

Title: SUSY Gauge Theory

Date: Jun 27, 2005 09:30 AM

URL: <http://pirsa.org/05060095>

Abstract:

SUSY gauge theories ($d=4$, non-perturbative)

Outline

I. Non-renormalization theorems

II. $N=1$ super-QCD : Seiberg duality (SCFT)

III. $N=2$ " : Seiberg-Witten theory (BPS)

IV. $N=4$ SYM : S-duality

SUSY gauge theories ($d=4$, non-perturbative)

Outline

I. Non-renormalization theorems

II. $N=1$ super-QCD : Seiberg duality (SCFT)

III. $N=2$ " : Seiberg-Witten theory (BPS)

IV. $N=4$ SYM : S-duality

I. NR things

- A. N=1 review
- B. Holomorphy $S^1 \times \mathbb{R}^3$
- C. NR then χ sf
- D. YM review
- E. NR then vsf
- F. Exactly marginal operators

A. N=1 reminder

chiral superfield $\xrightarrow{\text{complex}}$ $\xrightarrow{\text{Weyl}}$

$$\bar{\Phi}(x, \theta) \sim \phi(x) + \theta \psi + \dots \quad (\bar{D}_\alpha \bar{\Phi} = 0)$$

vector sxfld

$$V(x, \theta, \bar{\theta}) \sim \bar{\theta} \sigma^\mu \theta A_\mu + \bar{\theta} \theta \lambda + \dots \quad (V = \bar{V})$$

chiral F_{UV} strength

$$W_\alpha(x, \theta) \sim \lambda_\alpha + (\sigma^\mu \theta)_\alpha F_{\mu\nu} + \dots$$

$$F_{\mu\nu} = \frac{i}{2} (F_{\mu\nu} + i \tilde{F}_{\mu\nu})$$

$$\tilde{F}_{\mu\nu} = \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} F^{\rho\sigma}$$

gauge kinetic term

$$\frac{i\tau}{16\pi} (F)^2 + c.c.$$

$$\tau = \frac{\theta}{2\pi} + i \frac{4\pi}{g^2}$$

gauge invariance

$$e^{-V} \rightarrow e^{-i\bar{\Lambda}} e^{-V} e^{+i\Lambda}$$

$$W_\alpha \rightarrow e^{-i\Lambda} W_\alpha e^{+i\Lambda}$$

$$W_\alpha = -\frac{1}{4} \bar{D}^2 e^{-V} D_\alpha e^V$$

$$D^\alpha W_\alpha = \bar{D}_\alpha \bar{W}^{\dot{\alpha}}$$
 Bianchi

$N=1$ SQCD (SU(N_c))

$$\mathcal{L}_{\text{SQCD}} = \int d^4\theta \left(\bar{Q}^i (e^V)_i^j Q_{j\alpha} + \bar{\tilde{Q}}_i e^V \tilde{Q}^i \right) + \left(\int d^4\theta \left(\frac{i\tau}{16\pi} \text{tr}(W^2) \right) + \text{c.c.} \right)$$

$Q_{i\alpha}$ "quark" χ_{sf}

$\tilde{Q}^i_{\dot{\alpha}}$ "antiquark" χ_{sf}

$i = 1 \dots N_f$

$\alpha = 1 \dots N_c \rightarrow N_c \text{-rep}$

$\dot{\alpha} = 1 \dots N_c \rightarrow N_c \text{-rep}$

$(W_2)_n^L$

$(W_3)_n^R = 0$

$N=1$ SQCD (SU(N_c))

$$\mathcal{L}_{\text{SQCD}} = \int d^4\theta \left(\bar{Q}^i (e^V)_i^j Q_{j\alpha} + \bar{\tilde{Q}}_i e^V \tilde{Q}^i \right) + \left(\int d^4\theta \left(\frac{i\tau}{16\pi} \text{tr}(W^2) \right) + \text{c.c.} \right)$$

$Q_{i\alpha}$ "quark" χ_{sf}

$\tilde{Q}^i_{\dot{\alpha}}$ "antiquark" χ_{sf}

$i=1 \dots N_f$

$a=1 \dots N_c \rightarrow N_c - \text{up}$

\downarrow
 $\bar{N}_c - \text{up}$

$(W_2)_n^2$

$(W_1)_n^0 = 0$

$$\mathcal{L}_{\text{mass}} = \int d^4\theta \, m_i^j Q_j \tilde{Q}^i + \text{c.c.}$$

B. $N=1$ SUSY selection rule: holomorphy of \mathcal{N}

$$\mathcal{I}_{\text{eff}} = \int d^4\theta \mathcal{K}_{\text{eff}}(\Phi, \bar{\Phi}, m, \tau, \bar{m}, \bar{\epsilon})$$

$$+ \int_{\frac{k}{D^2}} d^2\theta \left[\mathcal{W}_{\text{eff}}(\bar{\Phi}, m, \tau, \bar{m}, \bar{\epsilon}) + \mathcal{L}_{\text{eff}}(\bar{\Phi}, \tau, \bar{m}, \bar{\epsilon}) + \text{tr}(\tilde{W}^2) \right] + c$$

+ higher-derivative terms...

