

Title: Astrobiology, homochirality, and the origin of life

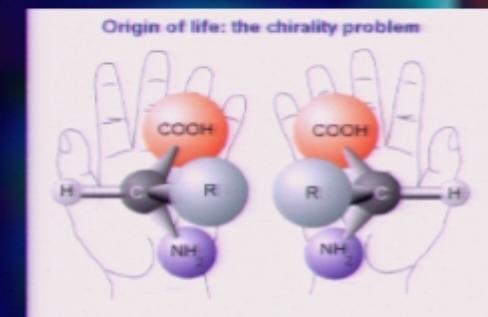
Date: Jun 01, 2005 02:00 PM

URL: <http://pirsa.org/05060055>

Abstract:

Astrobiology, homochirality, and the origin of life

Axel Brandenburg, Anja Andersen, Susanne Höfner, Martin Nilsson, Tuomas Multamäki
(all at *Nordita*)



If prerequisite for life:

Due to polarized light, electroweak force, magnetic fields, ...

If consequence of life:

Must have emerged during polymerization of first replicating molecules

Different at different places on Earth??

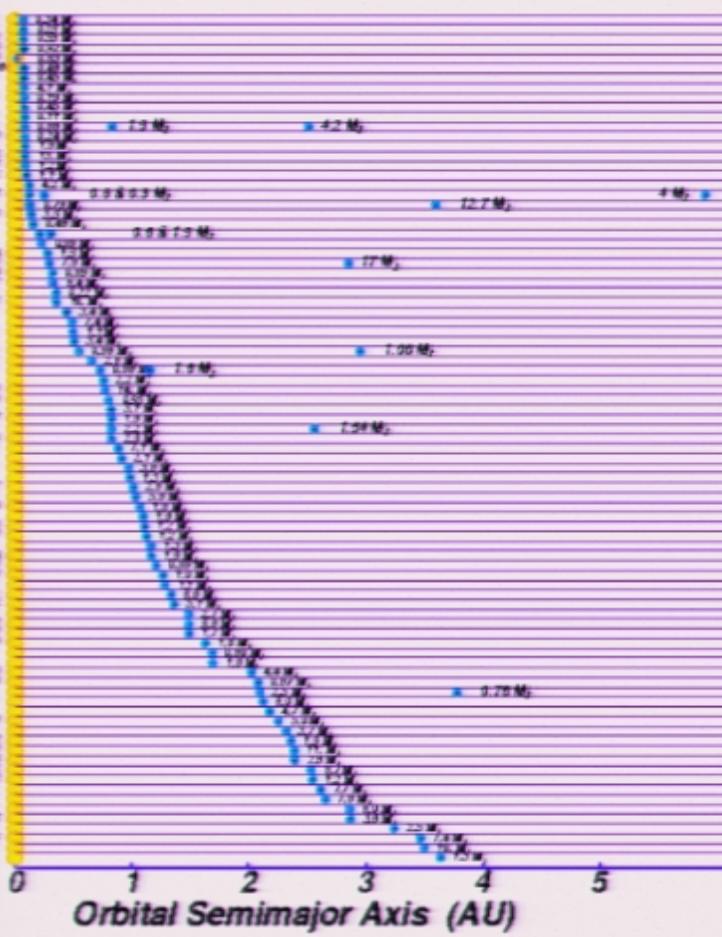
Reaction-diffusion-advection equation



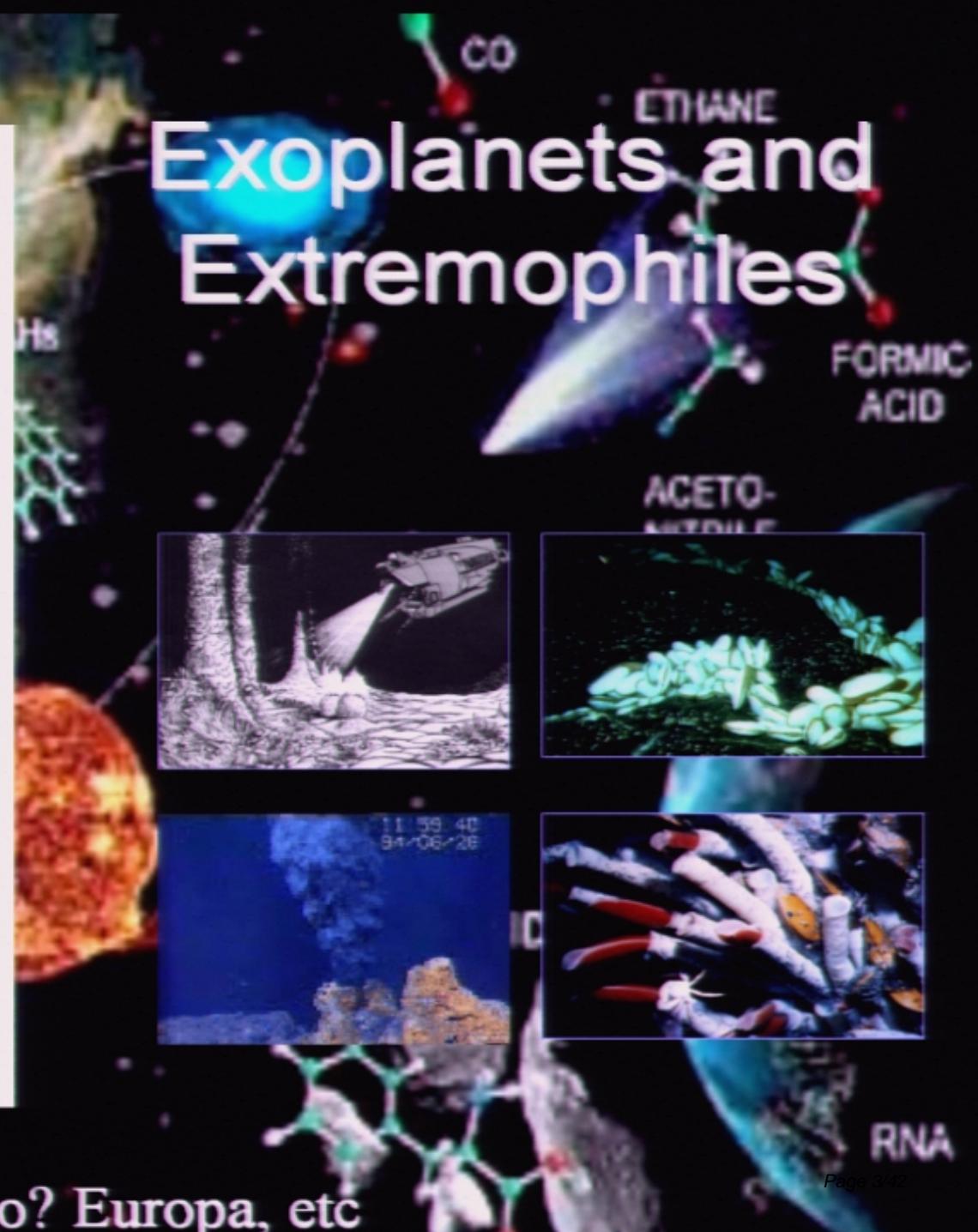
Int. J. Astrobiol. 3, 209 (2004), Orig. Life Evol. Biosph. (2005a,b),
Astrobiology (2005), see also arXiv: q-bio.BM/0401036, 0502008,

0505040, 0505041

By now 136 exoplanets...



Exoplanets and Extremophiles



time line

10^{10} yr after Big Bang

10^4 yr

10^6 yr

10^8 yr

10^9 yr

planetesimals

Sun ignites

All gas gone
Remaining dust settles

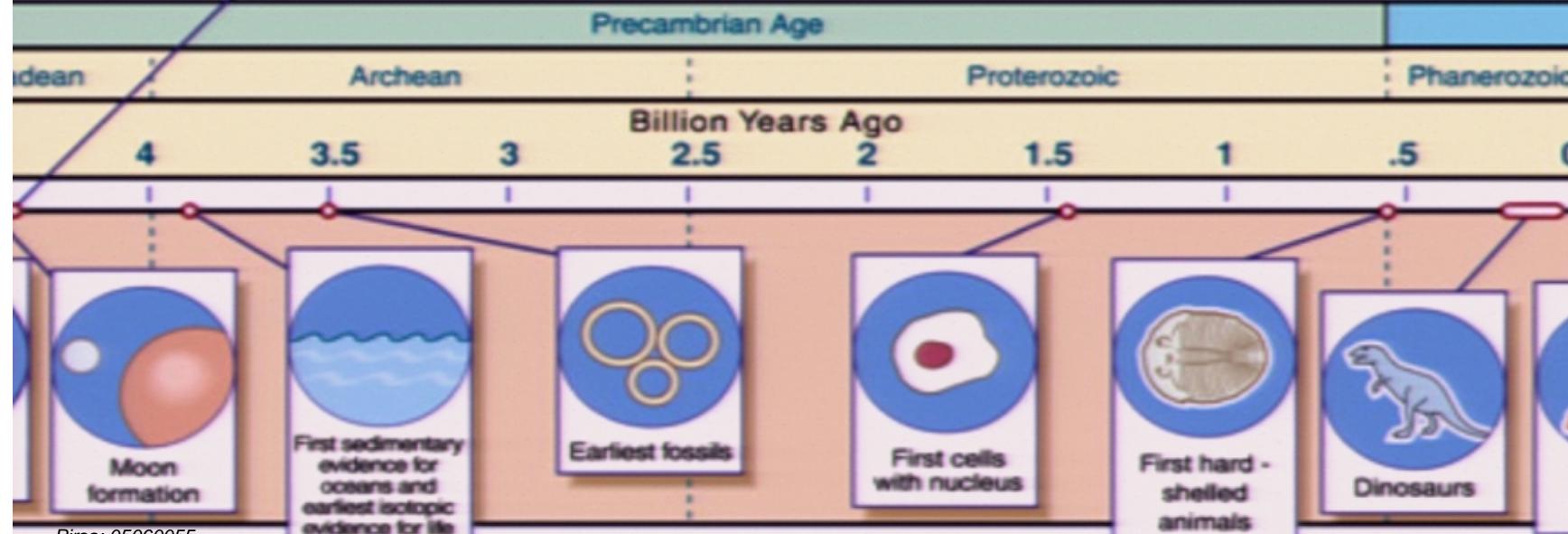
First life

Oldest Crystal Tells Story Of A Habitable Early Earth

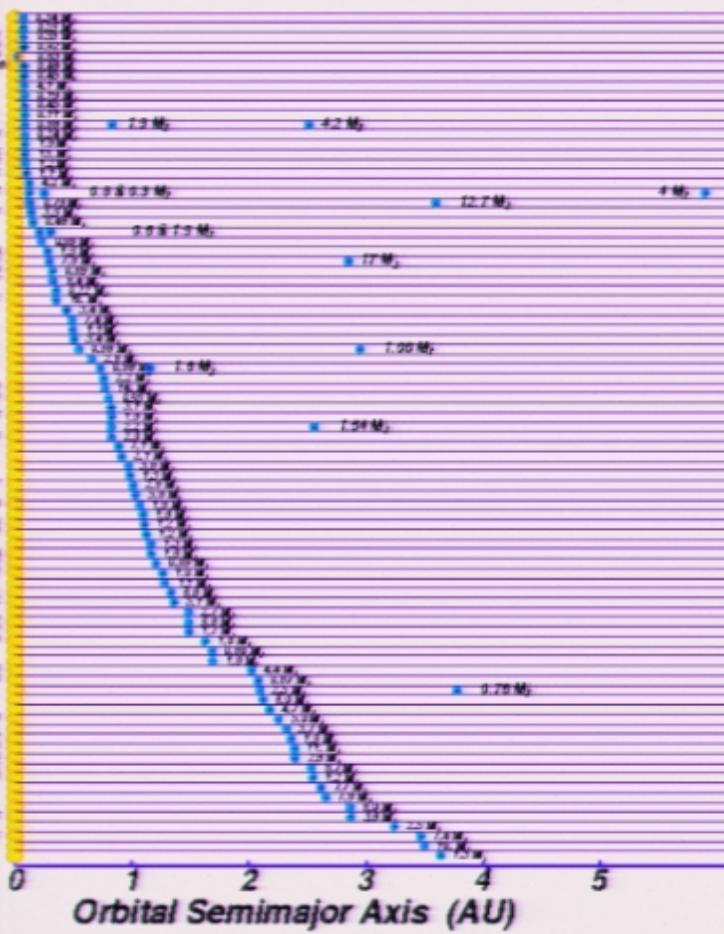


Newly discovered zircon crystal is the world's oldest known sample of a terrestrial material at an estimated 4.4 billion years old.

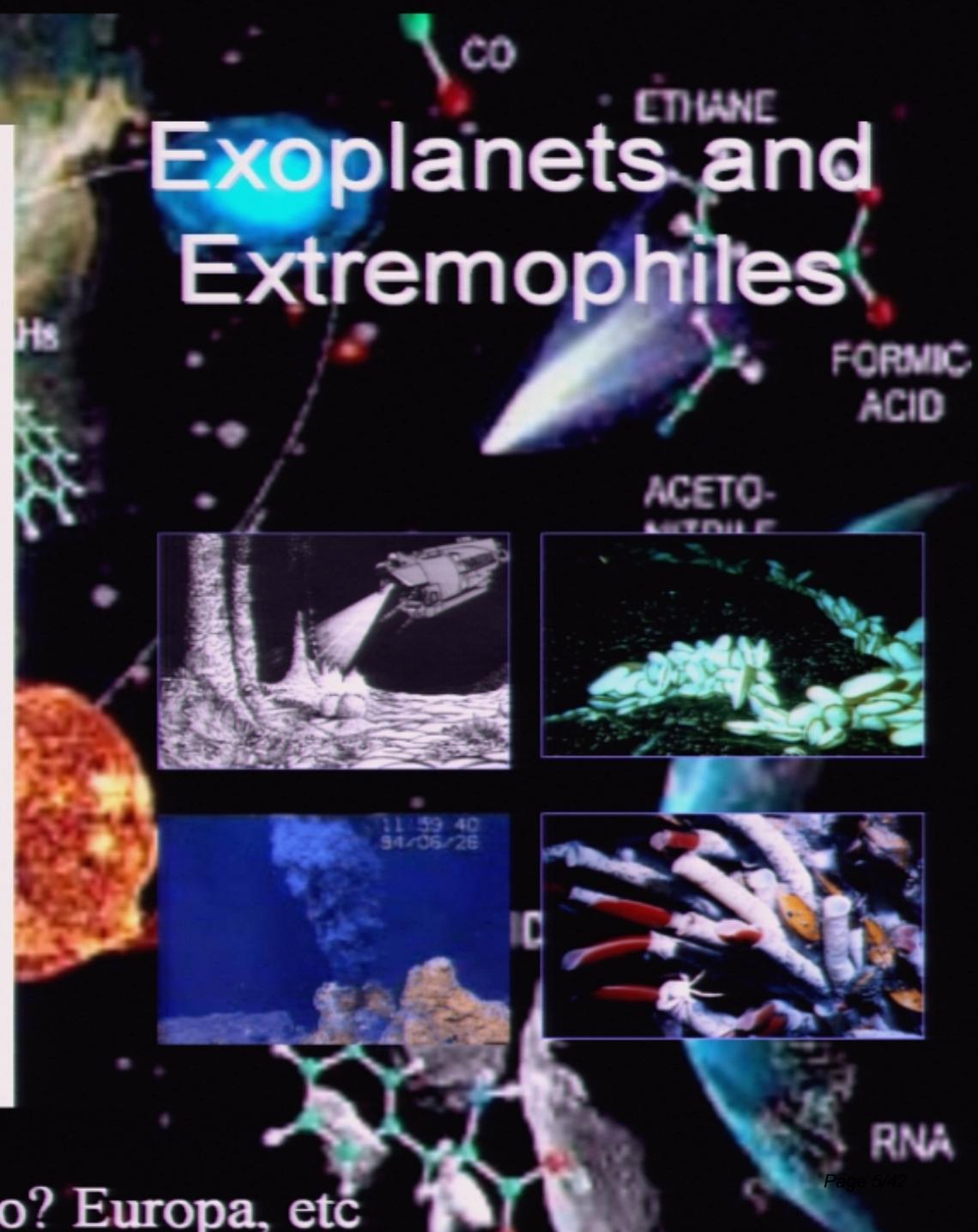
Chemical and isotopic analysis of this crystal suggest the presence of rocks formed at low temperatures and that the infant Earth cooled much faster after formation of the Moon than previously believed. Instead of being covered by a magma ocean, as conventional wisdom holds, the early Earth may instead have had continents, oceans, and a hydrosphere, all the conditions necessary for life.



