

Title: Quantum Geometry Phenomenology: A Discrete Machian Model

Date: Oct 31, 2004 09:00 AM

URL: <http://pirsa.org/04100044>

Abstract:

# A Discrete Machian Model

*Hypothesen sind Netze, nur der wird fangen der auswirft*

-Novalis

Seth Major

Quantum Gravity in the Americas  
Perimeter Institute  
31 October 2004

# A Discrete Machian Model

- Motivation:
  - Quantum Geometry Phenomenology, MDR
  - Discrete Spatial Geometry
- Assumptions
- Discrete Machian Model
- Constraining parameters with tests of LLI

## Quantum Geometry Phenomenology

### Can it be done?

- Modified dispersion relations (MDR) [Coleman, Glashow; Amelino-Camelia; Jacobson, Liberati, Mattingly; Alfaro, Morales-Tecotl, Urrutia; Myers, Pospelov; Konopka, SM, ...]

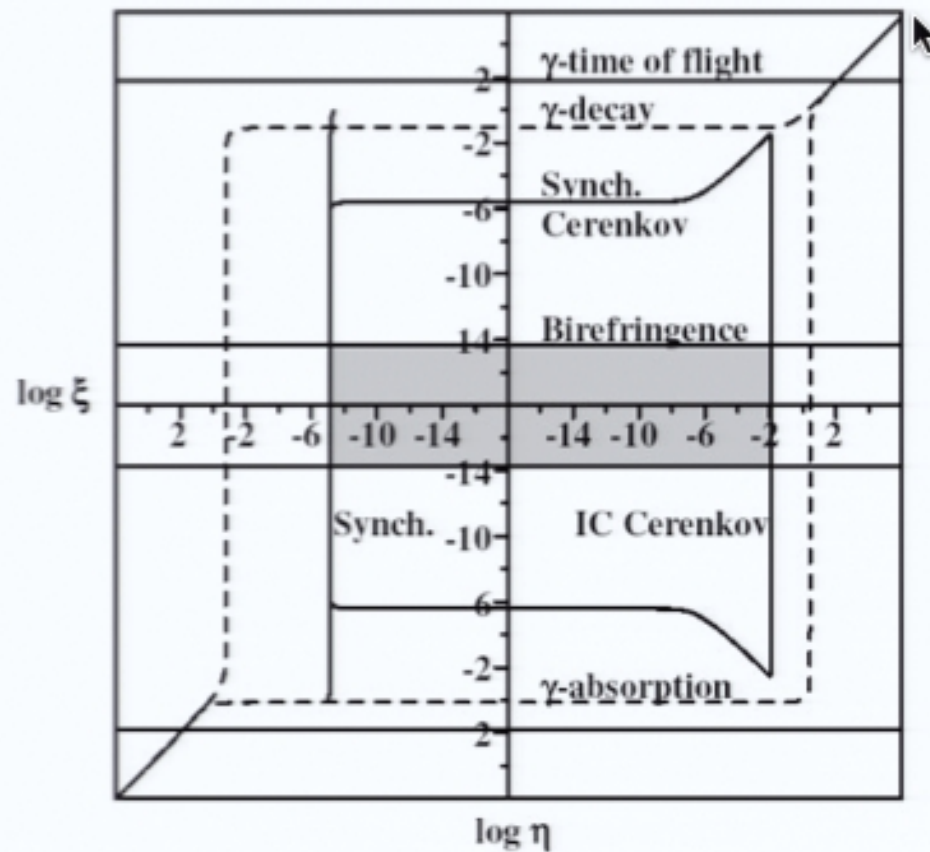
$$E_{\pm}^2 = p^2 + m^2 + \kappa_{\pm} \ell_P p^3$$

- Separate parameters for photons, fermions, and hadrons.
- Particle threshold kinematics is highly sensitive to cubic terms.
- Current astrophysical observation severely constrain modifications.

## Status of MDR 2004

- Myers & Pospelov: Effective field theory framework → 3 parameters for QED - helicities independent
- Jacobson, Liberati, Mattingly, Stecker: Spectacular new constraints from synchrotron radiation in SN remnants and polarization dependent dispersion from gamma ray burst GRB021206
- Collins, Perez, Sudarsky: A fine tuning problem → theoretical framework is sketchy
- Amelino-Camelia: Wait, wait! Should we be using *field theory at all* for QG phenomenology?

