

Title: Quantum limits to measuring spacetime geometry

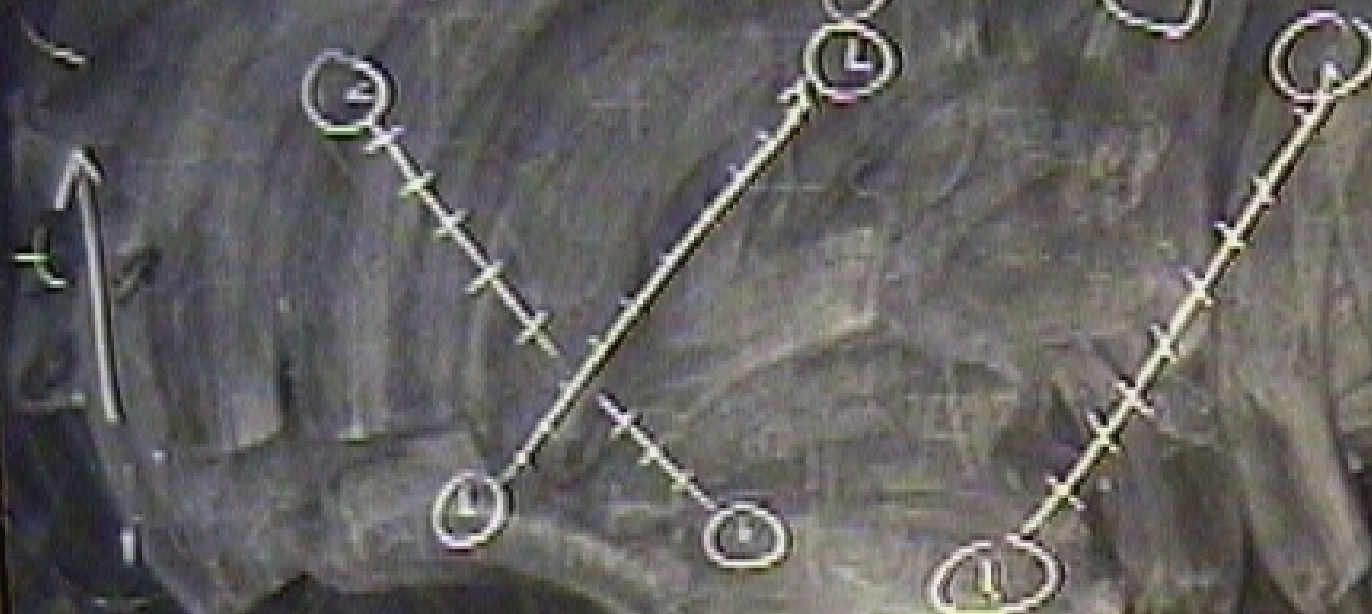
Date: Apr 13, 2005 02:00 AM

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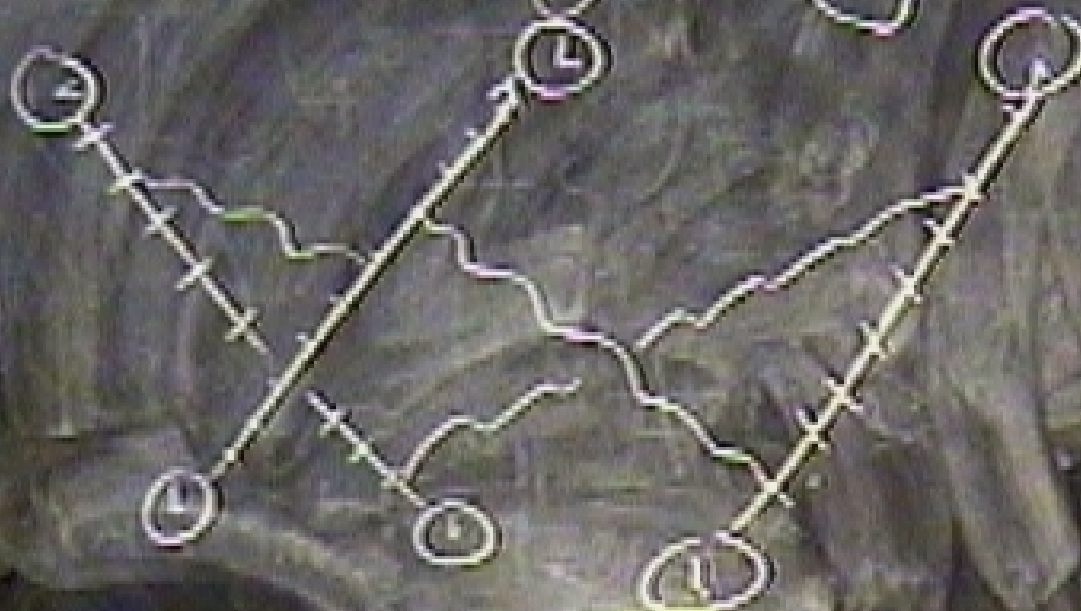
Abstract:

