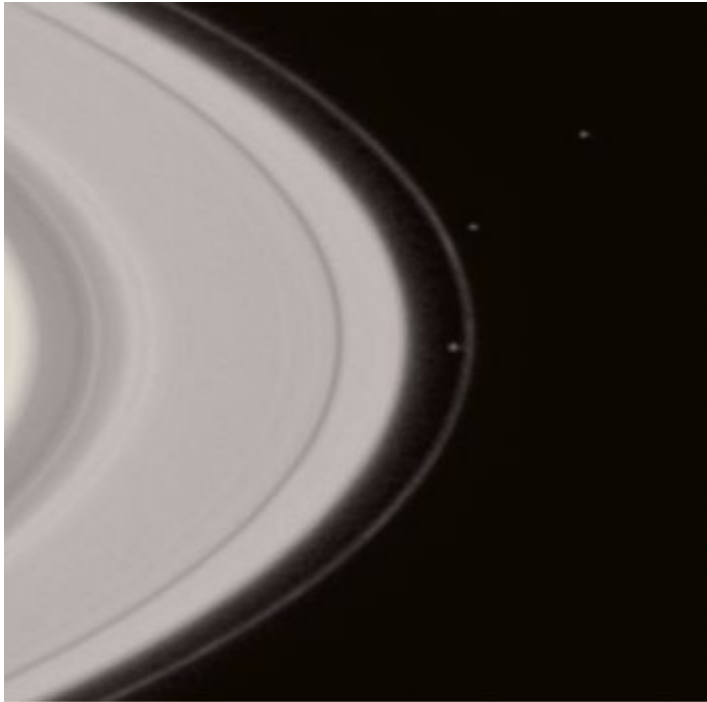




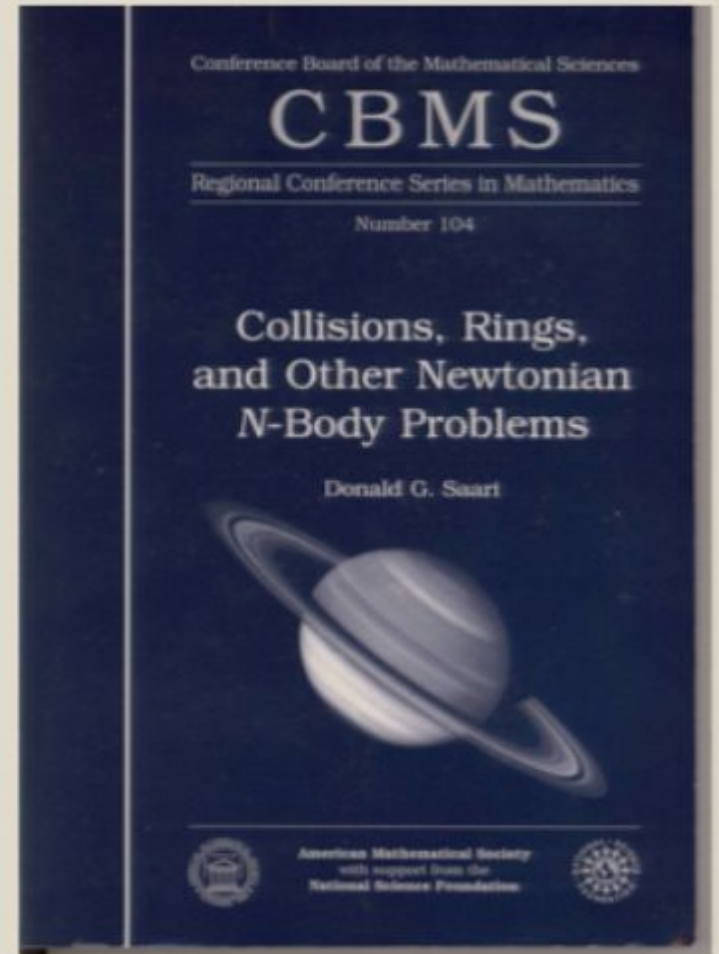
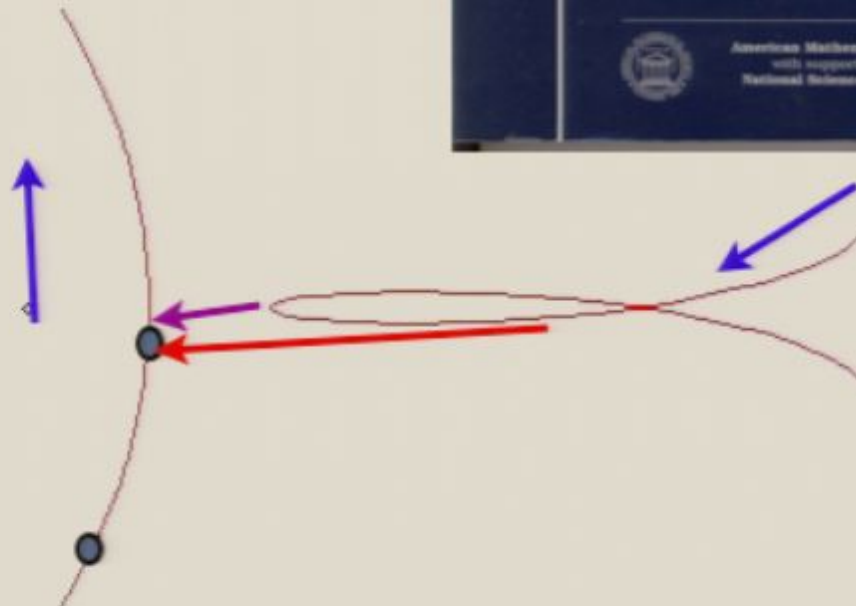


Chaos can appear  
in seemingly well-  
behaved settings





Chaos can appear  
in seemingly well-  
behaved settings



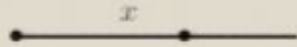
For the evolution of the universe,  
what do we think might happen?

For the evolution of the universe,  
what do we think might happen?