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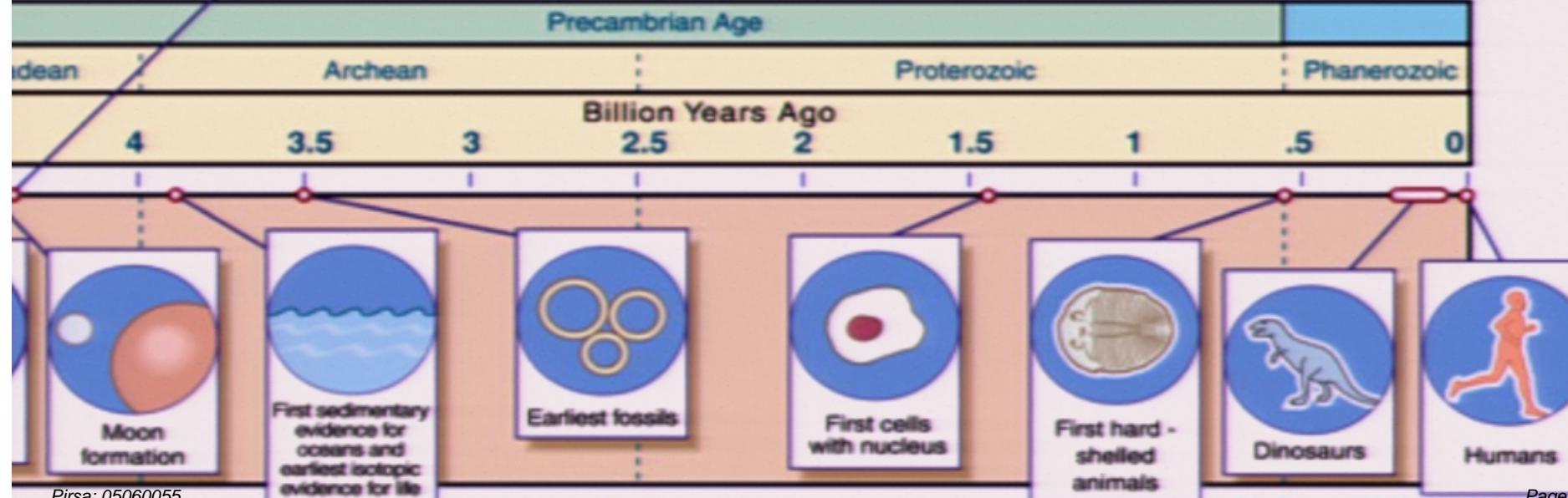
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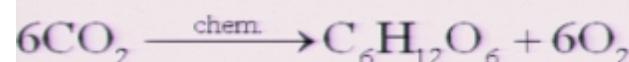
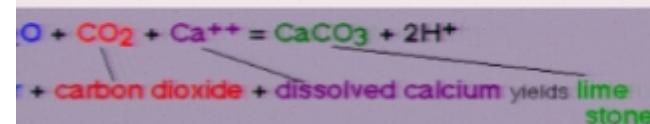
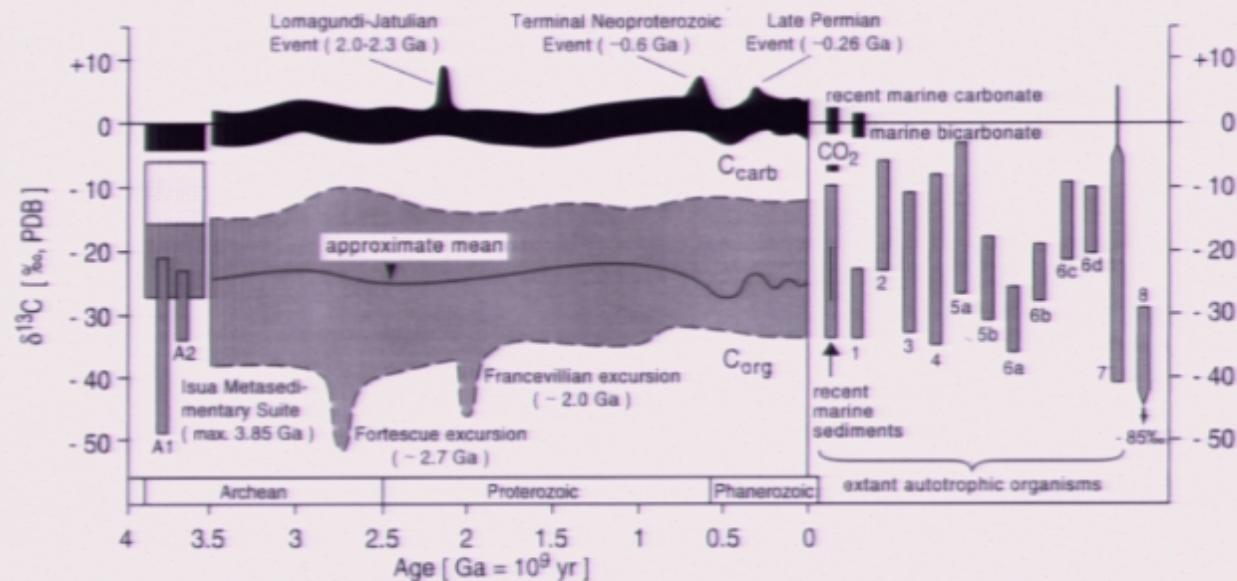
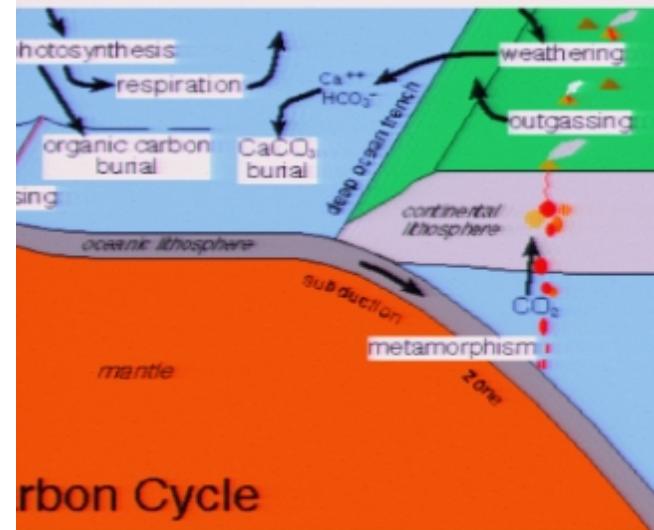
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Carbon cycle → Source of the atmosphere

Carbonic and organic carbon



Photosynthesis: 270 mW/m²
 Internal energy: 80 mW/m²
 Solar energy: 340 W/m²

(Rosing 2005)

time line

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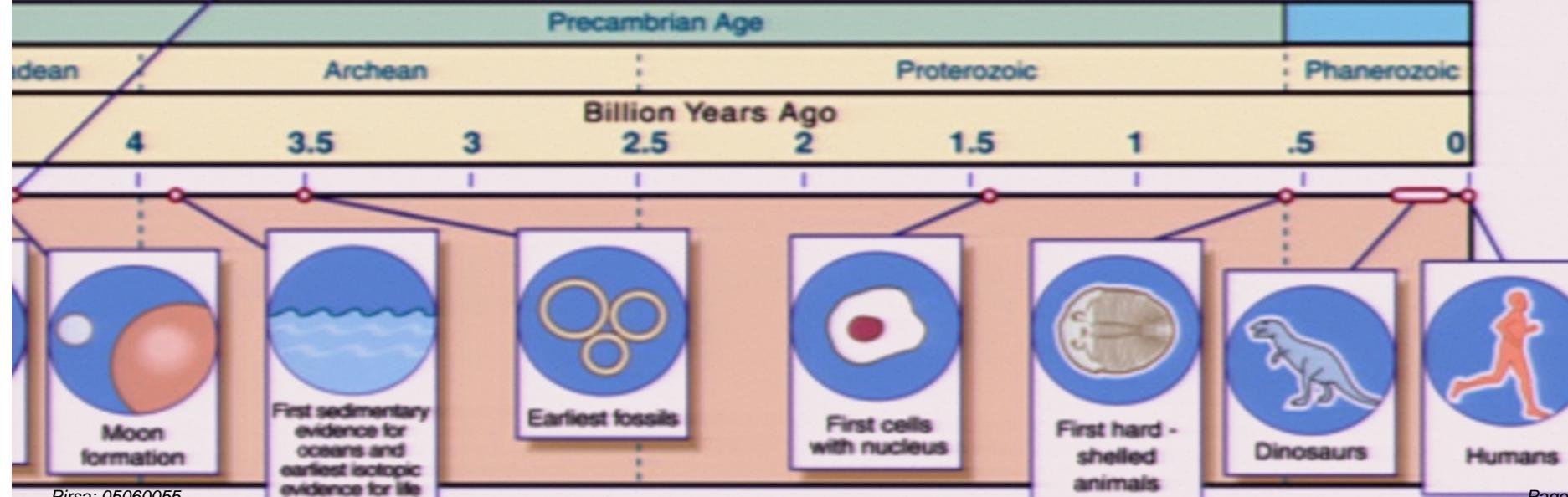
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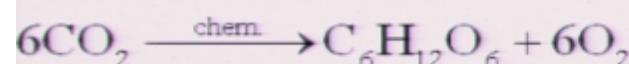
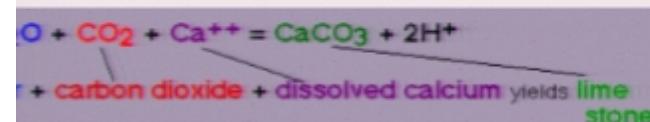
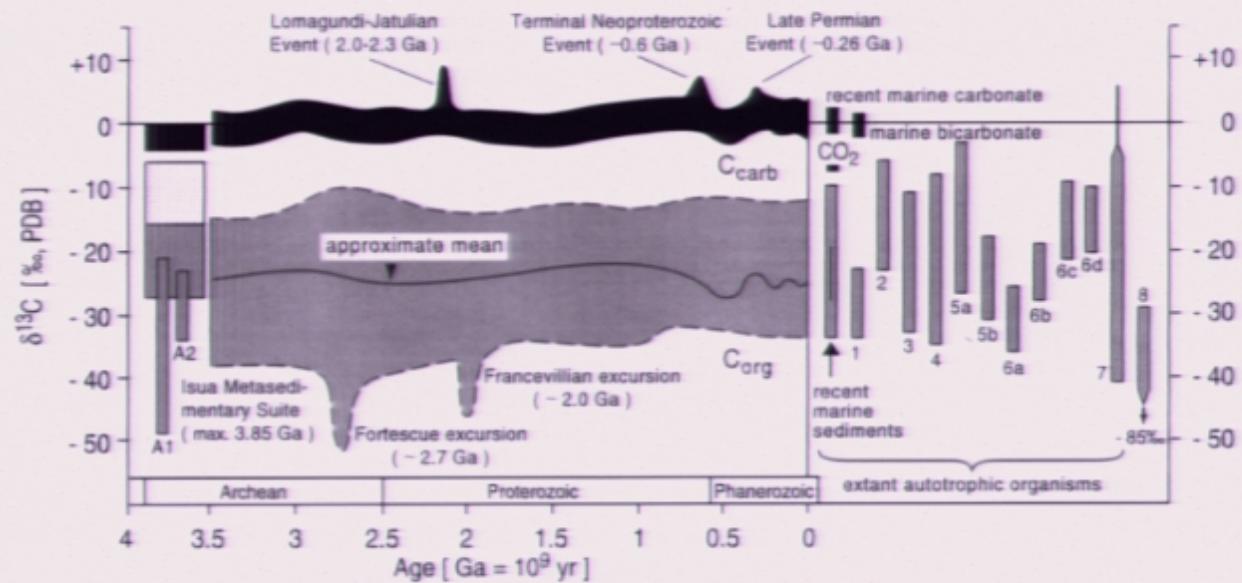
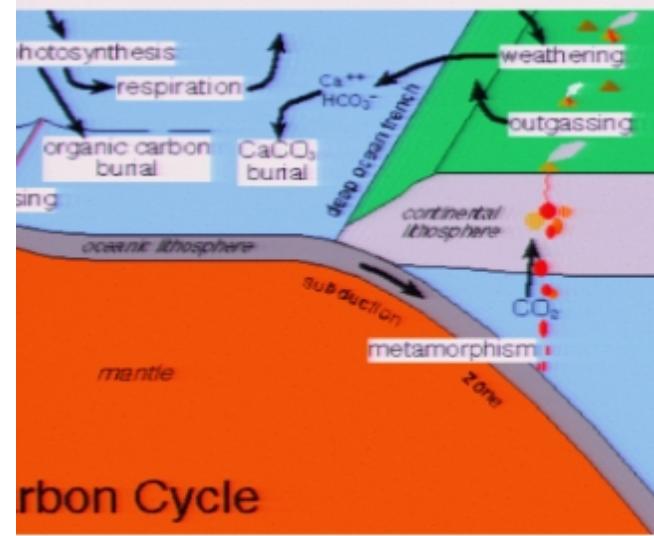
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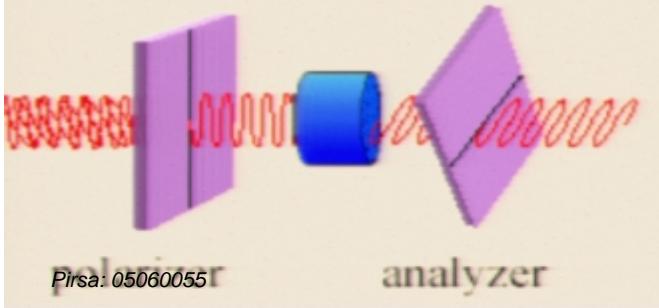
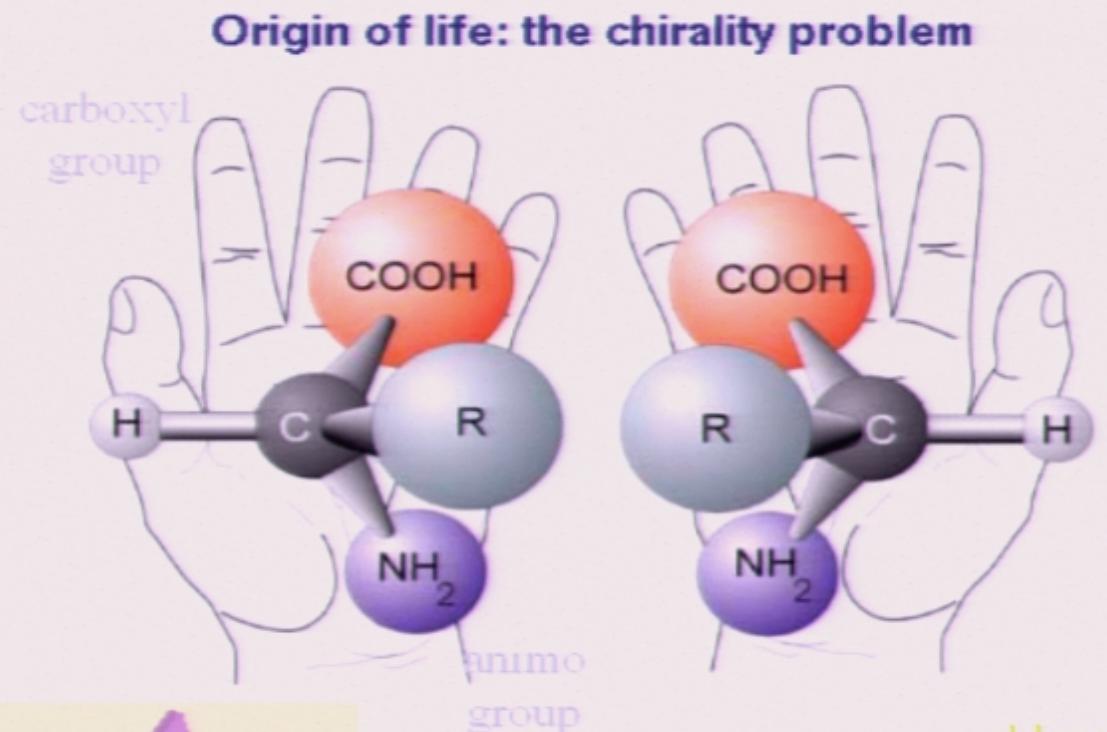
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(Rosing 2005)