Quantum limits to measuring
spacetime geometry.

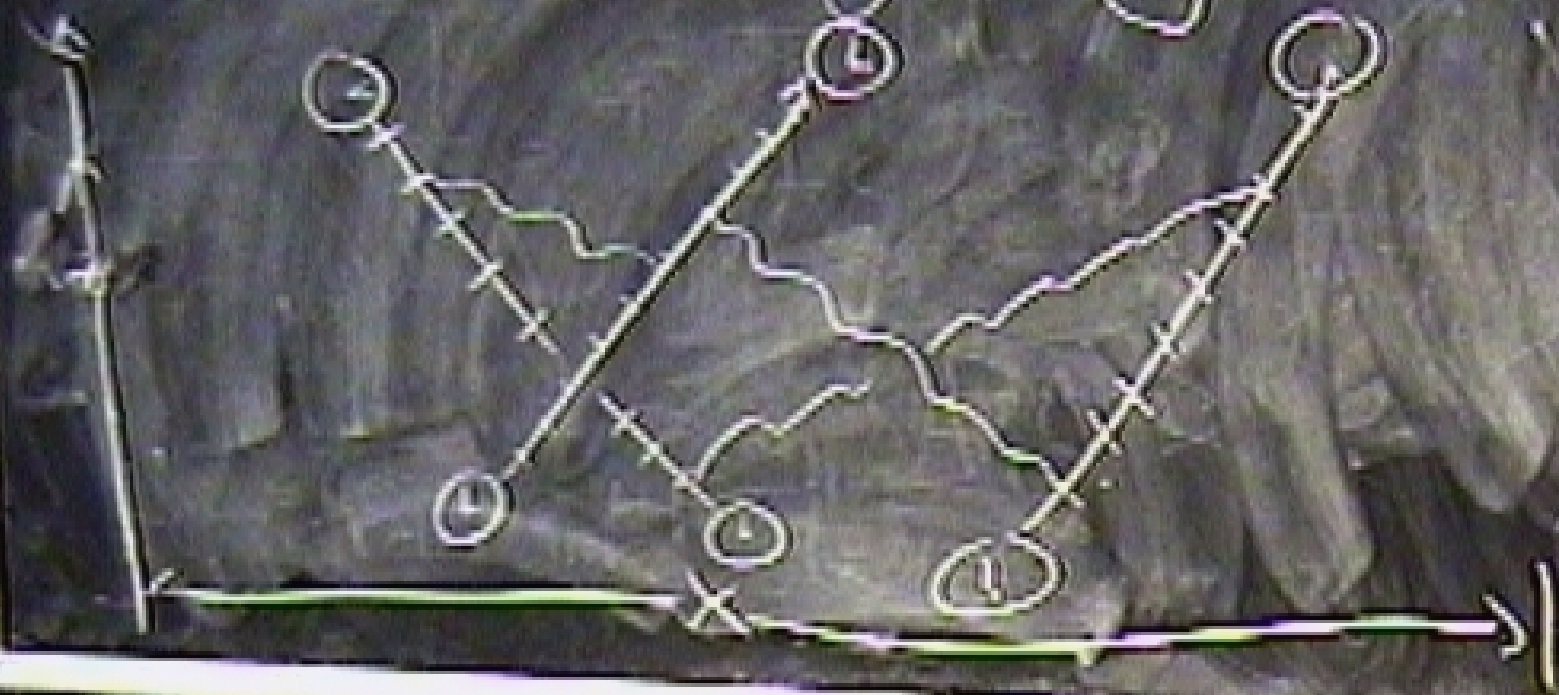
Quantum limits to measuring spacetime geometry



Quantum limits to measuring
spacetime geometry.