B. $N \geq 1$ SUSY selection rule: holomorphy of \mathcal{N}

$$\mathcal{L}_{\text{eff}} = \int d^4\theta \mathcal{K}_{\text{eff}}(\Phi, \bar{\Phi}, m, \tau, \bar{m}, \bar{c})$$

$$+ \int_{\frac{K}{D^2}} d^2\theta \left[\mathcal{W}_{\text{eff}}(\Phi, m, \tau, \bar{m}, \bar{c}) + \tau_{\text{eff}}(\Phi, \tau, \bar{c}, m, \bar{c}) + \text{tr}(\tilde{W}^2) \right] + \text{c.c.}$$

+ higher-derivative terms...

B. $N=1$ SUSY selection rule: holomorphy of \mathcal{W}

$$\mathcal{L}_{\text{eff}} = \int d^4\theta \mathcal{K}_{\text{eff}}(\Phi, \bar{\Phi}, m, \tau, \bar{m}, \bar{\epsilon})$$

$$+ \int_{\frac{K}{D^2}} d^4\theta \left[\mathcal{W}_{\text{eff}}(\bar{\Phi}, m, \tau, \bar{m}, \bar{\epsilon}) + \mathcal{L}_{\text{eff}}(\Phi, \tau, z, m, \bar{\epsilon}) + \text{tr}(\tilde{W}^2) \right] + c$$

+ higher-derivative terms...

N=1 SQCD (SU(N_c))

$$\sim -\frac{1}{4g^2} \text{tr} F^2 + \frac{\theta}{32\pi^2} \text{tr} F\tilde{F}$$

$$\mathcal{L}_{\text{SQCD}} = \int d^4\theta \left(\bar{Q}^i (e^V)_i^j Q_{j\alpha} + \bar{\tilde{Q}}_i e^V \tilde{Q}^i \right) + \left(\int d^2\theta \left(\frac{i\tau}{16\pi} \text{tr}(W^2) + \text{c.c.} \right) \right)$$

Q_{iα} "quark" χ_{SF}

$\bar{\tilde{Q}}^i_{\dot{\alpha}}$ "antiquark" χ_{SF}

i = 1 ... N_f

α = 1 ... N_c → N_c-rep

$\dot{\alpha} = 1 ... N_{\bar{c}}$

$$(W_2)_n^L$$

$$((W_2)_n^R = 0)$$

$$\mathcal{L}_{\text{mass}} = \int d^4\theta m_j^i Q_j^i \tilde{Q}^i + \text{c.c.}$$

B. $N=1$ SUSY selection rule: holomorphy of \mathcal{W}

$$\mathcal{I}_{\text{eff}} = \int d^4\theta \mathcal{K}_{\text{eff}}(\Phi, \bar{\Phi}, m, \tau, \bar{m}, \bar{\tau})$$

$$+ \int_{\frac{K}{D^2}} d^4\theta \left[\mathcal{W}_{\text{eff}}(\Phi, m, \tau, \bar{m}, \bar{\tau}) + \tau_{\text{eff}}(\Phi, \tau, \bar{m}, \bar{\tau}) + \text{tr}(\tilde{W}^2) \right] + \dots$$

+ higher-derivative terms...

$$\vec{E} = (q, 0, 0)$$

$$H = V(\vec{x}, \vec{x}) + \frac{\vec{p}^2}{2m} + q x_1 \Rightarrow \vec{E} \cdot \vec{x}$$

C. Xsf NR thm

A

• use R-symmetry \rightarrow Q_α transforms in non-trivial rep of R

N=1 SUSY : $U(1)_R$

$$Q_\alpha \rightarrow e^{-i\epsilon} Q_\alpha$$

$$\bar{Q}_\alpha \rightarrow e^{+i\epsilon} \bar{Q}_\alpha$$

$$[R, Q_\alpha] = -Q_\alpha$$

$$[R, \bar{Q}_\alpha] = +\bar{Q}_\alpha$$

$$Q_\alpha \sim \frac{\partial}{\partial \theta^\alpha} + \dots$$

$$R\left(\frac{\partial}{\partial \theta^\alpha}\right) = +1 = R(d\theta^\alpha)$$