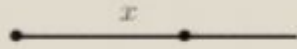


# Mathematics for the evolution of the universe

# Mathematics for the evolution of the universe

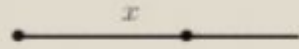


# Mathematics for the evolution of the universe



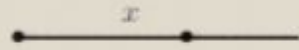
$$x'' = \frac{-1}{x^2}$$

# Mathematics for the evolution of the universe



$$x' x'' = \frac{-1}{x^2} x'$$

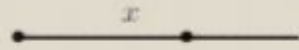
# Mathematics for the evolution of the universe



$$x' x'' = \frac{-1}{x^2} x'$$

$$\frac{1}{2} x'^2 = \frac{1}{x} + h$$

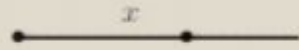
# Mathematics for the evolution of the universe



$$x' x'' = \frac{-1}{x^2} x'$$

$$\frac{1}{2} x'^2 = \frac{1}{x} + h$$

# Mathematics for the evolution of the universe

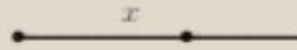


$$x' x'' = \frac{-1}{x^2} x'$$

$$\frac{1}{2} x'^2 = \frac{1}{x} + h$$

| h sign  | x behavior |
|---------|------------|
| $h < 0$ |            |
| $h = 0$ |            |
| $h > 0$ |            |

# Mathematics for the evolution of the universe



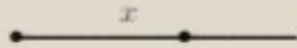
$$x' x'' = \frac{-1}{x^2} x'$$

$$\cancel{\frac{1}{2}x^2} = \frac{1}{x} + h$$

$$h < 0$$

| h sign  | x behavior |
|---------|------------|
| $h < 0$ |            |
| $h = 0$ |            |
| $h > 0$ |            |

# Mathematics for the evolution of the universe



$$\cancel{\frac{1}{2}x^2} = \frac{1}{x} + h$$

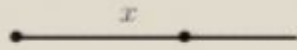
$$x' x'' = \frac{-1}{x^2} x'$$

$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior |
|---------|------------|
| $h < 0$ | $x = O(1)$ |
| $h = 0$ |            |
| $h > 0$ |            |



# Mathematics for the evolution of the universe



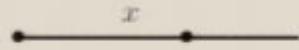
$$\cancel{\frac{1}{2}x^2} = \frac{1}{x} + h$$

$$x' x'' = \frac{-1}{x^2} x'$$

$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior |
|---------|------------|
| $h < 0$ | $x = O(1)$ |
| $h = 0$ |            |
| $h > 0$ |            |

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \frac{1}{x} + h$$

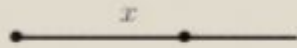
$$x' x'' = \frac{-1}{x^2} x'$$

$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior |
|---------|------------|
| $h < 0$ | $x = O(1)$ |
| $h = 0$ |            |
| $h > 0$ |            |

$$h > 0$$

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \cancel{\frac{1}{x}} + h$$

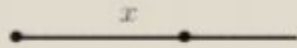
$$x' x'' = \frac{-1}{x^2} x'$$

$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior |
|---------|------------|
| $h < 0$ | $x = O(1)$ |
| $h = 0$ |            |
| $h > 0$ |            |

$$h > 0$$

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \cancel{\frac{1}{x}} + h$$

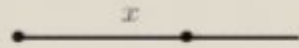
$$x' x'' = \frac{-1}{x^2} x'$$

$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior |
|---------|------------|
| $h < 0$ | $x = O(1)$ |
| $h = 0$ |            |
| $h > 0$ |            |

$$h > 0 \quad x' \geq \sqrt{2h}$$

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \frac{1}{x} + h$$

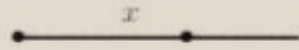
$$x' x'' = \frac{-1}{x^2} x'$$

$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior |
|---------|------------|
| $h < 0$ | $x = O(1)$ |
| $h = 0$ |            |
| $h > 0$ |            |

$$h > 0 \quad x' \geq \sqrt{2h}$$

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \frac{1}{x} + h$$

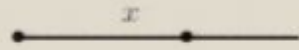
$$x' x'' = \frac{-1}{x^2} x'$$

$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior  |
|---------|-------------|
| $h < 0$ | $x = O(1)$  |
| $h = 0$ |             |
| $h > 0$ | $x \sim Bt$ |

$$h > 0 \quad x' \geq \sqrt{2h}$$

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \frac{1}{x} + h$$

$$x' x'' = \frac{-1}{x^2} x'$$

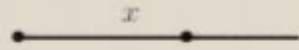
$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior  |
|---------|-------------|
| $h < 0$ | $x = O(1)$  |
| $h = 0$ |             |
| $h > 0$ | $x \sim Bt$ |

$$h > 0 \quad x' \geq \sqrt{2h}$$

$$h = 0$$

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \frac{1}{x} + \cancel{h}$$

$$x' x'' = \frac{-1}{x^2} x'$$

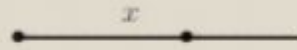
$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior  |
|---------|-------------|
| $h < 0$ | $x = O(1)$  |
| $h = 0$ |             |
| $h > 0$ | $x \sim Bt$ |

$$h > 0 \quad x' \geq \sqrt{2h}$$

$$h = 0$$

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \frac{1}{x} + \cancel{h}$$

$$x' x'' = \frac{-1}{x^2} x'$$

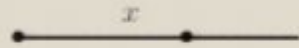
$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior  |
|---------|-------------|
| $h < 0$ | $x = O(1)$  |
| $h = 0$ |             |
| $h > 0$ | $x \sim Bt$ |

$$h > 0 \quad x' \geq \sqrt{2h}$$

$$h = 0 \quad \begin{aligned} x'^2 x &= 2 \\ x' x^{1/2} &= \sqrt{2} \end{aligned}$$

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \frac{1}{x} + \cancel{h}$$

$$x' x'' = \frac{-1}{x^2} x'$$

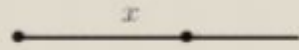
$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x = O(1)$        |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

$$h > 0 \quad x' \geq \sqrt{2h}$$

$$h = 0 \quad \begin{aligned} x'^2 x &= 2 \\ x' x^{1/2} &= \sqrt{2} \end{aligned}$$

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \frac{1}{x} + h$$

$$x' x'' = \frac{-1}{x^2} x'$$

$$h < 0 \quad x \leq \frac{1}{|h|}$$

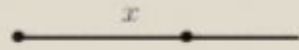
| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x = O(1)$        |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

$$h > 0 \quad x' \geq \sqrt{2h}$$

Need analytic tricks, but same for  
general two body problem

$$h = 0 \quad \begin{aligned} x'^2 x &= 2 \\ x' x^{1/2} &= \sqrt{2} \end{aligned}$$

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \frac{1}{x} + h$$

$$x' x'' = \frac{-1}{x^2} x'$$

$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x = O(1)$        |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

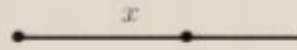
$$h > 0 \quad x' \geq \sqrt{2h}$$

Need analytic tricks, but same for general two body problem

$$h = 0 \quad \begin{aligned} x'^2 x &= 2 \\ x' x^{1/2} &= \sqrt{2} \end{aligned}$$

## Three bodies?

# Mathematics for the evolution of the universe



$$\frac{1}{2}x'^2 = \frac{1}{x} + h$$

$$x' x'' = \frac{-1}{x^2} x'$$

$$h < 0 \quad x \leq \frac{1}{|h|}$$

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x = O(1)$        |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

$$h > 0 \quad x' \geq \sqrt{2h}$$

Need analytic tricks, but same for general two body problem

$$h = 0 \quad \begin{aligned} x'^2 x &= 2 \\ x' x^{1/2} &= \sqrt{2} \end{aligned}$$

Three bodies? **Newton's Headache**

A setup?