Quantum limits to measuring
spacetime geometry.



resolution: # links + clicks

resolution: # ticks + clicks

OK*

resolution: # ticks + clicks

Δx

Two fundamental limits to measurement:

$$\Delta x \Delta p \geq \frac{\hbar}{2}$$

resolution: # ticks + clicks

Δx^*

Two fundamental limits to measurement:

$$\Delta x \Delta p \geq \frac{\hbar}{2} \Rightarrow \text{Standard Quantum Limit}$$



resolution: # ticks + clicks

Δt^*

Two fundamental limits to measurement:

$$\Delta x \Delta p \geq \frac{\hbar}{2} \Rightarrow \text{Standard Quantum Limit}$$

not a limit



$E-t$

$$\Delta E \Delta t \geq \hbar/2$$



E-t

$$\Delta E \Delta t \geq \hbar/2$$

time for clock to tick
detector to click
bit to flip

$E-t$

$$\Delta E \Delta t \geq \hbar/2$$

time for clock to tick
detector to click
bit to flip

$$\Delta E = \left(\langle \psi | H^2 | \psi \rangle - \langle \psi | H | \psi \rangle^2 \right)^{1/2}$$



$E-t$

$$\Delta E \Delta t \geq \hbar/2$$

time for clock to tick
detector to click
bit to flip

$$\Delta E = \left(\langle \psi | H^2 | \psi \rangle - \langle \psi | H | \psi \rangle^2 \right)^{1/2}$$

$\Delta t = \text{time for}$
 $|\psi\rangle \rightarrow |\psi^\perp\rangle$

-Kalligraphie)

$\Delta t = \text{time for}$
 $197 \rightarrow 147^{\frac{1}{2}}$

Heisenberg



E-t

$$\Delta E \Delta t \geq \frac{\hbar}{2}$$

time for clock to tick
detector to click
bit to flip

$$\Delta E = \left(\langle \psi | H^2 | \psi \rangle - \left(\langle \psi | H | \psi \rangle \right)^2 \right)^{1/2}$$

$\Delta t = \text{time for}$
 $\rightarrow |\psi\rangle$

Heisenberg

resolution: # ticks + clicks

Δx^*

Two fundamental limits to measurement:

$$\Delta x \Delta p \geq \frac{\hbar}{2} \Rightarrow \text{Standard Quantum Limit}$$

not a limit



Take care!

Heisenberg

$$E = \hbar \omega$$

Margolus-Levitin
thm.

$$E \Delta t \geq \frac{\hbar}{2}$$

$$E = \langle \psi | \hat{H} | \psi \rangle - E_0$$

↖ energy of ground state

Example: nuclear spin



$$E = 0$$

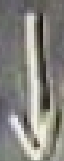


$$E = \pm \mu B = \pm \hbar \omega$$

Example: nuclear spin



$$E = 0$$



$$E = \hbar \mu B = \hbar \omega$$

$$\frac{1}{\sqrt{2}}(|\uparrow\uparrow\rangle + |\downarrow\downarrow\rangle)$$

$$\rightarrow \frac{1}{\sqrt{2}}(|\uparrow\uparrow\rangle + e^{-i\omega t} |\downarrow\downarrow\rangle)$$

Example: nuclear spin



$$E = 0$$

$$E = \hbar \omega = \hbar \gamma B$$

$$\frac{1}{\sqrt{2}}(|\uparrow\rangle + |\downarrow\rangle) \rightarrow \frac{1}{\sqrt{2}}(|\uparrow\rangle + e^{-i\omega t} |\downarrow\rangle)$$

$$\omega t = \frac{\pi}{2} = \frac{\hbar \omega}{E}$$

Example: nuclear spin



$$E = 0$$



$$E = \hbar \mu B = \hbar \omega$$



$$= \frac{1}{\sqrt{2}} (|\uparrow\rangle + |\downarrow\rangle)$$

$$\rightarrow \frac{1}{\sqrt{2}} (|\uparrow\rangle + e^{-i\omega t} |\downarrow\rangle)$$

$$= \frac{1}{\sqrt{2}} (|\uparrow\rangle - |\downarrow\rangle) = \langle \leftarrow$$