Aminoacids in proteins: left-handed sugars in DNA and RNA: right-handed

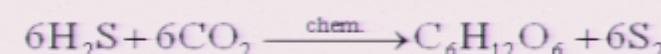


(Pasteur
1822-1895)



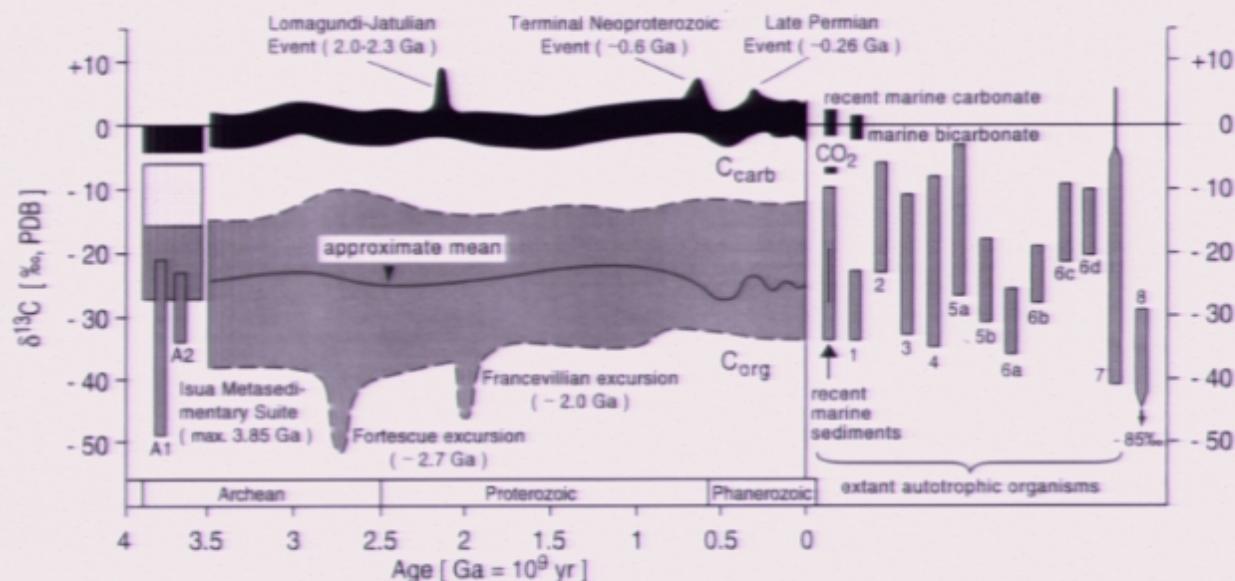
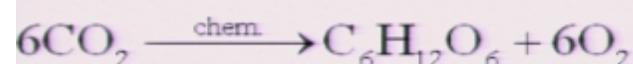
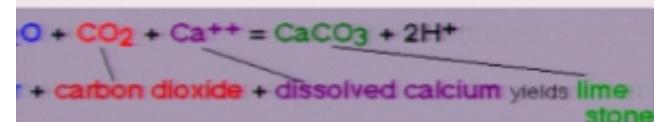
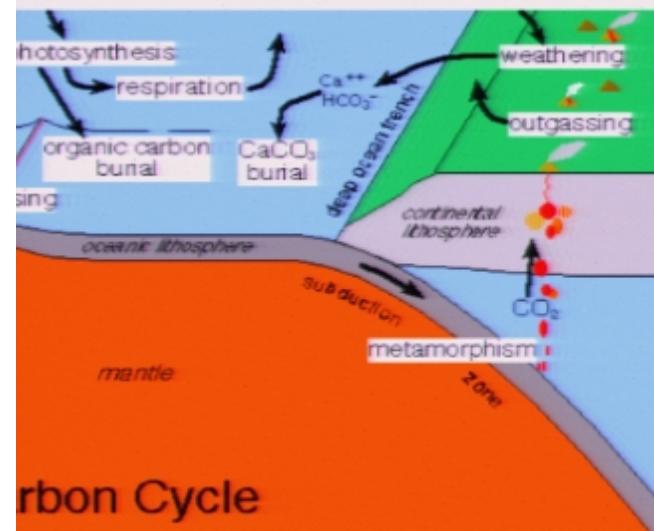
Presa: 05060055

Is chirality:
(i) prerequisite
(ii) consequence
of life?



Carbon cycle → Source of the atmosphere

Carbonic and organic carbon



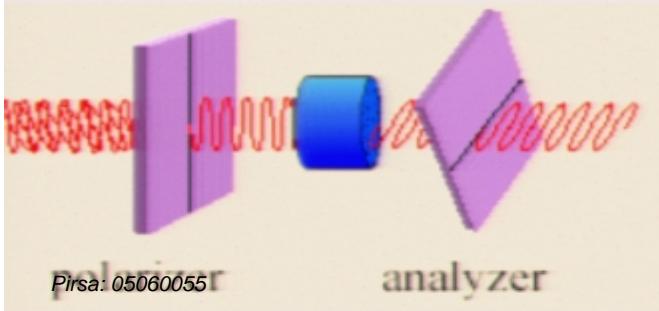
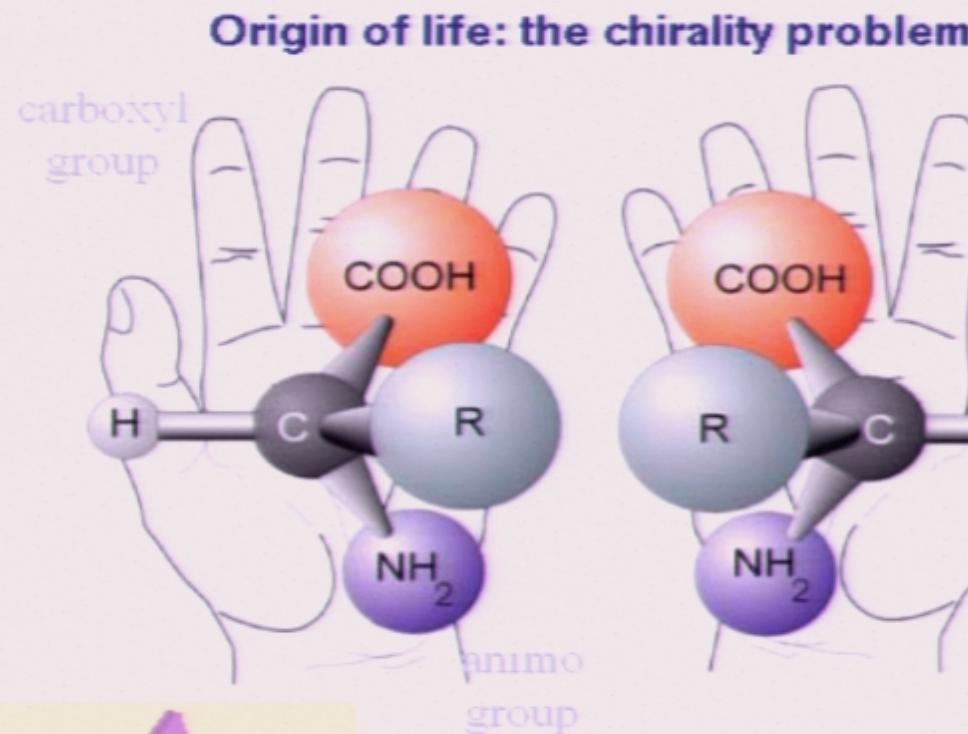
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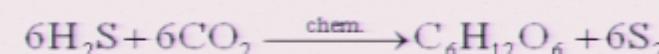


Louis Pasteur
(1822-1895)



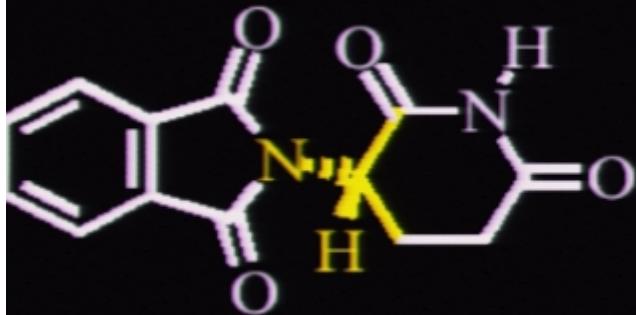
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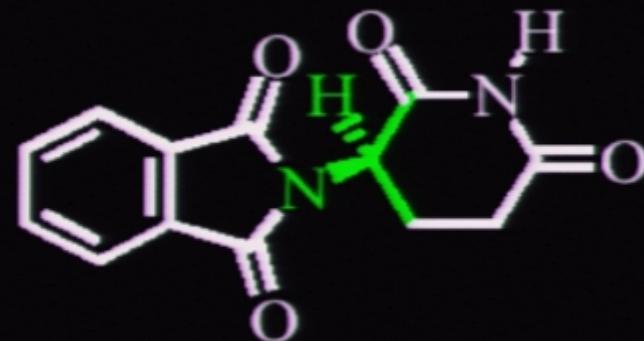


Contergan: was sold as racemic mixture

Thalidomide molecule



(S)-Contergan



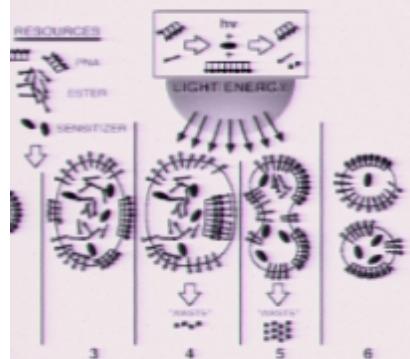
(R)-Contergan

causes misformations
abandoned in December 1961)

Cures morning sickness
during pregnancy

Chirality selection during polymerization of the first replicating molecule?

lipid world



Assen et al (2003)
Life 9, 269-316)

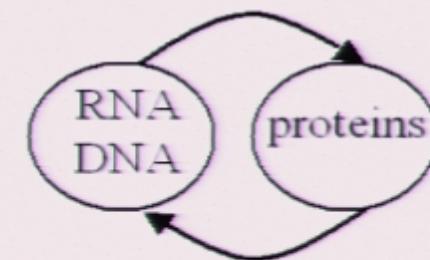
PNA world

achiral
↓
chiral

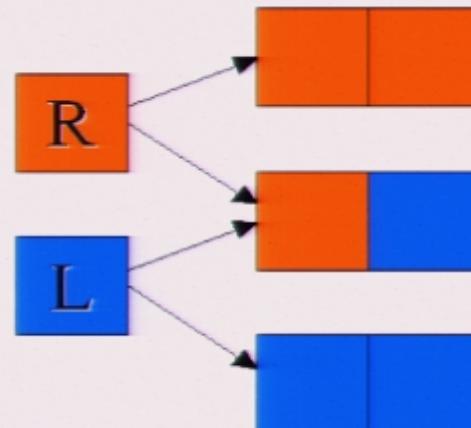
RNA world



dual world

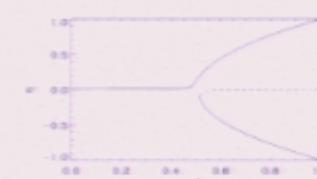


rization



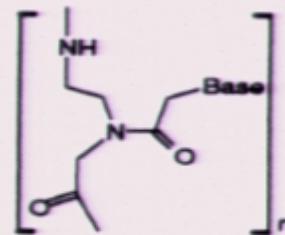
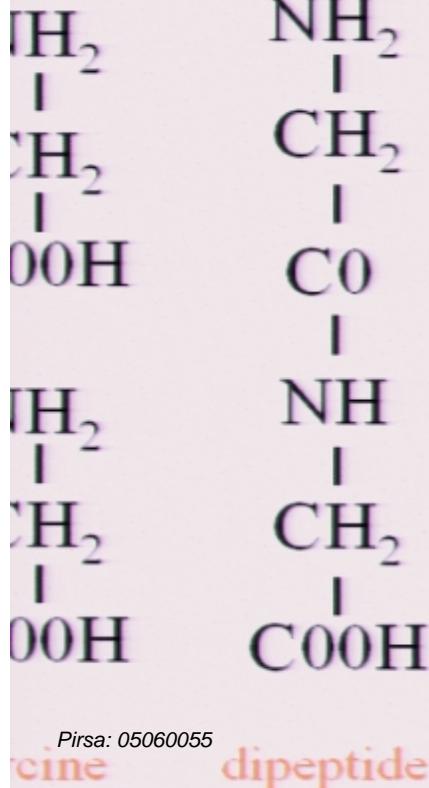
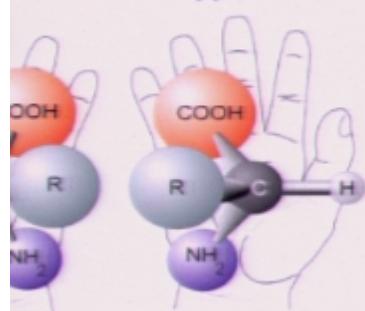
Isotactic polymer
(same chirality)