A setup?

*Acta Mathematica*

A setup?



*Acta Mathematica*

A setup?

*Acta Mathematica*



A setup?

*Acta Mathematica*



A setup?

*Acta Mathematica*



*Bernstein*

A setup?

*Acta Mathematica*

Solve the Newtonian  
N-body problem



*Weyl*

A setup?

*Acta Mathematica*

Solve the Newtonian  
**N-body** problem



*Hilbert*

A setup?

*Acta Mathematica*

Solve the Newtonian  
N-body problem

Restricted 3-body



*Brunsting*

A setup?

*Acta Mathematica*

Solve the Newtonian  
N-body problem

Restricted 3-body



*Burstein*

A setup?



*Acta Mathematica*



Solve the Newtonian  
N-body problem

Restricted 3-body

Errors



*W Killing*

A setup?



*Acta Mathematica*

Solve the Newtonian  
N-body problem

Restricted 3-body

Errors

And they  
were caused by  
**Chaos**



*Brunsting*

# Three-body problem -- and Newton's headache



# Three-body problem -- and Newton's headache

0

●

Jean Chazy, 1920s

0

# Three-body problem -- and Newton's headache

0

●

Jean Chazy, 1920s

0

***Expanding Universe,  
Maximum distance goes  
to Infinity***

# Three-body problem -- and Newton's headache

0

●

Jean Chazy, 1920s

0

Triangle inequality  
to make two two-body problems

***Expanding Universe,  
Maximum distance goes  
to Infinity***

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

***Expanding Universe,  
Maximum distance goes  
to Infinity***

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

***Expanding Universe,  
Maximum distance goes  
to Infinity***

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

Various combinations; e.g.

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

Various combinations; e.g.

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

$$\rho \sim t$$

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

Various combinations; e.g.

$$\rho \sim t$$

$$r \sim t,$$



# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

Various combinations; e.g.

$$\rho \sim t$$



$$r \sim t,$$

$$r \sim t^{2/3},$$

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

Various combinations; e.g.

$$\rho \sim t$$



$$\begin{aligned} r &\sim t, \\ r &\sim t^{2/3}, \\ r &= O(1) \end{aligned}$$

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

Various combinations; e.g.

$$\rho \sim t$$

$$\rho \sim t^{2/3}$$

$$\begin{aligned} r &\sim t, \\ r &\sim t^{2/3}, \\ r &= O(1) \end{aligned}$$

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

Various combinations; e.g.

$$\rho \sim t$$

$$\begin{aligned} r &\sim t, \\ r &\sim t^{2/3}, \\ r &= O(1) \end{aligned}$$

$$\rho \sim t^{2/3}$$

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

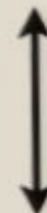
Various combinations; e.g.

$$\rho \sim t$$



$$\begin{aligned} r &\sim t, \\ r &\sim t^{2/3}, \\ r &= O(1) \end{aligned}$$

$$\rho \sim t^{2/3}$$



$$\begin{aligned} r &\sim t^{2/3}, \\ r &= O(1) \end{aligned}$$

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

$$r, \rho = O(1)$$

Various combinations; e.g.

$$\rho \sim t$$

$$\begin{aligned} r &\sim t, \\ r &\sim t^{2/3}, \\ r &= O(1) \end{aligned}$$

$$\rho \sim t^{2/3}$$

$$\begin{aligned} r &\sim t^{2/3}, \\ r &= O(1) \end{aligned}$$

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

$$r, \rho = O(1)$$

Various combinations; e.g.

$$\rho \sim t$$

$$\begin{aligned} r &\sim t, \\ r &\sim t^{2/3}, \\ r &= O(1) \end{aligned}$$

$$\rho \sim t^{2/3}$$

$$\begin{aligned} r &\sim t^{2/3}, \\ r &= O(1) \end{aligned}$$

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

$$r, \rho = O(1)$$

$$\limsup_{t \rightarrow \infty} \max(r_{i,j}) = \infty$$

$$\liminf_{t \rightarrow \infty} \max(r_{i,j}) < \infty$$

Various combinations; e.g.

$$\rho \sim t$$

$$r \sim t,$$

$$r \sim t^{2/3},$$

$$r = O(1)$$

$$\rho \sim t^{2/3}$$

$$r \sim t^{2/3},$$

$$r = O(1)$$

# Three-body problem -- and Newton's headache



Jean Chazy, 1920s

Triangle inequality  
to make two two-body problems

**Expanding Universe,  
Maximum distance goes  
to Infinity**

| h sign  | x behavior        |
|---------|-------------------|
| $h < 0$ | $x$ bounded       |
| $h = 0$ | $x \sim At^{2/3}$ |
| $h > 0$ | $x \sim Bt$       |

$$r, \rho = O(1)$$

$$\limsup_{t \rightarrow \infty} \max(r_{i,j}) = \infty$$

$$\liminf_{t \rightarrow \infty} \max(r_{i,j}) < \infty$$

Various combinations; e.g.