## Status of MDR 2004



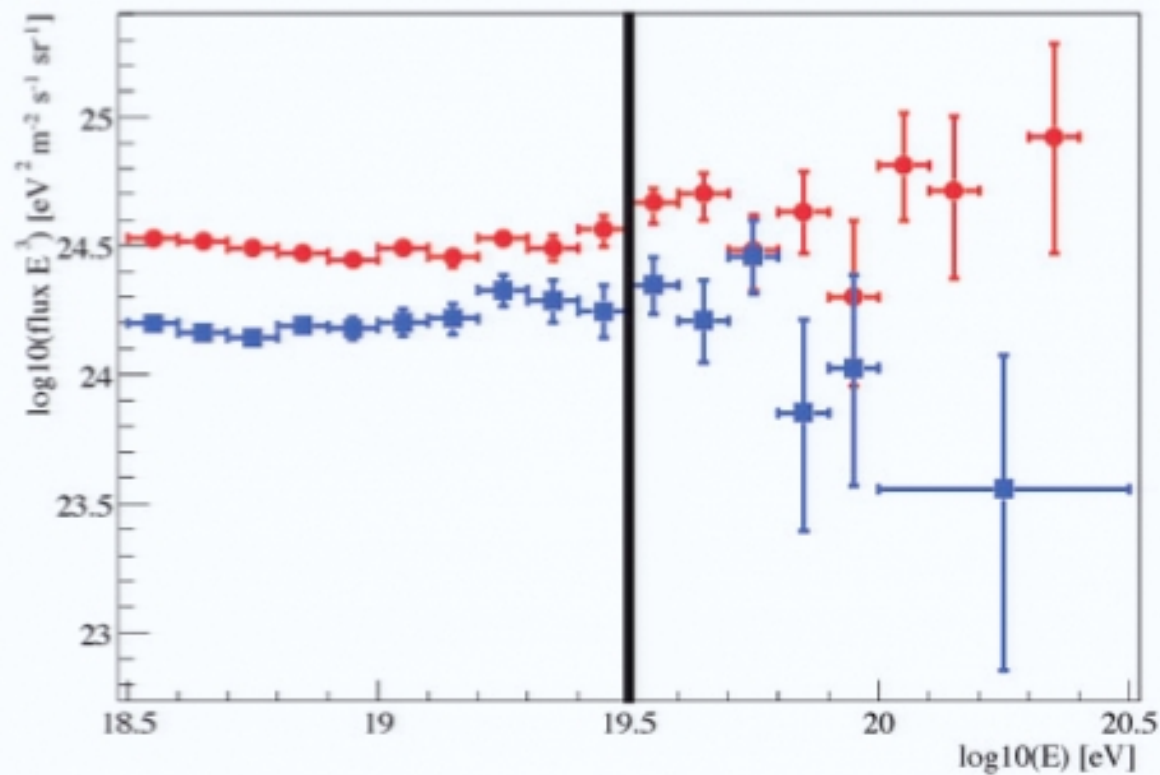
Allowed region is shaded

JLM arXiv:astro-ph/0309681



## Is the GZK cutoff observed?

De Marco *et. al.* see no cutoff at a  $2.5 \sigma$  level !



AGASA and HiRes  
Data

arXiv: astro-ph/0301497

## Quantum Geometry Phenomenology

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What about spatial geometry?

What about “breaking” **rotational invariance**?

No natural scale for onset of discreteness.