“waste”
(enantiomeric
cross-inhibition)

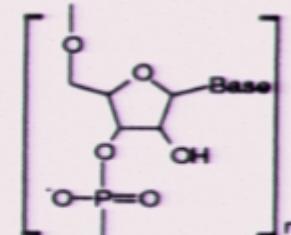


PNA world prior to RNA world

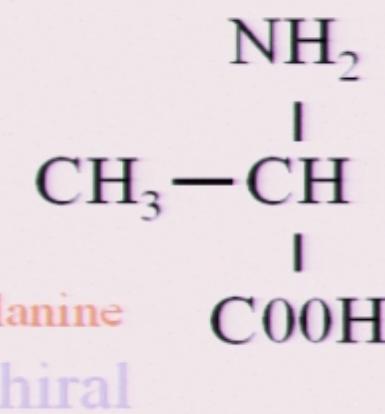
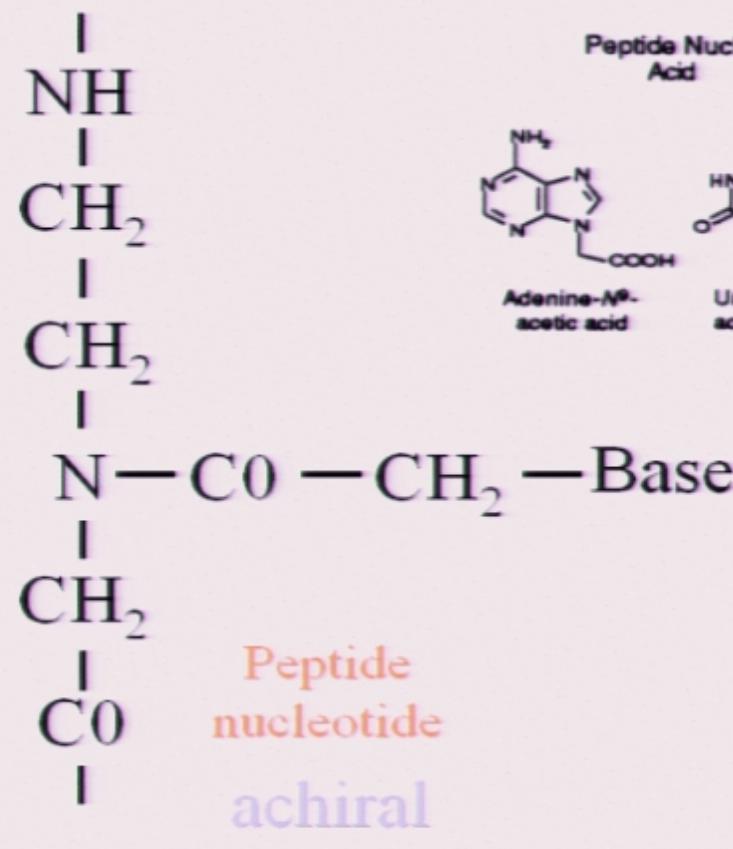
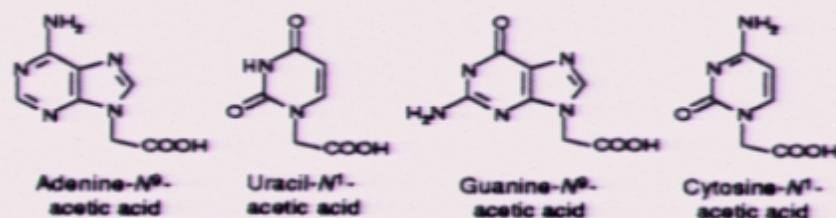
of life: the chirality problem



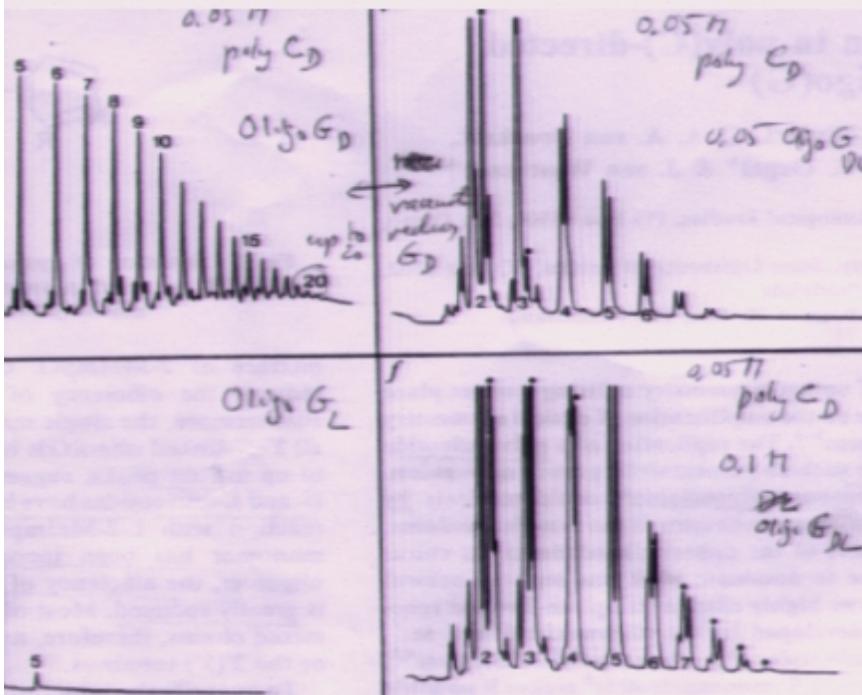
Peptide Nucleic Acid



RNA



Relevant experiments: nucleotides



→ Mononucleotides with wrong chirality terminate chain growth

— — ok

— — blue poisoned

ate-directed oligomerization
poly (C_D) → oligo (G_D)

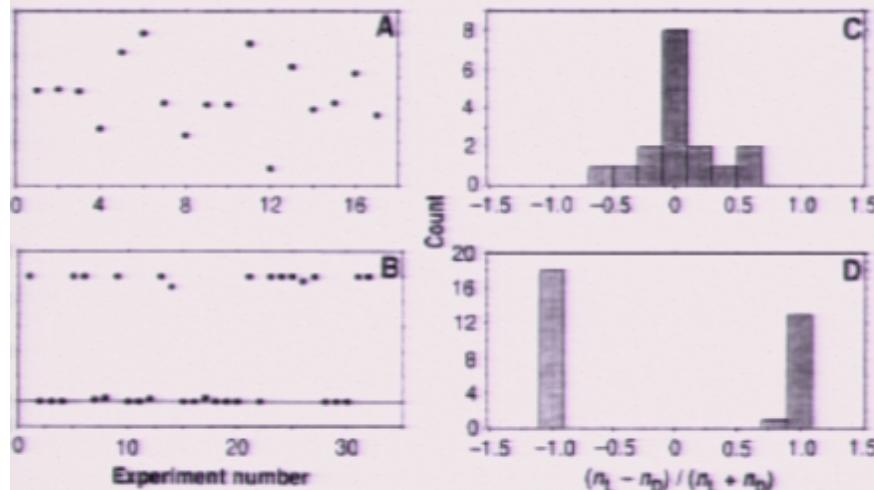
cytosine ↑

guanine ↑

(using HPLC)

→ *enantiomeric cross-inhibition*

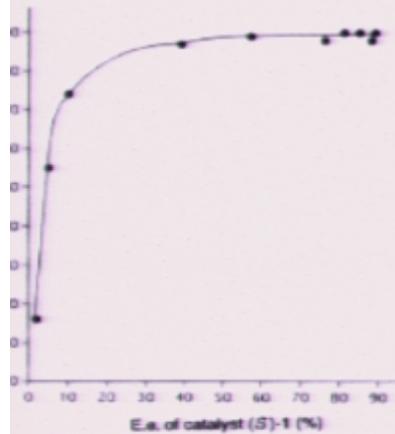
Relevant experiments: crystals



Crystal growth, many different nucleation sites: racemic mixture

Crystal growth with stirring:
primary nucleation suppressed

Kondepudi et al. (1990)



Alkanol with 2% e.e.
treated with
carboxylaldehyde
Soai et al. (1995)

→ competition important

→ autocatalytic
self-amplification

Frank (1953),
Goldanskii & Kuzmin
(1989), ...

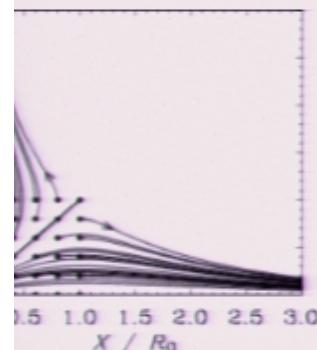
Simplistic models: trial and error?

Rank (1953)

$$= kR(1-L)$$

$$= kL(1-R)$$

st anti-catalyst



Ito & Hyuga (2004)

$$= kR^2(1-R-L) - \mu R$$

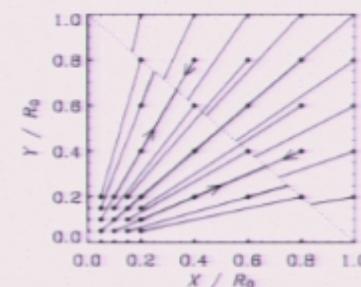
$$= kL^2(1-R-L) - \mu L$$

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unspecific quenching

$$\frac{dR}{dt} = kR(1-R-L)$$

$$\frac{dL}{dt} = kL(1-R-L)$$

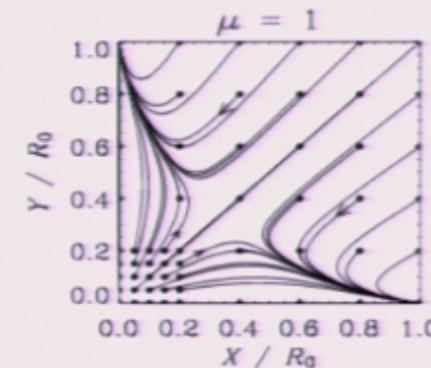


chemically unrealistic (Blackmond 2002)

Specific antagonism

$$\frac{dR}{dt} = kR(1-R-L) - \mu RL$$

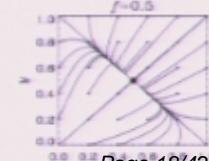
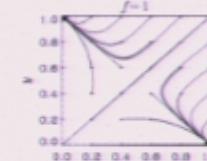
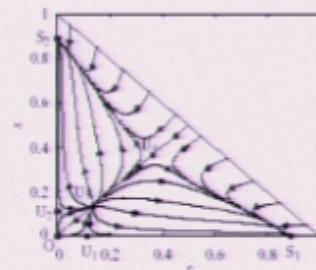
$$\frac{dL}{dt} = kL(1-R-L) - \mu RL$$



our model (BAHN 2005)

$$\frac{dR}{dt} = \frac{R^2}{R^2 + L^2} - (L+R)R$$

$$\frac{dL}{dt} = \frac{L^2}{R^2 + L^2} - (L+R)L$$



Polymerization model of Sandars

Orig. Life Evol. Biosph. **33**, 575-587 (2003)

ction for left-handed monomers



$$\frac{d}{dt}[L_n] = 2k_S[L_{n-1}][L_1]$$

Combined equations

$$\frac{d}{dt}[L_n] = -2k_S[L_1]([L_n] - [L_{n-1}])$$

$$\frac{d}{dt}[L_1] = Q_L - 2k_S[L_1] \sum_{n=1}^N [L_n]$$

s term for each constituent

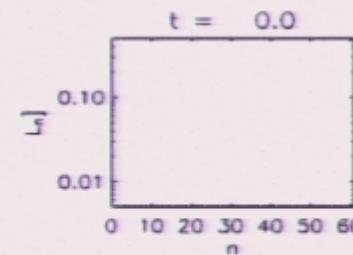
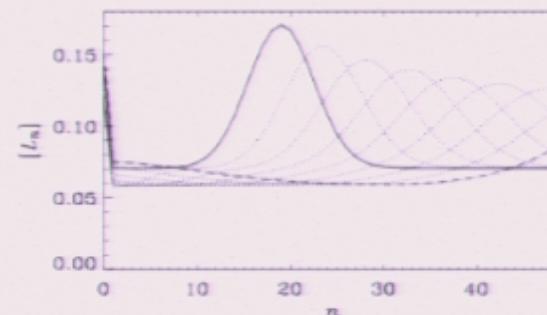
$$\frac{d}{dt}[L_{n-1}] = -2k_S[L_{n-1}][L_1]$$

$$\frac{d}{dt}[L_1] = -2k_S[L_{n-1}][L_1]$$

umber of left-handed
Building blocks const

$$\sum_{n=1}^N n[L_n] = \text{const}$$

(if $Q_L = 0$)



Coupling to substrate S

source of L_1 monomers \mathcal{Q}_L

$$\frac{d}{dt}[L_1] = \mathcal{Q}_L - 2k_s[L_1] \sum_{n=1}^N [L_n]$$

acts as a sink of substrate S

$$\frac{d}{dt}[S] = Q - (\mathcal{Q}_L + \mathcal{Q}_R)$$

-catalytic properties of polymers

$$\mathcal{Q}_L = k_c[S][L_N]$$

$$\mathcal{Q}_R = k_c[S][R_N]$$

Refinements

finite fidelity f

$$\mathcal{Q}_L = k_c[S] \left\{ \frac{1}{2}(1+f)C_L + \frac{1}{2}(1-f)C_R \right\}$$

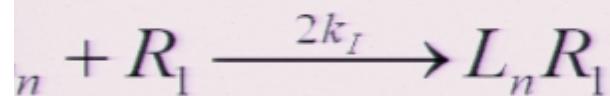
$$\mathcal{Q}_R = k_c[S] \left\{ \frac{1}{2}(1+f)C_R + \frac{1}{2}(1-f)C_L \right\}$$

Possible proposals for C_L (or C_R)

$$C_L = \begin{cases} [L_N] & \text{(Sandars 2003)} \\ [L_M] & \text{(another possibility)} \\ \sum [L_n] & \text{(Wattis \& Coveney 2004)} \\ \sum n[L_n] & \text{(our work)} \end{cases}$$