$$\rho \sim t$$

$$r \sim t,$$

$$r \sim t^{2/3},$$

$$r = O(1)$$

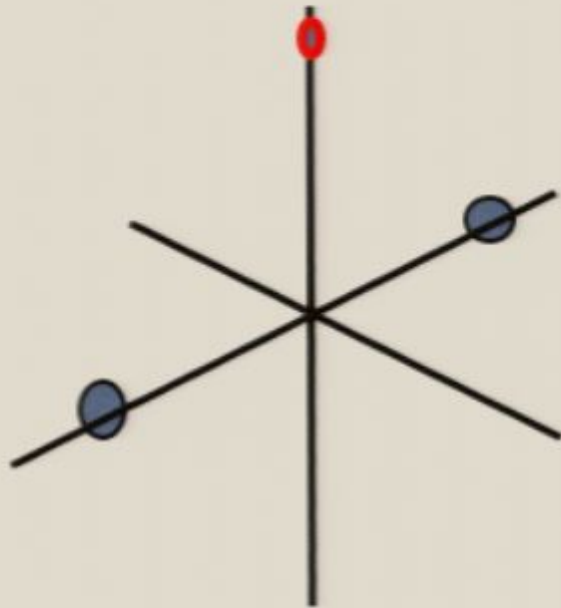
$$\rho \sim t^{2/3}$$

$$r \sim t^{2/3},$$

$$r = O(1)$$

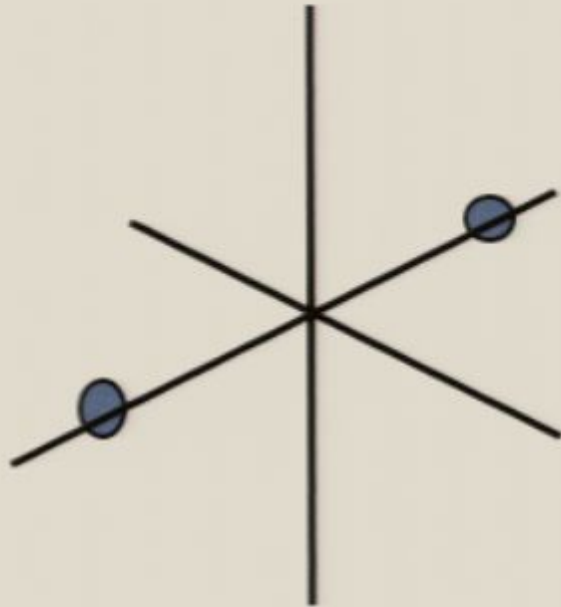
**Oscillatory Motion**

## ***K. Sitnikov and Oscillatory Motion, 1960***



**Circular**

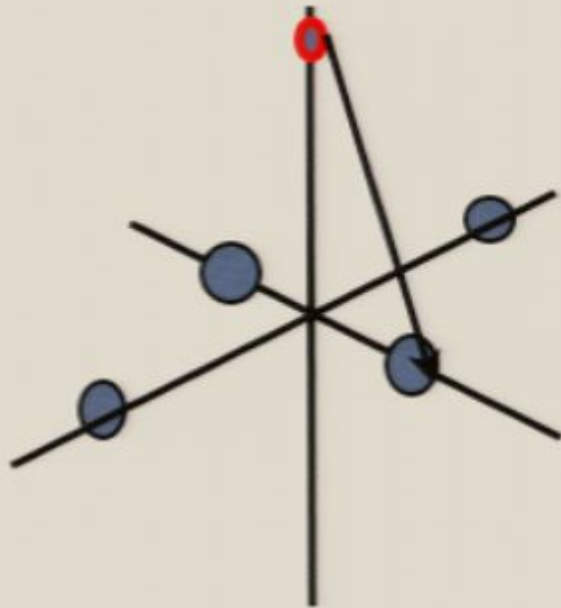
## ***K. Sitnikov and Oscillatory Motion, 1960***



**Circular**

**Elliptic**

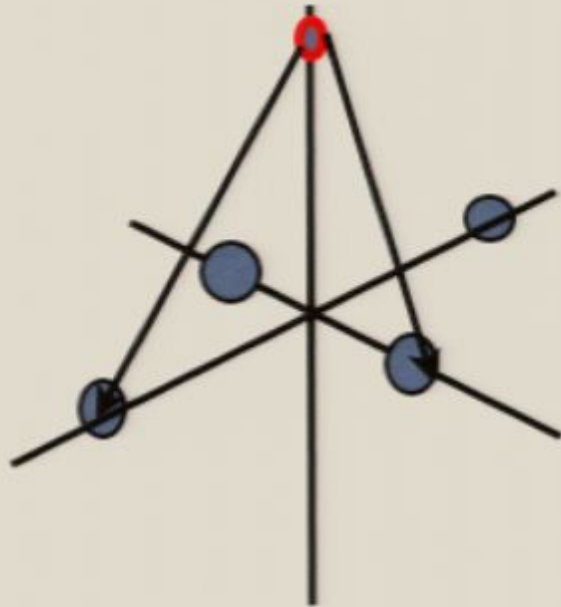
## ***K. Sitnikov and Oscillatory Motion, 1960***



**Circular**

**Elliptic**

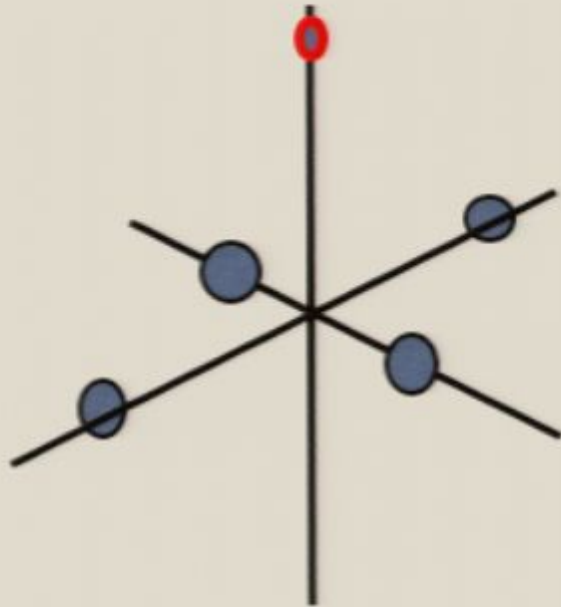
## ***K. Sitnikov and Oscillatory Motion, 1960***