C. Xsf NR thm

• use R-symmetry \rightarrow Q_α transforms in non-trivial rep of R

N=1 SUSY $U(1)_R$

$$Q_\alpha \rightarrow e^{-i\epsilon} Q_\alpha$$

$$\bar{Q}_{\dot{\alpha}} \rightarrow e^{+i\epsilon} \bar{Q}_{\dot{\alpha}}$$

$$[R, Q_\alpha] = -Q_\alpha$$

$$[R, \bar{Q}_{\dot{\alpha}}] = +\bar{Q}_{\dot{\alpha}}$$

$$Q_\alpha \sim \frac{\partial}{\partial \theta^\alpha} + \dots$$

$$R\left(\frac{\partial}{\partial \theta^\alpha}\right) = +1 = R(40^\circ)$$

$$R(\bar{Q}_{\dot{\alpha}}) = +1$$

\Rightarrow need

$$R(K) = 0$$

$$R(W) = +2$$

C. Xsf NR thm

• use R-symmetry \rightarrow Q_α transforms in non-trivial rep of R

N=1 SUSY $U(1)_R$

$$Q_\alpha \rightarrow e^{-i\epsilon} Q_\alpha$$

$$\bar{Q}_\alpha \rightarrow e^{+i\epsilon} \bar{Q}_\alpha$$

$$[R, Q_\alpha] = -Q_\alpha$$

$$[R, \bar{Q}_\alpha] = +\bar{Q}_\alpha$$

$$Q_\alpha \sim \frac{\partial}{\partial \theta^\alpha} + \dots$$

$$R\left(\frac{\partial}{\partial \theta^\alpha}\right) = +1 = R(40^\alpha)$$

$$R(\bar{Q}_\alpha) = +1$$

\Rightarrow need

$$R(K) = 0$$

$$R(W) = +2$$

$$R(\bar{\Phi}) = r$$

$$\bar{\Phi} = \phi + \theta^\alpha \psi_\alpha + \dots$$

$$R(\phi) = r$$

$$R(\psi) = r-1$$

C. χ sf NR thm

use R-symmetry $\rightarrow Q_\alpha$ transforms in non-trivial rep of R

$N=1$ susy $U(1)_R$

$$Q_\alpha \rightarrow e^{i\alpha} Q_\alpha$$

$$\bar{Q}_\alpha \rightarrow e^{-i\alpha} \bar{Q}_\alpha$$

$$[R, Q_\alpha] = -Q_\alpha$$

$$[R, \bar{Q}_\alpha] = +\bar{Q}_\alpha$$

$$Q_\alpha \sim \frac{\partial}{\partial \theta^\alpha} + \dots$$

$$R\left(\frac{\partial}{\partial \theta^\alpha}\right) = +1 = R(40^\circ)$$

$$R(\bar{\psi}^\alpha) = +1$$

\Rightarrow need

$$R(K) = 0$$

$$R(W) = +2$$

$$R(\bar{\psi}) = r$$

$$\bar{\psi} = \psi + \theta \psi + \dots$$

$$R(\psi) = r \Leftrightarrow \psi \rightarrow e^{i\alpha} \psi$$

$$R(\psi) = r-1 \quad \psi \rightarrow e^{i\alpha} \psi$$

$\neq R$

$$\Rightarrow \text{need}$$

$$R(K) = 0$$

$$R(W) = +2$$

$$R(\bar{\Phi}) = r$$

$$\bar{\Phi} = \phi + \theta \psi + \dots$$

$$R(\phi) = r \iff \phi \rightarrow e^{i\omega t} \phi$$

$$R(\psi) = r-1 \iff \psi \rightarrow e^{i(\omega-1)t} \psi$$

$$U(1)_R$$

$$\downarrow$$

$$R$$

$$U(1)_i \leftarrow \text{non-R}$$

$$\downarrow$$

$$U_i \leftarrow \text{operators}$$

$$R' = R + \sum_i a_i U_i \quad a_i \text{ arbitrary}$$

(Weinberg III)

'UV' theory @ scale μ_0 $\mathcal{W}_{\mu_0} = \mathcal{W}_{\mu_0}(\Phi_n)$

effective sup'd't @ scale $\mu < \mu_0$ $\mathcal{W}_\mu = \mathcal{W}_\mu(\Phi_n)$

(Weinberg III)

'UV' theory
@ scale μ_0

$$\mathcal{W}_{\mu_0} = \mathcal{W}_{\mu_0}(\Phi_n)$$

write as

$$= \gamma \mathcal{W}_{\mu_0}(\Phi_n)$$

	$U(1)_R$
γ	+2
Φ_n	0

effective superpotential
@ scale $\mu < \mu_0$

$$\mathcal{W}_\mu = \frac{1}{\gamma} g(\Phi_n) \\ = \gamma g(\Phi_n)$$

assume some
set of light def.