Including enantiomeric cross-inhibition

loss term for each constituent



$$\frac{d}{dt}[L_n R_1] = 2k_I [L_n][R_1]$$

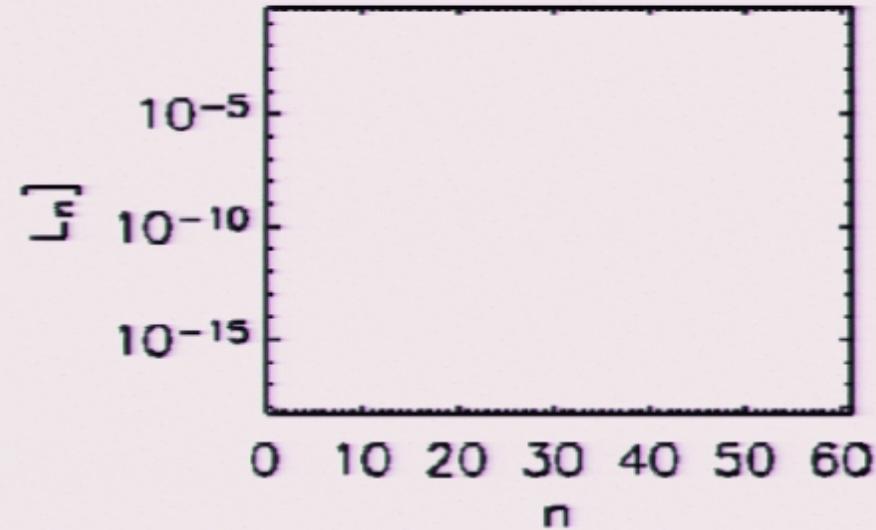
$$[L_n R_1] = \dots - 2k_I [R_1] ([L_n R_1] - [L_{n-1} R_1])$$

$$[R_1] = \dots - 2k_S [R_1] \sum_{n=1}^N [L_n R_1]$$

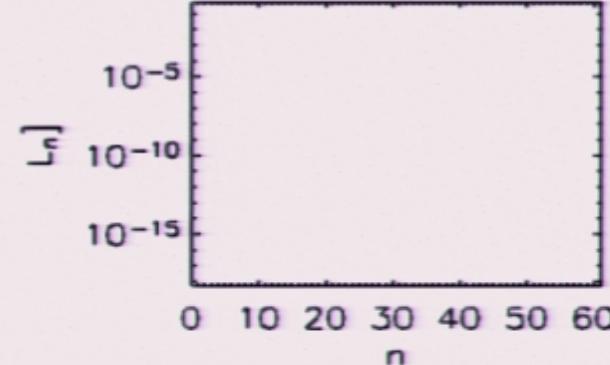
Stability

Racemic solution $\sim 2^{1-n}$

$$t = 0.0$$

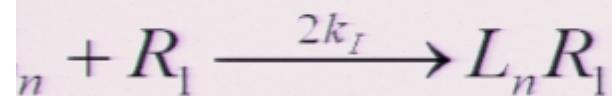


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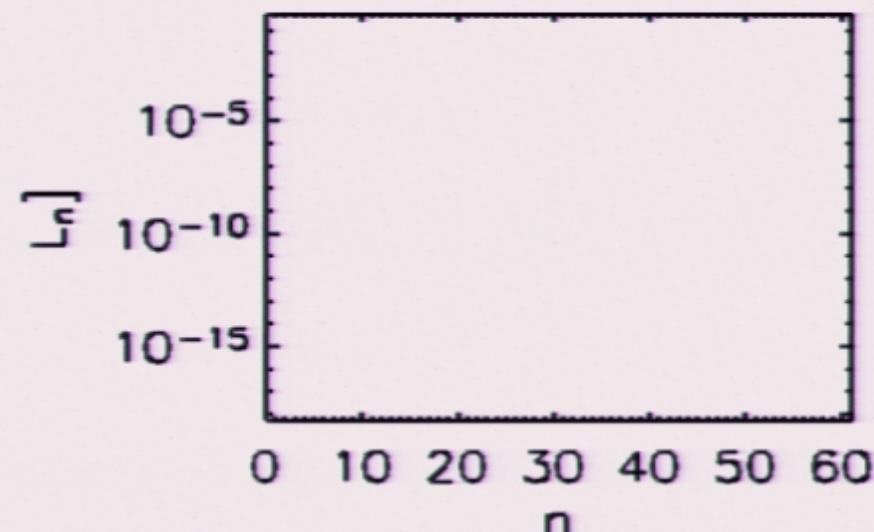
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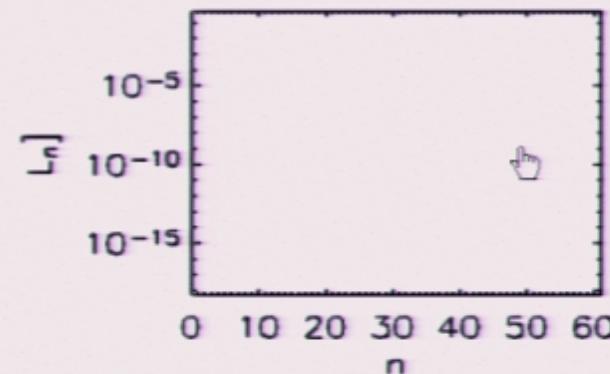
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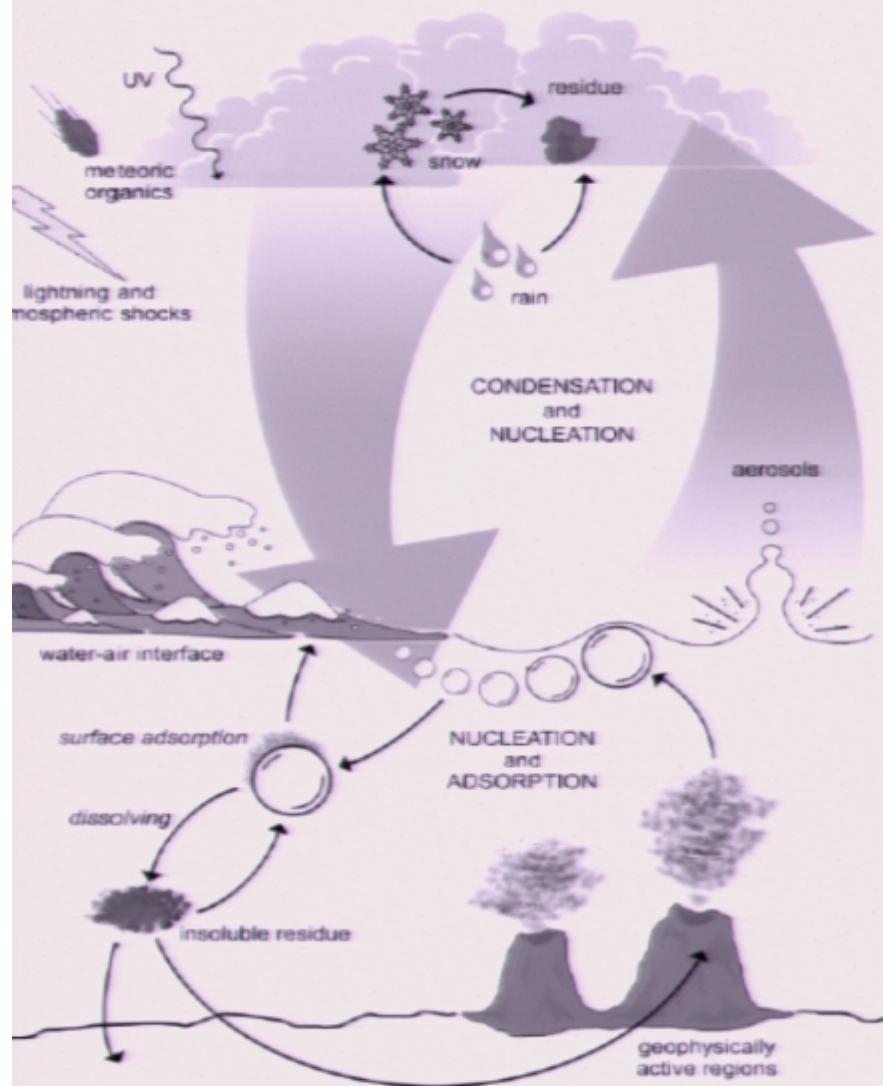
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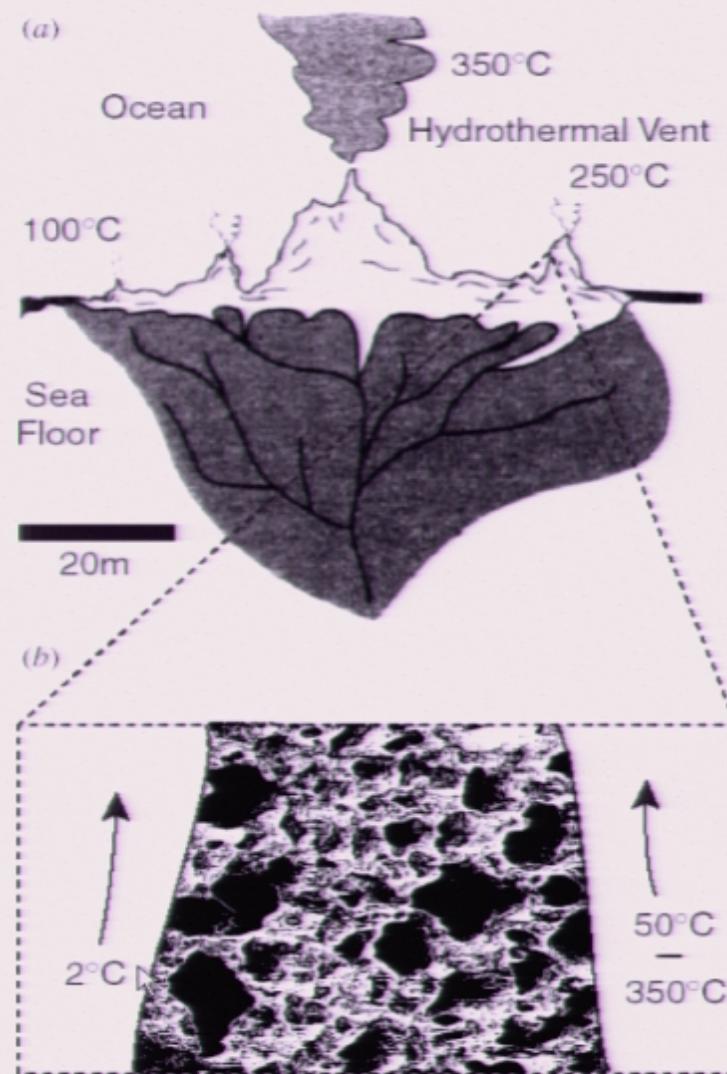


Where could this work?



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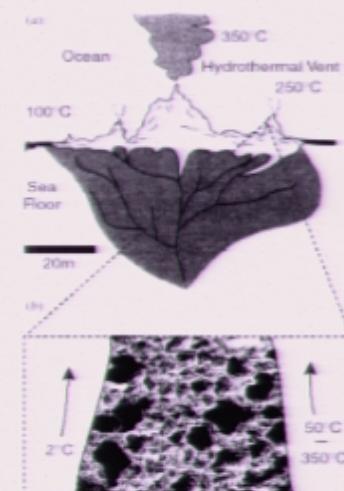
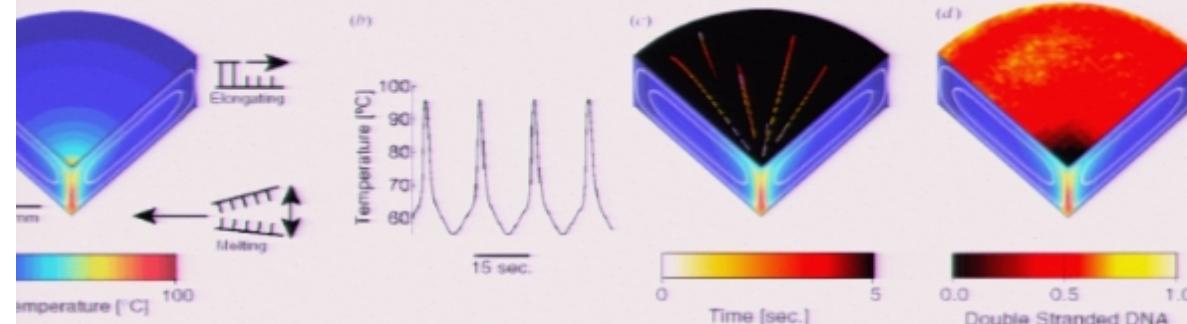
Lerman & Teng (2005)



Braun & Libchaber (2003)

How to drive the system?

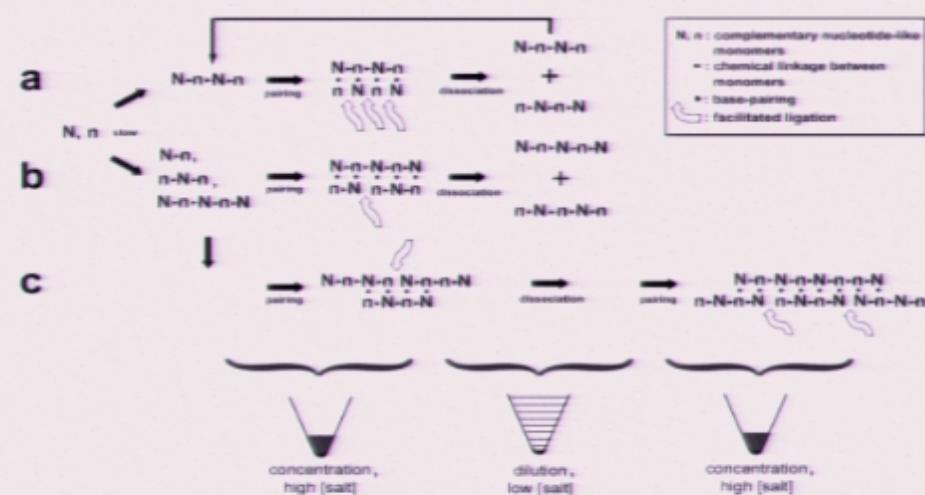
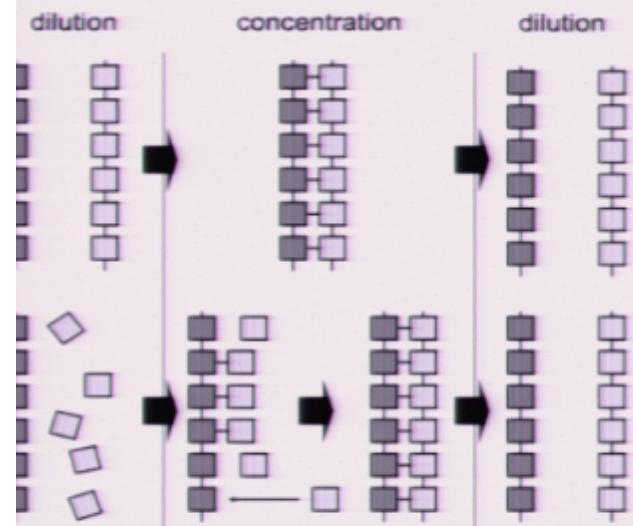
Hydrothermal vents



Braun & Libchaber (2003)

des

[NaCl]



Reduced equations

2-mode reduction

$$\frac{d}{dt}[L_2] = -2k_s[L_1]([L_2] - [L_1])$$

Adiabatic elimination
rapidly adjusting variables

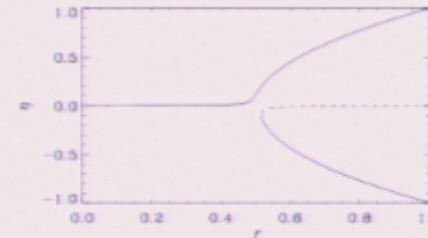
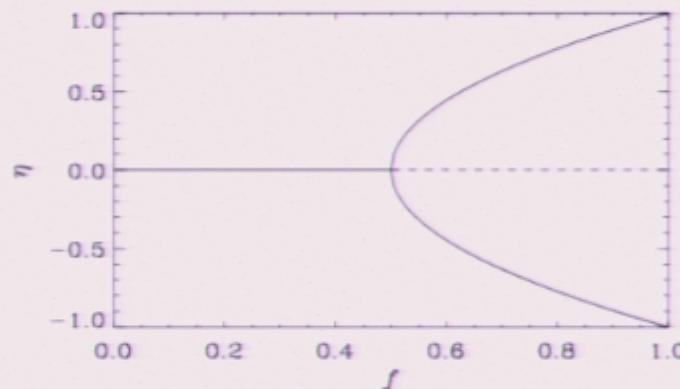
$$0 = -2k_s[L_1]([L_2] - [L_1])$$