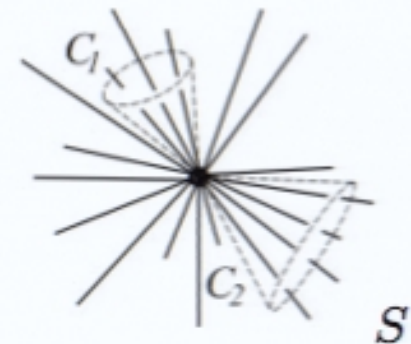
## Discrete spectra for Geometry

**Area\***:  $\hat{A}_S |s\rangle = a |s\rangle$

$$a = \ell_{eff}^2 \sum_{n=1}^N \sqrt{j_n(j_n + 1)}$$

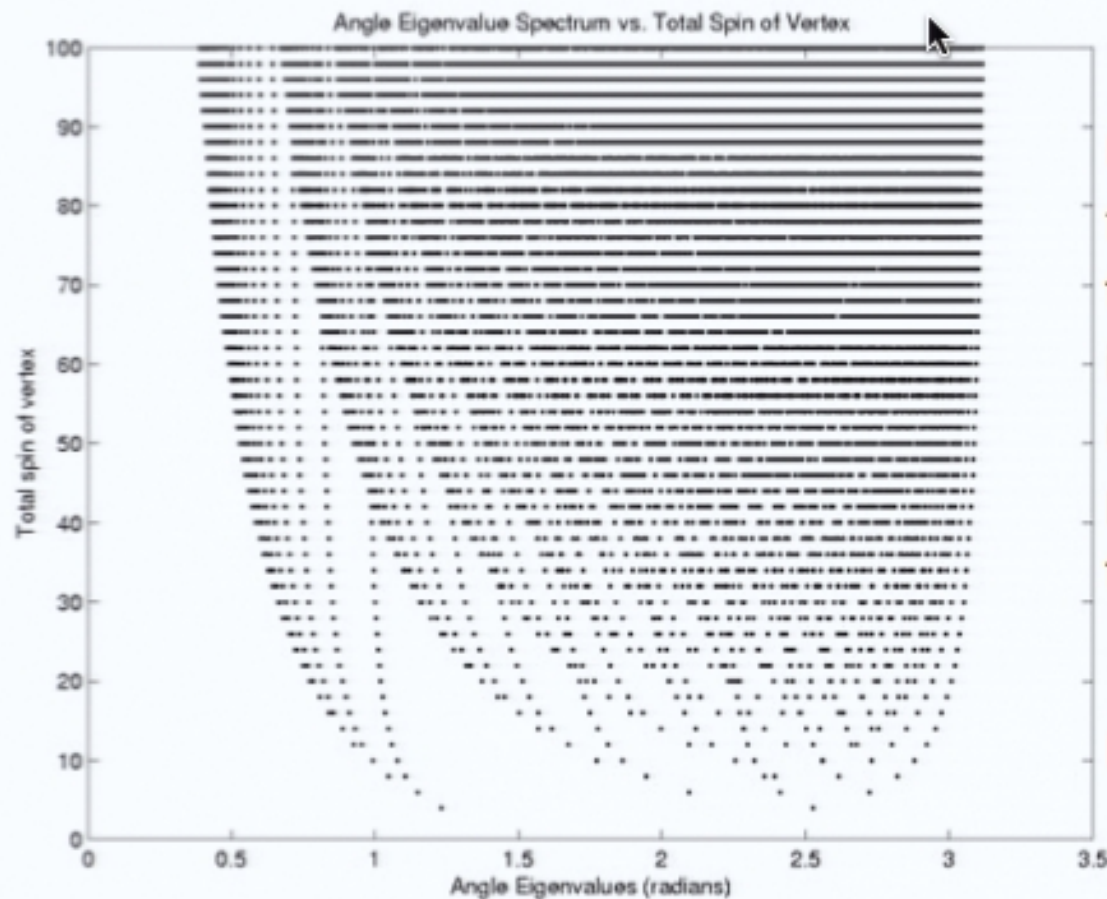
**Angle†**:  $\hat{\theta} |s\rangle = \theta |s\rangle$

$$\theta = \arccos \left( \frac{j_r(j_r + 1) - j_1(j_1 + 1) - j_2(j_2 + 1)}{2 [j_1(j_1 + 1) j_2(j_2 + 1)]^{1/2}} \right)$$



\*Rovelli, Smolin NPB **422** (1995) 593; Ashtekar, Lewandowski CQG **14** (1997) A43

# Discrete spectra for Geometry




Unless the graph is highly flocculent, the angle spectrum is coarse. OV studies suggest that continuum angles are only approximated for total vertex spins of

$$j \sim 10^{20} !$$

[SM, Siefert CQG **19** (2002) 2211]

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Model based on two assumptions: 

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(ii) Inertial mass is determined by the total distribution of matter

9-b

## A Discrete Machian Model

Cocconi-Salpeter (1958): Non-uniformities in matter distribution lead to **anisotropy of mass**  $m \rightarrow m_{ij}$ .

Quantitative form of Mach's principle:

$$m_{ij} \propto \frac{M(r)}{r^\nu}$$
$$m_{ij} \propto f(\theta)$$

$\theta$  between acceleration/velocity and preferred direction.

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**Idea:** As an experiment rotates with the Earth look for side-real variations in the structure of Zeeman transitions. A short calculation shows

$$\Delta E = \frac{\Delta m}{m} \bar{T} \langle P_2(\cos \theta) \rangle$$

Not all transitions are affected,  $\langle P_2 \rangle_{J=0} = 0$ , so one transition can act as a reference.

## A Discrete Machian Model

- Non-uniformities in matter variations cause deviations from isotropy
  - ↪ there is a local preferred direction
- Discrete random geometry cannot support a continuously changing angle
- If all deviations are equally likely, and the path does not depend on history, then “the preferred direction” performs a random walk
  - ↪ Diffusion on the space of directions



There is a local, stochastic preferred direction  $\hat{n}(t)$  .