let $\gamma \rightarrow 0$ becomes free \Rightarrow

$$g(\Phi_n) = \mathcal{W}_{\mu_0}(\Phi_n)$$

(up to conventional scalings)

set $\gamma = 1 \Rightarrow \mathcal{W}_\mu = \mathcal{W}_{\mu_0}(\Phi_n)$

$$\mathcal{I}_\mu = \int d^4\theta z_n \phi_n \bar{\phi}_n + \int d^4x \left(\frac{\mu}{\mu_0} \right)^{3-\Delta} \lambda \theta$$

$$\sigma_{N_\mu} = \lambda \theta, \quad \theta = \prod_n \phi_n^{\zeta_n} \quad \Delta = \sum_n \zeta_n$$

$$\sigma_{N_\mu} = \left(\frac{\mu}{\mu_0} \right)^{3-\Delta} \lambda \theta$$

$$\mathcal{I}_\mu = \int d^4\theta \, z_n \phi_n \bar{\phi}_n + \int d^4\theta \, \left(\frac{\mu_0}{\mu}\right)^{3-\Delta} \lambda \theta$$

rescale $\phi_n \rightarrow \phi_n^{cw} = \sqrt{z_n} \phi_n$ $\theta = \prod_n \phi_n^{r_n}$

$$d\mathcal{V}_\mu = \left(\frac{\mu_0}{\mu}\right)^{3-\Delta} \left(\prod_n z_n^{-r_n/2}\right) \lambda_r \theta^{cw}$$

$$\frac{d\lambda_r^{cw}}{\lambda_r^{cw}} = \lambda_r^{cw} \left(\Delta - 3 - \frac{1}{2} \sum_n r_n \gamma_n \right) \quad \gamma_n \equiv \frac{d \ln z_n}{d \ln \mu}$$

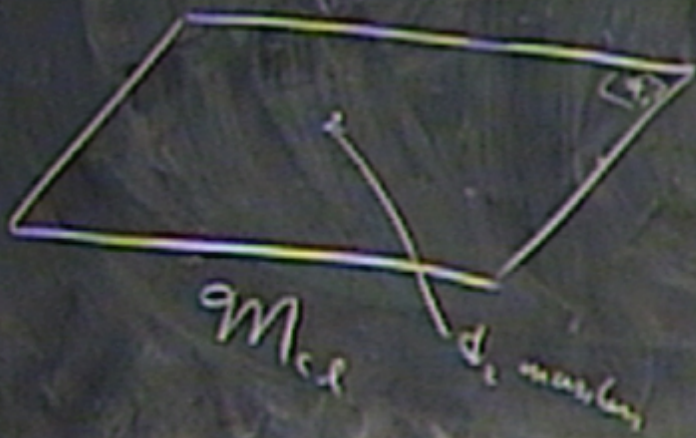
An example theory w/ 2 chfs

$$\phi_2 = 0 \quad \phi_1 \phi_2 = 0$$

• UV: $\mathcal{L}_{\mu_0} = \bar{\phi}_1 \phi_1 + \bar{\phi}_2 \phi_2$

$$\mathcal{W}_{\mu_0} = \lambda \phi_1 \phi_2^2$$

• classically: \Rightarrow moduli space $\mathcal{M}_{cl} = \{ \phi_2 = 0, \phi_1 \text{ arbitrary} = \mathbb{C} \}$
(of vacua)



• everywhere ϕ_1 massless
 ϕ_2 has mass $= |\lambda \phi_1|$

• Flow IR $\mu < \mu_0$

NR in $g_{\mu} = g_{\mu_0} \Rightarrow g_{\mu} \approx \mathcal{O}$

• 1-loop renormalization of \mathcal{K}_μ

for $\mu < \text{max } \phi_2 < \mu_0$



$$\mathcal{K}_\mu = \bar{\phi}_1 \phi_1 - (\#) \phi_1 \bar{\phi}_1 |\lambda|^2 \ln \left| \frac{\phi_1}{\mu_0} \right|^2$$

$$ds_\mu^2 = \left(\partial_\mu \partial_{\bar{\mu}} \mathcal{K}_\mu \right) d\phi_1 d\bar{\phi}_1$$



\mathcal{M}_{qu}

let $\gamma \rightarrow 0$ becomes free $\Rightarrow g(\underline{\Phi}_n) = \mathcal{W}_\mu(\underline{\Phi}_n)$

set $\gamma=1 \Rightarrow \mathcal{W}_\mu = \mathcal{W}_\mu(\underline{\Phi}_n)$