Quantitatively close to full model

$$\frac{dR}{dt} = \frac{R^2}{R^2 + L^2} - (L + R)R$$

$$\frac{dL}{dt} = \frac{L^2}{R^2 + L^2} - (L + R)L$$

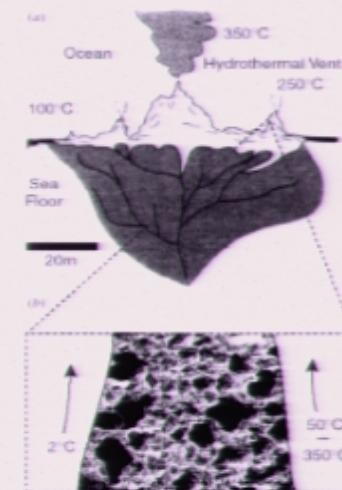
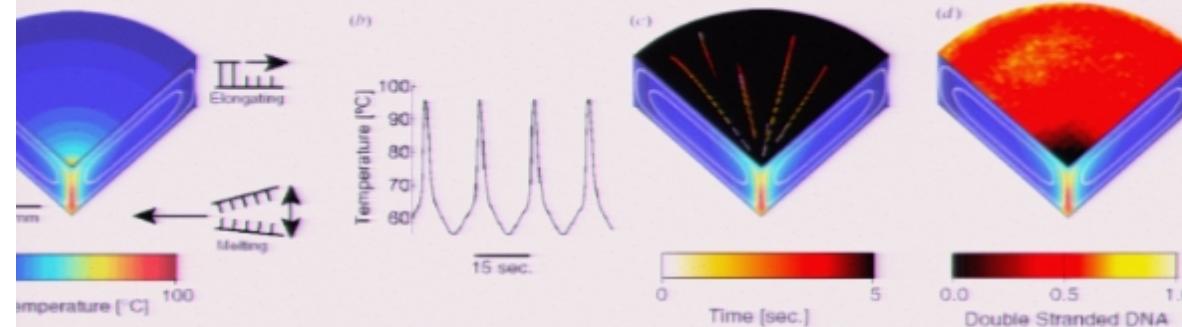
BAHN (Orig. Life Evol. Biosph. 2005)



Initial bias →

How to drive the system?

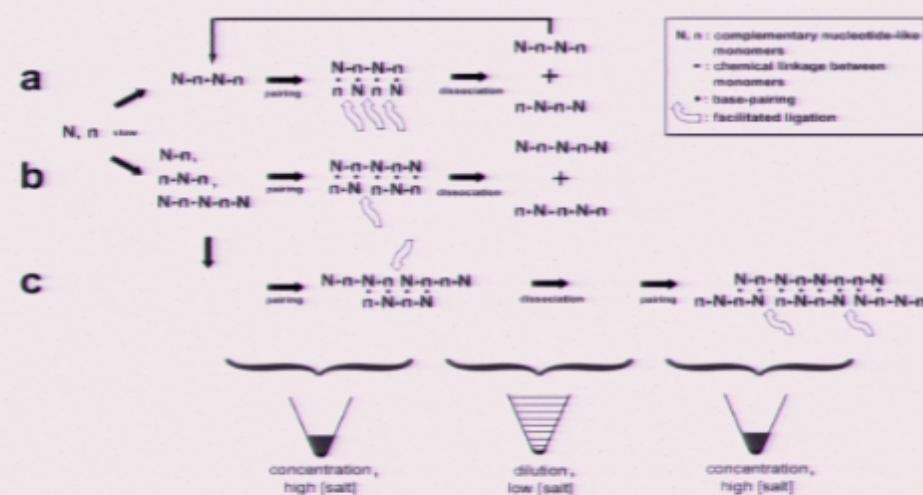
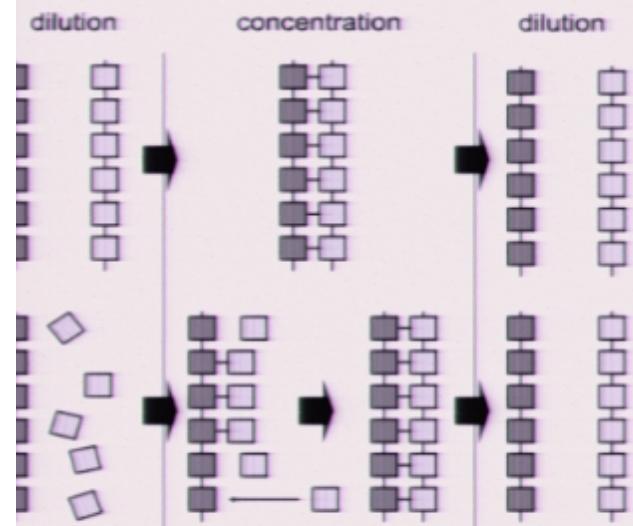
Hydrothermal vents



Braun & Libchaber (2003)

des

[NaCl]



Reduced equations

2-mode reduction

$$\frac{d}{dt}[L_2] = -2k_s[L_1]([L_2] - [L_1])$$

Adiabatic elimination
rapidly adjusting variables

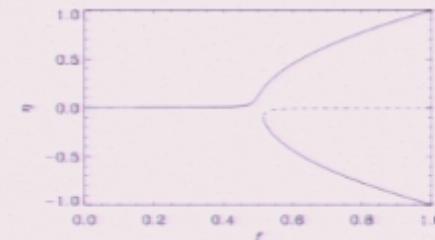
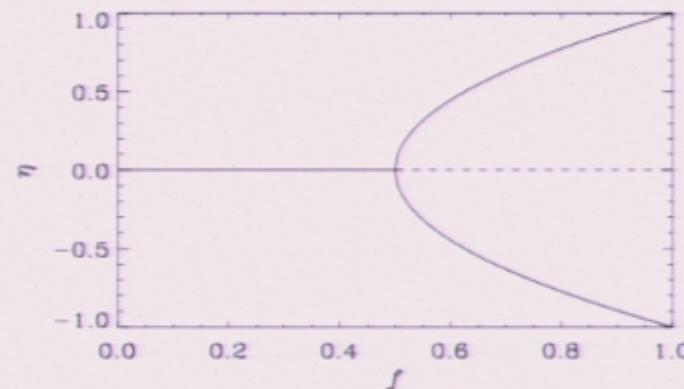
$$0 = -2k_s[L_1]([L_2] - [L_1])$$

Quantitatively close to full model

$$\frac{dR}{dt} = \frac{R^2}{R^2 + L^2} - (L + R)R$$

$$\frac{dL}{dt} = \frac{L^2}{R^2 + L^2} - (L + R)L$$

BAHN (Orig. Life Evol. Biosph. 2005)



Initial bias →

Spatially extended model

with Tuomas Multamäki, Int. J. Astrobiol. 3, 209 (2004)

reaction-diffusion equation

$$\frac{dR}{dt} = \frac{R^2}{R^2 + L^2} - (L+R)R + \kappa \nabla^2 R$$

$$\frac{dL}{dt} = \frac{L^2}{R^2 + L^2} - (L+R)L + \kappa \nabla^2 L$$

proto type: Fisher's equation

$$\dot{X} = X(1-X) + \kappa X''$$

propagating front solutions

$$r = f(x - ct)$$

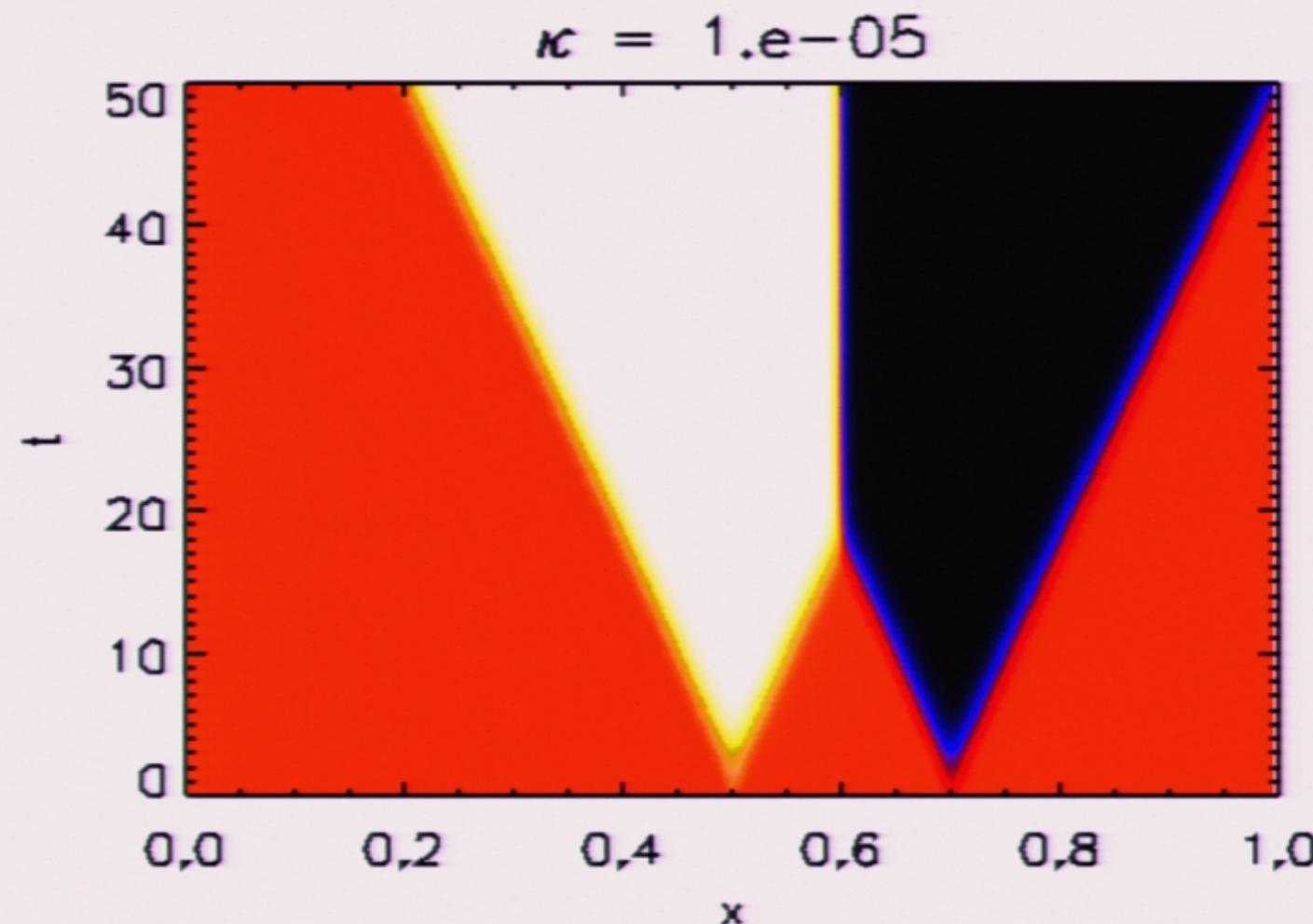
wave speed

$$c \geq 2\sqrt{\lambda \kappa}$$



Spread of the black death

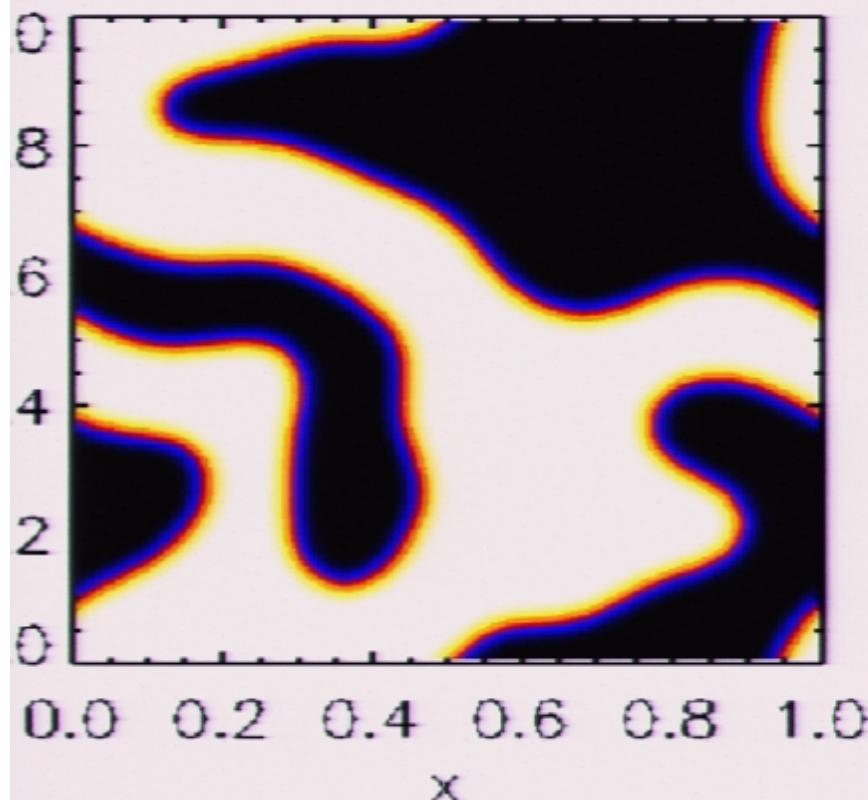
D model (reaction-diffusion equation)



Propagation into racemic environment

D model (reaction-diffusion equation)