**Circular**

**Elliptic**

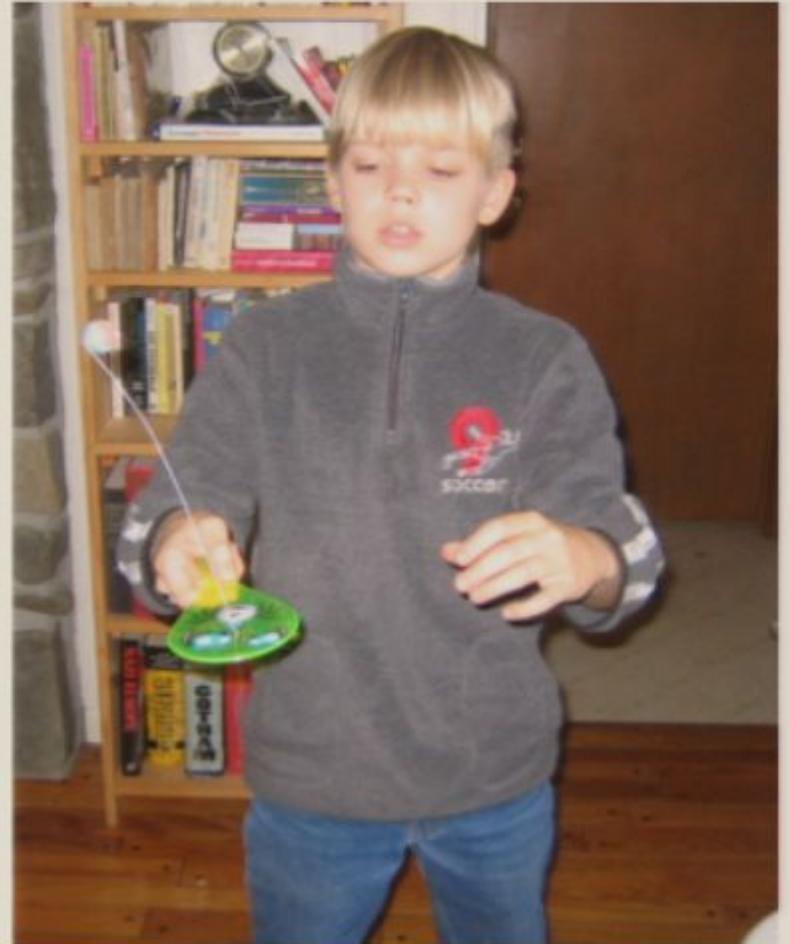
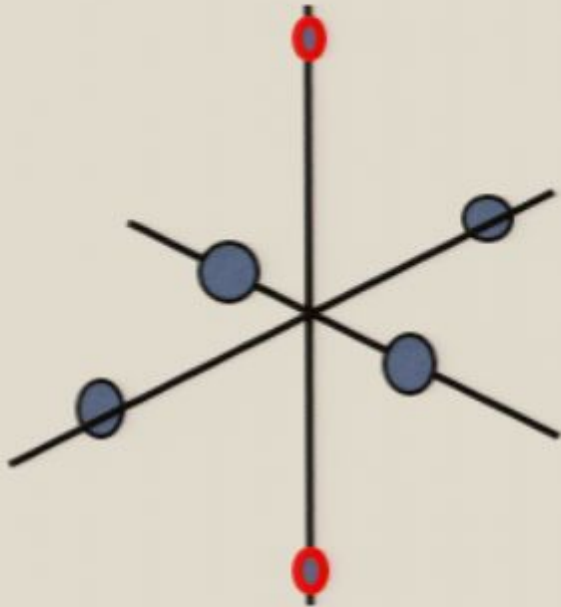
## ***K. Sitnikov and Oscillatory Motion, 1960***



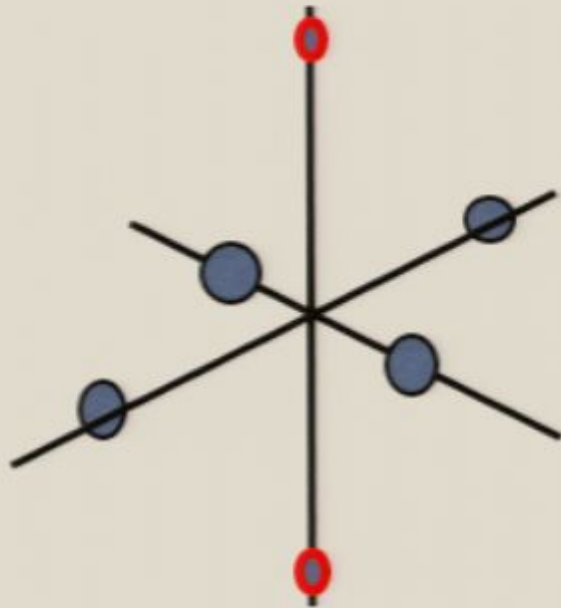
**Circular**

**Elliptic**

## **K. Sitnikov and Oscillatory Motion, 1960**



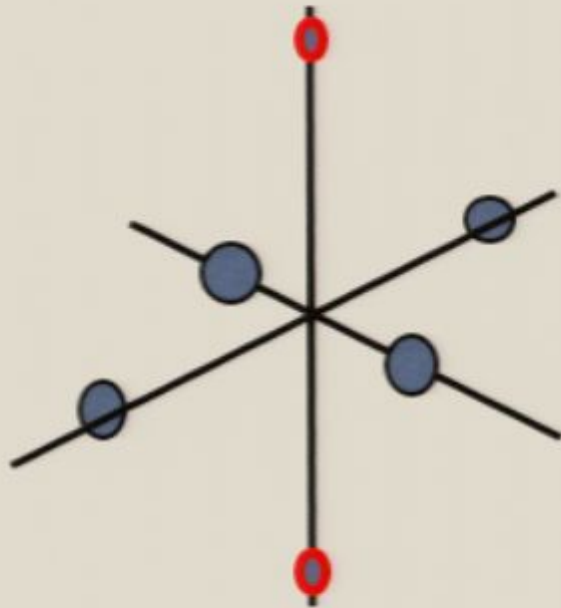
## ***K. Sitnikov and Oscillatory Motion, 1960***