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Suggests looking for a generalized Dirac equation (Lämmerzahl CQG 15 (1998) 13). In the non-relativistic limit

$$H\varphi = -\frac{1}{2m} \left( \delta^{ij} + \frac{\delta m^{ij}(t)}{m} + \frac{\delta \bar{m}_k^{ij}(t)}{m} \sigma^k \right) \nabla_i \nabla_j \varphi - A_j^i \sigma^j \nabla_i \varphi \\ + \frac{e}{2m} F^{ij} \left( K^k \epsilon_{ijk} + K_{ijk} \sigma^k \right) \varphi$$



## A Discrete Machian Model

Parameters of the model:

Scale of the anomalous mass tensor

$$\frac{\delta m_{ij}}{m} = \alpha \frac{M_{\text{non-uni}}}{r^\nu}$$

Diffusion constant:

$$\sigma = \frac{\delta^2}{4\Delta t}$$

$\delta$  is the angular scale and  $\Delta t$  is the time scale of the random walk

Can tests of local Lorentz invariance constrain the shape and scale of the discreteness?

## Tests of LLI

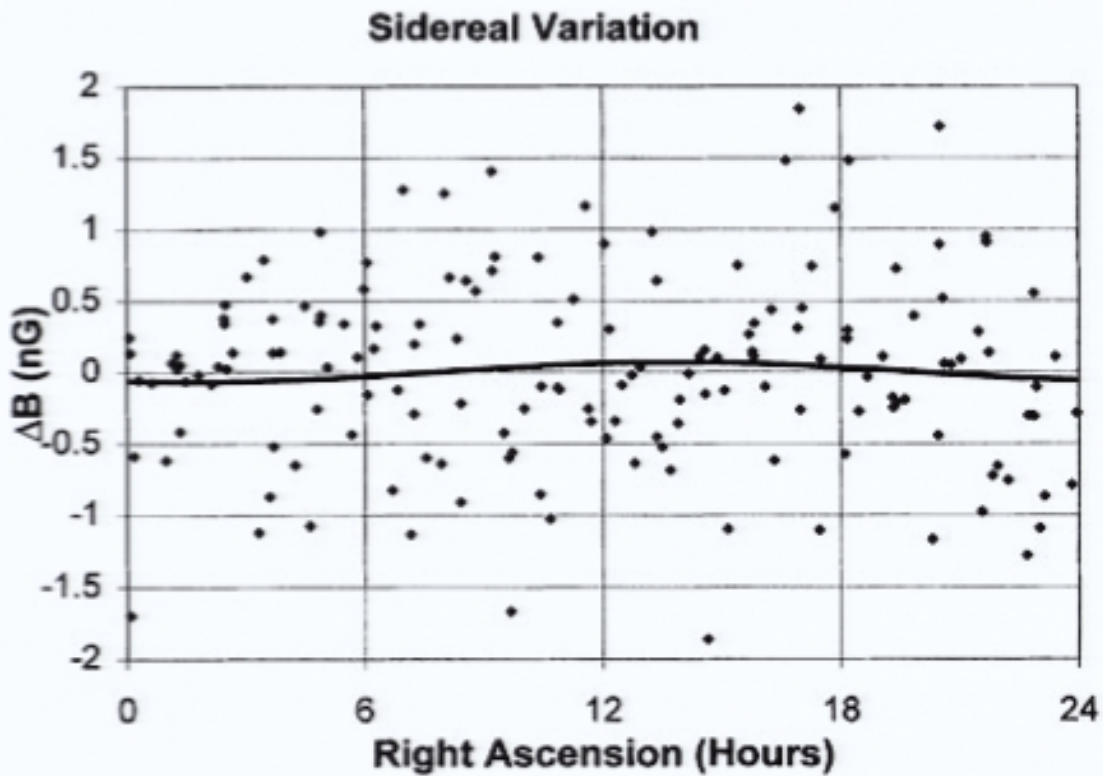
### • Hughes-Drever Experiments

- Search for Lorentz-violating Zeeman frequency shifts
- Some of the most precise null experiments ever performed.
- No sidereal variation have been observed.
  
- Modern versions: co-located magnetometers track relative phases and Larmor frequencies. Recent results show

$$\frac{\delta m}{m} \leq 10^{-23}$$

## Tests of LLI

Example data: Variation in **B** as function of sidereal day



From Berglund *et.al.*  
PRL 75 (1995) 1879.

## Summary: A Discrete Machian Model

- Three parameters  $\alpha, \sigma, \nu$
- Previous tests of LLI constrain parameters e.g.  $\alpha \leq 10^{-23}$

### To do:

- Relation to (generalized) Standard Model Extension (Kostelecky et. al.)
- Measurement schemes for discreteness **scale**,  $\delta$ , and **shape**.
- ...