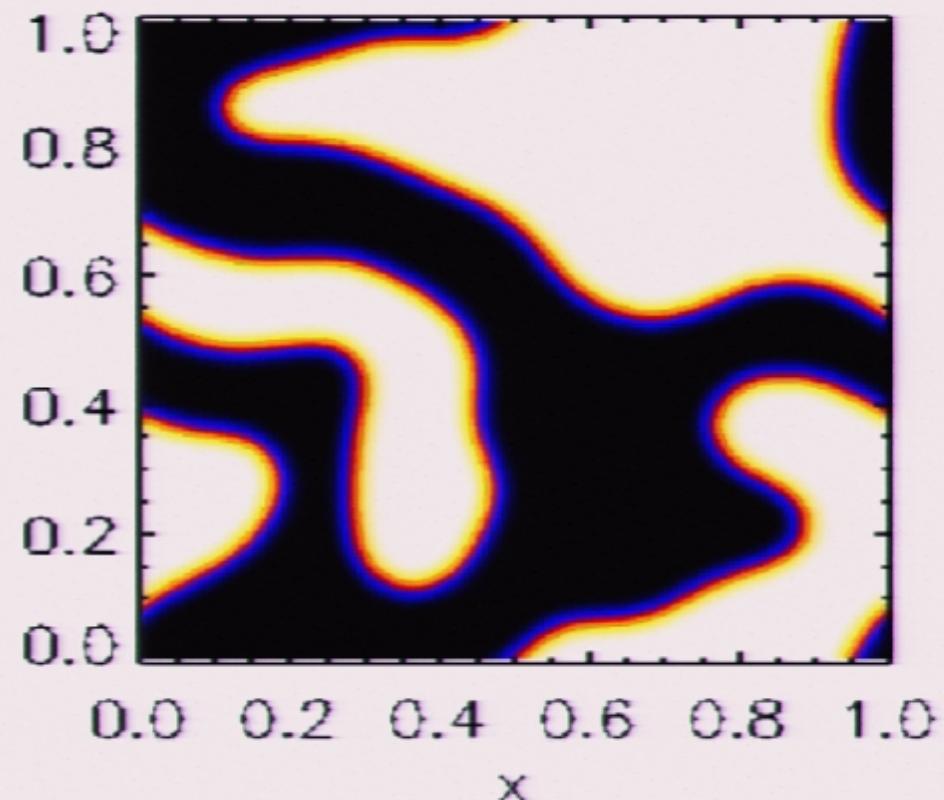
$t = 43.0$



R

short run

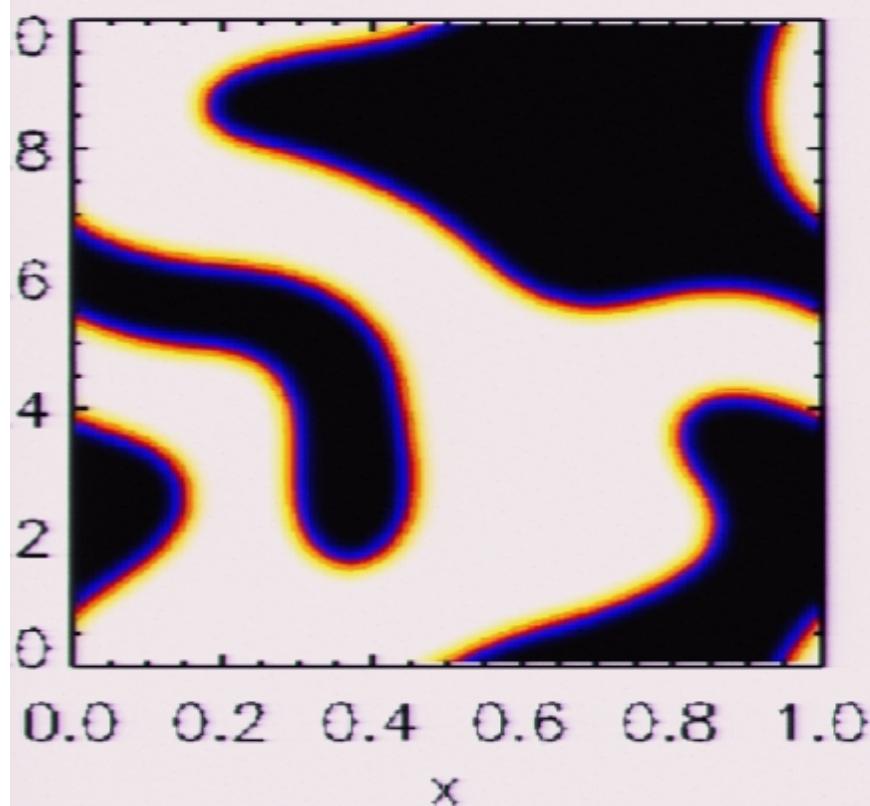
$\kappa = 1.e-04$



L

D model (reaction-diffusion equation)

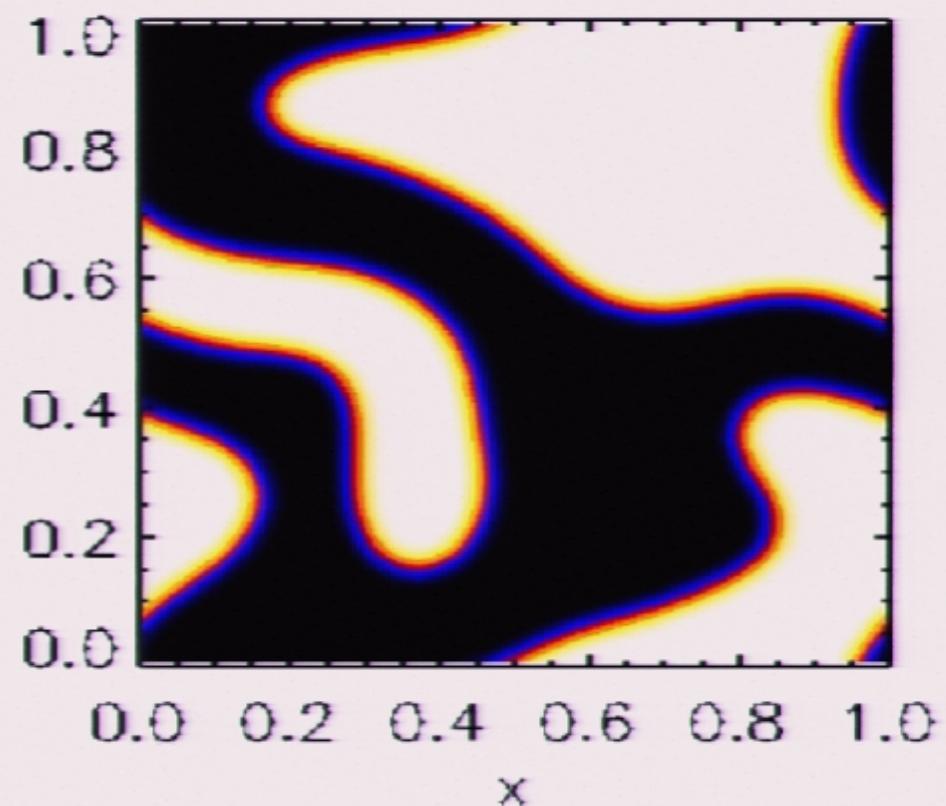
$t = 66.0$



R

short run

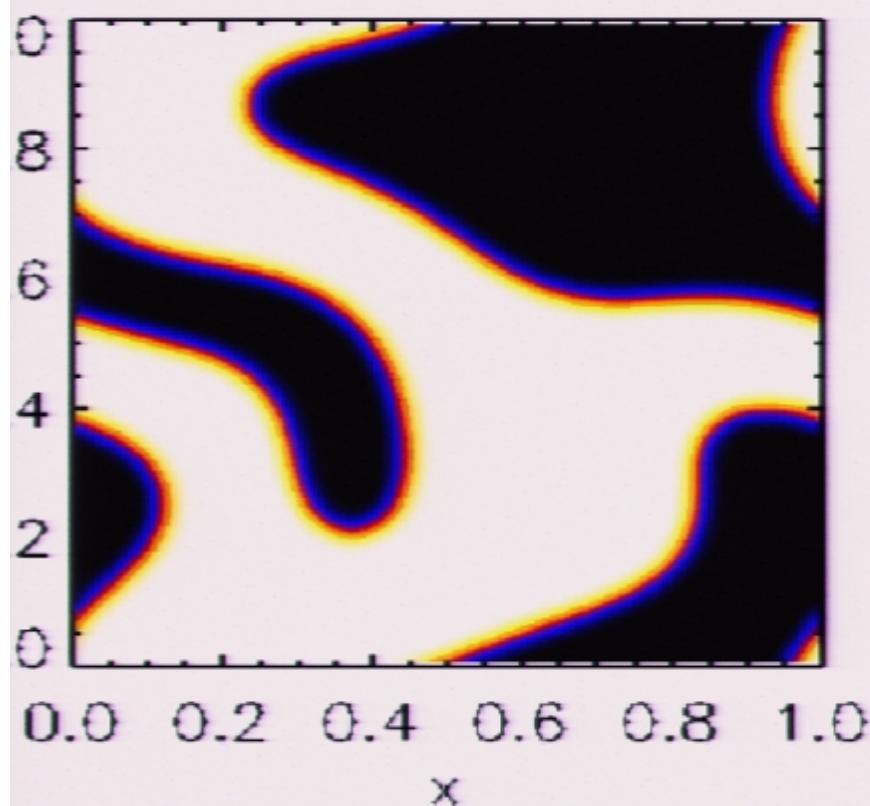
$\kappa = 1.e-04$



L

D model (reaction-diffusion equation)

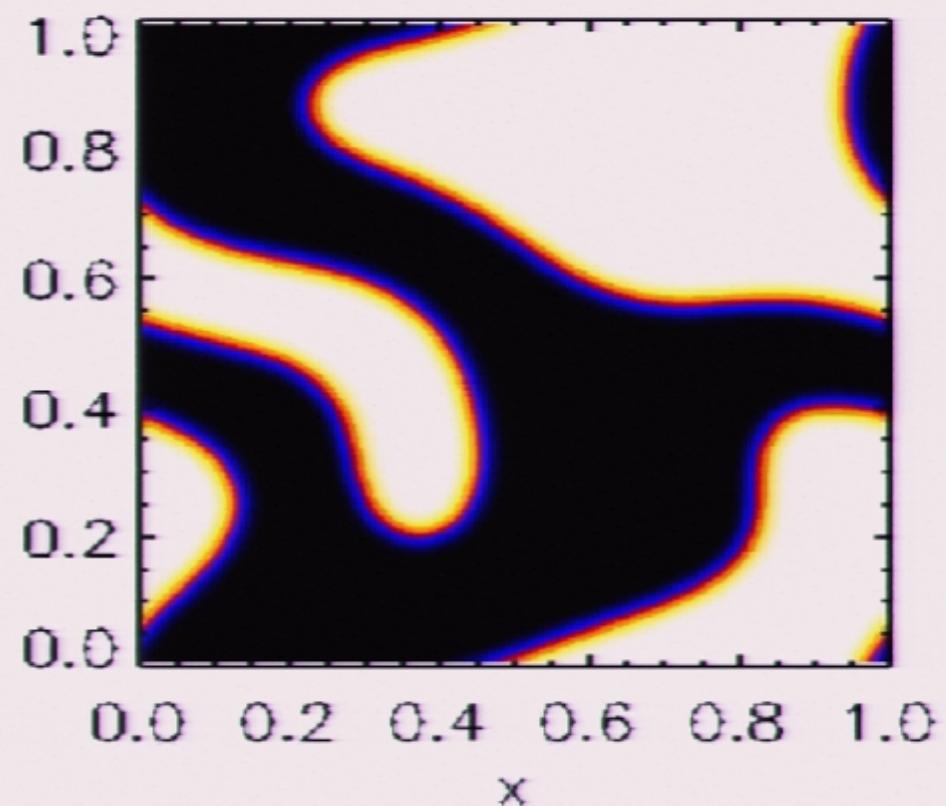
$t = 95.0$



R

short run

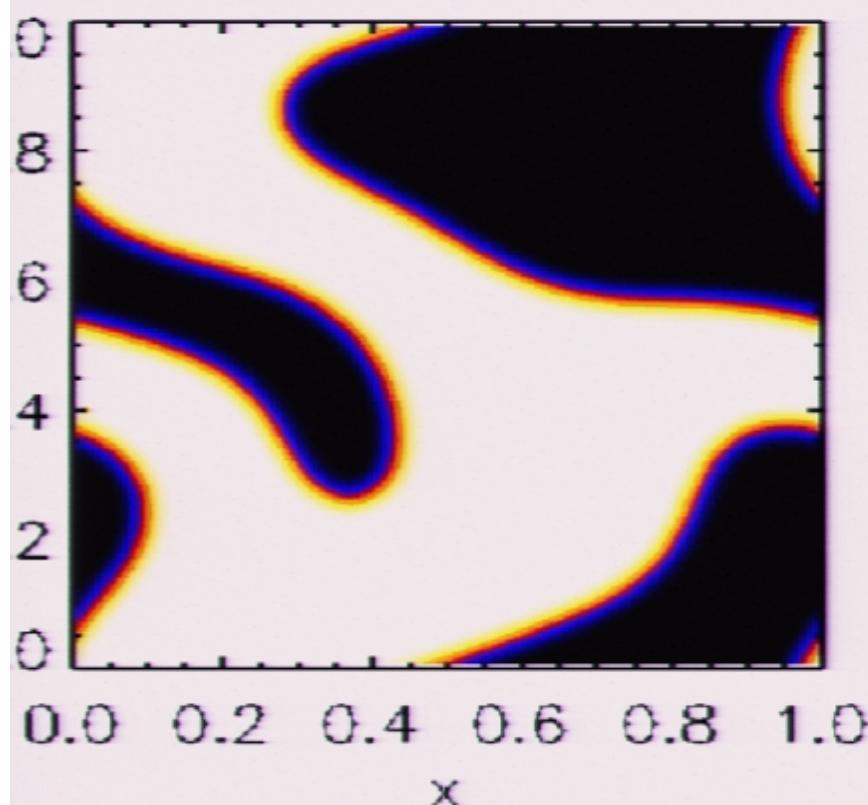
$\kappa = 1.e-04$



L

D model (reaction-diffusion equation)