**Circular**

**Elliptic**

## ***K. Sitnikov and Oscillatory Motion, 1960***

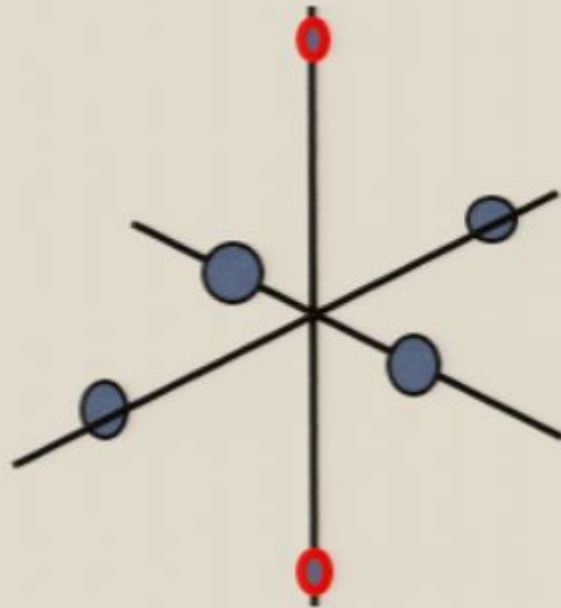


**Circular**

**Elliptic**

***Chaos***

## ***K. Sitnikov and Oscillatory Motion, 1960***



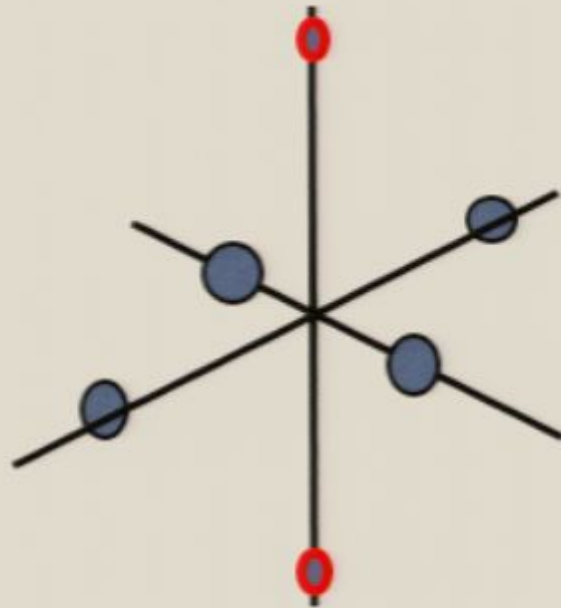
**Circular**

**Elliptic**

***Chaos***

**Oscillatory motion and chaos elsewhere**

## ***K. Sitnikov and Oscillatory Motion, 1960***



**Circular**

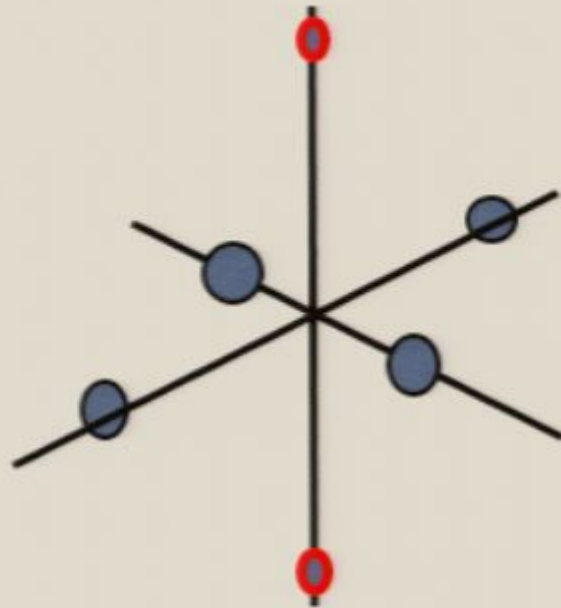
**Elliptic**

***Chaos***

## **Oscillatory motion and chaos elsewhere**

Saari and Xia--collinear case

## ***K. Sitnikov and Oscillatory Motion, 1960***



**Circular**

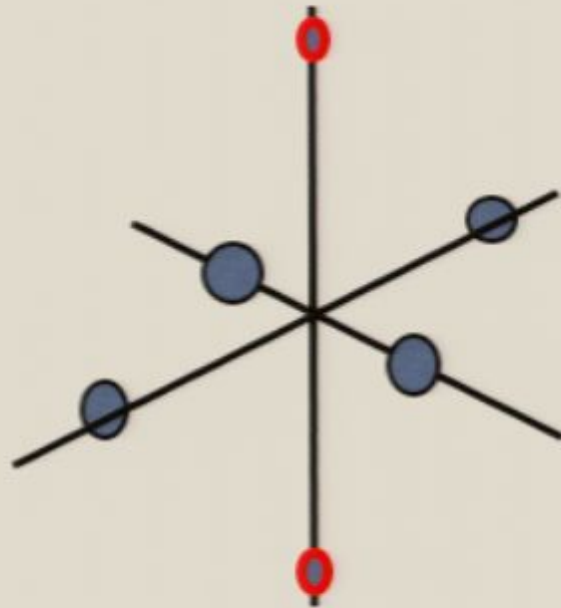
**Elliptic**

***Chaos***

## **Oscillatory motion and chaos elsewhere**

Saari and Xia--collinear case

## ***K. Sitnikov and Oscillatory Motion, 1960***



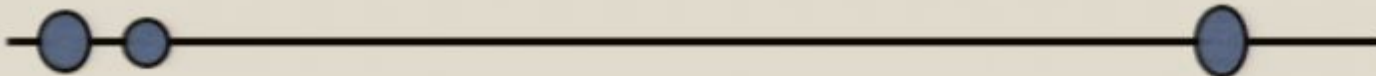
**Circular**

**Elliptic**

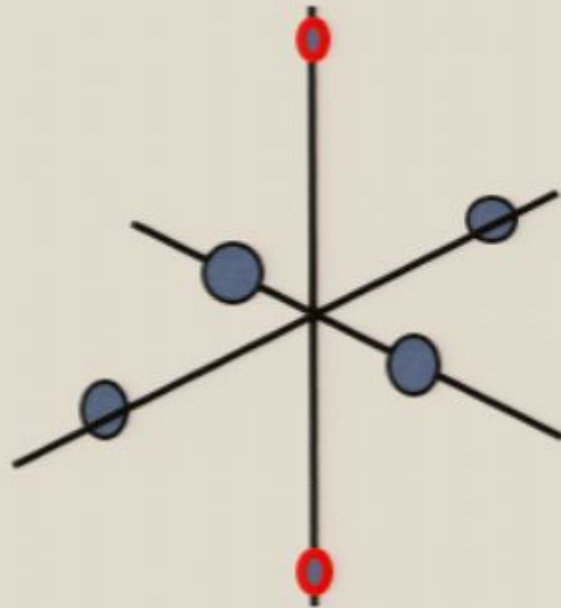
***Chaos***

## **Oscillatory motion and chaos elsewhere**

Saari and Xia--collinear case



## **K. Sitnikov and Oscillatory Motion, 1960**



**Circular**

**Elliptic**

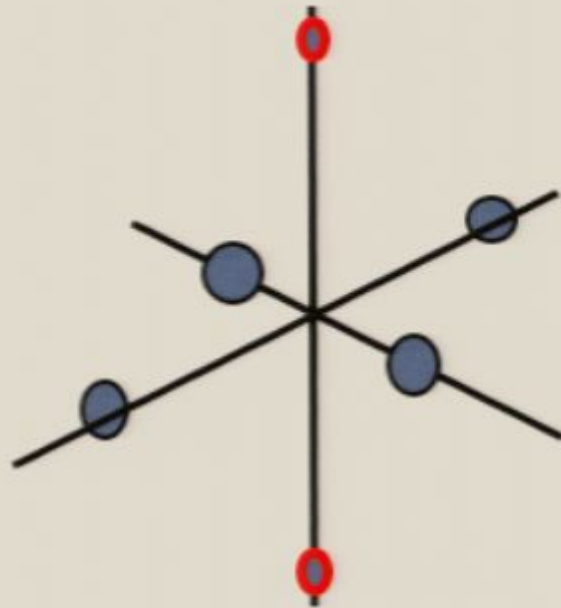
**Chaos**

## **Oscillatory motion and chaos elsewhere**

Saari and Xia--collinear case



## K. Sitnikov and Oscillatory Motion, 1960



**Circular**

**Elliptic**

**Chaos**

## Oscillatory motion and chaos elsewhere

Saari and Xia--collinear case



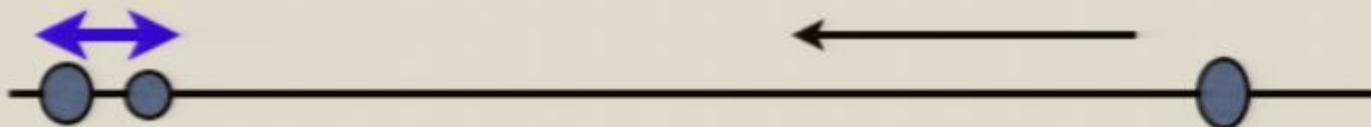
**R. McGehee**

## K. Sitnikov and Oscillatory Motion, 1960



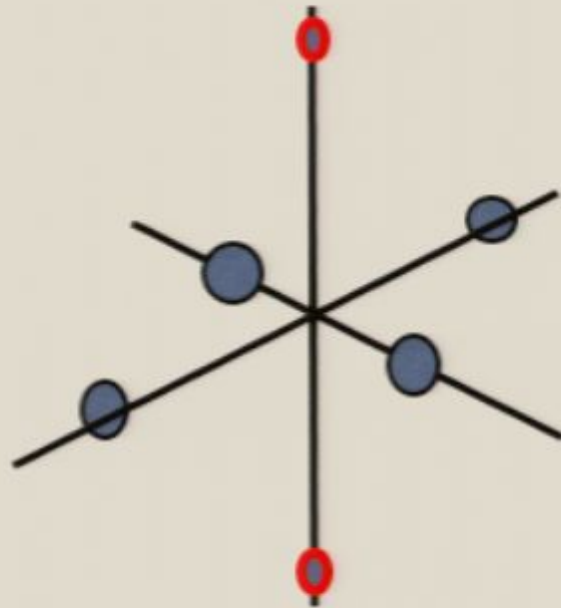
Osc

ere