$$\omega t = \frac{\pi}{4} = \frac{\hbar \pi}{E}$$

$$\langle E \rangle = \frac{E}{2} \quad \Delta E = E_k$$

$$\Delta t = \frac{\hbar \pi}{2 \langle E \rangle}$$

$$\langle E \rangle = \frac{E}{2} \quad \Delta E = E_h$$

$$\Delta t = \frac{h\pi}{2\langle E \rangle} = \frac{h\pi}{205}$$

$$M-L$$

(2) Atomic clocks

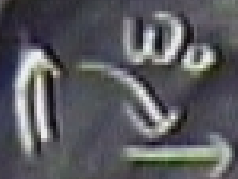


(2) Atomic clocks



$$\omega \approx$$

(2) Atomic clocks



$$\omega \approx \omega_0$$



accuracy $\rightarrow \Delta T \sim \frac{1}{\omega}$

n spins $\rightarrow \Delta T_n^c \sim \frac{1}{\sqrt{n}\omega}$

(2) Atomic clocks



$$\omega \approx \omega_0$$

accuracy $\rightarrow \Delta T \sim \frac{1}{\omega}$

n -spins $\rightarrow \Delta T_n \sim \frac{1}{\sqrt{n} \omega}$

(from Heisenberg $\Delta E_n = \sqrt{n} \Delta E_1$)

Suppose you have, rather than $| \rightarrow \rangle | \rightarrow \rangle$ $\rightarrow$
 $= \frac{1}{\sqrt{2}} (| \uparrow \downarrow \rangle + | \downarrow \uparrow \rangle) \dots \frac{1}{\sqrt{2}} (| \uparrow \downarrow \rangle)$

an entangled 'cat' state:

Suppose you have, rather than $| \rightarrow \rangle | \rightarrow \rangle | \rightarrow \rangle$
 $= \frac{1}{\sqrt{2}} \left(\begin{matrix} \uparrow & \downarrow \end{matrix} \right) \left(\begin{matrix} \uparrow & \downarrow \end{matrix} \right) \dots \frac{1}{\sqrt{2}} \left(\begin{matrix} \uparrow & \downarrow \end{matrix} \right)$

an entangled 'cat' state:

$$\frac{1}{\sqrt{2}} \left(| \uparrow \uparrow \dots \uparrow \rangle + | \downarrow \downarrow \dots \downarrow \rangle \right)$$

$$\rightarrow \frac{1}{\sqrt{2}} \left(| \uparrow \uparrow \dots \uparrow \rangle + e^{-i\text{not}} | \downarrow \downarrow \dots \downarrow \rangle \right)$$

Suppose you have, rather than $|1\rangle|1\rangle$ $\rightarrow$ $|1\rangle|1\rangle$
 $= \frac{1}{\sqrt{2}}(|\uparrow\downarrow\rangle + |\downarrow\uparrow\rangle) \cdot \frac{1}{\sqrt{2}}(|\uparrow\downarrow\rangle + |\downarrow\uparrow\rangle)$