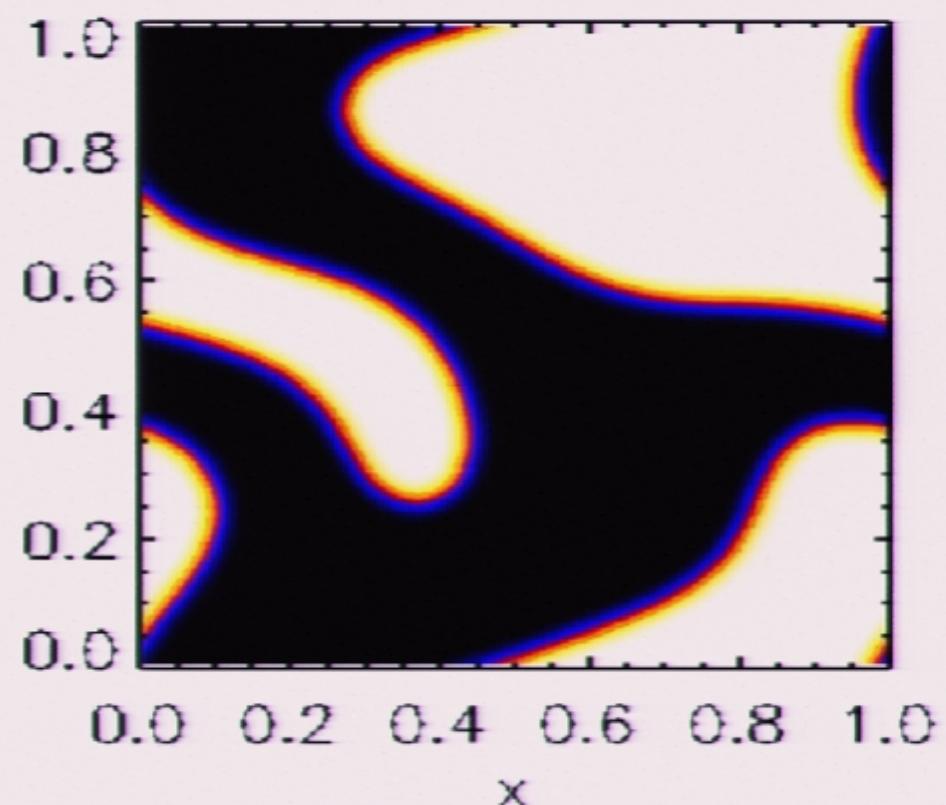
$t = 125.0$



R

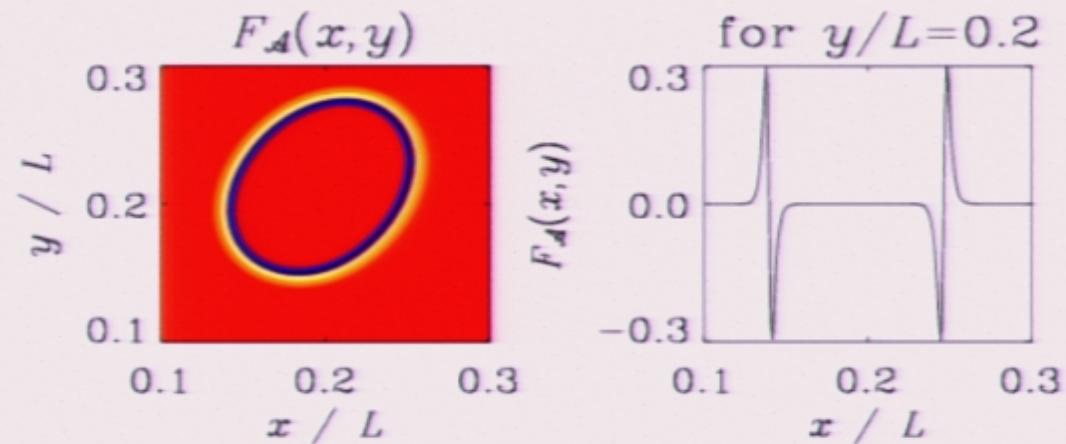
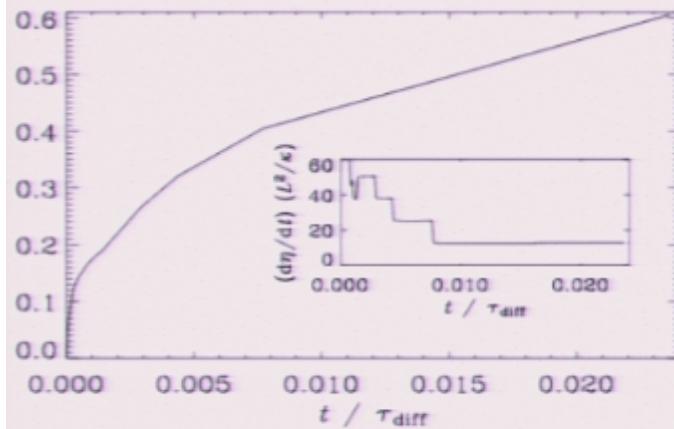
short run

$\kappa = 1.e-04$



L

Piecewise linear increase



Reduced equations

$$\frac{dR}{dt} = \frac{R^2}{R^2 + L^2} - (L + R)R$$

$$\frac{dL}{dt} = \frac{L^2}{R^2 + L^2} - (L + R)L$$

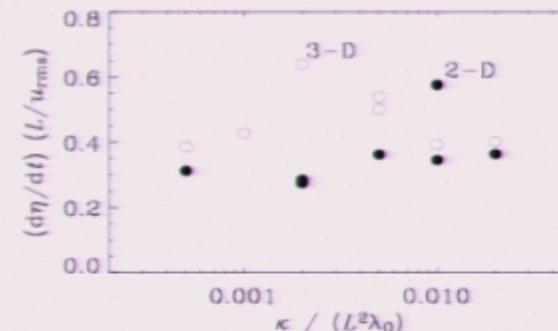
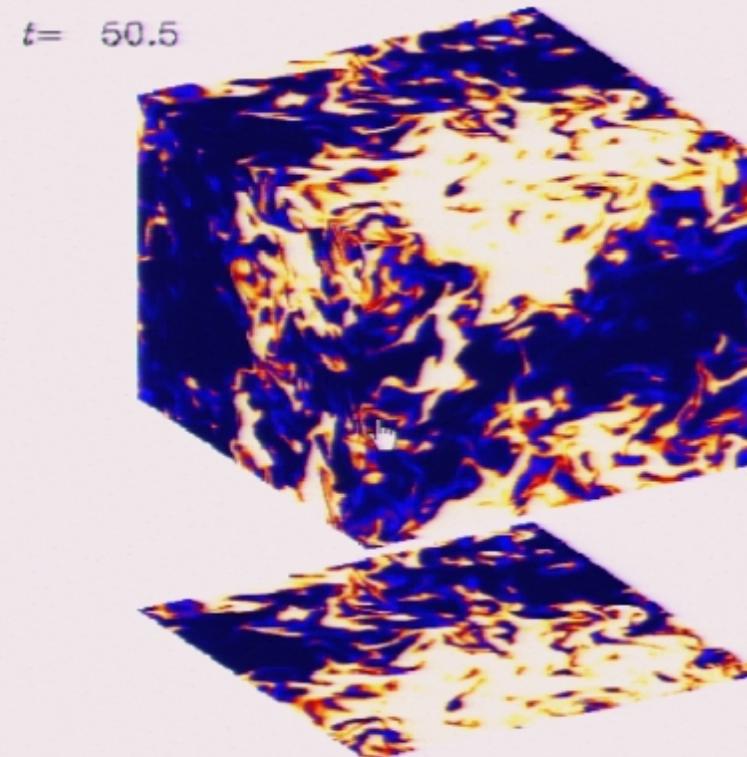
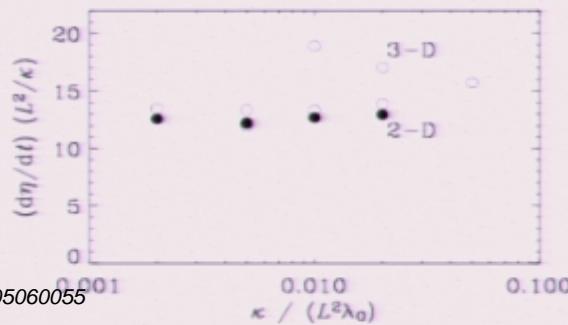
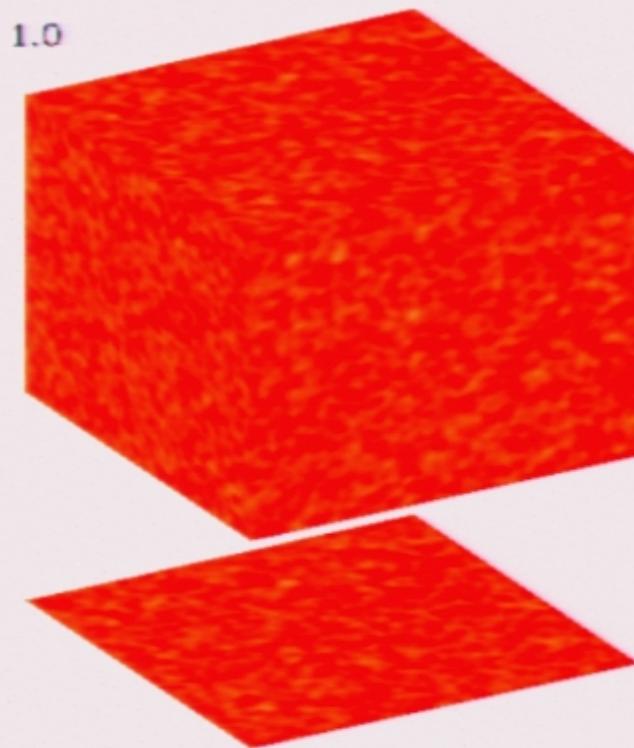
add/subtract:

$$\dot{S} = 1 - S^2 + \kappa \nabla^2 S$$

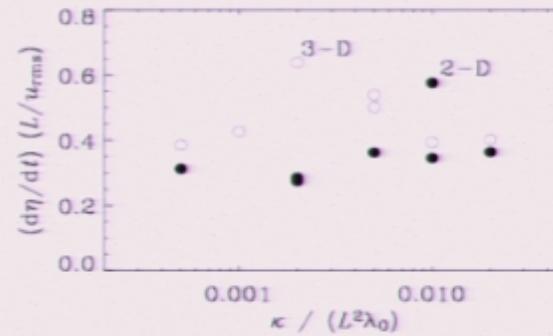
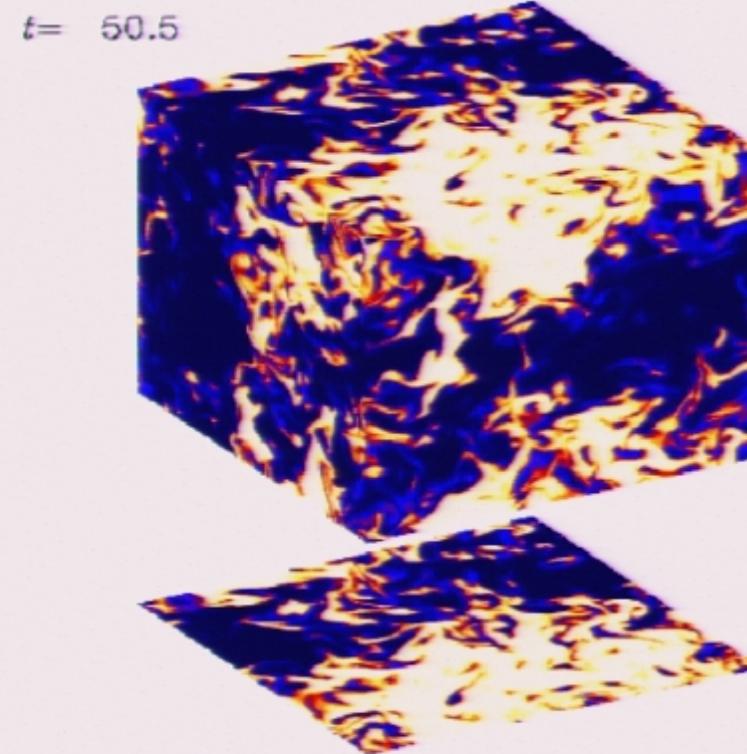
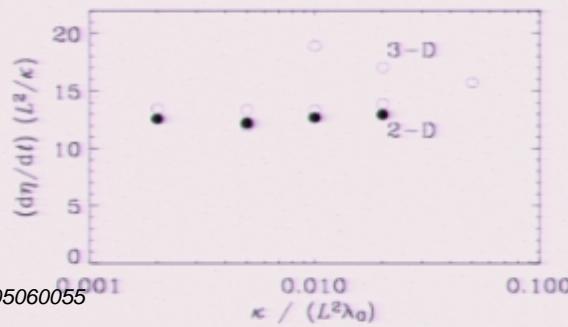
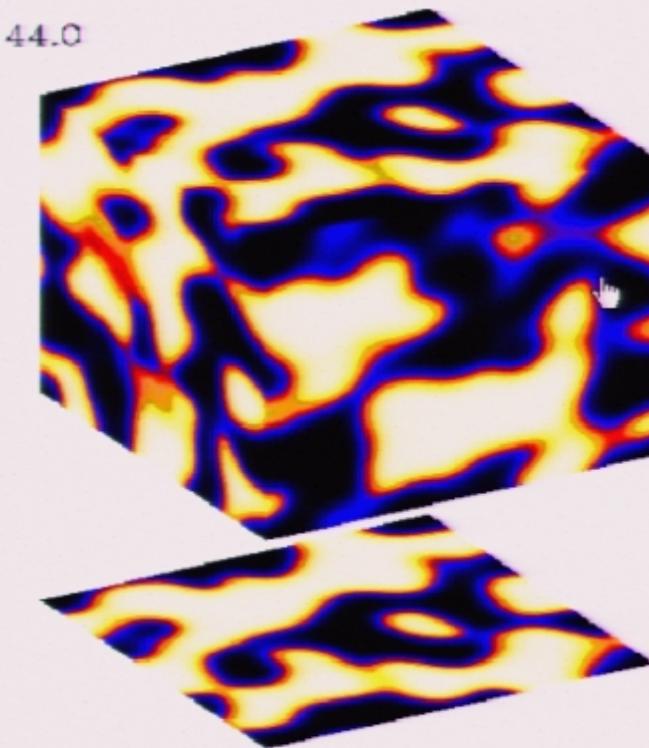
$$\dot{A} = \frac{2 - S^2 - A^2}{S^2 + A^2} SA + \kappa \nabla^2 A$$

$$\Rightarrow \frac{d\langle A \rangle}{dt} = \int \frac{1 - A^2}{1 + A^2} A dS$$

Effects of turbulence



Effects of turbulence





Pencil Code

Started in Sept. 2001 with Wolfgang Dobler

High order (6th order in space, 3rd order in time)

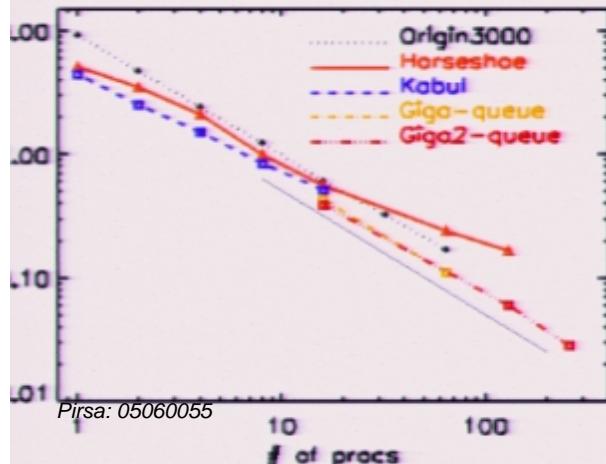
Cache & memory efficient

MPI, can run PacxMPI (across countries!)

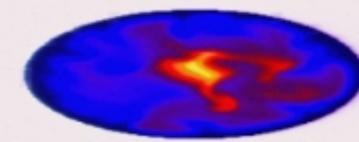
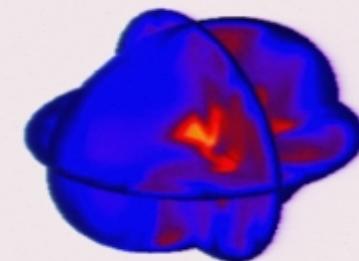
Maintained/developed by ~20 people (CVS!)

Automatic validation (over night or any time)

Max resolution so far 1024^3 , 256 procs



$t = 241.2$



- Isotropic turbulence
 - MHD, passive scl, CR
- Stratified layers
 - Convection, radiation
- Shearing box
 - MRI, dust, interstellar
- Sphere embedded in box
 - Fully convective stars
 - geodynamo
- Other applications
 - Homochirality
 - Spherical coordinates

Conclusions

- Polymerization model:
 - Based on measurable processes
 - Predicts wavelike chromatograms (HPLC)
- Reduction to accurate simplified model
 - Homochirality in space (earth, interstellar, etc)
 - Timescales 500 Myr; fossil evidence of spatially fragmented homochirality?
- Pencil Code: just google for it
 - Detailed manual, ...

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Click to add notes

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