**R. McGehee**

## K. Sitnikov and Oscillatory Motion, 1960



**Circular**

**Elliptic**

**Chaos**

## Oscillatory motion and chaos elsewhere

Saari and Xia--collinear case



**R. McGehee**

## ***Evolution for any number of bodies?***

## ***Evolution for any number of bodies?***

Saari, and later, Marchal and Saari

## ***Evolution for any number of bodies?***

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

Saari, and later, Marchal and Saari

## ***Evolution for any number of bodies?***

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3})$$

Saari, and later, Marchal and Saari

## ***Evolution for any number of bodies?***

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk}t^{2/3} \quad r_{jk} \sim B_{jk}t$$

## ***Evolution for any number of bodies?***

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk}t^{2/3} \quad r_{jk} \sim B_{jk}t \quad \text{OR ...}$$

## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk}t^{2/3} \quad r_{jk} \sim B_{jk}t \quad \text{OR ...}$$


0

## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$

00  
0

## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$



## Evolution for any number of bodies?

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$



Saari, and later, Marchal and Saari



Andromeda Galaxy  
GALEX



Andromeda Galaxy  
Visible light image (John Gleason)

## Evolution for any number of bodies?

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$

Saari, and later, Marchal and Saari



Andromeda Galaxy  
GALEX



Andromeda Galaxy  
Visible light image (John Gleason)

## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$



## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$



## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$

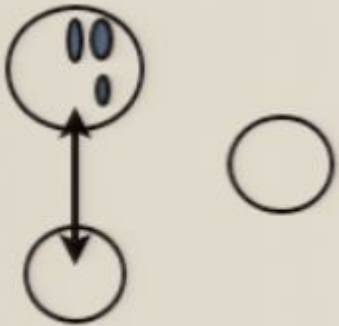


## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$

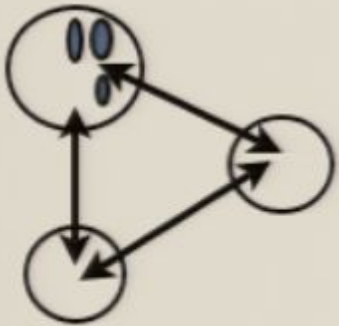


## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$

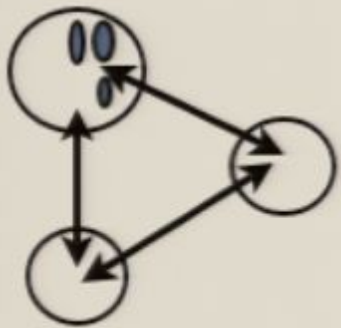


## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$



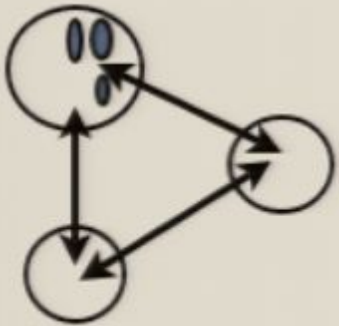
Cluster of galaxies

## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$

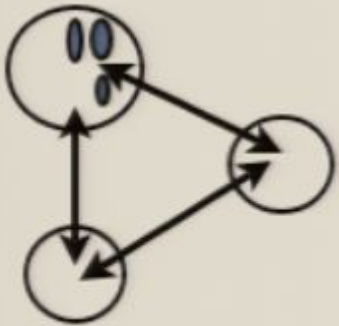


## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$

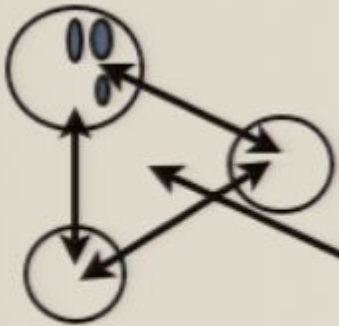


## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$

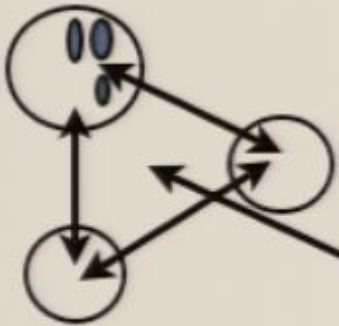


## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$

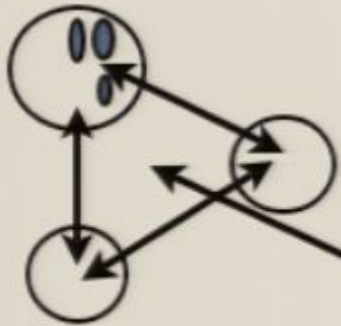


## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk}t^{2/3} \quad r_{jk} \sim B_{jk}t \quad \text{OR ...}$$



**Etc.**

Corresponds to what  
we observe

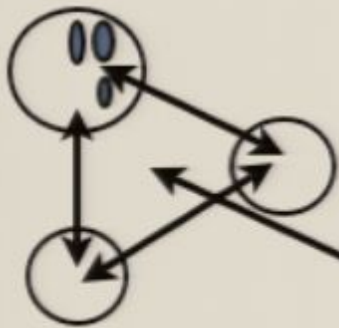


## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk}t^{2/3} \quad r_{jk} \sim B_{jk}t \quad \text{OR ...}$$



**Etc.**

Corresponds to what  
we observe

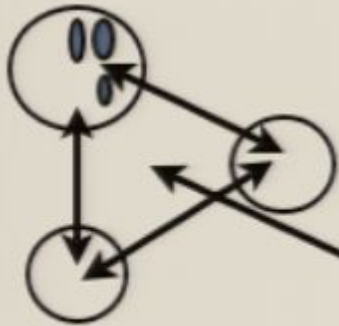


## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk} t^{2/3} \quad r_{jk} \sim B_{jk} t \quad \text{OR ...}$$



**Etc.**

Configurations

Corresponds to what  
we observe

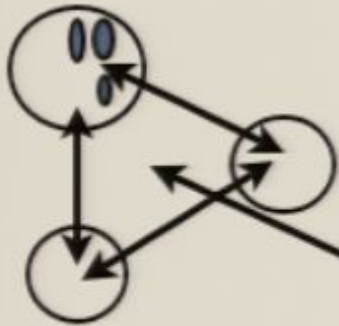


## Evolution for any number of bodies?

Saari, and later, Marchal and Saari

**Theorem:** As time goes to infinity,  
all mutual distances must satisfy

$$r_{jk} = O(t^{2/3}) \quad r_{jk} \sim A_{jk}t^{2/3} \quad r_{jk} \sim B_{jk}t \quad \text{OR ...}$$



Etc.

**The diameter of the universe goes  
to infinity faster than any multiple  
of  $t$**

Configurations

Corresponds to what  
we observe



**Theorem** (Saari and Xia) For  $n > 3$ , there exists a cantor set of examples where



**Theorem** (Saari and Xia) For  $n > 3$ , there exists a cantor set of examples where

$$\frac{\text{Diameter of system}}{t} \rightarrow \infty$$


**Theorem** (Saari and Xia) For  $n > 3$ , there exists a cantor set of examples where

$$\frac{\text{Diameter of system}}{e^{et}} \rightarrow \infty$$



**Theorem** (Saari and Xia) For  $n > 3$ , there exists a cantor set of examples where

$$\frac{\text{Diameter of system}}{e^{e^{e^{e^t}}}} \rightarrow \infty$$



**Theorem** (Saari and Xia) For  $n > 3$ , there exists a cantor set of examples where

$$\frac{\text{Diameter of system}}{e^{e^{e^{e^t}}}} \rightarrow \infty$$



**Theorem** (Saari and Xia) For  $n > 3$ , there exists a cantor set of examples where

$$\frac{\text{Diameter of system}}{e^{e^{e^{e^t}}}} \rightarrow \infty$$



**Theorem** (Saari and Xia) For  $n > 3$ , there exists a cantor set of examples where

$$\frac{\text{Diameter of system}}{e^{e^{e^{e^t}}}} \rightarrow \infty$$



**Theorem** (Saari and Xia) For  $n > 3$ , there exists a cantor set of examples where

$$\frac{\text{Diameter of system}}{e^{e^{e^{e^t}}}} \rightarrow \infty$$

Join us in investigating the N-body problem



**Theorem** (Saari and Xia) For  $n > 3$ , there exists a cantor set of examples where

$$\frac{\text{Diameter of system}}{e^{e^{e^{e^t}}}} \rightarrow \infty$$

Join us in investigating the N-body problem



**Thank you!**

$$-\frac{1}{r^2} + O\left(\frac{1}{r^2}\right)$$

**Theorem** (Saari and Xia) For  $n > 3$ , there exists a cantor set of examples where

$$\frac{\text{Diameter of system}}{e^{e^{e^{e^t}}}} \rightarrow \infty$$

Join us in investigating the N-body problem



**Thank you!**

$$-\frac{1}{r^2} + O\left(\frac{1}{r^2}\right)$$

$$\frac{1}{r} P$$

$$1 < P < 3$$

$$\frac{1}{r_2} + O\left(\frac{1}{r_2}\right)$$

$\frac{1}{r_P}$

$1 < P < 3$

$$\begin{array}{c}
 \text{INF} \\
 \frac{1}{r^2} + O\left(\frac{1}{r^2}\right) \\
 \hline
 \begin{array}{ccc}
 \textcircled{p=3} & \frac{1}{r^p} & \textcircled{1 < p < 3}
 \end{array}
 \end{array}$$





**Theorem** (Saari and Xia) For  $n > 3$ , there exists a cantor set of examples where

$$\frac{\text{Diameter of system}}{e^{e^{e^{et}}}} \rightarrow \infty$$

Join us in investigating the N-body problem



**Thank you!**