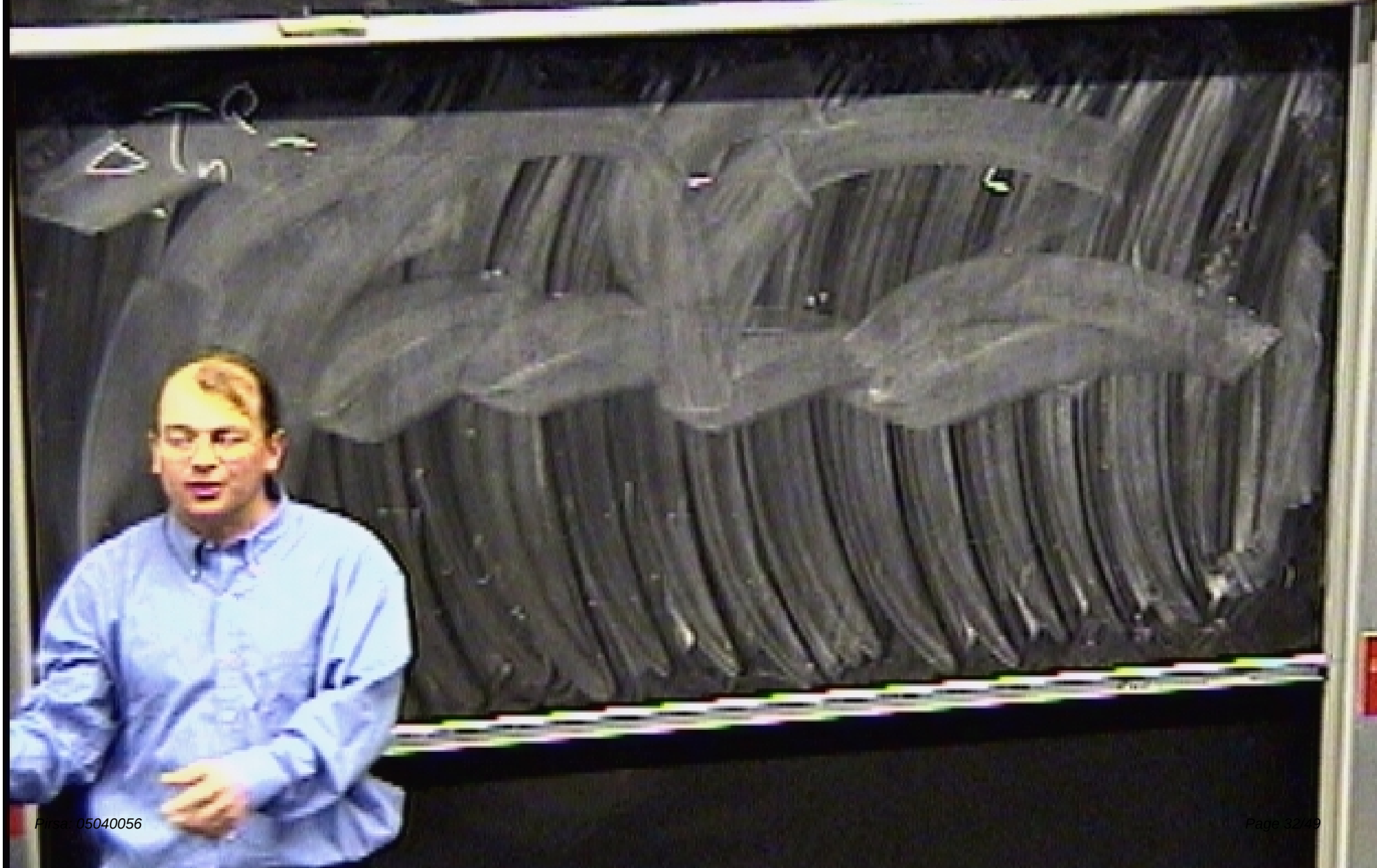
an entangled 'cat' states:

$$\frac{1}{\sqrt{2}}(|\uparrow\uparrow\rangle + |\downarrow\downarrow\rangle)$$

$$\rightarrow \frac{1}{\sqrt{2}}(|\uparrow\uparrow\rangle + e^{-i n \omega t} |\downarrow\downarrow\rangle)$$

$$E_n = n E_1, \quad \Delta E_n = n \Delta E$$

$$E_n = n\hbar\omega, \quad \Delta E_n = \hbar\omega$$



$$E_n = n\hbar\omega, \quad \Delta E_n = \hbar\omega$$

$$\Delta T_n = \frac{1}{n\omega}$$

$\hbar$ enhancement in accuracy

$$E_n = nE_1, \quad \Delta E_n = n\Delta E_1$$

$$\Delta T_n \approx \frac{1}{n\omega}, \quad \text{in enhancement in accuracy}$$

(3) time arrival measurements:

$$E_n = n E_1, \quad \Delta E_n = n \Delta E_1$$

$$\Delta T_n \sim \frac{1}{n \omega}, \quad \text{in enhancement in accuracy}$$

(3) time arrival measurements:

$$1 \text{ photon: } |\varphi\rangle = \int \varphi(\omega) |\omega, 1\rangle$$

$$E_n = n E_1, \quad \Delta E_n = n \Delta E$$

$$\Delta t_n \sim \frac{1}{n \omega}, \quad \sqrt{n} \text{ enhancement in accuracy}$$

(3) time arrival measurements:

1 photon: $|\varphi\rangle = \int \varphi(\omega) |\omega, 1\rangle d\omega$

n uncorrelated: $|\varphi\rangle |\varphi\rangle \dots |\varphi\rangle \rightarrow \Delta E_n \sim \sqrt{n} \Delta E$
 $\Delta T_n \sim \frac{1}{\sqrt{n}}$ shot noise limit

n entangled photons $\int \varphi(\omega) |\omega, 1\rangle_1 |\omega, 1\rangle_2 \dots |\omega, 1\rangle_n d\omega$

$$E_n = nE_1, \Delta E_n = n\Delta E_1$$

$$\Delta T_{arr}^Q = \frac{1}{n}$$

Greater accuracy $\rightarrow$ more energy

Greater accuracy $\rightarrow$ more energy
but too much energy $\Rightarrow$ black hole

$$R_s = \frac{2Gm}{c^2}, \quad m \leq \frac{c^2 R_s}{2G}$$

Greater accuracy $\rightarrow$ more energy
 but too much energy $\Rightarrow$ black hole

$$R_s = 2 \frac{GM}{c^2}$$

$$m \leq \frac{c^2 R_s}{2G} = \frac{c^2 \times}{4G}$$

$$E = mc^2 \leq \frac{xc^4}{4G}$$

$$\begin{aligned} \# \text{ ticks + clicks} &\leq t \cdot \frac{2L}{\pi h} \leq \frac{t}{\pi h} \times \frac{L^4}{L} = \frac{1}{\pi h} \frac{L^4}{L} \\ &= \frac{1}{\pi h} \frac{L^3}{L} \end{aligned}$$

$$\# \text{ ticks + clicks} \leq t \cdot \frac{2\pi}{\hbar} \leq \frac{t}{\hbar} \cdot \frac{\hbar}{4} \cdot \frac{1}{\hbar} = \frac{1}{4\hbar} \frac{t}{\hbar}$$

$$\text{Quantum geometric limit} \# \leq \frac{1}{4\hbar} \frac{t}{\hbar}$$

$$\# \text{ ticks + clicks} \leq t \cdot \frac{2E}{\pi \hbar} \leq \frac{t}{\pi \hbar} \cdot \frac{\hbar^2}{4 \epsilon^2} = \frac{1}{4\pi} \frac{t \epsilon^4}{\hbar \epsilon^2}$$

$$\text{Quantum geometric limit} \# \leq \frac{1}{4\pi} \frac{\hbar}{\epsilon^2} \frac{t}{\epsilon^2}$$

Examples t_p  1 tick

$$\# \text{ ticks + clicks} \leq t \cdot \frac{2\pi}{\hbar} \leq \frac{t}{\hbar} \cdot \frac{\hbar}{4\pi} = \frac{1}{4\pi} \frac{t}{\hbar}$$

Quantum geometric limit $\# \leq \frac{1}{4\pi} \frac{x \cdot t}{\hbar}$

Examples $t_p \square x_p$ 1 tick

$$\left(\frac{t}{t_p} \right)^2 \sim 10^{20}$$

Universe $x \sim ct$, $t \approx 10^{10}$ years $\sim 10^{17}$ sec.

$$\# \text{ ticks + clicks} \leq t \cdot \frac{2E}{\pi \hbar} \leq \frac{t \cdot 2 \cdot \hbar c^4}{\pi \hbar \cdot 4 \cdot G} = \frac{1}{4\pi} \frac{c^4}{\hbar G} t$$

ops/sec

Quantum geometric limit $\# \leq \frac{1}{4\pi} \frac{c^4}{\hbar G} t$



$$\left(\frac{t_p}{t_p} \right)^2 \sim 10^{20}$$

Universe $x \sim ct$, $t \sim 10^{10} \text{ yrs} \sim 10^{17} \text{ sec.}$

Ch. Holography : # bits in 3 volume
 $\leq \frac{\text{Area}}{l_P^2}$

ops in 4 volume
 $\leq \frac{\text{Area of external world sheet}}{l_P^2}$

Ch. holography : #bits in 3 volume
 $\leq \frac{\text{Area}}{l_P^2}$

ops in 4 volume
 $\leq \text{Area of } \underline{\text{extremal worldsheet}}$

instead of #ops $= \frac{E}{t} \epsilon$ $t_P l_P$
 $= \int T_M \, dV$

Ch. Holography : # bits in 3 volume
 $\leq \frac{\text{Area}}{l_P^2}$

ops in 4 volume
 $\leq$ Area of extended world sheet

instead of #ops = $\frac{E}{\hbar} t$ $t \sim l_P$
 $= \int T_{nn} dV$

$$E_n = n E_1, \quad \Delta E_n = n \Delta E_1$$

